

*Survey of India.*

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PROFESSIONAL PAPER—No. 16.

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THE  
EARTH'S AXES  
AND  
TRIANGULATION

BY

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PUBLISHED BY ORDER OF THE GOVERNMENT OF INDIA.



Dehra Dun:

PRINTED AT THE OFFICE OF THE TRIGONOMETRICAL SURVEY.

1918.

Price Rs. 4 or 5s. 4d.



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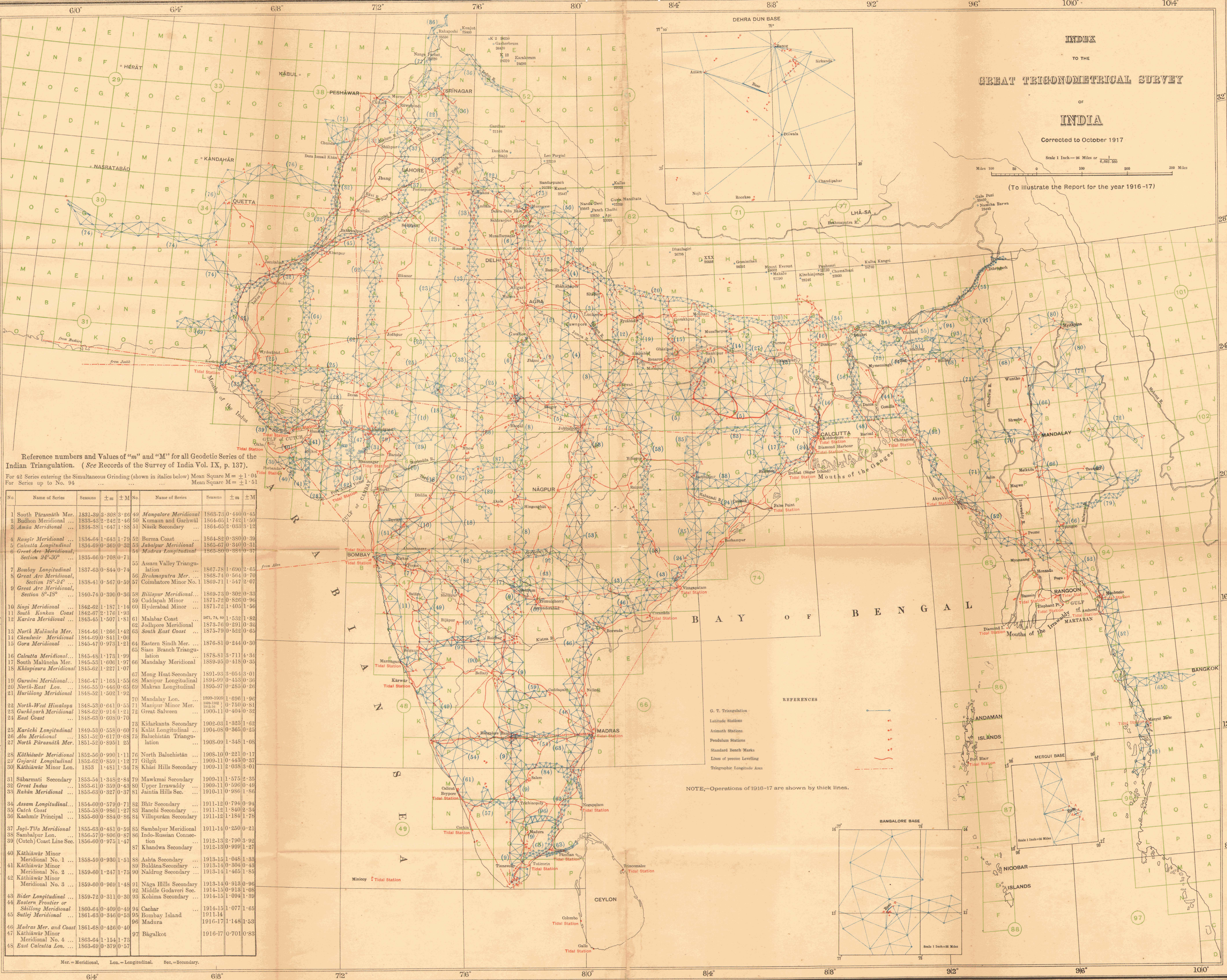


INDEX  
TO THE  
**GREAT TRIGONOMETRICAL SURVEY**  
OF  
**INDIA**

Corrected to October 1917

Scale 1 Inch = 96 Miles or 153.6 Kilometres

(To illustrate the Report for the year 1916-17)



Reference numbers and Values of "m" and "M" for all Geodetic Series of the Indian Triangulation. (See Records of the Survey of India Vol. IX, p. 137).

For 42 Series entering the Simultaneous Gridding (shown in italics below) Mean Square M = ± 1.04  
For Series up to No. 94 ... Mean Square M = ± 1.51

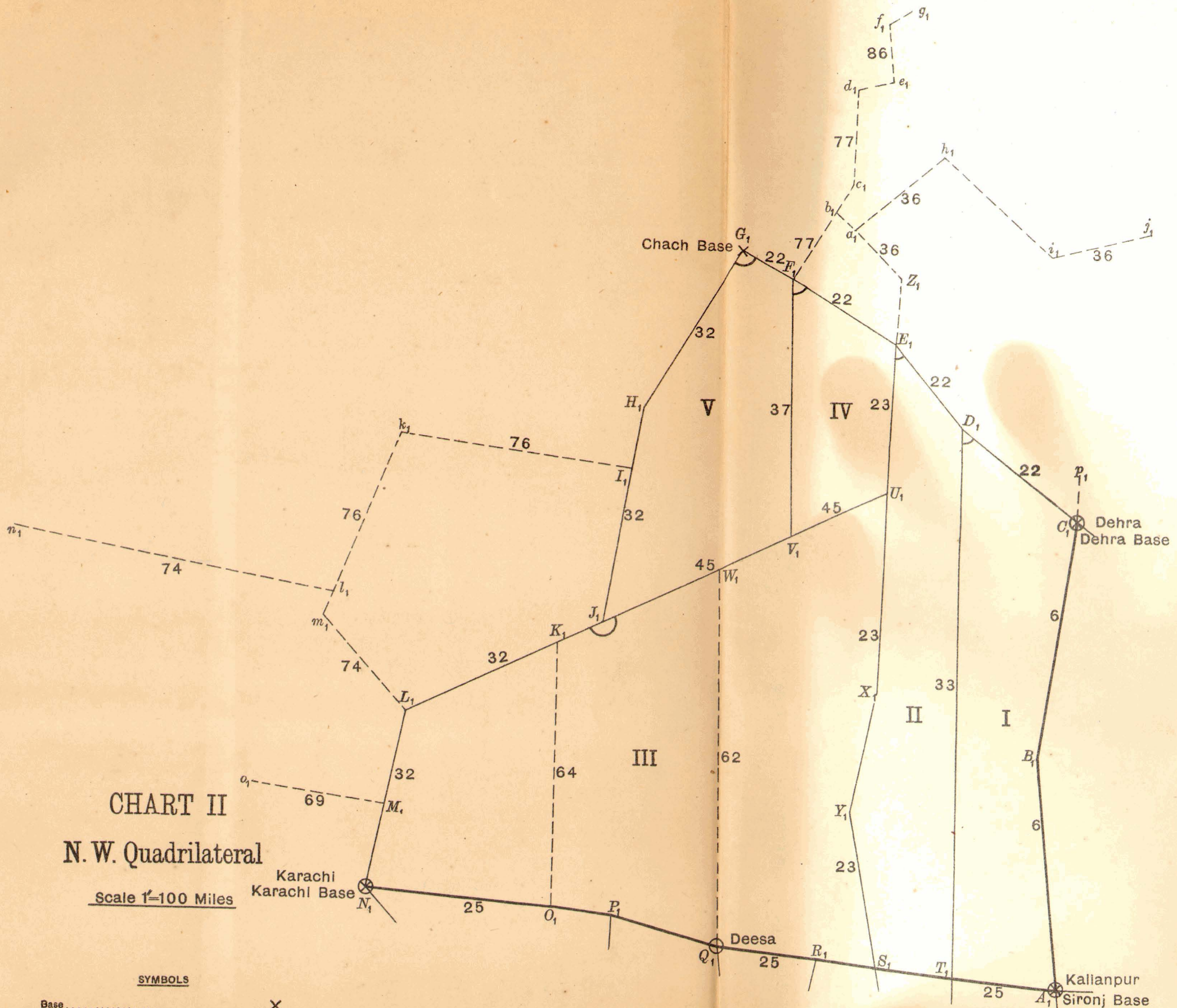
No.	Name of Series	Seasons	± m	± M	No.	Name of Series	Seasons	± m	± M
1	South Parasnath Mer.	1831-39	3.808	3.26	49	Mangalore Meridional	1863-73	0.440	0.43
2	Budhon Meridional ...	1833-43	2.242	2.46	50	Kumaon and Garhwal	1864-65	1.742	1.50
3	Amia Meridional ...	1834-38	1.647	1.88	51	Nasik Secondary ...	1864-65	2.033	3.12
4	Rangir Meridional ...	1834-64	1.643	1.76	52	Burma Coast	1864-82	0.380	0.39
5	Calcutta Longitudinal ...	1834-69	0.369	0.32	53	Jabalpur Meridional	1865-67	0.340	0.31
6	Great Arc Meridional, Section 24°-30° ...	1835-66	0.708	0.71	54	Madras Longitudinal	1865-80	0.384	0.37
7	Bombay Longitudinal, Section 15°-24° ...	1837-63	0.844	0.74	55	Assam Valley Triangulation	1867-78	1.690	2.65
8	Great Arc Meridional, Section 5°-15° ...	1838-41	0.567	0.50	56	Brahmaputra Mer. ...	1868-74	0.564	0.70
9	Great Arc Meridional, Section 5°-15° ...	1840-74	0.390	0.36	57	Coimbatore Minor No. 1	1869-71	1.547	2.07
10	Singri Meridional ...	1842-62	1.157	1.14	60	Hyderabad Minor ...	1871-72	1.405	1.56
11	South Konkan Coast	1842-67	2.176	1.93	61	Malabar Coast ...	1871-74	0.291	0.32
12	Karira Meridional ...	1843-45	1.507	1.81	62	Jodhpore Meridional	1873-76	0.291	0.32
13	North Malincha Mer.	1844-46	1.266	1.42	63	South East Coast ...	1873-79	0.522	0.63
14	Chandnur Meridional	1844-69	0.841	1.03	64	Eastern Sindh Mer. ...	1876-81	0.244	0.30
15	Gora Meridional ...	1845-47	0.973	1.21	65	Siam Branch Triangulation	1878-81	3.711	4.34
16	Calcutta Meridional ...	1845-48	1.173	1.99	66	Mandalay Meridional	1879-85	0.418	0.35
17	South Malincha Mer.	1845-53	1.606	1.97	67	Mong Heat Secondary	1891-93	3.054	3.01
18	Khampura Meridional	1845-62	1.227	1.07	68	Manipur Longitudinal	1894-99	0.453	0.36
19	Gurwani Meridional ...	1846-47	1.165	1.56	69	Makran Longitudinal	1895-97	0.285	0.26
20	North-East Lon. ...	1846-55	0.446	0.65	70	Mandalay Lon. ...	1899-1009	1.696	1.96
21	Havillong Meridional	1848-52	1.502	1.92	71	Manipur Minor Mer. ...	1899-1002	0.750	0.81
22	North-West Himalaya	1848-53	0.641	0.55	72	Great Salween ...	1900-11	0.404	0.32
23	Gurkagarh Meridional	1848-62	0.914	1.21	73	Kidarkanta Secondary	1902-03	1.323	1.62
24	East Coast ...	1848-63	0.608	0.70	74	Kalit Longitudinal ...	1904-08	0.365	0.25
25	Karachi Longitudinal	1849-53	0.558	0.60	75	Baluchistan Triangulation	1905-09	1.348	1.08
26	Abu Meridional ...	1851-52	0.617	0.63	76	North Baluchistan ...	1908-10	0.221	0.17
27	North Parasnath Mer.	1851-52	0.895	1.25	77	Gilgit	1909-11	0.443	0.37
28	Kathiawar Meridional	1852-56	0.990	1.11	78	Khasi Hills Secondary	1908-11	2.038	3.01
29	Gujarat Longitudinal	1852-62	0.839	1.12	79	Mawkmai Secondary	1909-11	1.575	2.35
30	Kathiawar Minor Lon.	1853	1.451	1.34	80	Upper Irrawaddy ...	1909-11	0.596	0.49
31	Sabarmati Secondary	1853-54	1.348	2.84	81	Jaintia Hills Sec. ...	1910-11	0.986	1.86
32	Great Indus ...	1853-61	0.359	0.43	82	Bhir Secondary ...	1911-12	0.794	0.94
33	Rahin Meridional ...	1853-63	0.327	0.37	83	Ranchi Secondary ...	1911-12	1.840	2.34
34	Assam Longitudinal ...	1854-60	0.579	0.71	84	Villupuram Secondary	1911-12	1.184	1.78
35	Cutch Coast ...	1855-58	0.956	1.27	85	Sambalpur Meridional	1911-14	0.250	0.21
36	Kashmir Principal ...	1855-60	0.884	0.86	86	Sambalpur Lon. ...	1856-57	0.306	0.37
37	Jag-Tila Meridional	1855-63	0.481	0.59	87	Indo-Russian Connection	1912-13	2.790	3.92
38	Sambalpur Lon. ...	1856-57	0.306	0.37	88	Khandwa Secondary	1912-13	0.999	1.27
39	(Cutch) Coast Line Sec.	1856-60	0.975	1.47	89	Ashta Secondary ...	1913-15	1.048	1.33
40	Kathiawar Minor	1858-59	0.930	1.51	90	Bhilai Secondary ...	1913-14	0.304	0.43
41	Kathiawar Minor Meridional No. 2	1859-60	1.247	1.75	91	Nalding Secondary ...	1913-14	1.465	1.83
42	Kathiawar Minor Meridional No. 3	1859-60	0.969	1.48	92	Naga Hills Secondary	1913-14	0.913	0.96
43	Bidar Longitudinal	1859-72	0.311	0.30	93	Middle Godavari Sec. ...	1914-15	0.913	1.08
44	Eastern Frontier or Shillong Meridional	1860-64	0.409	0.49	94	Kohima Secondary ...	1914-15	1.094	1.39
45	Sutlej Meridional ...	1861-63	0.346	0.53	95	Cachar	1914-15	1.077	1.65
46	Madras Mer. and Coast	1861-68	0.426	0.40	96	Bombay Island	1911-14		
47	Kathiawar Minor Meridional No. 4	1863-64	1.154	1.73	97	Madura	1916-17	1.148	1.53
48	East Calcutta Lon. ...	1863-69	0.379	0.57	98	Bagalgot	1916-17	0.701	0.83

Mer. = Meridional, Lon. = Longitudinal, Sec. = Secondary.

- REFERENCES
- G. T. Triangulation
  - Latitude Stations
  - Azimuth Stations
  - Pendulum Stations
  - Standard Bench Marks
  - Lines of precise Levelling
  - Telegraphic Longitude Area

NOTE.—Operations of 1916-17 are shown by thick lines.





**CHART II**  
**N. W. Quadrilateral**  
 Scale 1"=100 Miles

**SYMBOLS**

- Base ..... X
- Longitude ..... O
- Closing Point ..... A

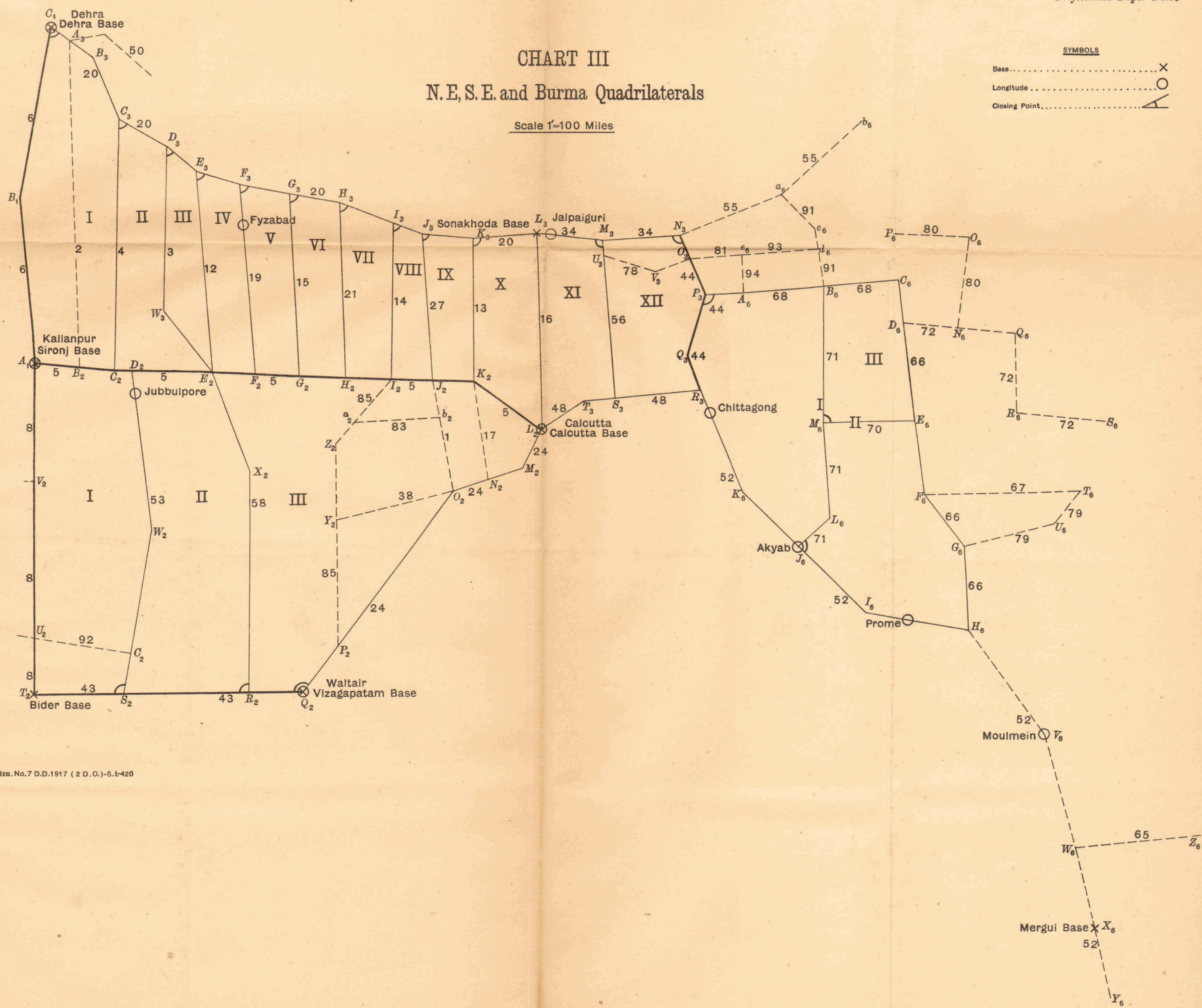


# CHART III N. E. S. E. and Burma Quadrilaterals

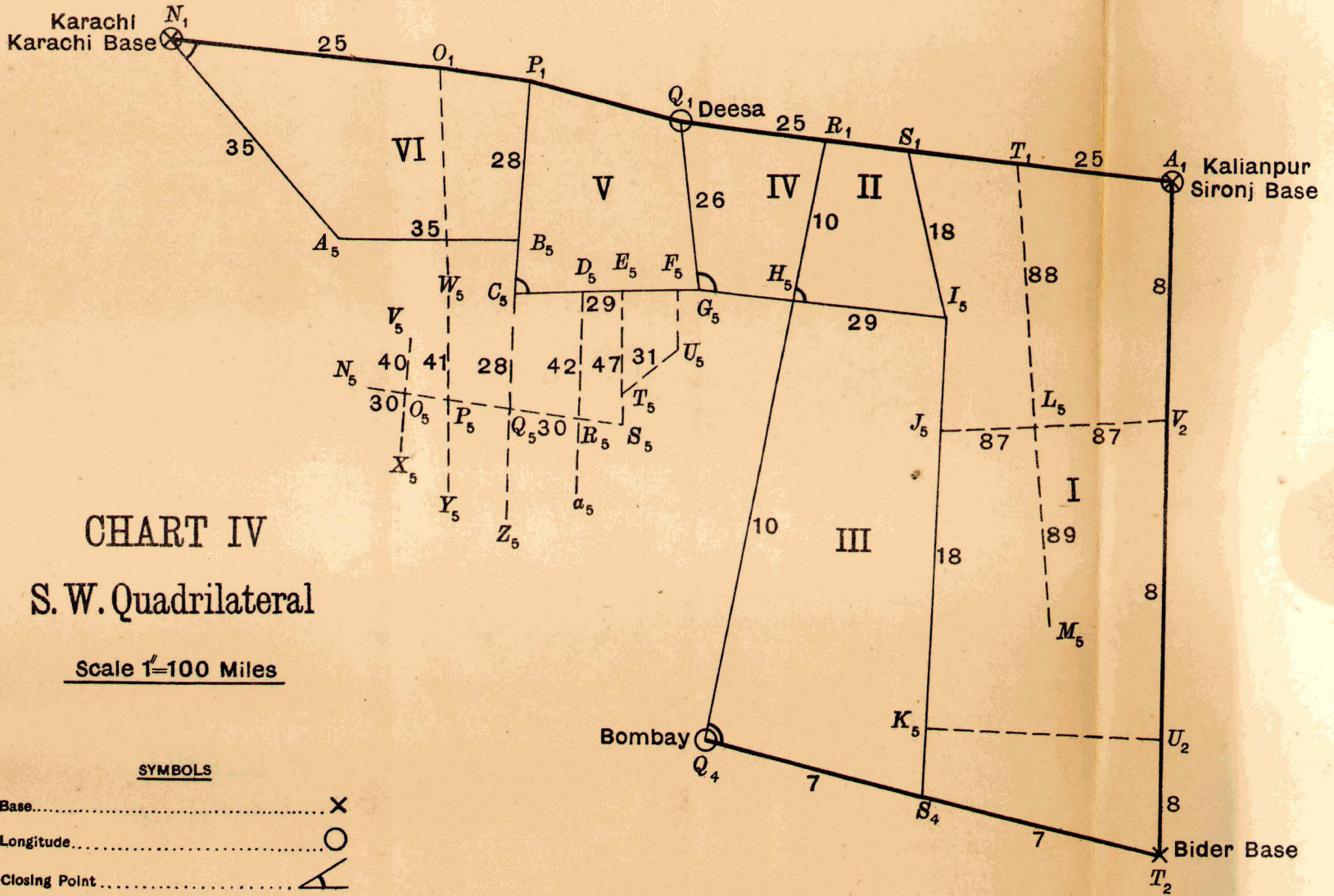
Scale 1"=100 Miles

### SYMBOLS

- Base.....X
- Longitude.....O
- Closing Point...../







# CHART IV

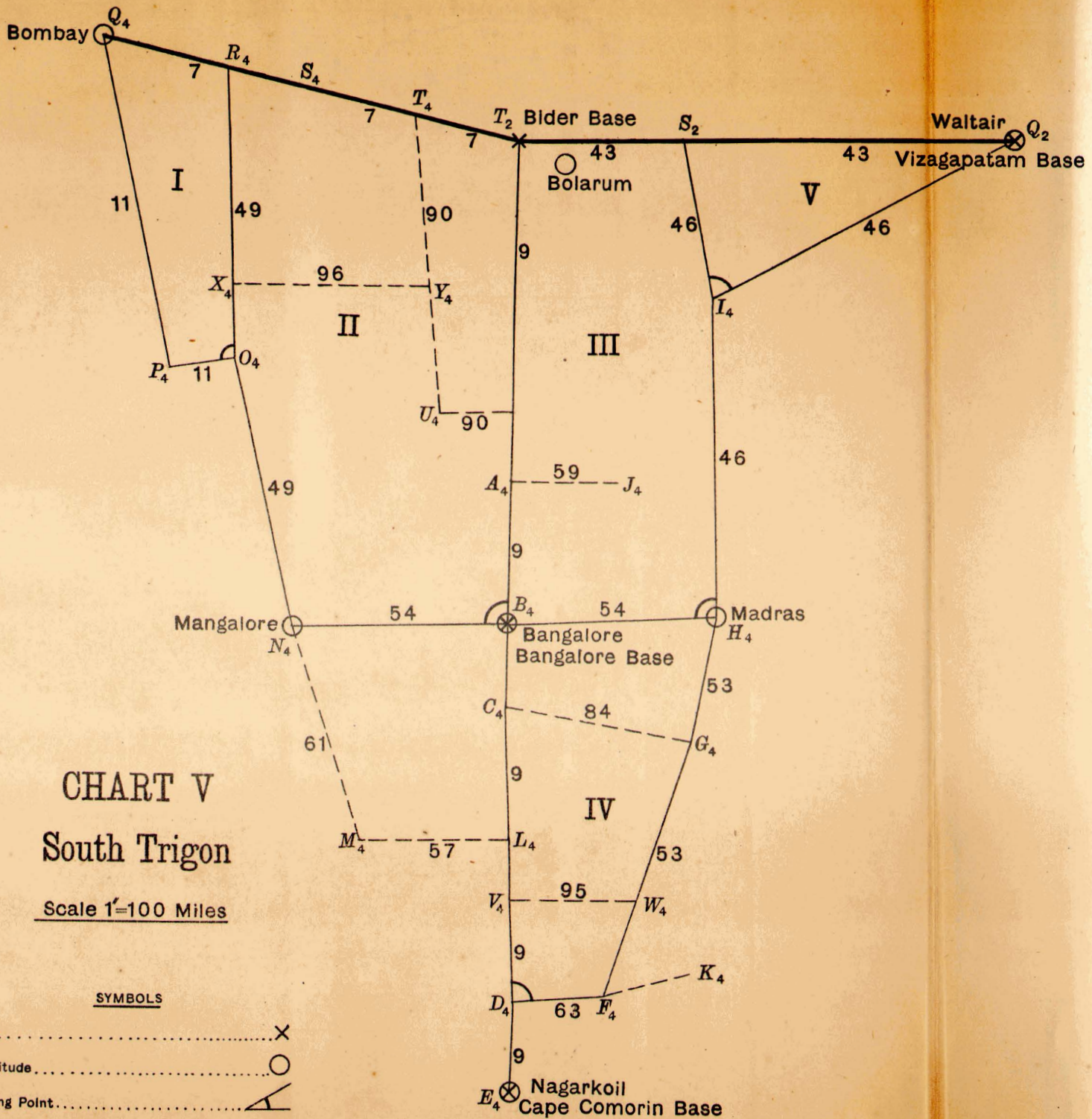
## S. W. Quadrilateral

Scale 1"=100 Miles

### SYMBOLS

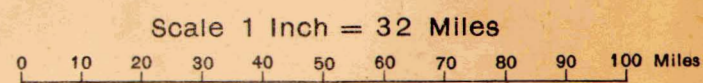
- Base..... X
- Longitude..... O
- Closing Point..... A







Gravity and Deflection Stations in Turkistan (Ferghana)



Astronomical stations are shown thus.....●  
with the number of the station and the deflections  
in meridian and prime vertical.

e.g.....2● + 22.5  
+ 3.7

The figures in red give the deflection in meridian

Gravity stations are shown thus.....● with the  
number of the station, the height in metres and  
the value of (g-y) in units of 10<sup>-3</sup> c.m.

e.g.....6● 4200m  
+ 169



## I N T R O D U C T I O N .

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The preparation of this work has extended over some four years and has been delayed by press of other work resulting mainly from a shortage of officers in the department due to the war. Its final completion has been very hurried as the author has been ordered to Mesopotamia. The end of Chapter VIII has been much abbreviated owing to there not being time to work out a sufficient number of cases to form the basis for a proper discussion. It has been decided however to publish the work at the stage it has reached rather than to wait an indefinite period for its amplification.

The origin of the research was the need which had arisen of converting geodetic results, obtained in India and referred to the now obsolete Everest spheroid, into terms of the best determined spheroid. The work soon showed that certain inconsistencies arose and that multiple values of the changes were found according to the route followed. The reason of this is that the observations of triangulation in India have been *adjusted* to fit the Everest spheroid and will not fit any other without *readjustment*. Had the spheroid of Everest been regarded merely as a reference figure and not, for purposes of triangulation, as identical with the geoid this difficulty would not have arisen: but I believe a similar method of treatment has been followed in all other countries. It would generally be impossible to avoid this treatment, as in the early stages of survey work values of deflections are not usually available. There is little reason why these deflections should not be determined roughly as the triangulation proceeds: and were this done all computations could be made quite correctly, and the deflection results would also be useful. The discrepancies involved, however, are not as large as the probable errors and offer what is more of a computation difficulty than a practical difficulty: and it is believed that the conclusions reached in the first four chapters overcome the difficulty quite satisfactorily for all practical geodetic purposes. The explanation of the discrepancy was by no means discovered at once and it is in Chapter V that the fundamental inconsistency is discussed.

Meanwhile other related subjects come under consideration and in Chapter VI reference is made to some of these. Here a complete analogy between adjustment of errors and small strains in a mechanical framework is shown to exist. This has led to the idea of "strength" of triangulation and gives a less abstract view of the adjustments by the method of least squares than was hitherto available. It enables one to picture in a tangible way the changes necessary, which will also be those most probable. A new method of adjustment of chains of triangulation has been developed, and applied to a number of Indian series.

In this connection a quantity  $M$  has been introduced as a criterion of the strength of triangulation. It is based on General Ferrero's quantity " $m$ ", but also takes cognisance of the length of sides and general formation of a series of triangulation. This quantity has been taken out numerically for all the Indian series.

The quantity  $M$  permits of probable errors not only of side and azimuth but also of latitude and longitude being expressed at any point of the triangulation. Application has been made to all the circuits of India and the closing errors actually met with are found in good accordance with the theoretical probable errors based on  $M$ .

The case of probable errors after adjustment has also been considered. This is much more troublesome and involves very heavy work in the form of solution of numerous equations. The value of the enquiry, however, will be considerable, as an answer can be found to questions such as,

how often should extra base lines or Laplace stations be introduced? It appears that Laplace stations should be just as numerous as extra bases: and that the observation of extra bases alone is only a half measure not likely to improve matters much. It may be roughly compared to closing a traverse circuit for northing and omitting to do so for easting. The improvement due to adjustment is but briefly considered owing to force of circumstances; but the necessary equations have been solved for the cases of the N.W. Quadrilateral of the Indian triangulation and this question will be resumed when it is possible to do so. As has been mentioned the process involves the solution of a large number of simultaneous linear equations; and this has to be done for a large number of values of the absolute term in these equations. It is believed that several novelties have been introduced in this connection which may be of general interest. This question is treated on pages 126-153.

In Chapter IX the results of all the deflection observations are expressed in terms of Helmert's spheroid; the azimuth observations all having been adjusted on the Laplace conditions, none of which had been made use of during the simultaneous adjustment of the triangulation. The quantities given also permit of easy reference to any other spheroid which may at any future time be adopted. Certain observations in Russian Turkistan have also been added, as by means of the recent Indo-Russian connection these can now be stated in terms of the Indian survey. It has also been considered convenient to give a tabular statement of all determinations of  $g$ , so as to make a complete statement of gravity, not only its direction, but also its intensity. Full details of the pendulum operations are to be found in Professional Papers Nos. 10, 15.

These quantities are of immediate interest when the question of the form of the geoid and the underlying reasons for that form are considered. A start had been made with this question, which I had hoped to include in this work, but which must now be held over. An approximate method of finding the underground density anomalies by means of Poisson's equations seems possible. This is independent of the usual isostatic hypotheses, and may throw light on the whole question of isostasy.

For convenience of reference a certain number of various determinations of the figure of the earth have been included.

The computations incidental to the preparation of this work have been very heavy. Those of the earlier chapters I to IV have been performed mostly by Babu Mukundananda Acharya and Babu Hem Chandra Banerji, B.A. The solution of equations in Chapter VIII has been entirely carried out by Babu Diwan Chand Nanda, to whom I am especially indebted for his industry and accuracy in a troublesome and monotonous piece of work. The data in Chapter IX have been compiled by Babu Surendranath Mitra, M.R.A.S.

Mr. Sarat Kumar Mukerji has been responsible for the printing of the whole letter press.

DEHRA DUN, }  
15th Sept. 1917. }

J. DE GRAAFF HUNTER.



## CHAPTER I.

First method of finding the changes of coordinates of triangulated points due to changes in axes of the terrestrial spheroid and coordinates of origin.

1. The subject of the first 16 sections of this chapter was printed in abstracted form in 1912 to draw attention to some of the difficulties of the problem, and with the hope that some light might be thrown on it at the Triennial Geodetic Conference held at Hamburg in that year.

The spheroid on which the triangulation of India has been adjusted is now believed to be considerably in error, as from the nature of the case was inevitable. Owing to possible deflection of the plumb line at the origin of the survey at Kalianpur, the values of latitude and azimuth at that point are somewhat in doubt. The problem for solution was to find the changes in latitude, longitude and azimuth of all triangulated points in India due to changes in the adopted values of the axes of the terrestrial spheroid and in the adopted coordinates of the origin of the Survey.

2. The new spheroid which is adopted in the first 16 sections of this chapter is defined by

$$a = \text{semi major axis} = 6378200 \text{ metres}$$

$$\epsilon = \text{compression or flattening} = \frac{1}{298.3} = \frac{a-b}{a}.$$

These values are given on page 173 of "*The Figure of the Earth and Isostasy, from Measurements in U.S.A.*" Washington 1909, where they are said to be Dr. Helmert's latest values.

Heretofore the axes used in the Survey of India are those due to Everest, known as "Everest's constants, first set". The numerical values are—

$$a = 20,922,931.80 \text{ feet}$$

$$\epsilon = \frac{1}{300.8}$$

All base lines of the Survey of India have been expressed in terms of the Indian ten-foot standard, known as bar *A*. The base lines were not reduced to standard British feet but were

given as some number of times  $\frac{A}{10}$  feet. In making use of Everest's constants we have accordingly been taking the semi major axis as

$$20,922,931 \cdot 80 \frac{A}{10} \text{ British feet.}$$

The value of  $A$  is given\* as  $3 \cdot 333,318,86 Y$ ,  $Y$  being the British standard yard.

We accordingly have  $\frac{A}{10} = 1 - \cdot 000,004,342$  from which it follows that the semi major axis which has actually been used in India is

$$a = 20,922,840 \cdot 95 \text{ British feet.}$$

Similarly

$$b = 20,853,284 \cdot 03 \quad ,, \quad ,,$$

3. Converting 6,378,200 metres into feet by means of the relation

$$1 \text{ metre} = 39 \cdot 370113 \text{ inches}$$

deduced by Benoit (see "*Rapport du Yard au mètre, Paris 1896*") we get as our new semi major axis 20,925,871·23 British feet, and denoting by  $\delta a$  and  $\delta b$  the corrections which have to be applied to the values used in the Survey of India, we have

$$\delta a = +3030 \cdot 28 \text{ feet} = 923 \cdot 63 \text{ metres}$$

and

$$\delta b = +2436 \cdot 78 \text{ feet} = 742 \cdot 73 \text{ metres}$$

Also since  $e^2 = \epsilon(2 - \epsilon)$  and  $\delta \epsilon = \frac{1}{298 \cdot 3} - \frac{1}{300 \cdot 8} = \cdot 000,027,86$  it follows that,

$$\delta e^2 = \cdot 000,055,54 \dagger$$

where  $e$  is the eccentricity.

4. It is now considered that the coordinates of the origin of the survey at Kalianpur require modification. Captain G. P. Lenox Conyngham R. E. observed a group of azimuths and latitudes round Kalianpur. His results gave the mean value reduced to Kalianpur

$$\text{Latitude } 24^\circ 7' 11'' \cdot 57 \qquad \text{Azimuth of Surantal } 190^\circ 27' 6'' \cdot 39.$$

The values heretofore adopted in the triangulation are,

$$\text{Latitude } 24^\circ 7' 11'' \cdot 26 \qquad \text{Azimuth of Surantal } 190^\circ 27' 5'' \cdot 10.$$

We have to apply corrections to the origin of  $+0'' \cdot 31$  in latitude and  $+1'' \cdot 29$  in azimuth. As regards the old value of azimuth a correction of  $-1'' \cdot 1$  was applied to the observed value by General Walker in order to make azimuths observed at other parts of the triangulation agree with geodetic values. We are now annulling this by reverting to an observed value of azimuth.

5. Accordingly it is necessary to investigate equations giving the change in coordinates due to the changes of both axes of the spheroid and of the latitude of the origin and of the azimuth of a ray through it, as exhibited in the following table

\* Account of the Operations of the G.T. Survey of India Vol. I, p. 28

† This corresponds to  $\epsilon = \frac{1}{300 \cdot 8}$ . Everest's actual value was  $\frac{1}{300 \cdot 8017}$  and the corresponding value of  $\delta e^2$  is  $\cdot 000,055,58$ .

TABLE I.

	Old value	New value
Longitude of Kalianpur ...	... ..	77° 39' 17"·57
Latitude of Kalianpur ...	24° 7' 11"·26	24° 7' 11"·57
Azimuth at Kalianpur of Surantal ...	190° 27' 5"·10	190° 27' 6"·39
Length of semi major axis ...	20,922,840·95 feet	20,925,871·23 feet (= 6,378,200 metres)
Length of semi minor axis ...	20,853,284·03 feet	20,855,720·81 feet (= 6,356,818, metres)
Compression ...	$\frac{1}{300\cdot8_{017}}$	$\frac{1}{298\cdot3}$

The latest value of longitude\* is merely given for convenience of reference. Any change in longitude of origin is of course immediately applicable to the whole of the triangulation by simple addition (or subtraction).

6. The old triangulation was adjusted, that is to say its apparent errors were distributed, by a process following the method of least squares as closely as was thought to be practicable in view of the great number of observed angles involved. Owing to the errors in the chosen values of the axes, the equations which the errors were made to satisfy were not quite correct. In the first place the spherical excesses of the several triangles were computed with uncorrect values of the axes: but, owing to the smallness of these spherical excesses, the change on this account is not appreciable to 0."01—the accuracy to which they were computed. None the less the error on this account being of a systematic kind—always of the same sign—will have had some small effect. With the "circuit equations" the case is less favourable. In following series of triangulation, which embrace much larger areas, the spherical excess becomes much more appreciable, and its value on the new spheroid differs from the old value by about one second in an area of 75 square degrees in Indian latitudes. This difference modifies the circuit equations. It is a smaller error than the errors generated in the triangulation, but is systematic.

The only theoretically accurate course would be to readjust all the triangulation. This would be a very large piece of work, and one object of the present paper is to avoid this labour by putting forward alternative methods, which will give the desired changes, with a departure from strict theoretical accuracy smaller than the errors due to fallible observations. The methods will also be applicable to any further changes that may be found desirable at any subsequent date.

7. The following notation is used

$a$  = semi major axis

$b$  = semi minor axis

$e$  = ellipticity or compression

$e$  = eccentricity

$\rho$  = radius of curvature to meridian =  $a(1 - e^2)(1 - e^2 \sin^2 \lambda)^{-\frac{3}{2}}$

$\nu$  = normal terminated by the minor axis =  $a(1 - e^2 \sin^2 \lambda)^{-\frac{1}{2}}$

which is the other principal radius of curvature

\* Account of the Operations of the G.T. Survey of India Vol. XVII p. xv.

- $\beta^2 = \nu/\rho$
- $\lambda =$  latitude
- $L =$  longitude
- $A =$  azimuth, measured from South by West
- $u =$  change in latitude due to change of origin and axes
- $v =$  „ longitude „ „ „
- $w =$  „ azimuth „ „ „
- $c =$  distance between points whose coordinates are  $\lambda, L$  and  $\lambda + \Delta \lambda, L + \Delta L$ .

As only very small values of  $c$  will be considered it is unnecessary to specify whether this distance is measured along a normal plain section or a geodesic line.

8. For small values of  $c$

$$\left. \begin{aligned} \Delta \lambda &= -\frac{c}{\rho} \cos A \\ \Delta L &= -\frac{c}{\nu} \frac{\sin A}{\cos \lambda} \\ \Delta A &= -\frac{c}{\nu} \sin A \tan \lambda \end{aligned} \right\} \dots \dots \dots (1)$$

Differentiating\* these equations with respect to  $A, \lambda, \rho, \nu$  the corresponding changes of  $\Delta \lambda, \Delta L, \Delta A$  are obtained: and remembering that  $\delta \Delta \lambda = \delta u$  and  $\delta \lambda = u$  etc.. we obtain

$$\left. \begin{aligned} \delta u &= \frac{c}{\rho} \cos A \cdot \frac{\delta \rho}{\rho} + \frac{c}{\rho} \sin A \cdot w \\ \delta v &= \frac{c}{\nu} \frac{\sin A}{\cos \lambda} \cdot \frac{\delta \nu}{\nu} - \frac{c}{\nu} \frac{\cos A}{\cos \lambda} \cdot w - \frac{c}{\nu} \frac{\sin A}{\cos^2 \lambda} \sin \lambda \cdot u \\ \delta w &= \frac{c}{\nu} \sin A \tan \lambda \cdot \frac{\delta \nu}{\nu} - \frac{c}{\nu} \cos A \tan \lambda \cdot w - \frac{c}{\nu} \sin A \sec^2 \lambda \cdot u \end{aligned} \right\} \dots \dots \dots (2)$$

These are three simultaneous partial equations from which  $u, v, w$  are to be determined. They express the small changes in  $u, v, w$  developed along a short (elementary) line in direction of azimuth  $A$ . Before they can be integrated it is necessary to define the route along which to travel. It might be supposed at first that the only important matter was the terminal points of the route: but it will be seen later that a different result is found from each route followed. The equations are not integrable in finite terms for all routes, and two special cases are now considered, firstly along a parallel of latitude and secondly along a meridian. These cases correspond to  $A = 90^\circ$  and  $A = 0$  respectively. A means of dealing with the general case of an oblique curvilinear ray is given later, §13 *et seq.*

9. *Case I, when  $A = 90^\circ$ .* In the case of a route along a parallel of latitude it is clear that  $\frac{c}{\nu \cos \lambda} = dL$  and equations (2) can accordingly be written

$$\left. \begin{aligned} -\frac{du}{dL} &= \frac{\nu}{\rho} \cos \lambda \cdot w \\ -\frac{dv}{dL} &= \frac{\delta \nu}{\nu} - \tan \lambda \cdot u \\ -\frac{dw}{dL} &= \sin \lambda \cdot \frac{\delta \nu}{\nu} - \sec \lambda \cdot u \end{aligned} \right\} \dots \dots \dots (3)$$

\* The quantities  $\frac{d\rho}{d\lambda}, \frac{d\nu}{d\lambda}$  were neglected as they contain the factor  $c^2$ .

Putting  $\beta^2 = \frac{\nu}{\rho}$  it follows at once from (3) that

$$\begin{aligned} \frac{d^2 u^2}{dL^2} &= -\frac{\nu}{\rho} \cos \lambda \cdot \frac{d\nu}{dL} \\ &= \beta^2 \sin \lambda \cos \lambda \cdot \frac{\delta \nu}{\nu} - \beta^2 u \end{aligned}$$

or  $\frac{d^2 u}{dL^2} + \beta^2 u = \frac{1}{2} \beta^2 \sin 2\lambda \frac{\delta \nu}{\nu}$  . . . . . (4)

The solution of this is

$$u = \frac{1}{2} \sin 2\lambda \frac{\delta \nu}{\nu} + P \cos (\beta L) + Q \sin (\beta L) \quad \dots \dots \dots (5)$$

where  $P$  and  $Q$  are constants. Using (3) and differentiating (5) it follows that

$$\begin{aligned} -\beta^2 \cos \lambda \cdot w &= \frac{du}{dL} = \beta \left\{ -P \sin (\beta L) + Q \cos (\beta L) \right\} \\ w &= \frac{1}{\beta \cos \lambda} \left\{ P \sin (\beta L) - Q \cos (\beta L) \right\} \quad \dots \dots \dots (6) \end{aligned}$$

To determine  $P$  and  $Q$  put  $L = 0$  in (5) and (6)

$$\left. \begin{aligned} u_0 &= \frac{1}{2} \sin 2\lambda \frac{\delta \nu}{\nu} + P \\ w_0 &= -\frac{Q}{\beta \cos \lambda} \end{aligned} \right\} \dots \dots \dots (7)$$

where the suffix zero indicates values at the beginning of the line.

Further, using (3) and (5)

$$-\frac{d\nu}{dL} = \frac{\delta \nu}{\nu} - \tan \lambda \left\{ \frac{1}{2} \sin 2\lambda \cdot \frac{\delta \nu}{\nu} + P \cos (\beta L) + Q \sin (\beta L) \right\}$$

whence

$$v - v_0 = -\frac{\delta \nu}{\nu} \cos^2 \lambda L + \frac{\tan \lambda}{\beta} \left\{ P \sin (\beta L) + Q (1 - \cos (\beta L)) \right\} \quad \dots \dots \dots (8)$$

longitude being measured from the starting point.

Expressing these equations in terms of seconds—they are at present in radian units—we write

$$\left. \begin{aligned} R'' &= \frac{1}{2} \sin 2\lambda \cdot \frac{\delta \nu}{\nu} \operatorname{cosec} 1'' \\ P'' &= u''_0 - R'' \\ Q'' &= -\beta w''_0 \cos \lambda \end{aligned} \right\} \dots \dots \dots (9)$$

and

$$\left. \begin{aligned} u'' &= R'' + P'' \cos (\beta L) + Q'' \sin (\beta L) \\ v'' &= v''_0 - R'' \cot \lambda \cdot \frac{L^\circ}{57.3} + \frac{\tan \lambda}{\beta} \left\{ P'' \sin (\beta L) + Q'' (1 - \cos (\beta L)) \right\} \\ w'' &= \frac{1}{\beta \cos \lambda} \left\{ P'' \sin (\beta L) - Q'' \cos (\beta L) \right\} \end{aligned} \right\} \quad \dots \dots (10)$$

Since  $\nu = a(1 - e^2 \sin^2 \lambda)^{-\frac{1}{2}}$  it follows from logarithmic differentiation that

$$\frac{\delta \nu}{\nu} = \frac{\delta a}{a} + \frac{\sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \frac{\delta e^2}{2}$$

Putting in the values of  $a, e, \delta a, \delta e^2$  from §§3, 4 and expanding

$$\frac{\delta \nu}{\nu} = \cdot 000, 144, 83 + \cdot 000, 027, 77 \sin^2 \lambda + \cdot 000, 000, 18 \sin^4 \lambda + \dots \quad (11)$$

10. *Case, II, when  $A = 0$ .* In the case of a route along a meridian it is clear that  $\frac{c}{\rho} = -d\lambda$  and equation (2) can accordingly be written

$$\left. \begin{aligned} \frac{du}{d\lambda} &= -\frac{\delta \rho}{\rho} \\ \frac{dv}{d\lambda} &= \frac{\rho}{\nu} \sec \lambda \cdot w \\ \frac{dw}{d\lambda} &= \frac{\rho}{\nu} \tan \lambda \cdot w \end{aligned} \right\} \dots \dots \dots (12)$$

Differentiating  $\rho = a(1 - e^2)(1 - e^2 \sin^2 \lambda)^{-\frac{3}{2}}$  logarithmically

$$\frac{\delta \rho}{\rho} = \frac{\delta a}{a} - \delta e^2 \left( \frac{1}{1 - e^2} - \frac{3}{2} \cdot \frac{\sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \right)$$

whence, expanding and putting in numerical values

$$\begin{aligned} \frac{\delta \rho}{\rho} &= \cdot 000, 144, 83 - \cdot 000, 055, 54 \left( 1 \cdot 00668 - \frac{3}{2} \sin^2 \lambda - \frac{3}{2} e^2 \sin^4 \lambda \dots \dots \dots \right) \\ &= \cdot 000, 088, 92 + \cdot 000, 083, 31 \sin^2 \lambda + \cdot 000, 000, 55 \sin^4 \lambda \\ &= \cdot 000, 130, 78 - \cdot 000, 041, 93 \cos 2\lambda + \cdot 000, 000, 07 \cos 4\lambda \dots \dots \dots (13) \end{aligned}$$

Integrating the first equation of (12)

$$\begin{aligned} u - u_0 &= - \int \frac{\delta \rho}{\rho} \cdot d\lambda \\ &= - \cdot 000, 130, 78 (\lambda - \lambda_0) + \left[ \cdot 000, 020, 97 \sin 2\lambda - \cdot 000, 000, 02 \sin 4\lambda \dots \right]_{\lambda_0}^{\lambda} \quad (14) \end{aligned}$$

From last equation of (12)

$$\frac{1}{w} \cdot \frac{dw}{d\lambda} = \frac{\rho}{\nu} \tan \lambda = \frac{1 - e^2}{1 - e^2 \sin^2 \lambda} \cdot \tan \lambda$$

Put  $y = \sin^2 \lambda$  and  $dy = 2 \sin \lambda \cos \lambda d\lambda$

Then

$$\begin{aligned} d \log w &= \frac{1 - e^2}{1 - e^2 y} \tan \lambda \cdot \frac{dy}{2 \sin \lambda \cos \lambda} \\ &= \frac{1}{2} \cdot \frac{1 - e^2}{1 - e^2 y} \cdot \frac{dy}{1 - y} \\ &= \frac{1}{2} \left( \frac{1}{1 - y} - \frac{e^2}{1 - e^2 y} \right) dy \end{aligned}$$

Integrating

$$\begin{aligned} \log w &= -\frac{1}{2} \log (1 - y) + \frac{1}{2} (1 - e^2 y) + \text{constant} \\ w &= K \sqrt{\frac{1 - e^2 y}{1 - y}} \quad \text{where } K \text{ is a constant} \end{aligned}$$

$$w = \frac{K \sqrt{1 - e^2 \sin^2 \lambda}}{\cos \lambda} = \frac{K a}{\nu \cos \lambda}$$

$$\therefore w = w_0 \cdot \frac{\nu_0 \cos \lambda_0}{\nu \cos \lambda} \dots \dots \dots (15)$$

Again  $\frac{dv}{d\lambda} = \frac{1 - e^2}{1 - e^2 \sin^2 \lambda} \sec \lambda \cdot K \frac{\sqrt{1 - e^2 \sin^2 \lambda}}{\cos \lambda}$   
 $\therefore v = K (1 - e^2) \int \frac{d\lambda}{\cos^2 \lambda \sqrt{1 - e^2 \sin^2 \lambda}}$

Putting  $x = \sin \lambda$  then  $dx = \cos \lambda d\lambda$  and  
 $v = K (1 - e^2) \int \frac{dx}{(1 - x^2)^{\frac{3}{2}} \sqrt{1 - e^2 x^2}}$

Now  $\frac{1}{(1 - x^2)^{\frac{3}{2}} \sqrt{1 - e^2 x^2}} = \frac{1}{(1 - x^2)^{\frac{3}{2}} \sqrt{1 - e^2 + e^2 (1 - x^2)}}$   
 $= \frac{1}{\sqrt{1 - e^2}} \cdot \frac{1}{(1 - x^2)^{\frac{3}{2}}} \left\{ 1 - \frac{1}{2} k (1 - x^2) + \frac{3}{8} k^2 (1 - x^2)^2 \dots \right\} \dots \dots (16)$

where  $k = \frac{e^2}{1 - e^2} = 0.006,682,2$ .

$$\therefore v = K \sqrt{1 - e^2} \int \left\{ \frac{dx}{(1 - x^2)^{\frac{3}{2}}} - \frac{1}{2} k \cdot \frac{1}{\sqrt{1 - x^2}} + \frac{3}{8} k^2 x^2 \sqrt{1 - x^2} \dots \dots \right\} dx$$

$$= K \sqrt{1 - e^2} \left[ \frac{x}{\sqrt{1 - x^2}} - \frac{1}{2} k \sin^{-1} x + \frac{3}{8} k^2 \left( \frac{1}{2} x \sqrt{1 - x^2} + \frac{1}{2} \sin^{-1} x \right) \dots \right]$$

$$= K \sqrt{1 - e^2} \left[ \tan \lambda - \frac{1}{2} k \lambda + \frac{3}{16} k^2 (\sin \lambda \cos \lambda + \lambda) \dots \dots \dots \right]$$

Collecting results and expressing them in terms of seconds we get

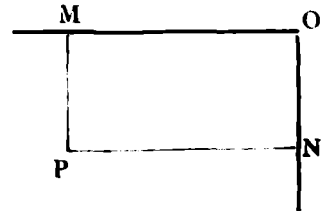
$$\left. \begin{aligned} u'' - u_0'' &= -000,130,78 (\lambda'' - \lambda_0'') + 4.325 (\sin 2\lambda - \sin 2\lambda_0) - 0035 (\sin 4\lambda - \sin 4\lambda_0) \\ v'' - v_0'' &= w_0'' \cdot \frac{\nu_0 \cos \lambda_0}{a} \sqrt{1 - e^2} \left[ \tan \lambda - 003,332,7 \lambda'' \sin 1'' + 000,004,2 \sin 2 \lambda \right]_{\lambda_0}^{\lambda} \\ w'' &= w_0'' \cdot \frac{\nu_0 \cos \lambda_0}{\nu \cos \lambda} \dots \dots \dots \end{aligned} \right\} (17)$$

The value of  $\log \sqrt{1 - e^2}$  is 1.9985538.

11. With the equations (10) and (17) we can now deduce the values of  $u, v, w$  for any point P. Starting from the origin O, we may compute along the parallel OM and find the values at M. Using these as initial values we can then proceed along the meridian MP and get values for P.

Or we may first proceed along meridian ON and then along the parallel NP.

The values arrived at by the two routes are not identical. This is inevitable. The discrepancy in azimuth is the change in spherical excess on the given area from the old to the new spheroid. We shall proceed to consider the discrepancies which occur. To study these



discrepancies the values of the changes were computed to five places of decimals. In the first place it was considered convenient to take as origin for this computation the point whose latitude and longitude were  $24^\circ$ ,  $78^\circ$  on the old spheroid. The double values of  $u$ ,  $v$ ,  $w$ , for this point, differ by a small amount and in view of what follows the following mean value of  $w$  was taken:—

$$w = \frac{\frac{w_x}{x} + \frac{w_y}{y}}{\frac{1}{x} + \frac{1}{y}} = \frac{yw_x + xw_y}{x+y}$$

The suffix  $x$  indicates the value found by route ONP and the suffix  $y$  " " " " " " " " OMP and  $x = \text{PN}$ ,  $y = \text{PM}$ , but  $x$  and  $y$  are always treated as positive. The values of  $u_x$ ,  $u_y$  and  $v_x$ ,  $v_y$  were practically identical.

Starting from this origin by means of our equations the values exhibited in the following three tables are obtained:—

TABLE II.

LATITUDE ( $u''$ ).

Lat.	Long.	$58^\circ$	$63^\circ$	$68^\circ$	$73^\circ$	$78^\circ$	$83^\circ$	$88^\circ$	$93^\circ$	$98^\circ$
$34^\circ$	$u_x$	-2.07723	-2.65069	-3.09166	-3.39675	-3.56361				
	$u_y$	-2.50858	-2.89432	-3.20019	-3.42385					
	$u_x - u_y$	+0.43135	+0.24363	+0.10853	+0.02710		0.00000			
$29^\circ$	$u_x$	-0.27462	-0.75738	-1.13306	-1.39881	-1.55258				
	$u_y$	-0.49755	-0.88329	-1.18916	-1.41282					
	$u_x - u_y$	+0.22293	+0.12591	+0.05610	+0.01401		0.00000			
$24^\circ$	$u_x$	+1.40241	+1.01667	+0.71080	+0.48714	+0.34738	+0.29261	+0.32323	+0.43903	+0.63911
	$u_y$									
	$u_x - u_y$									
$19^\circ$	$u_x$	+2.97072	+2.68699	+2.45452	+2.27510	+2.15009	+2.08045	+2.06672	+2.10900	+2.20697
	$u_y$	+3.20512	+2.81938	+2.51351	+2.28985		+2.09532	+2.12594	+2.24174	+2.44182
	$u_x - u_y$	-0.23440	-0.13239	-0.05899	-0.01475		-0.00000	-0.01487	-0.05922	-0.13274
$14^\circ$	$u_x$	+4.44968	+4.27179	+4.11549	+3.98199	+3.87299	+3.78724	+3.72751	+3.69353	+3.68557
	$u_y$	+4.92732	+4.54158	+4.23571	+4.01205		+3.81752	+3.84814	+3.96394	+4.16402
	$u_x - u_y$	-0.47764	-0.26979	-0.12022	-0.03006		-0.00000	-0.03028	-0.12063	-0.27041



TABLE III.  
LONGITUDE ( $v$ ).

		58°	63°	68°	73°	78°	83°	88°	93°	98°
31°	$v_x$	+11.82474	+8.92142	+5.98521	+3.03302	+0.06397				
	$v_y$	+11.55396	+8.70715	+5.84003	+2.95734					
	$v_x - v_y$	+0.27078	+0.21427	+0.14818	+0.07568	0.00000				
29°	$v_x$	+11.03498	+8.28493	+5.51411	+2.72776	-0.06874				
	$v_y$	+10.97521	+8.23535	+5.47878	+2.70942					
	$v_x - v_y$	+0.05977	+0.04958	+0.03533	+0.01834	0.00000				
24°	$v_x$	+10.45012	+7.80729	+5.15103	+2.18448	-0.18914	-2.86654	-5.54411	-8.21943	-10.88833
	$v_y$									
	$v_x - v_y$	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000
19°	$v_x$	+10.03969	+7.46470	+4.88195	+2.29303	-0.30049	-2.89693	-5.49462	-8.09189	-10.68705
	$v_y$	+9.96451	+7.41142	+4.84792	+2.27645		-2.88036	-5.46060	-8.03863	-10.61192
	$v_x - v_y$	+0.07518	+0.05328	+0.03403	+0.01658	0.00000	-0.01657	-0.03403	-0.05326	-0.07513
14°	$v_x$	+9.78031	+7.23867	+4.69380	+2.14557	-0.10531	-2.95831	-5.51288	-8.06847	-10.62452
	$v_y$	+9.50738	+7.03876	+4.56259	+2.08063		-2.89337	-5.38170	-7.86843	-10.35171
	$v_x - v_y$	+0.27293	+0.20011	+0.13121	+0.06494	+0.00000	-0.06494	-0.13118	-0.20004	-0.27281

TABLE IV.

AZIMUTH ( $w$ ).

		58°	63°	68°	73°	78°	83°	88°	93°	98°
34°	$w_x$	+8.78306	+6.98775	+5.13898	+3.25090	+1.33803				
	$w_y$	+5.83531	+4.75699	+3.64230	+2.49973					
	$w_x - w_y$	+2.94775	+2.23076	+1.49668	+0.75117	0.00000				
29°	$w_x$	+6.97539	+5.60208	+4.18593	+2.73775	+1.26862				
	$w_y$	+5.53262	+4.51023	+3.45336	+2.37006					
	$w_x - w_y$	+1.44277	+1.09185	+0.73257	+0.36769	0.00000				
24°	$w_x$	+5.29810	+4.31905	+3.30698	+2.26960	+1.21485	+0.15080	-0.91440	-1.97260	-3.01574
	$w_y$									
	$w_x - w_y$	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000
19°	$w_x$	+3.71867	+3.11339	+2.48428	+1.83619	+1.17400	+0.50282	-0.17217	-0.84588	-1.51311
	$w_y$	+5.11995	+4.17383	+3.19579	+2.19329		+0.14573	-0.88365	-1.90627	-2.91433
	$w_x - w_y$	-1.40128	-1.06044	-0.71151	-0.35710	0.00000	+0.35709	+0.71148	+1.06039	+1.40122
14°	$w_x$	+2.20928	+1.96355	+1.70277	+1.42897	+1.14420	+0.85068	+0.55065	+0.24640	-0.05974
	$w_y$	+4.99000	+4.06789	+3.11467	+2.13762		+0.14203	-0.86122	-1.85789	-2.84036
	$w_x - w_y$	-2.78072	-2.10434	-1.41190	-0.70865	0.00000	+0.70865	+1.41187	+2.10429	+2.78063

12. Denote the distances PN, PM (in any linear unit, not in angular units) by  $x$  and  $y$  then.

$$x = (L - L_0) \nu \cos \lambda$$

$$y = \int_{\lambda_0}^{\lambda} \rho d\lambda.$$

By inspection of the numbers shown in tables II, III, IV the following equations are found to be approximately true :

$$\left. \begin{aligned} u_x - u_y &= Ax^2y \\ v_x - v_y &= Bxy^2 \\ w_x - w_y &= Cxy \end{aligned} \right\} \dots \dots \dots (18)$$

where  $A, B, C$  are quantities varying slightly with the latitude, but which may be treated as constants with their mean values over any area with which we shall need to deal. The last equation simply expresses that the closing error in azimuth is equal to the change in spherical excess.

Now  $u_x - u_y$  is what we will call the "closing error in latitude" in proceeding round the circuit OMPN;  $v_x - v_y$  and  $w_x - w_y$  being corresponding quantities for longitude and azimuth. Over any elementary area

$$\left. \begin{aligned} dU &= d(u_x - u_y) = 2Ax dx dy \\ dV &= d(v_x - v_y) = 2By dx dy \\ dW &= d(w_x - w_y) = C dx dy \end{aligned} \right\} \dots \dots \dots (19)$$

By integrating over any area the closing error of the circuit enclosing that area is found.

To find the values of  $u, v, w$  then which would be obtained by proceeding along any route it is only necessary to find the values of  $u_x, v_x, w_x$  (or  $u_y, v_y, w_y$ ) and apply the closing error with the correct sign. Integrating (18) it follows for moderate areas,

$$\left. \begin{aligned} U &= 2A \bar{x}a \\ V &= 2B \bar{y}a \\ W &= Ca \end{aligned} \right\} \dots \dots \dots (20)$$

where  $a$  is the area of the circuit and  $\bar{x}, \bar{y}$  are the coordinates of its centre of gravity. We say for moderate areas because the coordinates  $x$  and  $y$  are curvilinear: but for the areas we shall require to apply the formulæ to,  $x$  and  $y$  may be treated as rectilinear coordinates.

13. By means of the above equations it is possible to find the result of the change of axes and origin as computed along any line of any curvature or along any route whatever, by computing first along a parallel and then along a meridian (or in the reverse order) and then applying the "closing errors" of the circuit formed by the line in question and the parallel and meridian

This, then, would solve the problem as far as solitary lines were concerned. When we come to a network of lines the case is different, for several values of the changes which occur at a point can be found corresponding to the several possible routes by which the point can be reached. In view of the fact that most of the triangulation of India is along meridian or parallel (see triangulation chart at end), the following procedure is suggested:—

- (1). Select central meridian and parallel for India (Burma will be dealt with separately). The selected meridian is  $78^\circ$  and the selected parallel  $24^\circ N$ .

(2). Assume the values of  $u, v, w$  found by the formulæ on these lines, which we will call axes, to be correct. We have then to distribute the closing errors in PM and PN. (see fig. §11).

(3). If PM is a meridional series the computations fixing the length PM depend only in a small measure on the size of the earth's axes. The way in which these axes have come in is through the spherical excess. In nearly all triangulation in the Survey of India, the spherical excess is such a small quantity that the change of axes proposed will not appreciably affect it (to  $0''\cdot01$ ). There is reason then for assuming that the length PM is correct. In the same way the length PN may be regarded as correct. If then  $u_y$  and  $v_x$  are taken for the changes in coordinates of P there should be no error to the first order: and as the values of  $u, v$  are so small the second order quantities may surely be neglected.

(4). This process would hold for the corners of circuits formed by meridian and longitudinal series, though some modification would be more correct for oblique series. In the Indian triangulation meridional and longitudinal series are the rule. Oblique series occur practically only along the coast of the Bay of Bengal and along the first range of the Himalayas. (See index chart of triangulation at end). As far as latitude and longitude are concerned we should not be committing much error in accepting the values of latitude and longitude,  $u_y$  and  $v_x$ .

(5). Now consider the azimuth change. This can be found from the change in position of two contiguous points. If we take two points originally on the same latitude whose changes are  $u_y$  and  $u_y + \frac{\delta u_y}{\delta L} dL$  the azimuth change on the line joining them is

$$\rho \frac{\delta u_y}{\delta L} dL \frac{1}{\nu \cos \lambda \cdot dL} = \frac{\rho}{\nu \cos \lambda} \cdot \frac{\delta u_y}{\delta L} = w_x$$

whereas the azimuth change deduced from two points originally on the same longitude is

$$\frac{\nu \cos \lambda}{\rho} \frac{\delta v_x}{\delta \lambda} = v_x$$

It has been seen that the azimuth closing errors is  $C \cdot x \cdot y$  where  $OM = x$   $ON = y$ . and  $C$  is a quantity which varies slightly with the latitude. Treating  $C$  as a constant and equal to its mean value over the area in consideration is permissible. This will be satisfied if the azimuth error is put into the lines PN and PM to amount proportional to their lengths.

This gives

$$\frac{\frac{w_x}{y} + \frac{v_x}{x}}{\frac{1}{y} + \frac{1}{x}} = \frac{xw_x + yv_x}{x + y}$$

as the best correction to the azimuth at P.

(6). The difference  $w_y - w_x$  is not to be regarded as an error contained in the value of the angle  $M P N$ . Its effects is to alter the curvature of the lines  $P M$  and  $P N$ .

(7). In the case then of a gridiron system of meridional and longitudinal series at regular intervals all of equal weight, it seems that the best values we could assign to the changes are  $u_y, v_x, \frac{xw_y + yw_x}{x + y}$ . In this case the next step in correcting the triangulation would be to find the changes of intermediate points on the series as follows:—

$$\frac{u_y}{c_x} \text{ for change in latitude along a meridian} \\ \text{longitude along a parallel}$$

and for the other coordinate and azimuth to simply interpolate between the terminal values.

(8). A difficulty arises when the actual points of the triangulation series are considered. For if the above formula for two adjacent points  $A$  and  $B$  is used, the difference of coordinates will not exactly give the correct change of azimuth. Adopting the rule of computing the azimuth from the coordinates, a different azimuth change on a ray going east from that found on a ray going south is arrived at. That is to say the actual change of  $w_x - w_y$  would be forced into the single angle formed by these rays.

To avoid this it appears better to take the azimuth change to be  $\frac{xw_y + yw_x}{x + y}$  and the change in coordinates of *one point only* to be given by  $u_y, v_x$ , and compute the coordinates of second adjacent point from this with the corrected azimuth (*i. e.* old azimuth  $+ \frac{xw_y + yw_x}{x + y}$ ) and the old value of the distance  $c$ .

The computation alluded to in (8) may be performed with tables such as are given in the "Auxiliary Tables"\* prepared for the new values of the axes: or we may at once deduce the changes in the position of the second point  $B$  by differentiation of the equations for  $\Delta\lambda, \Delta L$  and  $\Delta A$ .

The equations are:—

$$\Delta\lambda = -\frac{c}{\rho} \cos A - \frac{1}{2} \frac{c^2}{\rho\nu} \sin^2 A \tan \lambda = \delta_1\lambda + \delta_2\lambda$$

$$\Delta L = -\frac{c}{\nu} \frac{\sin A}{\cos \lambda} + \frac{1}{2} \frac{c^2}{\nu^2} \frac{\sin 2A \tan \lambda}{\cos \lambda} = \delta_1 L + \delta_2 L$$

$$\Delta A = -\frac{c}{\nu} \sin A \tan \lambda + \frac{1}{4} \frac{c^2}{\nu^2} (1 + 2 \tan^2 \lambda) = \delta_1 A + \delta_2 A$$

being equations (1) carried to an extra term in consideration of the larger value of  $c$  now contemplated.

Differentiating we have at once

$$\delta\Delta\lambda = \delta_1\lambda \left( -\frac{\delta\rho}{\rho} - \tan A.w \right) + \delta_2\lambda \left( -\frac{\delta\rho}{\rho} - \frac{\delta\nu}{\nu} + 2 \cot A.w + \frac{2}{\sin 2\lambda} u \right)$$

$$\delta\Delta L = \delta_1 L \left( -\frac{\delta\nu}{\nu} + \cot A.w + \tan \lambda u \right) + \delta_2 L \left\{ -\frac{2\delta\nu}{\nu} + 2 \cot 2A.w + (\cot \lambda + 2 \tan \lambda) u \right\}$$

\* Auxiliary Tables of the Survey of India, Dehra Dün, 1906.

$$\delta\Delta A = \delta_1 A \left( -\frac{\delta v}{v} + \cot A \cdot w + \frac{2}{\sin 2\lambda} u \right) + \delta_2 A \left( -\frac{2\delta v}{v} + \frac{4 \tan \lambda \sec^2 \lambda}{1 + 2 \tan^2 \lambda} u + 2 \cot 2A \cdot w \right)$$

In above  $u, v, w$  are the values found for one end of the base  $A$ : the values for the other end  $B$  are then  $u + \delta\Delta\lambda, v + \delta\Delta L, w + \delta\Delta A$ .

A third method is to reach  $B$  by proceeding first along the parallel  $AC$  through  $A$  and then down the meridian  $CB$  through  $B$ , by means of the formulæ (or tables) already given: and then by applying the closing error of the area  $ACB$ .

It appears, then, that the expressions  $u_y, v_x, \frac{xw_y + yw_x}{x + y}$  may be taken to represent the changes in latitude, longitude and azimuth respectively of any point in India (excluding Burma) with the restriction that adjacent points must be treated differently, the changes for the second point being deduced by one of the three methods just explained. On this basis the results may be given in convenient tabular form. They will represent the changes with accuracy considerably greater than the accuracy with which the points can be considered to be fixed in space by triangulation.

14. These values are believed to be satisfactory for all the purposes for which they can be used. As far as map producing goes the discrepancies are negligible. For geodetic purposes we require to know the absolute corrections to latitudes, longitudes and azimuths of a base where a junction is to be made with another survey—such as the Russian survey, or the Burma survey. We can do this as described in § 13 for one end of the base and then compute the coordinates of the other end of the base from a knowledge of its length. In the case of Burma the triangulation has not yet been adjusted. It will perhaps be adjusted with the new values of the axes and made to fit on to the most eastern series of the North-East Quadrilateral, *viz.*, the Shillong Meridional Series, after this has been corrected for change of axes.

We also wish to know corrections to triangulated latitudes or azimuths at stations where these quantities have also been observed astronomically, so as to know the actual plumb-line deflections. As regards latitude we have uncertainty of perhaps  $0''\cdot 1$  on account of axes change after leaving the central latitude by  $10^\circ$ , *i. e.* one part in 360,000 which is of the order of accuracy of our base-lines in India. The error generated in the triangulation must eventually be greater than this. The same argument holds as regards the azimuth, where the uncertainty of change due to change of axes, and due to error generated in triangulation are necessarily larger numbers when expressed in seconds of arc than occur in the latitude. The astronomic observations for azimuth are less precise, considered from point of view of plumb-line deflection, than the latitude observations. Apart from these considerations an error in plumb-line deflection in latitude of  $0''\cdot 1$  is of little account. In India we have plumb-line deflections of over  $50''$  and, at least at present, tenths of second are too minute to be taken account of in any discussion of deflections.

15. It seems then that the method sketched above is sufficiently precise for the geodetic uses to which the results can be put, and higher accuracy could not be applied with advantage to the results of triangulation. The method of § 13 is applicable to points which can be reached by either route (meridian or parallel) without the route departing out of the region of triangulation. Thus while it applies to all the triangulation in India which has been adjusted, it could not be fairly applied without modification to Burma, for this would imply the existence of triangulation across the Bay of Bengal. As the Burma triangulation remains to be adjusted, this does not matter and it will only be necessary to apply the method as far as the Shillong Meridional Series, which can be done very satisfactorily, the more so as our selected central latitude crosses this series.

16. The Survey of India was asked in 1912 by the Siamese Survey Department to furnish the best possible values of the coordinates of Bangkok. The way in which this has been done will serve as a good illustration of the method of using the closing error to determine the changes which occur along a route which is neither meridional nor longitudinal. As far as longitude  $90^\circ$  the route may be taken to follow the central parallel, latitude  $24^\circ$  (see triangulation chart at end). From there it proceeds along the Burma Coast Series down to latitude  $13^\circ 45'$ , and thence to Bangkok along latitude  $13^\circ 45'$ . In this case then we first compute along parallel  $24^\circ$  up to longitude  $98^\circ$ : we then proceed along meridian  $98^\circ$  down to latitude  $13^\circ 45'$ . The result at this point is found from tables II, III, IV by extrapolation to be

$$u_y = + 4.248$$

$$v_y = - 10.339$$

$$w_y = - 2.837$$

Now treating longitude  $98^\circ$  as axis from which  $x$  is measured, we evaluate the closing errors over the area between the Coast series and latitude  $24^\circ$  and meridian  $98^\circ$  and get

$$\Sigma Ax^2y = U = +.017$$

$$\Sigma Bxy^2 = V = -.026$$

$$\Sigma Cxy = W = +.426$$

Hence the changes at latitude  $13^\circ 45'$ , longitude  $98^\circ$ , as determined by the route following the Burma Coast Series, are  $u_y + U$ ,  $v_y + V$ ,  $w_y + W$ .

One further correction remains. The Bangkok Series which emanates from this point is expressed in "preliminary terms"—it was computed from preliminary values of the side from which it emanates. Later values of this side, found after the Coast Series had been computed from the preliminary value of the side, require the following changes to be applied to the beginning of the Bangkok Series, viz.

in	latitude	...	$-1''\cdot80$
	longitude	...	$-0''\cdot17$
	azimuth	...	$+5''\cdot50$

Combining these we arrive at the changes to be made at latitude  $13^\circ 45'$ , longitude  $98^\circ$ .

	...	...	Latitude	Longitude	Azimuth
By parallel and meridian route	...	...	$+4\cdot248$	$-10\cdot339$	$-2\cdot837$
Correction to bring into terms of Burma Coast Series route	...	...	$+0\cdot017$	$-0\cdot026$	$+0\cdot426$
Correction from "preliminary terms"	...	...	$-1\cdot80$	$-0\cdot17$	$+5\cdot50$
		Total	$+2\cdot47$	$-10\cdot54$	$+3\cdot09$

With these initial values by computing along parallel  $13^\circ 45'$  up to longitude  $100^\circ 33' 3''\cdot5$  the old value of the longitude of Phukhao Thong Station\* in Bangkok we get the changes

$$u = + 2''\cdot34$$

$$v = - 11''\cdot82$$

$$w = + 2''\cdot9$$

To bring into Greenwich terms the farther correction  $- 2' 27''\cdot18$  is required to the longitude, the final corrections being

$$+ 2''\cdot34 \text{ in latitude}$$

$$- 2' 39''\cdot00 \text{ in longitude}$$

$$+ 2''\cdot9 \text{ in azimuth}$$

\* Phukhao Thong Station is the most easterly triangulated point shown on the triangulation chart.

Owing to an unfortunate confusion of the quantities  $u_y, v_y, w_y$  with  $u_x, v_x, w_x$  the following corrections were wrongly supplied to the Royal Survey Department Siam in 1912

$$\begin{aligned} &+ 1'' \cdot 73 && \text{in latitude} \\ &- 2' 39'' \cdot 32 && \text{longitude} \\ &+ 5'' \cdot 7 && \text{azimuth} \end{aligned}$$

17. So far the particular case of definite numerical values of  $\frac{\delta a}{a}, \delta e^2, u_0, w_0$  has been considered. It is desirable to put the solution in a form in which the results of any desired change can be calculated rapidly.

Let  $u_1, u_2, u_3, u_4$ , be the changes in latitude respectively due to  $\delta a = 1000$  metres,  $\delta b = 1000$  metres,  $u_0 = 1''$ ,  $w_0 = 1''$  with corresponding notation for  $v$  and  $w$ . Since the quantities involved are small it is clear that

$$u = \delta a \cdot u_1 + \delta b \cdot u_2 + u_0 \cdot u_3 + w_0 u_4 \dots \dots \dots (18)$$

with similar equations for  $v$  and  $w$ : where  $da, db$  are expressed in kilometers and  $u_0, w_0$  in seconds. If values of  $u_1, u_2, u_3, u_4$  are tabulated, equation (18) will enable the quantities  $u, v, w$  to be evaluated for any desired case. Of the quantities  $\delta a, \delta b, \delta e^2$  any two may be regarded as independent, the third being determined from the result of differentiating  $b = a \sqrt{1-e^2}$  logarithmically.

i. e. 
$$\frac{\delta a}{a} - \frac{\delta b}{b} = \frac{1}{2} \frac{\delta e^2}{1-e^2} \dots \dots \dots (19)$$

When  $\delta a = \delta b = 1000$  metres = 39370·113 inches = 3280·843 feet

$$\frac{1}{2} \frac{\delta a}{a} \operatorname{cosec} 1'' = 16 \cdot 1718 \quad \text{of which the logarithm is } 1 \cdot 2087596$$

$$\frac{1}{2} \frac{\delta b}{b} \operatorname{cosec} 1'' = 16 \cdot 2258 \quad \text{,, ,, ,, } 1 \cdot 2102058$$

Also

$$\begin{aligned} \frac{\delta v}{v} &= \frac{\delta a}{a} + \frac{\sin^2 \lambda}{1-e^2 \sin^2 \lambda} \cdot \frac{\delta e^2}{2} \\ &= \frac{\delta a}{a} \left( 1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) - \frac{\delta b}{b} \cdot \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \dots \dots \dots \text{ by (19)} \end{aligned}$$

Hence equation (9) may be written, omitting the dashes

$$\left. \begin{aligned} R &= \sin 2\lambda \left\{ 16 \cdot 1718 \left( 1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) \delta a - 16 \cdot 2258 \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \delta b \right\} \\ P &= u_0 - R \\ Q &= w_0 \beta \cos \lambda \end{aligned} \right\} \dots \dots \dots (20)$$

where  $\delta a, \delta b$  are expressed in kilometres and  $u_0$  and  $w_0$  in seconds.

18. Along a parallel of latitude

Case I, when  $\delta a = 1, \delta b = 0, u_0 = 0, w_0 = 0$

By (20) 
$$\left. \begin{aligned} R &= 16 \cdot 1718 \sin 2\lambda \left( 1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) = -P \\ Q &= 0 \end{aligned} \right\} \dots \dots \dots (21)$$

and (10) becomes

$$\left. \begin{aligned} u &= R (1 - \cos (\beta L)) \\ v - v_0 &= -R \left( \frac{L^0 \cot \lambda}{57 \cdot 3} + \frac{1}{\beta} \tan \lambda \sin (\beta L) \right) \\ w &= -\frac{R}{\beta} \sec \lambda \sin (\beta L) \end{aligned} \right\} \dots (22)$$

where  $L$  is expressed in degrees.

Case II, when  $\delta a = 0, \delta b = 1, u_0 = 0, w_0 = 0$

From (20)

$$R = -16 \cdot 2258 \sin 2\lambda \cdot \frac{(1 - e^2) \sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \dots (23)$$

and equations (22) hold for this case also.

Case III, when  $\delta a = 0, \delta b = 0, u_0 = 1, w_0 = 0$

In this case  $\frac{\delta v}{v} = 0$  so that

$$\left. \begin{aligned} R &= 0 \\ v' &= u_0 = 1 \\ Q &= 0 \end{aligned} \right\} \dots (24)$$

and

$$\left. \begin{aligned} u &= \cos (\beta L) \\ v - v_0 &= \frac{1}{\beta} \tan \lambda \sin (\beta L) \\ w &= \frac{1}{\beta} \sec \lambda \sin (\beta L) \end{aligned} \right\} \dots (25)$$

Case IV, when  $\delta a = 0, \delta b = 0, u_0 = 0, w_0 = 1$

$$\left. \begin{aligned} R &= P = 0 \\ Q &= -\beta \cos \lambda \end{aligned} \right\} \dots (26)$$

and

$$\left. \begin{aligned} u &= -\beta \cos \lambda \sin (\beta L) \\ v - v_0 &= -\sin \lambda (1 - \cos (\beta L)) \\ w &= \cos (\beta L) \end{aligned} \right\} \dots (27)$$

19. From (10) and (19)

$$\begin{aligned} \frac{\delta \rho}{\rho} &= \frac{\delta a}{a} - 2 \left( \frac{\delta a}{a} - \frac{\delta b}{b} \right) \left( 1 - \frac{3}{2} \frac{(1 - e^2) \sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \right) \\ &= \frac{\delta a}{a} - 2 \left( \frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 1 - \frac{3}{2} (1 - e^2) \sin^2 \lambda (1 + e^2 \sin^2 \lambda + \dots) \right\} \\ &= \frac{\delta a}{a} - 2 \left( \frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 1 - \frac{3}{2} (1 - e^2) \left( \frac{1 - \cos 2\lambda}{2} + \frac{e^2}{8} (3 - 4 \cos 2\lambda + \cos 4\lambda) \dots \right) \right\} \end{aligned}$$

Hence from (12)

$$\begin{aligned} u - u_0 &= - \int_{\lambda_0}^{\lambda} \frac{\delta \rho}{\rho} d\lambda \\ &= -\frac{\delta a}{a} (\lambda - \lambda_0) + 2 \left( \frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 0 \cdot 2513 (\lambda - \lambda_0) + 0 \cdot 3750 (\sin 2\lambda - \sin 2\lambda_0) - 0 \cdot 0003 (\sin 4\lambda - \sin 4\lambda_0) \right\} \end{aligned}$$



Expressing in seconds and introducing  $\delta a$ ,  $\delta b$  expressed in kilometres and  $\lambda, \lambda_0$  in degrees this becomes

$$u - u_0 = \delta a \left\{ -0.2807 (\lambda^\circ - \lambda_0^\circ) + 24.26 (\sin 2\lambda + \sin 2\lambda_0) - 0.019 (\sin 4\lambda - \sin 4\lambda_0) \right\} \\ + \delta b \left\{ -0.2847 (\lambda^\circ - \lambda_0^\circ) - 24.34 (\sin 2\lambda - \sin 2\lambda_0) + 0.019 (\sin 4\lambda - \sin 4\lambda_0) \right\} \dots (28)$$

20. *Along a meridian*

Case I, when  $\delta a = 1, \delta b = 0, u_0 = 0, w_0 = 0$

By (28) and (17)

$$\left. \begin{aligned} u &= -0.2807 (\lambda^\circ - \lambda_0^\circ) + 24.26 (\sin 2\lambda - \sin 2\lambda_0) - 0.02 (\sin 4\lambda - \sin 4\lambda_0) \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \dots (29)$$

Case II, when  $\delta a = 0, \delta b = 1, u_0 = 0, w_0 = 0$

By (28) and (17)

$$\left. \begin{aligned} u &= -0.2847 (\lambda^\circ - \lambda_0^\circ) - 24.34 (\sin 2\lambda - \sin 2\lambda_0) + 0.02 (\sin 4\lambda - \sin 4\lambda_0) \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \dots (30)$$

Case III, when  $\delta a = 0, \delta b = 0, u_0 = 1, w_0 = 0$

Then

$$\left. \begin{aligned} u &= u_0 = 1 \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \dots \dots \dots (31)$$

Case IV, when  $\delta a = 0, \delta b = 0, u_0 = 0, w_0 = 1$

Then

$$\left. \begin{aligned} u - u_0 &= 0 \\ v - v_0 &= \frac{v_0 \cos \lambda_0}{a} \cdot \sqrt{1 - e^2} \left[ \tan \lambda - 0.000,058,2\lambda^c + 0.000,004,2 \sin 2\lambda \right]_{\lambda_0}^{\lambda} \\ w &= \frac{v_0 \cos \lambda_0}{v \cos \lambda} \dots \dots \dots \end{aligned} \right\} \dots (32)$$

21. With the equations of §18, 20 changes can be computed at any point for any case. There are two possible routes. As an example consider Case I. We can first proceed along a parallel to the appropriate longitude. In proceeding thence along a meridian we have to apply Cases I, III and IV (because the values given by (22) now become initial values), and so find the values  $u_y, v_y, w_y$ : secondly we can proceed along a meridian and afterwards along a parallel and so find the second set of values  $u_x, v_x, w_x$ .

These computations have been made and the results for every degree square corner, so far as concerns the Indian triangulation, are exhibited in the following tables.







Case I.— $\delta\alpha = 1$  km. Values of  $v_x - v_y$  in seconds. TABLE VIII.

Lat.	P o s i t i v e										N e g a t i v e																						
	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°				
36°	0.128	0.118	0.107	0.097	0.086	0.075	0.064	0.053	0.042	0.030	0.019	0.008	0.004	0.015	0.027	0.038	0.050	0.062	0.074	0.086	0.098	0.110	0.122	0.134	0.146	0.158	0.170	0.182	0.194	0.206			
35	0.104	0.095	0.087	0.078	0.070	0.061	0.052	0.043	0.034	0.025	0.015	0.005	0.003	0.012	0.022	0.031	0.040	0.049	0.058	0.067	0.076	0.085	0.094	0.103	0.112	0.121	0.130	0.139	0.148	0.157	0.166		
34	0.082	0.076	0.069	0.062	0.056	0.049	0.041	0.034	0.027	0.020	0.012	0.005	0.003	0.010	0.017	0.025	0.032	0.039	0.046	0.053	0.060	0.067	0.074	0.081	0.088	0.095	0.102	0.109	0.116	0.123	0.130		
33	0.064	0.059	0.054	0.049	0.043	0.038	0.032	0.027	0.021	0.015	0.010	0.004	0.002	0.008	0.014	0.020	0.026	0.032	0.038	0.044	0.050	0.056	0.062	0.068	0.074	0.080	0.086	0.092	0.098	0.104	0.110		
32	0.049	0.045	0.041	0.037	0.033	0.029	0.025	0.020	0.016	0.012	0.007	0.003	0.001	0.006	0.010	0.015	0.019	0.023	0.027	0.031	0.035	0.039	0.043	0.047	0.051	0.055	0.059	0.063	0.067	0.071	0.075		
31	0.035	0.033	0.030	0.027	0.024	0.021	0.018	0.015	0.012	0.009	0.005	0.002	0.001	0.004	0.008	0.011	0.014	0.017	0.020	0.023	0.026	0.029	0.032	0.035	0.038	0.041	0.044	0.047	0.050	0.053	0.056		
30	0.025	0.023	0.021	0.019	0.017	0.015	0.013	0.010	0.008	0.006	0.004	0.002	0.001	0.003	0.005	0.007	0.009	0.011	0.013	0.015	0.017	0.019	0.021	0.023	0.025	0.027	0.029	0.031	0.033	0.035	0.037		
29	0.016	0.015	0.014	0.013	0.011	0.010	0.008	0.007	0.005	0.004	0.002	0.001	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011	0.012	0.013	0.014	0.015	0.016	0.017	0.018	0.019		
28	0.010	0.009	0.008	0.008	0.007	0.005	0.005	0.004	0.003	0.002	0.001	0.001	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011	0.012	0.013	0.014	0.015	0.016	0.017	0.018		
27	0.005	0.005	0.004	0.004	0.004	0.003	0.003	0.002	0.002	0.001	0.001	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011	0.012	0.013	0.014	0.015	0.016	0.017	0.018		
26	0.002	0.002	0.002	0.002	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001		
25	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
24	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
23	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
22	0.004	0.002	0.003	0.002	0.002	0.002	0.002	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
21	0.007	0.005	0.005	0.005	0.004	0.004	0.003	0.003	0.002	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
20	0.011	0.010	0.009	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000		
19																																	
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Values of  $w_z$  in seconds.

Case I. —  $\delta\alpha = 1$  km.

TABLE IX.

Long. Lat.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
	P o s i t i v e												N e g a t i v e																	
36°	4.750	4.347	3.943	3.538	3.132	2.725	2.317	1.903	1.499	1.089	0.679	0.269	0.142	0.552	0.982	1.372	1.781													
35	4.587	4.130	3.792	3.402	3.012	2.620	2.238	1.835	1.441	1.047	0.653	0.253	0.136	0.531	0.925	1.319	1.713													
34	4.394	4.021	3.648	3.273	2.877	2.521	2.143	1.765	1.387	1.008	0.628	0.249	0.131	0.511	0.820	1.209	1.618													
33	4.229	3.871	3.511	3.150	2.789	2.438	2.083	1.699	1.335	0.970	0.605	0.230	0.128	0.492	0.857	1.222	1.598													
32	4.073	3.727	3.381	3.034	2.685	2.336	1.986	1.636	1.285	0.934	0.592	0.230	0.121	0.473	0.825	1.176	1.527	1.878	2.228	2.577	2.926	3.274	3.620	3.966	4.310					
31	3.925	3.692	3.258	2.923	2.588	2.251	1.914	1.577	1.239	0.800	0.561	0.222	0.117	0.466	0.795	1.131	1.472	1.809	2.148	2.483	2.819	3.154	3.488	3.821	4.153					
30	3.784	3.464	3.142	2.819	2.485	2.171	1.846	1.520	1.104	0.868	0.541	0.214	0.113	0.440	0.767	1.093	1.419	1.745	2.070	2.395	2.719	3.042	3.364	3.685	4.005					
29	3.653	3.342	3.032	2.720	2.408	2.095	1.781	1.467	1.152	0.837	0.522	0.207	0.100	0.425	0.740	1.055	1.370	1.684	1.998	2.311	2.624	2.936	3.246	3.556	3.865					
28	3.628	3.229	2.929	2.628	2.326	2.024	1.721	1.417	1.113	0.819	0.504	0.200	0.105	0.410	0.715	1.019	1.323	1.627	1.930	2.232	2.534	2.835	3.136	3.435	3.733	4.030	4.326	4.621	4.914	
27	3.410	3.121	2.831	2.540	2.249	1.957	1.654	1.370	1.076	0.782	0.488	0.193	0.102	0.307	0.601	0.985	1.279	1.573	1.866	2.158	2.450	2.741	3.032	3.321	3.609	3.896	4.182	4.467	4.751	
26	3.301	3.021	2.740	2.450	2.170	1.883	1.610	1.326	1.042	0.757	0.472	0.187	0.000	0.394	0.680	0.954	1.238	1.522	1.806	2.089	2.371	2.653	2.934	3.214	3.493	3.771	4.047	4.323	4.598	
25	3.193	2.927	2.655	2.382	2.109	1.835	1.560	1.235	1.009	0.733	0.457	0.181	0.095	0.372	0.618	0.924	1.200	1.475	1.750	2.024	2.298	2.571	2.843	3.114	3.385	3.654	3.922	4.180	4.455	
24	3.102	2.839	2.576	2.311	2.046	1.780	1.513	1.246	0.970	0.711	0.444	0.176	0.003	0.361	0.628	0.896	1.164	1.431	1.697	1.963	2.229	2.494	2.758	3.021	3.283	3.544	3.805	4.064	4.322	
23	3.013	2.757	2.501	2.244	1.987	1.729	1.470	1.210	0.951	0.691	0.431	0.170	0.000	0.350	0.610	0.870	1.130	1.390	1.640	1.907	2.165	2.422	2.678	2.931	3.189	3.442	3.695	3.946	4.197	
22	2.930	2.682	2.433	2.183	1.932	1.681	1.429	1.177	0.925	0.672	0.419	0.166	0.087	0.341	0.594	0.847	1.099	1.351	1.603	1.855	2.106	2.356	2.605	2.854	3.101	3.348	3.593	3.838	4.082	
21	2.854	2.612	2.370	2.128	1.882	1.637	1.392	1.147	0.901	0.654	0.418	0.161	0.085	0.332	0.578	0.821	1.070	1.316	1.561	1.806	2.050	2.294	2.537	2.779	3.021	3.261	3.500	3.738	3.976	
20	2.784	2.548	2.311	2.074	1.836	1.597	1.358	1.119	0.879	0.638	0.398	0.158	0.082	0.324	0.561	0.804	1.044	1.284	1.523	1.762	2.000	2.238	2.475	2.711	2.946	3.181	3.414	3.646	3.878	
19					1.793	1.560	1.328	1.092	0.853	0.624	0.389	0.154	0.091	0.316	0.551	0.786	1.020	1.254	1.488	1.721	1.954									
18					1.754	1.526	1.298	1.069	0.840	0.610	0.370	0.151	0.079	0.309	0.530	0.759	0.998	1.227	1.456	1.684	1.911									
17					1.719	1.496	1.272	1.047	0.823	0.595	0.373	0.147	0.078	0.303	0.528	0.753	0.978	1.202	1.420	1.650	1.873									
16					1.687	1.468	1.248	1.028	0.807	0.587	0.366	0.145	0.076	0.298	0.519	0.739	0.960	1.180	1.400	1.619	1.838									
15									0.794	0.577	0.359	0.142	0.075	0.292	0.500	0.728	0.943													
14									0.782	0.568	0.354	0.140	0.074	0.288	0.501	0.715	0.929													
13									0.771	0.560	0.349	0.138	0.073	0.284	0.494	0.705	0.914													
12									0.761	0.553	0.345	0.137	0.072	0.280	0.489	0.697	0.905													
11									0.753	0.547	0.341	0.135	0.071	0.277	0.483	0.689	0.895													
10									0.746	0.542	0.338	0.134	0.071	0.275	0.479	0.683	0.887													
9									0.740	0.539	0.335	0.133	0.070	0.272	0.475	0.677	0.880													
8									0.735	0.534	0.333	0.132	0.069	0.271	0.472	0.673	0.874													



























The closing errors.

22. The tables just given exhibit the "closing errors" or differences between  $u_x v_x w_x$  and  $u_y v_y w_y$  respectively. The formulæ for these differences  $u_x - u_y$  etc. for the four cases will now be considered separately and approximate expressions found for them.

Case I, when  $\delta a = 1, \delta b = 0, u_0 = 0, w_0 = 0$

Along the parallel OM the changes at M are given by equations (21) and (22) in which suffix zero may be added to  $\beta, R, \lambda$  to indicate that it applies to latitude  $\lambda_0$ . Apply the changes at M to the case of a meridian: it is necessary to consider cases I, III, IV of § 20 and the following equations are deduced:

$$\left. \begin{aligned} u_y &= \left[ -0.2807\lambda + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} + R_0(1 - (\cos \beta_0 L)) \\ v_y - v_0 &= -R_0 \left( \frac{L^{\circ} \cot \lambda_0}{57.3} + \frac{1}{\beta_0} \tan \lambda_0 \sin(\beta_0 L) \right) \\ &\quad - \frac{R_0 v_0}{\beta_0 a} \sqrt{1 - e^2} \sin(\beta_0 L) \left[ \tan \lambda - 0.000,058, 2\lambda^{\circ} + 0.000,004, 2\sin 2\lambda \right]_{\lambda_0}^{\lambda} \dots (33) \\ w_y &= -\frac{R_0 v_0}{\beta_0 v} \sec \lambda \sin(\beta_0 L) \dots \dots \dots \\ R_0 &= 16.1718 \sin 2\lambda_0 \left( 1 + \frac{\sin^2 \lambda_0 (1 - e^2)}{1 - e^2 \sin^2 \lambda_0} \right) \dots \dots \dots \end{aligned} \right\}$$

Next proceeding along ON, the changes at N are given by (29), and applying these initial values to the parallel NP the following equations are formed by consideration of cases I and III of § 18

$$\left. \begin{aligned} u_x &= R(1 - \cos(\beta L)) + \cos(\beta L) \left[ -0.2807\lambda^{\circ} + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ v_x - v_0 &= -R \left( \frac{L^{\circ} \cot \lambda}{57.3} + \frac{1}{\beta} \tan \lambda \sin(\beta L) \right) \\ &\quad + \frac{1}{\beta} \tan \lambda \sin(\beta L) \left[ -0.2807\lambda^{\circ} + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ w_x &= -\frac{R}{\beta} \sec \lambda \sin(\beta L) + \frac{1}{\beta} \sec \lambda \sin(\beta L) \left[ -0.2807\lambda^{\circ} + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \end{aligned} \right\} (34)$$

From (33) and (34) it follows that

$$\left. \begin{aligned} u_x - u_y &= \left[ R(1 - \cos(\beta L)) \right]_{\lambda_0}^{\lambda} + (1 - \cos(\beta L)) \left[ 0.2807\lambda^{\circ} - 24.26\sin 2\lambda + 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ v_x - v_y &= -\left[ R \left( \frac{L^{\circ} \cot \lambda}{57.3} + \frac{1}{\beta} \tan \lambda \sin(\beta L) \right) \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{1}{\beta} \tan \lambda \sin(\beta L) \left[ -0.280\lambda L^{\circ} + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{R_0 v_0}{\beta_0 a} \sqrt{1 - e^2} \sin(\beta_0 L) \left[ \tan \lambda - 0.000,058, 2\lambda^{\circ} + 0.000,004, 2\sin 2\lambda \right]_{\lambda_0}^{\lambda} \dots (35) \\ w_x - w_y &= -\frac{R}{\beta} \sec \lambda \sin(\beta L) + \frac{1}{\beta} \sec \lambda \sin(\beta L) \left[ -0.2807\lambda^{\circ} + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{R_0 v_0}{\beta_0 v} \sec \lambda \sin(\beta_0 L) \end{aligned} \right\}$$

23. Equations (35) may be simplified and written in approximate form if terms depending on  $e^2$  are neglected. The closing errors will still be expressed with sufficient accuracy. Then  $\beta$  becomes unity and "a" may be substituted for  $\nu$ . Denote  $\lambda - \lambda_0$  by  $\theta$ . In what follows  $u_x$  etc. are expressed in seconds and  $\lambda, L, \theta$  are expressed in radians except when the degree mark is added—thus  $\lambda^\circ$ , and  $\lambda^\circ/\lambda = 57 \cdot 3$ . The successive terms of (35) are taken one by one and reduced.  $\doteq$  denotes approximate equality.

Case I, when  $\delta a = 1, \delta b = 0, u_0 = 0, v_0 = 0$

$$\text{Here } \frac{R}{16 \cdot 17} \doteq \sin 2\lambda (1 + \sin^2 \lambda) = \frac{3}{2} \sin 2\lambda - \frac{1}{4} \sin 4\lambda$$

$$\begin{aligned} u_x - u_y &\doteq (1 - \cos L) \left[ 16 \cdot 17 \left( \frac{3}{2} \sin 2\lambda - \frac{1}{4} \sin 4\lambda \right) + 16 \cdot 08\lambda - 24 \cdot 26 \sin 2\lambda + 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ &= (1 - \cos L) \left[ 16 \cdot 08\lambda - 4 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ &= 16 \cdot 1 (1 - \cos L) \left\{ \theta - \frac{1}{2} \sin 2\theta \cos 2(\lambda + \lambda_0) \right\} \\ &= \cdot 281 \theta^\circ (1 - \cos L) \left\{ 1 - \frac{\sin 2\theta}{2\theta} \cos 2(\lambda + \lambda_0) \right\} \dots \dots \dots (36) \end{aligned}$$

$$\frac{R}{16 \cdot 17} \cot \lambda = 2 \cos^2 \lambda (1 + \sin^2 \lambda) = \frac{5}{4} + \cos 2\lambda - \frac{1}{4} \cos 4\lambda$$

$$\frac{R}{16 \cdot 17} \tan \lambda = 2 \sin^2 \lambda (1 + \sin^2 \lambda) = \frac{7}{4} - 2 \cos 2\lambda + \frac{1}{4} \cos 4\lambda$$

$$\begin{aligned} \left[ -R \left( \frac{L^\circ \cot \lambda}{57 \cdot 3} + \tan \lambda \sin L \right) \right]_{\lambda_0}^\lambda &= -16 \cdot 17 \frac{L^\circ}{57 \cdot 3} \left[ \frac{5}{4} + \cos 2\lambda - \frac{1}{4} \cos 4\lambda + \frac{\sin L}{L} \left( \frac{7}{4} - 2 \cos 2\lambda + \frac{1}{4} \cos 4\lambda \right) \right]_{\lambda_0}^\lambda \\ &= + \cdot 2823 L^\circ \left\{ 2 \left( 1 - 2 \frac{\sin L}{L} \right) \sin \theta \sin (\lambda + \lambda_0) - \frac{1}{2} \left( 1 - \frac{\sin L}{L} \right) \sin 2\theta \sin 2(\lambda + \lambda_0) \right\} \\ &= -0 \cdot 0098 L^\circ \theta^\circ \left\{ \left( 2 \frac{\sin L}{L} - 1 \right) \frac{\sin \theta}{\theta} \sin (\lambda + \lambda_0) + \frac{1}{2} \left( 1 - \frac{\sin L}{L} \right) \frac{\sin 2\theta}{2\theta} \sin 2(\lambda + \lambda_0) \right\} \end{aligned}$$

$$\begin{aligned} \frac{1}{\beta} \tan \lambda \sin (\beta L) \left[ -0 \cdot 2807 \lambda^\circ + 24 \cdot 26 \sin 2\lambda - 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ \doteq \theta^\circ \tan \lambda \sin L \left\{ -0 \cdot 2807 + \cdot 8469 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) - \cdot 0014 \frac{\sin 2\theta}{2\theta} \cos 2(\lambda + \lambda_0) \right\} \\ \doteq -0 \cdot 0049 L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ 1 - 3 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) \right\} \end{aligned}$$

$$\begin{aligned} \frac{R_0 v_0}{\beta_0 a} \sqrt{1 - e^2} \sin (\beta_0 L) \left[ \tan \lambda - 0 \cdot 000,058, 2\lambda^\circ + 0 \cdot 000,004, 2 \sin 2\lambda \right]_{\lambda_0}^\lambda \\ \doteq 16 \cdot 17 \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \sin L \left\{ \frac{\sin \theta}{\cos \lambda \cos \lambda_0} - 0 \cdot 000,058, 2\theta^\circ + 0 \cdot 000,008, 4 \cos (\lambda + \lambda_0) \sin \theta \right\} \\ \doteq + 0 \cdot 0049 L^\circ \theta^\circ \frac{\sin L}{L} \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \left\{ \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0 \cdot 00334 \right\} \end{aligned}$$

Hence

$$\begin{aligned}
 v_x - v_y &= +0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ 2 \left( \frac{L}{\sin L} - 2 \right) \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} + \left( 1 - \frac{L}{\sin L} \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{\theta} \right. \\
 &\quad \left. + \left( -1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right) \tan \lambda \right. \\
 &\quad \left. + \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \left( \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0.00334 \right) \right\} \\
 &= 0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left[ -\tan \lambda - 0.00334 \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \right. \\
 &\quad \left. + \frac{\sin \theta}{\theta} \left\{ 2 \left( \frac{L}{\sin L} - 2 \right) \sin (\lambda + \lambda_0) + 3 \cos (\lambda + \lambda_0) \tan \lambda + 2 \sin \lambda_0 (1 + \sin^2 \lambda_0) \sec \lambda \right\} \right. \\
 &\quad \left. + \left( 1 - \frac{L}{\sin L} \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right]
 \end{aligned}$$

Now  $\frac{\sin \theta}{\theta} \doteq 1$  and  $\frac{L}{\sin L} \doteq 1$  the error being about .01 when  $\theta$  or  $L = 15^\circ$ : hence  $\frac{\sin \theta}{\theta}$  or

$\frac{\sin 2 \theta}{2 \theta}$  may be treated as unity when multiplied by  $\frac{L}{\sin L} - 1$ . It follows that

$$v_x - v_y \doteq +0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \tan \lambda \left( \cos (\lambda + \lambda_0) - 1 \right) - 0.00334 \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \right. \\
 \left. + 2 \sec \lambda \sin^3 \lambda_0 + \left( \frac{L}{\sin L} - 1 \right) \left( 2 \sin (\lambda + \lambda_0) - \sin 2 (\lambda + \lambda_0) \right) \right\} \quad (37)$$

$$\begin{aligned}
 -\frac{R}{\beta} \sec \lambda \sin (\beta L) + \frac{R_0 \nu_0}{\beta_0 \nu} \sec \lambda \sin (\beta_0 L) &\doteq -\sec \lambda \sin L (R - R_0) \\
 &= -0.2823 L^\circ \frac{\sin L}{L} \sec \lambda \left[ \frac{3}{2} \sin 2 \lambda - \frac{1}{4} \sin 4 \lambda \right]_{\lambda_0}^\lambda \\
 &= -0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} - \cos 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right\} \\
 \frac{1}{\beta} \sec \lambda \sin (\beta L) \left[ -0.2807 \lambda^\circ + 24.26 \sin 2 \lambda - 0.02 \sin 4 \lambda \right]_{\lambda_0}^\lambda \\
 &\doteq \sec \lambda \sin L \left\{ -0.2807 \theta^\circ + 48.52 \cos (\lambda + \lambda_0) \sin \theta \right\} \\
 &= -0.0049 L^\circ \theta^\circ \sec \lambda \frac{\sin L}{L} \left\{ 1 - 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
 \end{aligned}$$

Hence

$$v_x - v_y \doteq -0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ 1 - \cos 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right\} \quad \dots \quad (38)$$

24. Case II, when  $\delta a = 0, \delta b = 1, u_0 = 0, w_0 = 0$ . Here

$$\begin{aligned}
 R &= -16.2258 \frac{(1 - e^2) \sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \sin 2 \lambda \\
 &\doteq -8.11 (\sin 2 \lambda - \frac{1}{2} \sin 4 \lambda)
 \end{aligned}$$

Equations (35) hold for this case if we use the above value of  $R$  and change the quantity  $-0.2807 \lambda + 24.26 \sin 2 \lambda - 0.02 \sin 4 \lambda$  into  $-0.2847 \lambda - 24.34 \sin 2 \lambda + 0.02 \sin 4 \lambda$ .

Then

$$\begin{aligned}
 u_x - u_y &\doteq (1 - \cos L) \left[ -8 \cdot 11 (\sin 2\lambda - \frac{1}{2} \sin 4\lambda) + 16 \cdot 3 \lambda + 24 \cdot 26 \sin 2\lambda - 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
 &= (1 - \cos L) \left[ 16 \cdot 3 \lambda + 16 \cdot 1 \sin 2\lambda + 4 \cdot 04 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
 &= 283 \theta^\circ (1 - \cos L) \left\{ 1 + 2 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) + \frac{\sin 2\theta}{2\theta} \cos 2 (\lambda + \lambda_0) \right\} \dots \dots \dots (39)
 \end{aligned}$$

$$\frac{R}{16 \cdot 23} \cot \lambda = -2 \sin^3 \lambda \cos \lambda \cot \lambda = -2 \sin^2 \lambda \cos^2 \lambda = -\frac{1}{4} (1 - \cos 4\lambda)$$

$$\frac{R}{16 \cdot 23} \tan \lambda = -2 \sin^3 \lambda \cos \lambda \tan \lambda = -2 \sin^4 \lambda = -\frac{1}{4} (3 - 4 \cos 2\lambda + \cos 4\lambda)$$

$$\begin{aligned}
 \therefore \left[ -R \left( \frac{L^\circ \cot \lambda}{57 \cdot 3} + \tan \lambda \sin L \right) \right]_{\lambda_0}^{\lambda} &= \frac{16 \cdot 23}{4} \cdot \frac{L^\circ}{57 \cdot 3} \left[ 1 - \cos 4\lambda + \frac{\sin L}{L} (3 - 4 \cos 2\lambda + \cos 4\lambda) \right] \\
 &= 0 \cdot 0049 L^\circ \theta^\circ \left\{ \left( 1 - \frac{\sin L}{L} \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} + 2 \frac{\sin L}{L} \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
 \end{aligned}$$

$$\begin{aligned}
 \frac{1}{\beta} \tan \lambda \sin \beta L &\left[ -0 \cdot 2847 \lambda^\circ - 24 \cdot 34 \sin 2\lambda + 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
 &= \tan \lambda \sin L \theta^\circ \left\{ -0 \cdot 2847 - 0 \cdot 8497 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} + 0 \cdot 0014 \cos 2 (\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\
 &\doteq -L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ +0 \cdot 00497 + 0 \cdot 0148 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} - 0 \cdot 00002 \cos 2 (\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\
 &\doteq -0 \cdot 0050 L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ 1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
 \end{aligned}$$

$$\begin{aligned}
 \frac{R_0 \nu_0}{\beta_0 a} \sqrt{1 - e^2} \sin (\beta_0 L) &\left[ \tan \lambda - 0 \cdot 000,058, 2\lambda^\circ + 0 \cdot 000,004, 2 \sin 2\lambda \right]_{\lambda_0}^{\lambda} \\
 &\doteq -16 \cdot 23 \sin 2\lambda_0 \sin^2 \lambda_0 \sin L \left\{ \frac{\sin \theta}{\cos \lambda \cos \lambda_0} - 0 \cdot 000058 2\theta^\circ + 0 \cdot 0000084 \cos (\lambda + \lambda_0) \sin \theta \right\} \\
 &= -0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \sin 2\lambda_0 \sin^2 \lambda_0 \left\{ \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0 \cdot 00334 \right\}
 \end{aligned}$$

Hence

$$\begin{aligned}
 v_x - v_y &\doteq 0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \left( \frac{L}{\sin L} - 1 \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} + 2 \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right. \\
 &\quad \left. - \left( 1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right) \tan \lambda \right. \\
 &\quad \left. - \sin 2\lambda_0 \sin^2 \lambda_0 \left( -0 \cdot 00334 + \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} \right) \right\} \\
 &\doteq 0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \frac{\sin \theta}{\theta} \left( \sin 2\lambda_0 \cos \lambda_0 \sec \lambda - \cos (\lambda + \lambda_0) \tan \lambda \right) - \tan \lambda \right. \\
 &\quad \left. + \left( \frac{L}{\sin L} - 1 \right) \sin 2 (\lambda + \lambda_0) + 0 \cdot 00334 \sin 2\lambda_0 \sin^2 \lambda_0 \right\}. \quad (40)
 \end{aligned}$$

$$\begin{aligned}
 -\frac{R}{\beta} \sec\lambda \sin(\beta L) + \frac{R_0 \nu_0}{\beta_0 \nu} \sec\lambda \sin(\beta_0 L) &= -\sec\lambda \sin L (R - R_0) \\
 &= +16 \cdot 23 \sec\lambda \sin L \left[ \frac{1}{2} \sin 2\lambda - \frac{1}{4} \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
 &= 0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \sec\lambda \left\{ \cos(\lambda + \lambda_0) \frac{\sin \theta}{\theta} - \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\
 \frac{1}{\beta} \sec\lambda \sin(\beta L) &\left[ -0 \cdot 2847\lambda - 24 \cdot 34 \sin 2\lambda + 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
 &\doteq \sec\lambda \sin L \left\{ -0 \cdot 2847\theta^\circ - 48 \cdot 68 \cos(\lambda + \lambda_0) \sin \theta \right\} \\
 &\doteq -0 \cdot 0050 L^\circ \theta^\circ \sec\lambda \frac{\sin L}{L} \left\{ 1 + 3 \cos(\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
 \end{aligned}$$

Hence

$$w_x - w_y = -0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \sec\lambda \left\{ 1 + 2 \cos(\lambda + \lambda_0) \frac{\sin \theta}{\theta} + \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \dots \dots \dots (41)$$

25. The closing errors for cases III and IV have also been considered and are practically zero. This is at once evident also from the computed values of  $u_x u_y$  etc. which agree to at least 0·001 of a second. It is otherwise clear that there would be no closing error on a sphere caused by moving the origin: and accordingly the effect on a spheroid must vanish with  $e^2$  and accordingly have  $e^2$  as a factor. In considering closing errors then it is only necessary to take cases I and II into account, and this may be done by means of equations (36) to (41). The form of these equations explains how the closing errors found in tables II, III, IV approximately satisfied the empirical relations (18). The relations would not have been equally satisfactory for case I and case II considered independently.

In the case of Indian triangulation  $\theta^\circ$  only exceeds  $8^\circ$  for values of  $L^\circ$  between  $-7^\circ$  and  $-1^\circ$  and is greater than  $-8^\circ$  for values of  $L^\circ$  between  $-5^\circ$  and  $+3^\circ$ : so that we can always consider one of the quantities,  $\theta^\circ$  or  $L^\circ$ , numerically less than  $8^\circ$ . Closing errors for the elementary area  $dL d\lambda$  are now deduced from the equations (36) to (41). In what follows  $L$  is treated as identical with  $\sin L$ .

Putting  $U_1$  for  $u_x - u_y$  in case I,  $U_2$  for  $u_x - u_y$  in case II etc. we have, omitting small terms

$$dU_1 \doteq 16 \cdot 1 \sin L (1 - \cos 4\lambda) dL d\lambda \dots \dots \dots (42)$$

$$dU_2 \doteq 16 \cdot 1 \sin L (1 + 2 \cos 2\lambda + \cos 4\lambda) dL d\lambda \dots \dots \dots (43)$$

$$\begin{aligned}
 dV_1 \doteq &-16 \cdot 17 \cos L \left\{ -2 \sin 2\lambda + \sin 4\lambda + 4 \sin 2\lambda - \sin 4\lambda \right\} dL d\lambda \\
 &+ \cos L \left\{ \sec^2 \lambda (-16 \cdot 08\lambda + 24 \cdot 26 \sin 2\lambda) + \tan \lambda (-16 \cdot 08 + 48 \cdot 52 \cos 2\lambda) \right\} dL d\lambda \\
 &+ 16 \cdot 17 \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \cos L \left\{ \sec^2 \lambda - 0 \cdot 00334 \right\} dL d\lambda \\
 \doteq &16 \cdot 1 \cos L \left\{ \sec^2 \lambda \left( -\lambda + \frac{1}{2} \sin 2\lambda + \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \right) + \tan \lambda (3 \cos 2\lambda - 1) - 2 \sin 2\lambda \right\} dL d\lambda \\
 \doteq &16 \cdot 1 \cos L \left\{ \sec^2 \lambda (0 \cdot 871 - \lambda + \frac{1}{2} \sin 2\lambda) + \sin 2\lambda - 4 \tan \lambda \right\} dL d\lambda. \\
 \doteq &16 \cdot 1 \cos L \left\{ \sec^2 \lambda (0 \cdot 871 - \lambda) + \sin 2\lambda - \tan \lambda \right\} dL d\lambda \dots \dots \dots (44)
 \end{aligned}$$

$$\begin{aligned}
 dV_2 &\doteq 16 \cdot 23 \cos L \left\{ \sin 4\lambda + 2 \sin 2\lambda - \sin 4\lambda \right\} dL d\lambda \\
 &\quad + \cos L \left\{ \sec^2 \lambda (-16 \cdot 3\lambda - 24 \cdot 34 \sin 2\lambda) + \tan \lambda (-16 \cdot 3 - 48 \cdot 68 \cos 2\lambda) \right\} dL d\lambda \\
 &\quad - 16 \cdot 23 \sin 2\lambda_0 \sin^2 \lambda_0 \cos L \sec^2 \lambda dL d\lambda \\
 &\doteq 16 \cdot 2 \cos L \left\{ 2 \sin 2\lambda + \sec^2 \lambda (-\lambda - \frac{3}{2} \sin 2\lambda - \sin 2\lambda_0 \sin^2 \lambda_0) + \tan \lambda (-1 - 3 \cos 2\lambda) \right\} dL d\lambda \\
 &\doteq -16 \cdot 2 \cos L \left\{ \sec^2 \lambda (0 \cdot 125 + \lambda + \frac{3}{2} \sin 2\lambda) + 2 \tan \lambda - \sin 2\lambda \right\} dL d\lambda \quad \dots \dots \dots (45) \\
 &\doteq -16 \cdot 2 \cos L \left\{ \sec^2 \lambda (0 \cdot 125 + \lambda) + 5 \tan \lambda - \sin 2\lambda \right\} dL d\lambda
 \end{aligned}$$

Putting  $\lambda_0 = 24^\circ 7' 11'' \cdot 26$

$$\begin{aligned}
 dW_1 &\doteq -32 \cdot 34 \cos L \cos \lambda (1 + 3 \sin^2 \lambda) dL d\lambda \\
 &\quad + \cos L \left\{ \sin \lambda \sec^2 \lambda (-16 \cdot 1\lambda + 24 \cdot 26 \sin 2\lambda) + \sec \lambda (-16 \cdot 1 + 48 \cdot 5 \cos 2\lambda) \right\} dL d\lambda \\
 &\doteq -16 \cdot 1 \cos L \left\{ 2 \cos \lambda (1 + 3 \sin^2 \lambda) + \sin \lambda \sec^2 \lambda (\lambda - \frac{3}{2} \sin 2\lambda) + \sec \lambda (1 - 3 \cos 2\lambda) \right\} dL d\lambda \\
 &\doteq -16 \cdot 1 \cos L \left\{ \lambda \tan \lambda + 1 + 5 \cos^2 \lambda - 6 \cos^4 \lambda \right\} \sec \lambda dL d\lambda \quad \dots \dots \dots (46)
 \end{aligned}$$

$$\begin{aligned}
 dW_2 &\doteq 16 \cdot 2 \cos L \times 6 \cos \lambda \sin^2 \lambda dL d\lambda \\
 &\quad + \cos L \left\{ \sin \lambda \sec^2 \lambda (-16 \cdot 3\lambda - 24 \cdot 34 \sin 2\lambda) + \sec \lambda (-16 \cdot 3 - 48 \cdot 68 \cos 2\lambda) \right\} dL d\lambda. \\
 &\doteq -16 \cdot 2 \cos L \left\{ -6 \cos \lambda \sin^2 \lambda + \sec \lambda (\lambda \tan \lambda + \frac{3}{2} \tan \lambda \sin 2\lambda + 1 + 3 \cos 2\lambda) \right\} dL d\lambda \\
 &\doteq -16 \cdot 2 \cos L \sec \lambda \left\{ \lambda \tan \lambda + 1 - 3 \cos^2 \lambda + 6 \cos^4 \lambda \right\} dL d\lambda \quad \dots \dots \dots (47)
 \end{aligned}$$

By means of equations (42) to (47) it is possible to find the changes  $u, v, w$  at  $P$  as computed by any route. For

$$u = u_y + \int dU$$

the integration being taken over the area between the desired route (upper limit) and the central parallel and the meridian through  $P$ : and  $u_y$  being the quantity to be found by properly combining the four cases. This obviously does not get rid of the multiple values obtainable for  $u, v, w$  according to the route followed, but if any route has special advantages, results of following it become available. One such route is the geodesic, or the shortest path between any point and the origin. There is something to be said in favour of following this route, and the subject of the geodesic is accordingly considered in some detail in the following chapter, where a direct method of finding the quantities  $u, v, w$  along a geodesic is also made use of.

In concluding this chapter it may be pointed out that the equations (42) to (47) enable the differences of the values of  $u, v, w$  to be rapidly estimated. This makes it clear at once how far it is a matter of importance to strictly adhere to any route that may be selected: for the difference in values that will be found by any two routes is the closing error.

## CHAPTER II.

### Geodesics on a Spheroid.

1. It is now necessary to develop some properties and relations of a geodesic in order that the changes of coordinates due to change of axes may be computed along geodesics. A fundamental relation of a geodesic on a conicoid is\*

$$pD = \text{constant} \dots \dots \dots (1)$$

where  $p$  is the perpendicular from the centre on the tangent plane at a point and  $D$  is the semidiameter of the quadric parallel to the tangent to the curve at the same point. In the case of a spheroid there is symmetry about the polar axis. In the figure  $ZOX$  is the equatorial plane and  $YCY'$  any meridian. Let  $P$  be any point on a geodesic: then the plane through  $O$  parallel to the tangent plane at  $P$  is the plane  $DOX$ , where  $OD$  is the diameter conjugate to  $OP$ . If  $\phi$  is the eccentric angle of  $P$  so that the coordinates of  $P$  are  $O, a \cos\phi, b \sin\phi$  then

$$Od^2 = a^2 \sin^2\phi + b^2 \cos^2\phi$$

For a geodesic proceeding from  $P$  in azimuth  $A$ , the semidiameter parallel to it is  $OQ$  where  $DOQ = 180^\circ - A$  and so

$$D = \left( \frac{\sin^2 A}{a^2} + \frac{\cos^2 A}{a^2 \sin^2\phi + b^2 \cos^2\phi} \right)^{-\frac{1}{2}}$$

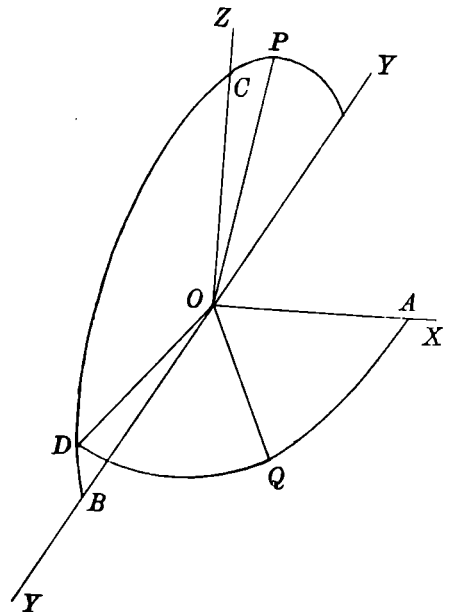
$$= \left( \frac{a^2 (a^2 \sin^2\phi + b^2 \cos^2\phi)}{a^2 + \sin^2 A \cos^2\phi (b^2 - a^2)} \right)^{\frac{1}{2}}$$

Also

$$p^2 = \frac{1}{\frac{\cos^2\phi}{a^2} + \frac{\sin^2\phi}{b^2}}$$

hence

$$p^2 D^2 = \frac{a^4 b^2}{a^2 + \sin^2 A \cos^2\phi (b^2 - a^2)}$$



which is constant along a geodesic. It follows that

$$\sin A \cos\phi = \text{constant} = k = \sin A_0 \dots \dots \dots (2)$$

along any geodesic,  $A_0$  being the azimuth of the geodesic on crossing the equator.

\* *Geometry of Three Dimensions* by George Salmon, 3rd edition, § 397.



2. Take two consecutive points on a spheroid and let  $A$  be the azimuth of the elementary line joining them. Then if the latitudes and longitudes are  $\lambda, L$ ;  $\lambda + d\lambda, L + dL$  it follows that

$$\tan A = \frac{\nu \cos \lambda \, dL}{\rho \, d\lambda} \dots \dots \dots (3)$$

where  $\rho$  is the radius of curvature of the meridian and  $\nu$  is the normal terminated by the minor axis.

The relation between the latitude  $\lambda$  and the eccentric angle or "reduced latitude"  $\phi$  is

$$\tan \phi = \frac{b}{a} \tan \lambda = \sqrt{1-e^2} \tan \lambda = (1 - \epsilon) \tan \lambda \dots \dots \dots (4)$$

Differentiating (2) logarithmically

$$\frac{dA}{d\phi} = \tan A \tan \phi \dots \dots \dots (5)$$

and by (4)

$$\frac{d\phi}{d\lambda} = \frac{b}{a} \frac{\cos^2 \phi}{\cos^2 \lambda} \dots \dots \dots (6)$$

Multiplying (5) and (6)

$$\begin{aligned} \frac{dA}{d\lambda} &= \frac{b}{a} \tan A \tan \phi \frac{\cos^2 \phi}{\cos^2 \lambda} = \frac{b^2}{a^2} \frac{\tan A \tan \lambda}{\cos^2 \lambda (1 + \frac{b^2}{a^2} \tan^2 \lambda)} \\ &= \frac{b^2}{a^2} \frac{\tan A \tan \lambda}{1 - e^2 \sin^2 \lambda} = \frac{\rho}{\nu} \tan A \tan \lambda \dots \dots \dots (7) \end{aligned}$$

Equation (3) may be written

$$L = \int \frac{\rho}{\nu \cos \lambda} \cdot \tan A \, d\lambda \dots \dots \dots (8)$$

and the integration of this will now be performed.

$$\begin{aligned} \cot^2 A &= \operatorname{cosec}^2 A - 1 = \frac{\cos^2 \phi}{k^2} - 1 \dots \dots \text{by (2)} \\ &= \frac{1}{k^2 \left(1 + \frac{b^2}{a^2} \tan^2 \lambda\right)} - 1 = \frac{\cos^2 \lambda}{k^2 (1 - e^2 \sin^2 \lambda)} - 1 \end{aligned}$$

$$\therefore \tan A = \pm \frac{k \sqrt{1 - e^2 \sin^2 \lambda}}{\sqrt{1 - k^2 - (1 - k^2 e^2) \sin^2 \lambda}} = \pm \frac{k}{\sqrt{1 - k^2}} \frac{\sqrt{1 - e^2 \sin^2 \lambda}}{\sqrt{1 - a^2 \sin^2 \lambda}}$$

where  $a^2 = \frac{1 - k^2 e^2}{1 - k^2}$  and the + sign is taken for 1st and 4th quadrants and the minus sign for the 2nd and 3rd quadrants.

Hence by (8)

$$\begin{aligned} L &= \pm \frac{k(1 - e^2)}{\sqrt{1 - k^2}} \int \frac{1}{\sqrt{1 - e^2 \sin^2 \lambda} \sqrt{1 - a^2 \sin^2 \lambda} \cos \lambda} \, d\lambda \\ &\text{since } \frac{\rho}{\nu} = \frac{1 - e^2}{1 - e^2 \sin^2 \lambda} \end{aligned}$$

Put  $x$  for  $\sin \lambda$ ; then

$$L = \pm \frac{k(1 - e^2)}{\sqrt{1 - k^2}} \int \frac{dx}{\sqrt{1 - e^2 x^2} \sqrt{1 - a^2 x^2} (1 - x^2)} \dots \dots \dots (9)$$

This is an elliptic integral which cannot be integrated exactly: but it may be developed in a series of integrable terms as follows.

Put  $1-x^2 = y^2$ , then

$$\begin{aligned} \frac{1}{\sqrt{1-e^2x^2}} \cdot \frac{1}{1-x^2} &= \frac{1}{y^2\sqrt{1-e^2+e^2y^2}} = \frac{1}{\sqrt{1-e^2}} \cdot \frac{1}{y^2\sqrt{1+\beta^2y^2}} \\ &= \frac{1}{\sqrt{1-e^2}} \left\{ \frac{1}{y^2} - \frac{1}{2}\beta^2 + \frac{1 \cdot 3}{2^2 \cdot 2}\beta^4y^2 \dots \dots \right\} \dots \dots (10) \end{aligned}$$

where  $\beta^2 = e^2/(1-e^2)$ : hence

$$\int \frac{dx}{\sqrt{1-e^2x^2}\sqrt{1-a^2x^2}(1-x^2)} = \frac{1}{\sqrt{1-e^2}} \int \frac{dx}{\sqrt{1-a^2x^2}} \left\{ \frac{1}{1-x^2} - \frac{1}{2}\beta^2 + \frac{3}{8}\beta^4(1-x^2) \dots \right\} \dots (11)$$

$$\text{Finally put} \quad \sin\theta = ax = a \sin\lambda \dots \dots \dots (12)$$

Then

$$\int \frac{dx'}{\sqrt{1-a^2x^2}} = \frac{1}{a} \int \frac{\cos\theta d\theta}{\cos\theta} = \frac{\theta}{a} \dots \dots \dots (13)$$

$$\begin{aligned} \int \frac{dx}{(1-x^2)\sqrt{1-a^2x^2}} &= \frac{1}{a} \int \frac{d\theta}{1-\sin^2\theta} = a \int \frac{d \tan \theta}{a^2+a^2-1|\tan^2\theta} \\ &= \frac{1}{\sqrt{a^2-1}} \tan^{-1} \left( \frac{\sqrt{a^2-1}}{a} \tan\theta \right) \dots \dots \dots (14) \end{aligned}$$

The remaining terms of (11) may be dealt with by the formula of reduction (17) now deduced.

$$\begin{aligned} u_n &= \int \frac{x^n dx}{\sqrt{1-a^2x^2}} = \frac{1}{a^2} \int \frac{x^{n-2}(1-\overline{1-a^2x^2}) dx}{\sqrt{1-a^2x^2}} \\ &= \frac{1}{a^2} \cdot u_{n-2} - \frac{1}{a^2} \int x^{n-2}\sqrt{1-a^2x^2} dx \dots \dots \dots (15) \end{aligned}$$

Integrating by parts

$$\begin{aligned} u_n &= -\frac{1}{a^2} \int x^{n-1} d\sqrt{1-a^2x^2} \\ &= -\frac{1}{a^2} x^{n-1}\sqrt{1-a^2x^2} + \frac{n-1}{a^2} \int x^{n-2}\sqrt{1-a^2x^2} dx \dots (16) \end{aligned}$$

Multiplying (15) by  $(n-1)$  and adding to (16)

$$nu_n = \frac{n-1}{a^2} \cdot u_{n-2} - \frac{1}{a^2} \cdot x^{n-1}\sqrt{1-a^2x^2} \dots \dots \dots (17)$$

Hence

$$\int \frac{1-x^2}{\sqrt{1-a^2x^2}} dx = \frac{\theta}{a} \left(1 - \frac{1}{2a^2}\right) + \frac{1}{2a^2} x \sqrt{1-a^2x^2} \dots \dots \dots (18)$$

and from (9), (11), (13), (14) and (18)

$$\begin{aligned} L-L' &= \pm \frac{k\sqrt{1-e^2}}{\sqrt{1-k^2}} \left[ \frac{1}{\sqrt{a^2-1}} \tan^{-1} \left( \frac{\sqrt{a^2-1}}{a} \tan\theta \right) - \frac{1}{2}\beta^2 \cdot \frac{\theta}{a} \right. \\ &\quad \left. + \frac{3}{8}\beta^4 \left\{ \frac{\theta}{a} \left(1 - \frac{1}{2a^2}\right) + \frac{1}{2a^2} \sin\lambda \cos\theta \right\} \dots \dots \right] \dots (19) \end{aligned}$$

in which

$$\left. \begin{aligned} \theta &= \sin^{-1}(a \sin \lambda) \\ a^2 &= \frac{1 - k^2 e^2}{1 - k^2} \\ \beta^2 &= \frac{e^2}{1 - e^2} \end{aligned} \right\} \dots \dots \dots (20)$$

Hence

$$\begin{aligned} \sin \theta &= \sqrt{\frac{1 - k^2 e^2}{1 - k^2}} \sin \lambda \\ \tan \theta &= \frac{\sqrt{1 - k^2 e^2} \sin \lambda}{\sqrt{1 - k^2 - (1 - k^2 e^2) \sin^2 \lambda}} = \frac{\sqrt{1 - k^2 e^2} \tan \lambda}{\sqrt{1 - k^2 - k^2 \tan^2 \lambda (1 - e^2)}} \\ &= \sqrt{\frac{1 - k^2 e^2}{1 - e^2}} \frac{\tan \phi}{\sqrt{1 - k^2 \sec^2 \phi}} \quad \text{by (4)} \\ &= \pm \sqrt{\frac{1 - k^2 e^2}{1 - e^2}} \tan \phi \sec A \dots \dots \dots (21) \end{aligned}$$

$$\therefore \frac{\sqrt{a^2 - 1}}{a} \tan \theta = \pm k \tan \phi \sec A = \pm \tan A \sin \phi \dots \dots \dots (22)$$

Since  $\theta$  is given by (20) we may always arrange that it shall be in 1st quadrant and the sign in (22) must be taken accordingly.

Put

$$\tan \psi = \pm \tan A \sin \phi \dots \dots \dots (23)$$

$\psi$  being always in the first quadrant.

Then (19) may be written

$$L - L' = \pm \left[ \pm \psi - \frac{e^2}{2} (1 + \sqrt{1 + k^2}) \frac{e^2}{8} + \dots \right] k \theta + \frac{3e^4}{16} (1 + \dots) k \sqrt{1 - k^2} \sin \lambda \cos \theta + \dots ] \dots \dots (24)$$

Now by (2) it follows that

$$\tan A \sin \phi = \frac{\sin A_0 \sec \phi}{\sqrt{1 - \sin^2 A_0 \sec^2 \phi}} \sin \phi = \frac{\sin A_0 \tan \phi}{\sqrt{\cos^2 A_0 - \sin^2 A_0 \tan^2 \phi}} = \frac{\tan A_0 \tan \phi}{\sqrt{1 - \tan^2 A_0 \tan^2 \phi}}$$

Hence

$$\psi = \pm \tan^{-1} (\tan A \sin \phi) = \pm \sin^{-1} (\tan A_0 \tan \phi) \dots \dots \dots (25)$$

Neglecting terms involving  $e^4$  (23) may be written

$$[L] = \pm \left[ \pm \psi - \frac{ke^2}{2} \theta \right] \dots \dots \dots (26)$$

where  $\psi$  and  $\theta$  are both in first quadrant and are defined by (25) and (20) respectively. This result is correct to nearest second for the terrestrial spheroid.

The rules for the double sign outside bracket are + 1st and 4th quadrant of azimuth  
 - 2nd and 3rd quadrant ,,  
 and for double signs before  $\psi$  + 1st and 2nd quadrant ,,  
 - 3rd and 4th quadrant ,,

The quantity  $\theta$  may also be found from (21) which may be written

$$\tan \theta = \frac{1}{k} \sqrt{\frac{1-k^2 e^2}{1-e^2}} \cdot \tan \psi \dots \dots \dots (27)$$

3. To solve (26) for  $A_0$  take the first approximation to  $A_0$ , namely  $A_1$  such that

$$\left[ \sin^{-1}(\tan A_1 \tan \phi) \right]_{\phi'}^{\phi} = L - L' \dots \dots \dots (28)$$

and for brevity put  $\tan A_1 \tan \phi = x$ ,  $\tan A_1 \tan \phi' = y$  and  $L - L' = \theta$

Then (26) becomes

$$\begin{aligned} \sin^{-1} x - \sin^{-1} y &= \theta \dots \dots \dots (29) \\ \therefore x \sqrt{1-y^2} - y \sqrt{1-x^2} &= \sin \theta \end{aligned}$$

Squaring and transposing

$$x^2 + y^2 - 2x^2 y^2 - \sin^2 \theta = 2xy \sqrt{(1-x^2)(1-y^2)}$$

Squaring again

$$\begin{aligned} \therefore \sin^4 \theta - 2 \sin^2 \theta (x^2 + y^2 - 2x^2 y^2) + (x^2 - y^2)^2 &= 0 \\ \text{or } (x^2 - y^2)^2 + 4x^2 y^2 \sin^2 \theta - 2(x^2 + y^2) \sin^2 \theta + \sin^4 \theta &= 0 \end{aligned}$$

and putting this into factors

$$(x^2 + y^2 + 2xy \cos \theta - \sin^2 \theta) (x^2 + y^2 - 2xy \cos \theta - \sin^2 \theta) = 0$$

Substituting for  $x$  and  $y$  it follows that

$$\tan^2 A_1 \left( \tan^2 \phi + \tan^2 \phi' \pm 2 \tan \phi \tan \phi' \cos (L - L') \right) = \sin^2 (L - L') \dots \dots (30)$$

The double signs have been introduced by the process of squaring and it is necessary to return to (28) to decide which signs give the required solution.

First supposing  $\tan \phi$  and  $\tan \phi'$  of the same sign and  $\phi > \phi'$ : then from (28) changing the sign of  $\tan \phi'$  diminishes the value of  $\tan A_1$ : hence by (30) we see that the lower sign must be taken. The same is true if  $\tan \phi$  and  $\tan \phi'$  are of opposite sign, and as  $\phi$  and  $\phi'$  are interchangeable this shows that the lower sign in (29) must always be taken.

Again if  $\phi > \phi'$  the sign of  $\tan A$  is the same as that of  $L - L'$ , and if  $\phi < \phi'$  the sign of  $\tan A$  is opposite to that of  $L - L'$ .

Hence we may write (30)

$$\tan A_1 = \pm \sqrt{\frac{\sin(L-L')}{\tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos L-L'}} \dots \dots \dots (31)$$

the upper or lower sign being taken according as  $\phi >$  or  $<$   $\phi'$ .

Also  $\tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos \overline{L-L'}$

$$\begin{aligned} &= (1-e^2) \left\{ \tan^2 \lambda + \tan^2 \lambda' - 2 \tan \lambda \tan \lambda' \cos \overline{L-L'} \right\} \text{ by (4)} \\ &= (1-e^2) \left( \tan \lambda - \tan \lambda' \right)^2 \left\{ 1 + \frac{4 \tan \lambda \tan \lambda'}{(\tan \lambda - \tan \lambda')^2} \sin^2 \frac{L-L'}{2} \right\} \\ &= (1-e^2) \left( \frac{\sin (\lambda - \lambda')}{\cos \lambda \cos \lambda'} \right)^2 \sec^2 \omega \end{aligned}$$

where  $\tan^2 \omega = \sin^2 \frac{L-L'}{2} \cdot \frac{\sin 2\lambda \sin 2\lambda'}{\sin^2 (\lambda - \lambda')}$

so that finally

$$\left. \begin{aligned} \tan A_1 &= + \frac{\sin(L-L') \cos \omega \cos \lambda \cos \lambda'}{\sqrt{1-e^2} \sin(\lambda-\lambda')} \\ \tan \omega &= \frac{\sin \frac{1}{2}(L-L')}{\sin(\lambda-\lambda')} \frac{\sin 2\lambda \sin 2\lambda'}{\sin^2(\lambda-\lambda')} \end{aligned} \right\} \dots \dots \dots (32)$$

where and  $\cos \omega$  is taken positive.

4. Denote by  ${}_1A$  the approximate value of  $A$  which corresponds to  $A_1$  which is an approximate value of  $A_0$ .

Then

$$\left. \begin{aligned} \sin {}_1A \cos \phi &= \sin A_1 \\ {}_1A + \delta_1 A &= A \end{aligned} \right\} \dots \dots \dots (33)$$

and Hence

$$\begin{aligned} \tan A_1 &= \frac{\sin {}_1A \cos \phi}{\sqrt{1 - \sin^2 {}_1A \cos^2 \phi}} = \frac{1}{\sqrt{\sec^2 \phi \operatorname{cosec}^2 {}_1A - 1}} \\ &= \frac{1}{\sqrt{\sec^2 \phi \cot^2 {}_1A + \tan^2 \phi}} \end{aligned}$$

By (30)

$$\begin{aligned} \sin^2(L-L') (\tan^2 \phi + \sec^2 \phi \cot^2 {}_1A) &= \tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos(L-L') \\ \cot^2 {}_1A \sec^2 \phi \sin^2(L-L') &= \left\{ \tan \phi \cos(L-L') - \tan \phi' \right\}^2 \\ \therefore \tan {}_1A &= + \frac{\sec \phi \sin(L-L')}{\tan \phi \cos(L-L') - \tan \phi'} \dots \dots \dots (34) \end{aligned}$$

Similarly

$$\tan {}_1A' = - \frac{\sec \phi' \sin(L-L')}{\tan \phi' \cos(L-L') - \tan \phi} \dots \dots \dots (35)$$

By differentiating (33) logarithmically with regard to  $A_1$  and  ${}_1A$  we get the relation between  $\delta A_1$  and  $\delta_1 A$  as follows:

$$\cot {}_1A \delta_1 A = \cot A_1 \delta A_1 \dots \dots \dots (36)$$

Equations (34) and (35) correspond to the ordinary equations of spherical trigonometry to which they reduce if the eccentric angles  $\phi, \phi'$  are replaced by latitudes  $\lambda, \lambda'$ .

5. Suppose next that

$$A_0 = A_1 + \delta A_1 \dots \dots \dots (37)$$

where  $\delta A_1$  gives a second approximation to  $A_0$ . Then by (26) neglecting terms in  $e^4$ , it follows that

$$\begin{aligned} \frac{e^2 k}{2} (\theta - \theta') &= \pm \delta [\psi] = \delta \left[ \sin^{-1} (\tan A_1 \tan \phi) \right] \phi \\ &= \delta A_1 \left[ \frac{\sec^2 A_1 \tan \phi}{\sqrt{1 - \tan^2 A_1 \tan^2 \phi}} \right] \phi \dots \dots \dots (38) \end{aligned}$$

With notation of (20)

$$\sin \theta = a \sin \lambda = \sqrt{\frac{1-e^2 \sin^2 A_1}{1-\sin^2 A_1}} \sin \lambda = \sec A_1 \sin \lambda \sqrt{1-e^2 \sin^2 A_1}$$

since  $k = \sin A_0 = \sin A_1$

$$\begin{aligned} \therefore \cos^2 \theta &= 1 - \sin^2 \lambda \sec^2 A_1 (1 - e^2 \sin^2 A_1) \\ &= \cos^2 \lambda - (1 - e^2) \tan^2 A_1 \sin^2 \lambda \end{aligned}$$

$$\therefore \frac{\sec^2 A_1 \tan \phi}{\sqrt{1 - \tan^2 A_1 \tan^2 \phi}} = \frac{\sqrt{1 - e^2} \sec^2 A_1 \tan \lambda}{\sqrt{1 - (1 - e^2) \tan^2 A_1 \tan^2 \lambda}} = \sqrt{1 - e^2} \sec^2 A_1 \sin \lambda \sec \theta$$

\(\therefore\) by (38)

$$\delta A_1 = \frac{e^2 \sin A_1 \cos^2 A_1}{2\sqrt{1-e^2}} \frac{(\theta - \theta')}{\left[ \sin \lambda \sec \theta \right]_{\lambda}^{\lambda'}}$$

Now

$$\sin \lambda \sec \theta = \frac{1}{a} \tan \theta$$

$$\therefore \delta A_1 = \frac{e^2 \sin 2 A_1}{4} \sqrt{\frac{1 - e^2 \sin^2 A_1}{1 - e^2}} \cdot \frac{\theta - \theta'}{\tan \theta - \tan \theta'} \left. \vphantom{\frac{e^2 \sin 2 A_1}{4}} \right\} \dots \dots \dots (39)$$

in which

$$\theta = \sin^{-1} \left( \sin \lambda \sec A_1 \sqrt{1 - e^2 \sin^2 A_1} \right)$$

For computation

$$\begin{aligned} \sin \theta &= \sin \phi \sec A_1 \left\{ 1 + \frac{e^2}{2} (\cos^2 \phi - \sin^2 A_1) \right\} && \text{since } \sin \lambda = \sin \phi \sqrt{1 - e^2 \cos^2 \phi} \\ &= \sin \phi \sec A_1 \left( 1 + \frac{e^2}{2} \cos^2 \phi \cos^2 A_1 \right) && \text{since } \sin A_1 = \sin A \cos \phi \end{aligned}$$

Let  $\sin \theta_1 = \sin \phi \sec A_1$  and  $\theta = \theta_1 + \delta \theta$

$$\tan \theta_1 = \frac{\sin \phi}{\sqrt{\cos^2 A_1 - \sin^2 \phi}} = \frac{\sin \phi}{\sqrt{\cos^2 \phi - \sin^2 A \cos^2 \phi}} = \tan \phi \sec A$$

and  $\delta \theta = \frac{e^2}{2} \cos^2 \phi \cos^2 A \tan \theta_1 = \frac{e^2}{4} \sin 2 \phi \cos A$ .

With a given value of the larger quantity  $\theta$ ,  $\theta'$  say  $\theta$  it is clear that  $\frac{\theta - \theta'}{\tan \theta - \tan \theta'}$  is greater

the smaller the value of  $\theta'$ ; its maximum value accordingly is  $\frac{\theta}{\tan \theta}$  and the maximum value of this quantity occurs when  $\theta = 0$ , when it becomes unity. It is quite clear then that  $\delta A_1$  cannot exceed  $\frac{1}{4} e^2 \sin 2A$ , i. e.  $6' \cdot \sin 2A_1$  in the case of the terrestrial spheroid where  $e^2 = \frac{1}{150}$ .

6. If  $ds$  is the length of an elementary line

$$\rho d\lambda = -ds \cos A$$

$$\therefore s = -\int \rho \sec A d\lambda$$

On a geodesic

$$\cos A = \sqrt{1 - \sin^2 A_0 \sec^2 \phi} = \cos A_0 \sqrt{1 - \tan^2 A_0 \tan^2 \phi}$$

$$\rho = \frac{a(1 - e^2)}{(1 - e^2 \sin^2 \lambda)^{\frac{3}{2}}}; \quad \tan \lambda = \frac{\tan \phi}{\sqrt{1 - e^2}}$$

$$\sin \lambda = \frac{\sin \phi}{\sqrt{1 - e^2 \cos^2 \phi}}; \quad 1 - e^2 \sin^2 \lambda = \frac{1 - e^2}{1 - e^2 \cos^2 \phi}; \quad \cos \lambda = \cos \phi \sqrt{\frac{1 - e^2}{1 - e^2 \cos^2 \phi}}$$

whence

$$\rho = \frac{a}{\sqrt{1 - e^2}} (1 - e^2 \cos^2 \phi)^{\frac{3}{2}}$$

Also  $\frac{d\lambda}{\sin \lambda \cos \lambda} = \frac{d\phi}{\sin \phi \cos \phi}$

$$d\lambda = d\phi \frac{\sqrt{1 - e^2}}{1 - e^2 \cos^2 \phi}$$

$$\begin{aligned} \therefore s &= - \int \frac{a(1 - e^2 \cos^2 \phi)^{\frac{3}{2}}}{\sqrt{1 - e^2}} \cdot \frac{\sec A_0}{\sqrt{1 - \tan^2 A_0 \tan^2 \phi}} \cdot \frac{\sqrt{1 - e^2}}{1 - e^2 \cos^2 \phi} \cdot d\phi \\ &= -a \sec A_0 \int \sqrt{\frac{1 - e^2 \cos^2 \phi}{1 - \tan^2 A_0 \tan^2 \phi}} \cdot d\phi \end{aligned}$$

Put  $\sin \phi = x \quad d\phi = \frac{dx}{\sqrt{1 - x^2}}$

$$\begin{aligned} s &= -a \sec A_0 \int \sqrt{\frac{1 - e^2(1 - x^2)}{1 - \tan^2 A_0 x^2 / (1 - x^2)}} \cdot \frac{dx}{\sqrt{1 - x^2}} \\ &= -a \int \sqrt{\frac{1 - e^2(1 - x^2)}{\cos^2 A_0 x^2 - x^2}} \cdot dx = -a \sqrt{1 - e^2} \int \frac{\sqrt{1 + \beta^2 x^2}}{\sqrt{\cos^2 A_0 - x^2}} \cdot dx \end{aligned}$$

where  $\beta^2 = \frac{e^2}{1 - e^2}$

Put then  $x = \cos A_0 \sin \chi = \sin \phi$

$$\begin{aligned} s &= -a \sqrt{1 - e^2} \int \frac{\chi \sqrt{1 + \beta^2 \cos^2 A_0 \sin^2 \chi}}{\cos A_0 \cos \chi} \cdot \cos A_0 \cos \chi d\chi \\ &= -a \sqrt{1 - e^2} \int \frac{\chi}{\cos \chi} \sqrt{1 + h^2 \sin^2 \chi} d\chi \quad \text{where } h^2 = \beta^2 \cos^2 A_0 = \frac{e^2 \cos^2 A_0}{1 - e^2} \\ &= -a \sqrt{1 - e^2} \left[ \chi + \frac{1}{2} h^2 \int \sin^2 \chi d\chi - \frac{1}{8} h^4 \int \sin^4 \chi d\chi \dots \dots \right] \frac{\chi}{\cos \chi} \end{aligned}$$

Now  $\int \sin^n \chi d\chi = -\frac{1}{n} \cos \chi \sin^{n-1} \chi + \frac{(n-1)}{n} \int \sin^{n-2} \chi d\chi$

$\therefore \int \sin^2 \chi d\chi = -\frac{1}{2} \cos \chi \sin \chi + \frac{\chi}{2}$

$$\int \sin^4 \chi d\chi = -\frac{1}{4} \cos \chi \sin^3 \chi - \frac{3}{4} \int \sin^2 \chi d\chi$$

$$= -\frac{1}{4} \cos \chi \sin^3 \chi - \frac{3}{8} \cos \chi \sin \chi + \frac{3}{8} \chi$$

etc.

$$\therefore s = -a \sqrt{1-e^2} \left[ \chi \left( 1 + \frac{1}{4} h^2 - \frac{3}{64} h^4 \dots \right) - \frac{h^2}{4} \cos \chi \sin \chi \left( 1 - \frac{3}{16} h^2 \dots \right) + \frac{h^4}{32} \cos \chi \sin^3 \chi \left( 1 \dots \dots \right) - \dots \dots \right] \chi' \dots (40)$$

where

$$\left. \begin{aligned} \sin \chi &= \frac{\sin \phi}{\cos A_0} \\ h^2 &= \frac{e^2 \cos^2 A_0}{1-e^2} \end{aligned} \right\} \dots \dots \dots (41)$$

Otherwise since  $\sin A \cos \phi = \sin A_0$

$$\sin \chi = \frac{\sin \phi \sin A_0}{\sin A \cos \phi \cos A_0} = \tan \phi \tan A_0 \operatorname{cosec} A$$

Also  $\tan^2 A_0 = \frac{\sin^2 A \cos^2 \phi}{1 - \sin^2 A \cos^2 \phi} = \frac{\sin^2 A}{\tan^2 \phi + \cos^2 A}$

and  $\sin \chi = \pm \frac{\tan \phi}{\sqrt{\tan^2 \phi + \cos^2 A}}$

$$\tan \chi = \pm \tan \phi \sec A \dots \dots \dots (42)$$

7. To facilitate reductions, tables are now given enabling the conversion from  $\lambda$  to  $\phi$  and *vice versa* to be easily performed. They have been computed from formula  $\lambda - \phi = \frac{1}{2} \frac{\epsilon \sin^2 \lambda}{1 - \epsilon \sin^2 \lambda}$  which is readily deducible from (4).

Table XXI.

$\lambda$	$\phi - \lambda$	$\lambda$	$\phi - \lambda$	$\lambda$	$\phi - \lambda$	$\lambda$	$\phi - \lambda$
0	0 0'0	10	-1 57'3	20	-3 40'5	30	-4 57'1
1	0 12'0	11	-2 8'4	21	-3 49'5	31	-5 3'0
2	0 23'8	12	-2 19'4	22	-3 58'2	32	-5 8'5
3	0 35'8	13	-2 30'3	23	-4 6'7	33	-5 13'5
4	0 47'7	14	-2 41'0	24	-4 14'9	34	-5 18'2
5	0 59'5	15	-2 51'4	25	-4 22'8	35	-5 22'5
6	-1 11'3	16	-3 1'8	26	-4 30'4	36	-5 26'4
7	-1 23'0	17	-3 11'9	27	-4 37'6	37	-5 29'9
8	-1 34'5	18	-3 21'6	28	-4 44'5	38	-5 33'1
9	-1 45'9	19	-3 31'2	29	-4 50'9	39	-5 35'8
10	-1 57'3	20	-3 40'5	30	-4 57'1	40	-5 38'2

Table XXII.

$\phi$	$\lambda - \phi$	$\phi$	$\lambda - \phi$	$\phi$	$\lambda - \phi$	$\phi$	$\lambda - \phi$
0	+0 0'0	10	+1 57'7	20	+3 41'1	30	+4 57'6
1	+0 12'0	11	+2 8'8	21	+3 50'1	31	+5 3'5
2	+0 23'9	12	+2 19'8	22	+3 58'8	32	+5 8'9
3	+0 35'9	13	+2 30'7	23	+4 7'3	33	+5 13'9
4	+0 47'9	14	+2 41'5	24	+4 15'5	34	+5 18'6
5	+0 59'7	15	+2 51'9	25	+4 23'4	35	+5 22'8
6	+1 11'5	16	+3 2'3	26	+4 30'9	36	+5 26'7
7	+1 23'3	17	+3 12'4	27	+4 38'1	37	+5 30'2
8	+1 34'8	18	+3 22'1	28	+4 45'0	38	+5 33'3
9	+1 46'2	19	+3 31'7	29	+4 51'4	39	+5 36'0
10	+1 57'7	20	+3 41'1	30	+4 57'6	40	+5 38'4



Values of  $A_0, k, s$  together with certain of the quantities by means of which they are computed are now given in tabular form for geodesics passing through the origin and points  $L, \phi$ , for values of  $L'-L$  differing by  $4^\circ$  from 0 to  $24^\circ$  and for values of  $\phi$  from  $10^\circ$  to  $38^\circ$ . It is clear that for longitudes east of the origin the value of  $A$  is  $360 -$  (its value in table) : and that for  $s$  there is no change.

TABLE XYIII.

$\phi$	$L'-L$	$0^\circ$	$4^\circ$	$8^\circ$	$12^\circ$	$16^\circ$	$20^\circ$	$24^\circ$
$38^\circ$	$180^\circ - A_1$	0 0 0	11 40 32.2	21 58 5.6	30 12 43.8	36 27 5.0	41 2 45.3	44 23 53.7
	$-\delta A_1$	0 0	1 37.2	2 41.6	3 6.7	3 2.2	2 48.9	2 27.1
	$180^\circ - A_0$	0 0 0	11 42 9	22 0 47	30 15 51	36 30 7	41 5 34	44 26 21
	$\log k$	$-\infty$	$\bar{1}.3071363$	$\bar{1}.5738213$	$\bar{1}.7024180$	$\bar{1}.7744081$	$\bar{1}.8177511$	$\bar{1}.8451916$
	$h^2$	0.006682	0.006407	0.005743	0.004985	0.004318	0.003795	0.003407
$s/b$	0.243714	0.250921	0.271391	0.302418	0.341099	0.385090	0.432724	
$34^\circ$	$180^\circ - A_1$	0 0 0	16 46 26.0	30 1 52.5	39 8 30.0	45 6 17.2	48 59 53.3	51 34 13.2
	$-\delta A_1$	0 0	2 14.0	3 21.5	3 20.0	2 53.0	2 21.7	1 52.7
	$180^\circ - A_0$	0 0 0	16 48 40	30 5 14	39 11 50	45 9 10	49 2 15	51 36 6
	$\log k$	$-\infty$	$\bar{1}.4612245$	$\bar{1}.7001131$	$\bar{1}.8007114$	$\bar{1}.8506405$	$\bar{1}.8780268$	$\bar{1}.8941560$
	$h^2$	0.006682	0.006123	0.005003	0.004013	0.003323	0.002872	0.002578
$s/b$	0.173823	0.184246	0.212464	0.252555	0.299755	0.351166	0.405139	
$30^\circ$	$180^\circ - A_1$	0 0 0	27 11 14.3	43 2 45.7	50 56 59.8	55 2 12.4	57 17 10.9	58 34 48.1
	$-\delta A_1$	0 0	3 24.5	3 27.8	2 36.8	1 51.6	1 18.3	0 54.9
	$180^\circ - A_0$	0 0 0	27 14 39	43 6 14	50 59 37	55 4 4	57 18 29	58 35 43
	$\log k$	$-\infty$	$\bar{1}.6606593$	$\bar{1}.8346251$	$\bar{1}.8904627$	$\bar{1}.9137238$	$\bar{1}.9250392$	$\bar{1}.9312075$
	$h^2$	0.006682	0.005282	0.003562	0.002647	0.002191	0.001949	0.001814
$s/b$	0.103942	0.121206	0.162319	0.213923	0.270030	0.328293	0.387702	
$26^\circ$	$180^\circ - A_1$	0 0 0	52 54 33.6	60 59 16.4	62 59 19.7	63 41 3.3	63 56 27.7	
	$-\delta A_1$	0 0	2 46.9	1 7.7	0 32.1	0 16.4	0 6.5	
	$180^\circ - A_0$	0 0 0	52 57 21	61 0 24	62 59 52	63 41 20	63 56 34	
	$\log k$	$-\infty$	$\bar{1}.9020954$	$\bar{1}.9418474$	$\bar{1}.9498721$	$\bar{1}.9525018$	$\bar{1}.9534486$	
	$h^2$	0.006682	0.002425	0.001570	0.001378	0.001313	0.001259	
$s/b$	0.034077	0.072033	0.131409	0.193359	0.255090	0.318858		
$22^\circ$	$180^\circ - A_1$	0 0 0	53 31 3.3	62 24 14.3	64 41 45.0	65 31 8.4	65 50 30.9	65 56 48.9
	$-\delta A_1$	0 0	3 5.1	1 19.0	0 38.7	0 20.4	0 9.4	0 2.1
	$180^\circ - A_0$	0 0 0	53 34 8	62 25 33	64 42 24	65 31 29	65 50 49	65 56 51
	$\log k$	$-\infty$	$\bar{1}.9055652$	$\bar{1}.9476361$	$\bar{1}.9562317$	$\bar{1}.9591081$	$\bar{1}.9602121$	$\bar{1}.9605528$
	$h^2$	0.006682	0.002357	0.001432	0.001220	0.001147	0.001119	0.001110
$s/b$	0.035775	0.073722	0.133777	0.196608	0.260182	0.324032	0.387973	
$18^\circ$	$180^\circ - A_1$	0 0 0	29 19 11.0	46 25 6.2	55 5 13.3	59 41 41.7	62 18 45.2	63 52 30.2
	$-\delta A_1$	0 0	4 2.5	4 7.9	3 11.5	2 21.1	1 43.1	1 15.0
	$180^\circ - A_0$	0 0 0	29 23 14	46 29 14	55 8 25	59 44 3	62 20 28	63 53 45
	$\log k$	$-\infty$	$\bar{1}.6908226$	$\bar{1}.8604705$	$\bar{1}.9141068$	$\bar{1}.9363608$	$\bar{1}.9473002$	$\bar{1}.9532744$
	$h^2$	0.006682	0.005073	0.003168	0.002183	0.001697	0.001440	0.001294
$s/b$	0.105616	0.124192	0.168012	0.222619	0.281789	0.343134	0.405645	
$14^\circ$	$180^\circ - A_1$	0 0 0	19 22 52.8	34 31 14.1	44 51 50.8	51 41 49.6	56 15 27.3	59 21 59.3
	$-\delta A_1$	0 0	3 8.3	4 27.9	4 27.1	3 54.8	3 17.9	2 43.8
	$180^\circ - A_0$	0 0 0	19 26 1	34 35 42	44 56 18	51 45 44	56 18 45	59 24 43
	$\log k$	$-\infty$	$\bar{1}.5220722$	$\bar{1}.7541739$	$\bar{1}.8490169$	$\bar{1}.8951186$	$\bar{1}.9201629$	$\bar{1}.9349366$
	$h^2$	0.006682	0.005943	0.004528	0.003348	0.002560	0.002056	0.001730
$s/b$	0.157964	0.187488	0.219676	0.254754	0.317253	0.374042	0.433406	
$10^\circ$	$180^\circ - A_1$	0 0 0	14 27 13.6	27 2 2.7	36 58 9.7	44 26 58.7	50 0 9.0	54 8 7.8
	$-\delta A_1$	0 0	2 29.4	4 5.5	4 41.4	4 38.9	4 18.7	3 51.3
	$180^\circ - A_0$	0 0 0	14 29 43	27 6 8	37 2 51	44 31 38	50 4 28	54 11 59
	$\log k$	$-\infty$	$\bar{1}.3984612$	$\bar{1}.6585649$	$\bar{1}.7799407$	$\bar{1}.8458708$	$\bar{1}.8847263$	$\bar{1}.9090537$
	$h^2$	0.006682	0.006264	0.005295	0.004257	0.003396	0.002752	0.002287
$s/b$	0.245270	0.254192	0.279254	0.316635	0.362531	0.414102	0.469464	

TABLE XXIV.

$\phi$	$L'-L$	4°	8°	12°	16°	20°	24°	$L'-L$	4°	8°	12°	16°	20°	24°
38°	$\psi$	9 18 48.0	18 24 47.6	27 7 19.7	35 19 17.2	42 57 7.6	50 0 29.7	$\psi'$	5 18 13.3	10 23 38.1	15 5 35.6	19 16 57.0	22 54 14.4	25 57 2.0
	$\theta$	39 2 46	41 41 34	45 32 6	50 2 52	54 49 40	59 36 57	$\theta'$	24 39 44	28 8 25	29 12 42	30 30 56	32 46 56	34 50 54
	$-x$	38 57 23.4	41 36 40.7	45 27 50.1	49 59 15.0	54 46 36.4	59 34 23.7	$-x'$	24 35 33.6	26 4 32.5	28 0 8.0	30 27 41.7	32 43 58.1	34 48 9.4
	$180^\circ - A$	14 54 57	28 24 7	39 45 34	49 0 53	56 31 21	62 41 0	$180^\circ - A'$	12 49 59	24 13 58	33 20 48	40 38 49	46 2 4	50 3 31
34°	$\psi$	11 45 31.0	23 0 15.0	33 22 16.8	42 41 47.9	50 58 58.9	58 19 40.8	$\psi'$	7 44 54.6	14 59 2.1	21 20 26.0	26 30 22.0	30 55 57.1	34 16 3.1
	$\theta$	35 49 35	40 19 53	46 14 27	52 30 24	58 34 31	64 13 30	$\theta'$	26 15 46	28 0 24	31 46 34	35 20 40	38 28 29	41 2 20
	$-x$	35 44 36.7	40 15 39.3	46 11 1.0	52 27 39.7	58 32 19.7	64 11 40.3	$-x'$	25 11 43.3	28 5 50.3	31 43 20.4	35 17 50.0	38 26 4.7	41 0 8.1
	$180^\circ - A$	20 25 3	37 12 26	49 40 0	58 46 54	65 37 29	70 57 58	$180^\circ - A'$	19 27 50	33 17 47	43 47 37	50 55 60	55 47 0	59 6 56
30°	$\psi$	17 17 39.5	32 42 25.8	45 27 48.4	55 45 11.5	64 6 17.8	71 1 40.6	$\psi'$	13 17 1.7	24 41 9.8	33 25 54.1	39 42 30.6	44 3 9.4	46 57 53.6
	$\theta$	34 17 29	43 16 20	52 38 4	60 51 33	67 47 47	73 39 40	$\theta'$	27 20 35	33 63 30	40 23 14	45 24 11	49 0 29	51 28 28
	$-x$	34 13 16.3	43 13 17.5	52 35 54.0	60 49 57.2	67 46 36.8	73 38 49.5	$-x'$	27 16 53.2	33 65 30.9	40 20 59.8	45 22 18.5	48 59 49.2	51 26 56.8
	$180^\circ - A$	31 54 40	52 5 42	63 48 14	71 12 6	78 21 23	80 14 48	$180^\circ - A'$	30 5 9	49 26 27	58 18 49	63 51 53	67 9 30	69 10 9
26°	$\psi$	40 15 21.9	61 39 31.7	73 9 48.7	80 31 41.5	85 55 33.0		$\psi'$	36 14 41.9	53 39 12.2	61 7 51.0	64 29 4.6	65 52 17.6	
	$\theta$	46 43 33	64 45 38	74 55 13	81 30 7	86 20 26		$\theta'$	42 36 4	57 14 20	63 51 15	68 51 1	68 5 36	
	$-x$	46 41 28.7	64 44 36.1	74 54 37.8	81 29 39.5	86 20 19.5		$-x'$	42 33 59.6	57 13 6.2	63 50 18.6	68 60 9.7	68 4 50.3	
	$180^\circ - A$	62 37 43	76 41 49	82 26 38	85 49 2	88 12 42		$180^\circ - A'$	60 55 52	73 17 55	77 20 15	78 59 35	79 30 20	
22°	$\psi$	33 11 16.7	50 41 8.8	58 45 26.4	62 34 8.8	64 17 12.1	64 51 13.0	$\psi'$	37 11 57.8	58 42 30.0	70 47 27.9	78 36 50.4	81 20 35.0	88 55 15.0
	$\theta$	39 8 33	54 2 38	61 16 15	64 43 37	66 17 29	66 48 31	$\theta'$	43 21 54	61 42 9	73 31 30	79 37 11	84 60 18	89 0 54
	$x$	39 6 34.5	54 1 28.4	61 15 21.6	64 42 51.2	66 16 46.1	66 47 49.1	$x'$	43 19 53.1	61 41 7.8	72 30 64.2	79 36 40.8	84 50 7.7	89 0 52.3
	$A$	60 11 58	72 56 44	77 11 49	78 59 49	79 46 27	80 1 36	$A'$	61 46 9	76 5 18	81 55 9	85 19 31	87 41 19	89 33 39
18°	$\psi$	10 32 37.5	20 0 48.5	27 48 16.9	33 50 3.8	38 18 48.4	41 32 18.6	$\psi'$	14 33 19.3	28 2 12.2	39 50 22.2	46 52 51.1	58 22 16.3	65 36 27.9
	$\theta$	20 49 13	28 42 15	32 45 22	37 50 19	41 45 18	44 37 51	$\theta'$	27 56 41	36 19 59	45 30 30	53 58 35	61 24 15	67 51 24
	$x$	20 46 19.0	28 40 3.8	32 43 39.5	37 48 53.7	41 44 7.0	44 36 44.2	$x'$	27 53 5.0	36 17 23.8	45 28 38.0	53 57 11.3	61 23 12.4	67 50 37.4
	$A$	31 3 42	49 41 20	59 37 49	65 15 0	68 38 23	70 46 15	$A'$	32 30 12	52 34 32	63 59 4	71 2 53	75 54 39	79 31 50
14°	$\psi$	5 2 47.9	9 54 8.1	14 24 21.7	18 26 45.7	21 57 51.5	24 56 51.9	$\psi'$	9 3 30.7	17 55 33.6	26 26 30.3	34 29 36.7	42 1 25.3	49 1 7.3
	$\theta$	14 54 25	17 7 37	20 0 55	23 2 6	25 62 56	28 24 27	$\theta'$	25 40 9	29 43 42	35 11 35	41 12 57	47 18 42	53 14 3
	$x$	14 51 52.8	17 5 26.5	19 59 3.3	23 0 32.1	25 51 33.0	28 23 11.6	$x'$	25 36 10.8	29 40 21.6	35 8 52.9	41 10 46.4	47 16 56.3	53 12 36.0
	$A$	23 3 13	35 48 50	46 42 59	54 2 48	59 2 33	62 31 26	$A'$	21 22 1	33 28 36	50 40 4	59 19 43	55 40 1	70 30 19
10°	$\psi$	2 36 46.0	5 10 38.7	7 38 55.8	9 69 16.4	12 9 46.7	14 9 4.2	$\psi'$	6 37 29.5	31 12 5.9	19 41 6.8	26 2 11.0	32 13 24.1	38 13 24.9
	$\theta$	10 21 50	11 16 40	12 35 31	14 7 14	15 43 7	17 17 14	$\theta'$	24 57 38	27 18 20	30 45 21	34 54 30	39 27 20	44 11 29
	$x$	10 19 56.4	11 14 55.3	12 33 58.2	14 6 51.2	15 41 52.9	17 16 7.5	$x'$	24 53 31.7	27 14 38.1	30 42 0.1	34 51 46.9	39 25 1.4	44 9 31.7
	$A$	14 43 26	27 33 20	37 43 4	45 24 11	51 8 33	55 26 39	$A'$	15 54 28	29 55 34	41 16 49	50 9 57	57 6 57	62 38 32

For the case  $L'-L=0$  it is clear that

$$\psi = \psi' = 0; \theta = \lambda; \theta' = \lambda'; \chi = \pm \phi; \chi' = \pm \phi'$$

8. It may be of interest to find the expression for the azimuthal angle of a vertical plane at the origin which passes through any given point on the earth, so that the difference of this and the geodesic may be studied.

The spheroid may be expressed

$$\frac{x^2 + y^2}{a^2} + \frac{z^2}{b^2} = 1 \dots \dots \dots (43)$$

and  $P$  and  $Q$  two points on surface

$$P \quad a \cos \phi, 0, b \sin \phi$$

$$Q \quad a \cos \phi' \cos L, a \cos \phi' \sin L, b \sin \phi'$$

Tangent plane at  $P$  is

$$\frac{x \cos \phi}{a} + \frac{z \sin \phi}{b} = 1 \dots \dots \dots (44)$$

The vertical plane at  $P$  which passes through  $Q$  also is

$$lx + my + nz = 1$$

subject to the conditions

$$\begin{aligned}
 la \cos\phi + nb \sin\phi &= 1 && \text{since } P \text{ is in it} \\
 la \cos\phi' \cos L + ma \cos\phi' \sin L + nb \sin\phi' &= 1 && \text{since } Q \text{ is in it} \\
 \frac{l \cos\phi}{a} + \frac{n \sin\phi}{b} &= 0 && \text{since it is perpendicular to (44)}
 \end{aligned}$$

$$\therefore \frac{la \cos\phi}{a^2} = \frac{nb \sin\phi}{-b^2} = \frac{la \cos\phi + nb \sin\phi}{a^2 - b^2} = \frac{1}{a^2 - b^2} \dots \dots \dots (45)$$

Also

$$\begin{aligned}
 \frac{1}{a^2 - b^2} &= \frac{la}{a^2 \sec\phi} = \frac{nb}{-b^2 \operatorname{cosec}\phi} = \frac{la \cos\phi' \cos L + nb \sin\phi' - 1}{a^2 \sec\phi \cos\phi' \cos L - b^2 \operatorname{cosec}\phi \sin\phi' - a^2 + b^2} \\
 &= \frac{ma \cos\phi' \sin L}{a^2 \left(1 - \frac{\cos\phi' \cos L}{\cos\phi}\right) - b^2 \left(1 - \frac{\sin\phi'}{\sin\phi}\right)} \dots \dots \dots (46)
 \end{aligned}$$

The azimuthal angle of *Q* from *P* as determined by the vertical plane through *P* is the angle this plane makes with *ZOX* or otherwise it is the angle between the normal to this plane, whose direction cosines are proportional to *lmn*, and the axis *OY*.

The angle accordingly is  $\cos^{-1} \frac{m}{\sqrt{l^2 + m^2 + n^2}} = \cot^{-1} \frac{m}{\sqrt{l^2 + n^2}} = \psi$  say

$$\therefore \tan \psi = \frac{\sqrt{l^2 + n^2}}{m} \dots \dots \dots (47)$$

Now

$$\frac{l}{a \sec\phi} = \frac{n}{-b \operatorname{cosec}\phi} = \frac{1}{a^2 - b^2} = \frac{\sqrt{l^2 + n^2}}{\sqrt{a^2 \sec^2\phi + b^2 \operatorname{cosec}^2\phi}}$$

$$\begin{aligned}
 \therefore \sqrt{l^2 + n^2} &= \frac{a}{\sin\phi \cos\phi} \sqrt{\sin^2\phi + (1 - e^2) \cos^2\phi} \cdot \frac{1}{a^2 e} \\
 &= \frac{\sqrt{1 - e^2 \cos^2\phi}}{ae^2 \sin\phi \cos\phi} \dots \dots \dots (48)
 \end{aligned}$$

and by (46)

$$\begin{aligned}
 m &= \frac{1 - \frac{\cos\phi' \cos L}{\cos\phi} - (1 - e^2) \left(1 - \frac{\sin\phi'}{\sin\phi}\right)}{ae^2 \cos\phi' \sin L} \\
 &= \left\{ \frac{\sin\phi'}{\sin\phi} - \frac{\cos\phi' \cos L}{\cos\phi} + e^2 \left(1 - \frac{\sin\phi'}{\sin\phi}\right) \right\} / ae^2 \cos\phi' \sin L
 \end{aligned}$$

$$\begin{aligned}
 \therefore \tan \psi &= \frac{\sqrt{1 - e^2 \cos^2\phi}}{ae^2 \sin\phi \cos\phi} \cdot \frac{ae^2 \cos\phi' \sin L}{\frac{\sin\phi'}{\sin\phi} - \frac{\cos\phi' \cos L}{\cos\phi} + e^2 \left(1 - \frac{\sin\phi'}{\sin\phi}\right)} \\
 &= \frac{2 \cos\phi' \sin L \sqrt{1 - e^2 \cos^2\phi}}{\sin 2\phi \left\{ \frac{\sin\phi'}{\sin\phi} - \frac{\cos\phi' \cos L}{\cos\phi} + e^2 \left(1 - \frac{\sin\phi'}{\sin\phi}\right) \right\}} \dots \dots (49)
 \end{aligned}$$

Substitute for  $\phi$  in terms of  $\lambda$  by means of (4)

$$\tan \phi = \frac{1}{\sqrt{1 - e^2}} \tan \lambda; \quad \frac{\sin \phi}{\sqrt{1 - e^2} \sin \lambda} = \frac{\cos \phi}{\cos \lambda} = \frac{1}{\sqrt{1 - e^2 \sin^2 \lambda}}$$

Hence the value of  $\tan \psi$  is

$$\frac{\frac{\cos \lambda' \sin L}{\sqrt{1-e^2 \sin^2 \lambda}} \sqrt{\frac{1-e^2}{1-e^2 \sin^2 \lambda}}}{\frac{(\sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L) \sqrt{1-e^2}}{\sqrt{(1-e^2 \sin^2 \lambda)(1-e^2 \sin^2 \lambda')}} + \frac{e^2 \cos \lambda \sqrt{1-e^2}}{\sqrt{1-e^2 \sin^2 \lambda}} \left( \frac{\sin \lambda}{\sqrt{1-e^2 \sin^2 \lambda}} - \frac{\sin \lambda'}{\sqrt{1-e^2 \sin^2 \lambda'}} \right)}$$

from which it follows that

$$\tan \psi = \cos \lambda' \sin L / \left\{ \sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L + e^2 \cos \lambda \left( \sin \lambda \sqrt{\frac{1-e^2 \sin^2 \lambda'}{1-e^2 \sin^2 \lambda}} - \sin \lambda' \right) \right\}. \quad (50)$$

$\psi$  being the azimuthal angle of  $\phi$  from  $P$ .

The case of a sphere is found by putting  $e=0$ , when (50) becomes

$$\tan \psi_0 = \frac{\cos \lambda' \sin L}{\sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L} \dots \dots \dots (51)$$

which is the ordinary formula.

Let  $\psi = \psi_0 + \delta\psi$ : then  $\cot \overline{\psi_0 + \delta\psi} - \cot \psi_0 = -\frac{\sin \delta \psi}{\sin \psi \sin \psi_0}$

$$-\delta\psi = e^2 \frac{\sin^2 \psi_0 \cos \lambda}{\sin L \cos \lambda'} \left\{ \sin \lambda \sqrt{\frac{1-e^2 \sin^2 \lambda'}{1-e^2 \sin^2 \lambda}} - \sin \lambda' \right\} \dots \dots \dots (52)$$

This formula gives the correction to be applied to the azimuthal angle, found from the formula for a sphere, to obtain the spheroidal azimuth.

### CHAPTER III.

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#### Changes of coordinates of triangulated points, due to changes in axes of the terrestrial spheroid, calculated along geodesics.

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1. Consider a geodesic on any surface, and let  $A B C$  be three consecutive points on it. Then  $A B C$  is the osculating plane at  $B$  and from the fundamental property of the geodesic it contains the normal to the surface at  $B$ . This shows that measured from  $B$  the azimuths of  $A$  and  $C$  differ by two right angles. It is possible then to describe a geodesic on a surface of unknown form by fulfilling this condition: and, to take a practical case, a traverse along a geodesic can be observed on the earth without knowing its figure if access to a level surface is possible. It follows that if there is a geodesic on one surface which has been selected as representing the earth and it is desired to change to another surface, the geodesic on the first surface will, on transfer to the second surface, remain a geodesic. This property makes it possible to differentiate along a geodesic with respect to the constants of the first surface and so to find relations between changes in these constants and the quantities defining the position of points.

This fact will now be made use of in connection with the equations of Chapter II for the case of a spheroid. Now it has been shown in Chapter I that the effect of slightly changing the latitude and azimuth at the origin may be computed, and that the result is practically independent of the route followed—values of  $u_x$  and  $u_y$  etc. being identical to nearer than 0.001 of a second: and the resulting changes of latitude, longitude and azimuth for unit changes at the origin are given in tables XVII—XX. These values then are equally applicable to a geodesic and so it is only necessary to consider the effects of changes in  $a$  and  $b$ .

It is however found convenient not to alter the constant  $k$  of the geodesic; and since this is equal to  $\sin A \cos \phi$ , this is given effect to by not changing the values of  $A$  and  $\phi$  at the origin. It is moreover more convenient in dealing with geodesics on a spheroid to make use of the eccentric angle, or reduced latitude  $\phi$ , in place of the latitude  $\lambda$ . Now the relation between  $\lambda$  and  $\phi$  involves  $e^2$  and so if  $\phi$  is unchanged at the origin while  $e^2$  undergoes a change it follows that  $\lambda$  must also change at the origin. In the solution which follows the effect of this origin change of  $\lambda$  occurs: but as its amount can be found from tables XVII, XVIII there is no difficulty in removing it.

2. For convenience of reference equations (26), (25), (27), (2), (40), (42), (41) of Chapter II are repeated.

$$[L] = \pm \left[ \pm \psi - \frac{ke^2}{2} \theta \right] \dots \dots \dots (1)$$

$$\psi = \pm \tan^{-1} (\tan A \sin \phi) = \pm \sin^{-1} (\tan A_0 \tan \phi) \dots \dots \dots (2)$$

$$\tan \theta = \frac{1}{k} \sqrt{\frac{1-k^2e^2}{1-e^2}} \tan \psi \dots \dots \dots (3)$$

$$k = \sin A \cos \phi = \sin A_0 \dots \dots \dots (4)$$

$$s = -a \sqrt{1-e^2} \left[ \chi \left( 1 + \frac{1}{4} h^2 \dots \right) - \frac{h^2}{8} \sin 2\chi \left( 1 - \dots \right) + \dots \right] \frac{\chi}{\chi'} \dots (5)$$

$$\tan \chi = \pm \tan \phi \sec A = \pm \frac{1}{k} \tan \psi \dots \dots \dots (6)$$

$$h^2 = \frac{e^2 \cos^2 A_0}{1-e^2} \dots \dots \dots (7)$$

$$\sin \chi = \sin \phi \sec A_0 \dots \dots \dots (8)$$

The signs occurring in (1) and (2) are to be determined as explained in Chapter II. The sign of  $\chi$  is determined by (8). It is best to consider  $a$  and  $e^2$  as independent variables and  $b$  as dependent on these. At the end there is no difficulty in passing to the case of  $a$  and  $b$  considered as independent and  $e^2$  as the dependent variable.

3. Suppose then that  $a$  and  $e^2$  are changed, while the azimuth and reduced latitude  $\phi$  of the origin remain unchanged. Values at the origin will be denoted by dashes.

Differentiating (4) and keeping  $k$  constant

$$\delta k = 0 = \cos A \cos \phi \delta A - \sin A \sin \phi \delta \phi$$

Changes of *reduced latitude*, longitude and azimuth, in keeping with the notation of Chapter I, will be denoted by  $u_1, v, w$ : and the above equation may be written

$$w = \tan A \tan \phi. u_1 \dots \dots \dots (9)$$

Differentiating (1) and remembering that  $v' = 0$

$$\pm v = \left[ \delta \psi \right] - \frac{ke^2}{2} \left[ \delta \theta \right] - \frac{k \delta e^2}{2} \left[ \theta \right] \dots \dots \dots (10)$$

By differentiating logarithmically  $\tan \psi = \pm \tan A \sin \phi$  which is the same as (2): and by (3) and (6)

$$\frac{\delta \psi}{\sin \psi \cos \psi} = \frac{w}{\sin A \cos A} + u_1 \cot \phi = \frac{\delta \chi}{\sin \chi \cos \chi} = \frac{\delta \theta}{\sin \theta \cos \theta} - (1-k^2) \frac{\delta e^2}{2} \equiv x \dots (11)$$

any of these expressions being denoted by  $x$ . This quantity  $x$  vanishes at the origin.

Finally differentiate (5) keeping  $s$  constant, replacing  $\delta \chi$  by means of (11) and  $a \sqrt{1-e^2}$  by  $b$

$$\begin{aligned} \frac{s}{b} \left( -\frac{\delta a}{a} + \frac{\delta e^2}{2(1-e^2)} \right) &= -\frac{x}{2} \sin 2\chi \left\{ \left( 1 + \frac{h^2}{4} + \dots \right) - \frac{h^2}{4} \cos 2\chi (1 \dots) \dots \right\} - \frac{\delta h^2}{8} \left[ 2\chi (1 \dots) - \sin 2\chi (\dots) \right] \\ &\doteq -\frac{x}{2} \sin 2\chi \left\{ 1 + \frac{h^2}{4} (1 - \cos 2\chi) \right\} - \frac{\delta h^2}{8} \left[ 2\chi - \sin 2\chi \right] \dots \dots \dots (12) \end{aligned}$$

where

$$\delta h^2 = \frac{\cos^2 A_0}{(1-e^2)^2} \delta e^2 \dots \dots \dots (13)$$

Equations (12) and (13) serve to determine  $x$  in terms of  $\frac{\delta a}{a}$  and  $\delta e^2$ . For the other quantities from (9) and (11) it follows that

$$w = \tan A \tan \phi \cdot u_1 = \frac{x \tan A}{\sec^2 A + \cot^2 \phi} \dots \dots \dots (14)$$

and from (10) and (11)

$$\pm v = \frac{x}{2} \sin 2\psi - \frac{ke^2}{4} \left\{ [\sin 2\theta] (1-k^2) \frac{\delta e^2}{2} + x \sin 2\theta \right\} - \frac{k\delta e^2}{2} [\theta] \dots \dots \dots (15)$$

In all the above equations square brackets indicate that the quantity enclosed has to be taken between limits.

4. It remains to give the relation between  $u$  and  $u_1$   
From (4) of Chapter II

$$\tan \phi = \sqrt{1-e^2} \tan \lambda$$

Differentiating this logarithmically

$$\frac{u_1}{\sin \phi \cos \phi} = \frac{u}{\sin \lambda \cos \lambda} - \frac{1}{2} \frac{\delta e^2}{1-e^2}$$

or

$$u = \frac{\sin 2\lambda}{\sin 2\phi} \cdot u_1 + \frac{1}{4} \frac{\delta e^2}{1-e^2} \sin 2\lambda$$

Now

$$\begin{aligned} \frac{\sin 2\lambda}{\sin 2\phi} &= \frac{\sqrt{1-e^2}}{1-e^2 \cos^2 \phi} = 1 + \frac{e^2}{2} (2\cos^2 \phi - 1) \\ &= 1 + \frac{e^2}{2} \cos 2\phi \end{aligned}$$

so that

$$u = \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 + \frac{1}{4} \frac{\delta e^2}{1-e^2} \sin 2\phi \dots \dots \dots (16)$$

At the origin  $u_1' = 0$  : hence

$$u' = \frac{1}{4} \frac{\delta e^2}{1-e^2} \sin 2\phi' \dots \dots \dots (17)$$

5. It may be noticed that by this method the changes  $u, v, w$  appear to be found without any integration, whereas in Chapter I simultaneous differential equations occurred which had to be solved. The case under consideration is a particular case of the general equations (2) of Chapter I. The decision to follow a geodesic introduces a relation by which these equations can be reduced to total differential equations: and the integration of these equations would lead to the same results as may be obtained from equations (12) to (17). The same results will be seen to be obtainable by application to values of  $u_x, v_x, w_x$  of the appropriate closing errors. For the case now under consideration the equations formed in Chapter II give the results of integration: and so no further integration is necessary.

6. In making use of the equations (12) to (17) two cases are considered in which

(i)  $\delta a = 1$  km. and  $\delta e^2 = 0$

(ii)  $\delta a = 0$  and  $\delta e^2 = \cdot 0001$

The first of these corresponds to a combination of cases I and II of Chapter I, while the second corresponds to a combination of cases II and III. This arrangement simplifies computation and there is no difficulty in deriving cases I and II when the computations are complete.

As no azimuthal change is being made at the origin it is clear that there is symmetry about a central meridian. In Chapter II values of  $\psi, \theta, \chi, A, \frac{s}{b}, h^2, k$  (*vide* tables XXIII, XXIV) have already been given for every  $4^\circ$  of  $\phi$  from  $10^\circ$  to  $38^\circ$  and for longitude differences of  $4^\circ$  from  $4^\circ$  to  $24^\circ$ . With the help of these the values of  $u, v, w$  exhibited in the following two tables have been found. A double sign is prefixed to  $v$  and  $w$  and of these the upper or lower is to be taken according as the point is west or east of the origin. The results are given to three places of decimals as found by the computations: but the last figure is liable to error, which is not sufficiently large to be practically important for the present purpose.

TABLE XXV.

$$\delta a = 1 \text{ km.}, (\delta e^2 = 0) \quad \delta b = \frac{b}{a} = \cdot 9967 \text{ km.}$$

$\phi$	$L-L'$	4	8	12	16	20	24
38	$u_1$	- 7·834	- 7·710	- 7·510	- 7·227	- 6·863	- 6·415
	$u$	- 7·840	- 7·716	- 7·516	- 7·233	- 6·868	- 6·420
	$\pm v$	+ 2·644	+ 5·280	+ 7·914	+ 10·532	+ 13·143	+ 15·733
	$\pm w$	+ 1·630	+ 3·258	+ 4·892	+ 6·499	+ 8·106	+ 9·703
34	$u_1$	- 5·577	- 5·468	- 5·281	- 5·019	- 4·683	- 4·270
	$u$	- 5·584	- 5·475	- 5·288	- 5·025	- 4·689	- 4·275
	$\pm v$	+ 2·198	+ 4·996	+ 7·489	+ 9·967	+ 12·437	+ 14·894
	$\pm w$	+ 1·400	+ 2·799	+ 4·195	+ 5·586	+ 6·971	+ 8·348
30	$u_1$	- 3·325	- 3·222	- 3·052	- 2·812	- 2·502	- 2·123
	$u$	- 3·331	- 3·227	- 3·057	- 2·817	- 2·506	- 2·127
	$\pm v$	+ 2·334	+ 4·768	+ 7·143	+ 9·516	+ 11·871	+ 14·228
	$\pm w$	+ 1·195	+ 2·300	+ 3·581	+ 4·769	+ 5·951	+ 7·128
26	$u_1$	- 1·071	- 0·977	- 0·820	- 0·603	- 0·322	+ 0·024
	$u$	- 1·073	- 0·979	- 0·822	- 0·604	- 0·323	+ 0·024
	$\pm v$	+ 2·294	+ 4·586	+ 6·873	+ 9·152	+ 11·435	+ 13·718
	$\pm w$	+ 1·009	+ 2·016	+ 3·019	+ 4·024	+ 5·025	+ 6·024
22	$u_1$	+ 1·185	+ 1·268	+ 1·409	+ 1·606	+ 1·860	+ 2·172
	$u$	+ 1·188	+ 1·271	+ 1·412	+ 1·610	+ 1·864	+ 2·177
	$\pm v$	+ 2·225	+ 4·447	+ 6·666	+ 8·881	+ 11·088	+ 13·291
	$\pm w$	+ 0·836	+ 1·671	+ 2·504	+ 3·338	+ 4·164	+ 4·991
18	$u_1$	+ 3·439	+ 3·513	+ 3·639	+ 3·814	+ 4·044	+ 4·319
	$u$	+ 3·449	+ 3·522	+ 3·649	+ 3·824	+ 4·055	+ 4·331
	$\pm v$	+ 2·172	+ 4·343	+ 6·512	+ 8·674	+ 10·834	+ 12·981
	$\pm w$	+ 0·673	+ 1·346	+ 2·018	+ 2·689	+ 3·360	+ 4·024
14	$u_1$	+ 5·695	+ 5·760	+ 5·869	+ 6·024	+ 6·223	+ 6·467
	$u$	+ 5·712	+ 5·777	+ 5·886	+ 6·042	+ 6·241	+ 6·486
	$\pm v$	+ 2·135	+ 4·271	+ 6·401	+ 8·536	+ 10·660	+ 12·777
	$\pm w$	+ 0·519	+ 1·037	+ 1·554	+ 2·070	+ 2·587	+ 3·100
10	$u_1$	+ 7·952	+ 8·008	+ 8·102	+ 8·232	+ 8·404	+ 8·616
	$u$	+ 7·977	+ 8·033	+ 8·127	+ 8·258	+ 8·430	+ 8·643
	$\pm v$	+ 2·114	+ 4·229	+ 6·343	+ 8·451	+ 10·559	+ 12·662
	$\pm w$	+ 0·369	+ 0·737	+ 1·105	+ 1·472	+ 1·840	+ 2·206



TABLE XXVI.

$$\delta a = 0; (\delta e^2 = 0.0001); \delta b = -\frac{a^2 \delta e^2}{2b} = -0.3200 \text{ km.}; u_0 = 3'' \cdot 872.$$

$\phi$	$L - L'$	4°	8°	12°	16'	20°	24°
38°	$u_1$	+ 1.837	+ 1.807	+ 1.757	+ 1.689	+ 1.600	+ 1.493
	$u$	+ 6.881	+ 6.850	+ 6.800	+ 6.733	+ 6.643	+ 6.536
	$\pm v$	- 0.094	- 0.187	- 0.279	- 0.370	- 0.455	- 0.540
	$\pm w$	- 0.382	- 0.763	- 1.142	- 1.519	- 1.890	- 2.258
34°	$u_1$	+ 1.364	+ 1.336	+ 1.289	+ 1.223	+ 1.139	+ 1.037
	$u$	+ 6.185	+ 6.157	+ 6.111	+ 6.045	+ 5.960	+ 5.857
	$\pm v$	- 0.060	- 0.121	- 0.182	- 0.240	- 0.295	- 0.346
	$\pm w$	- 0.342	- 0.684	- 1.024	- 1.362	- 1.696	- 2.027
30°	$u_1$	+ 0.845	+ 0.818	+ 0.775	+ 0.713	+ 0.633	+ 0.536
	$u$	+ 5.350	+ 5.324	+ 5.280	+ 5.218	+ 5.138	+ 5.041
	$\pm v$	- 0.083	- 0.067	- 0.100	- 0.131	- 0.160	- 0.186
	$\pm w$	- 0.304	- 0.607	- 0.909	- 1.208	- 1.506	- 1.800
26°	$u_1$	+ 0.282	+ 0.257	+ 0.216	+ 0.158	+ 0.084	- 0.007
	$u$	+ 4.381	+ 4.357	+ 4.315	+ 4.258	+ 4.184	+ 4.096
	$\pm v$	- 0.011	- 0.024	- 0.033	- 0.041	- 0.049	- 0.056
	$\pm w$	- 0.265	- 0.530	- 0.794	- 1.057	- 1.317	- 1.579
22°	$u_1$	- 0.321	- 0.344	- 0.382	- 0.435	- 0.504	- 0.587
	$u$	+ 3.294	+ 3.271	+ 3.233	+ 3.180	+ 3.111	+ 3.027
	$\pm v$	+ 0.008	+ 0.015	+ 0.023	+ 0.032	+ 0.043	+ 0.055
	$\pm w$	- 0.227	- 0.454	- 0.679	- 0.904	- 1.127	- 1.349
18°	$u_1$	- 0.960	- 0.980	- 1.015	- 1.063	- 1.126	- 1.201
	$u$	+ 2.098	+ 2.078	+ 2.043	+ 1.995	+ 1.932	+ 1.856
	$\pm v$	+ 0.024	+ 0.046	+ 0.068	+ 0.092	+ 0.116	+ 0.143
	$\pm w$	- 0.188	- 0.375	- 0.563	- 0.749	- 0.935	- 1.119
14°	$u_1$	- 1.628	- 1.646	- 1.677	- 1.720	- 1.775	- 1.844
	$u$	+ 0.812	+ 0.794	+ 0.762	+ 0.719	+ 0.664	+ 0.596
	$\pm v$	+ 0.036	+ 0.069	+ 0.102	+ 0.134	+ 0.169	+ 0.205
	$\pm w$	- 0.148	- 0.296	- 0.444	- 0.591	- 0.738	- 0.884
10°	$u_1$	- 2.322	- 2.337	- 2.364	- 2.401	- 2.450	- 2.508
	$u$	- 0.547	- 0.562	- 0.590	- 0.627	- 0.676	- 0.734
	$\pm v$	+ 0.042	+ 0.083	+ 0.122	+ 0.162	+ 0.202	+ 0.244
	$\pm w$	- 0.108	- 0.215	- 0.322	- 0.429	- 0.536	- 0.642

By the help of tables XVII and XVIII of Chapter I and their extensions (*vide* tables XXXV, XXXVI below) the effect of  $u_0$  is eliminated: the figures in these tables give by interpolation values of  $u$  for each point in table XXVI, and these multiplied by  $-3 \cdot 872$  are applied to the corresponding figures in XXVI, leaving residuals due to case  $\delta b = -0.3200$ . The residuals are multiplied by  $-\frac{1}{.32}$  and the result is the case II,—  $\delta b = 1, \delta a = 0$ . This case is then multiplied by  $-0.9967$  and applied to table XXV, leaving as a residual case I,—  $\delta a = 1, \delta b = 0$ . From these tables values corresponding to even degrees of  $\lambda$  are interpolated and the results are exhibited in tables XXVII, XXVIII.

TABLE XXVII.

 $\delta a = 1 \text{ km}, \delta b = 0.$ 

$\lambda$	$L-L'$	$0^\circ$	$4^\circ$	$8^\circ$	$12^\circ$	$16^\circ$	$20^\circ$	$24^\circ$
$38^\circ$	$u$	+ 1.524	+ 1.567	+ 1.684	+ 1.873	+ 2.148	+ 2.492	+ 2.923
	$\pm v^*$	0.000	+ 3.007	+ 6.006	+ 8.996	+ 11.965	+ 14.935	+ 17.871
	$\pm w$	0.000	+ 1.504	+ 3.005	+ 4.495	+ 5.972	+ 7.436	+ 8.873
$34^\circ$	$u$	+ 1.608	+ 1.645	+ 1.755	+ 1.942	+ 2.203	+ 2.534	+ 2.944
	$\pm v$	0.000	+ 2.878	+ 5.750	+ 8.609	+ 11.458	+ 14.294	+ 17.114
	$\pm w$	0.000	+ 1.346	+ 2.688	+ 4.019	+ 5.339	+ 6.645	+ 7.930
$30^\circ$	$u$	+ 1.253	+ 1.291	+ 1.467	+ 1.579	+ 1.828	+ 2.150	+ 2.541
	$\pm v$	0.000	+ 2.769	+ 5.521	+ 8.277	+ 11.024	+ 13.753	+ 16.474
	$\pm w$	0.000	+ 1.218	+ 2.436	+ 3.637	+ 4.836	+ 6.013	+ 7.168
$26^\circ$	$u$	+ 0.490	+ 0.525	+ 0.631	+ 0.803	+ 1.044	+ 1.351	+ 1.730
	$\pm v$	0.000	+ 2.670	+ 5.334	+ 7.988	+ 10.645	+ 13.290	+ 15.926
	$\pm w$	0.000	+ 1.119	+ 2.229	+ 3.330	+ 4.422	+ 5.503	+ 6.563
$22^\circ$	$u$	- 0.646	- 0.606	- 0.506	- 0.338	- 0.103	+ 0.194	+ 0.562
	$\pm v$	0.000	+ 2.592	+ 5.170	+ 7.751	+ 10.325	+ 12.885	+ 15.447
	$\pm w$	0.000	+ 1.038	+ 2.066	+ 3.089	+ 4.101	+ 5.095	+ 6.068
$18^\circ$	$u$	- 2.111	- 2.071	- 1.972	- 1.809	- 1.581	- 1.288	- 0.932
	$\pm v$	0.000	+ 2.523	+ 5.034	+ 7.539	+ 10.040	+ 12.535	+ 15.018
	$\pm w$	0.000	+ 0.971	+ 1.941	+ 2.899	+ 3.844	+ 4.776	+ 5.686
$14^\circ$	$u$	- 3.850	- 3.812	- 3.715	- 3.559	- 3.333	- 3.049	- 2.701
	$\pm v$	0.000	+ 2.454	+ 4.906	+ 7.344	+ 9.782	+ 12.216	+ 14.637
	$\pm w$	0.000	+ 0.927	+ 1.843	+ 2.753	+ 3.651	+ 4.534	+ 5.395
$10^\circ$	$u$	- 5.812	- 5.776	- 5.679	- 5.526	- 5.306	- 5.028	- 4.681
	$\pm v$	0.000	+ 2.392	+ 4.785	+ 7.163	+ 9.539	+ 11.916	+ 14.289
	$\pm w$	0.000	+ 0.890	+ 1.773	+ 2.643	+ 3.504	+ 4.355	+ 5.177

\* + or - according as point is west or east of origin.

The justification for this method of computing along a geodesic is that a point is reached by the shortest available route from the origin. Now in practice a point's position relative to the origin is determined by a network of triangulation: and in general the geodesic takes a more or less medial course through this network. When this is not the case a modified procedure becomes desirable, and this occurs in the Burma triangulation. Geodesics from Kaliānpur to Lower Burma cross the Bay of Bengal. As a single geodesic is thus unsatisfactory the next simplest course seems to be to follow a geodesic from Kaliānpur to Chittagong and then, taking a different geodesic, to proceed to any desired point. To give effect to this, recomputation of quantities occurring above would be necessary: but by a slight modification this can be avoided. Choose as first geodesic the one from Kaliānpur which cuts the meridian  $91^\circ 39' 17'' \cdot 57$  i.e.  $14^\circ$  east of this meridian of the origin at the same latitude as the origin,  $24^\circ 7' 11''$ . This point may be regarded as a second

TABLE XXVIII.

$\delta b = 1 \text{ km}, \delta a = 0.$

$\lambda$	$L-L'$	$0^\circ$	$4^\circ$	$8^\circ$	$12^\circ$	$16^\circ$	$20^\circ$	$24^\circ$
$35^\circ$	$u$	- 9.383	- 9.386	- 9.379	- 9.368	- 9.362	- 9.341	- 9.324
	$\pm v^*$	0.000	- 0.367	- 0.735	- 1.096	- 1.453	- 1.816	- 2.166
	$\pm w$	0.000	+ 0.120	+ 0.243	+ 0.371	+ 0.507	+ 0.645	+ 0.800
$34^\circ$	$u$	- 7.205	- 7.203	- 7.204	- 7.205	- 7.203	- 7.198	- 7.194
	$\pm v$	0.000	- 0.384	- 0.761	- 1.132	- 1.508	- 1.877	- 2.244
	$\pm w$	0.000	+ 0.049	+ 0.102	+ 0.163	+ 0.229	+ 0.304	+ 0.391
$30^\circ$	$u$	- 4.587	- 4.591	- 4.597	- 4.605	- 4.615	- 4.626	- 4.638
	$\pm v$	0.000	- 0.388	- 0.766	- 1.144	- 1.522	- 1.898	- 2.266
	$\pm w$	0.000	- 0.027	- 0.054	- 0.068	- 0.083	- 0.082	- 0.065
$26^\circ$	$u$	- 1.556	- 1.560	- 1.572	- 1.587	- 1.613	- 1.639	- 1.673
	$\pm v$	0.000	- 0.379	- 0.753	- 1.123	- 1.503	- 1.868	- 2.241
	$\pm w$	0.000	- 0.114	- 0.220	- 0.322	- 0.410	- 0.496	- 0.560
$22^\circ$	$u$	+ 1.848	+ 1.838	+ 1.821	+ 1.794	+ 1.756	+ 1.712	+ 1.657
	$\pm v$	0.000	- 0.369	- 0.727	- 1.092	- 1.453	- 1.808	- 2.169
	$\pm w$	0.000	- 0.206	- 0.401	- 0.595	- 0.776	- 0.948	- 1.096
$18^\circ$	$u$	+ 5.588	+ 5.568	+ 5.543	+ 5.507	+ 5.453	+ 5.391	+ 5.310
	$\pm v$	0.000	- 0.353	- 0.695	- 1.032	- 1.374	- 1.710	- 2.048
	$\pm w$	0.000	- 0.301	- 0.601	- 0.890	- 1.169	- 1.432	- 1.680
$14^\circ$	$u$	+ 9.596	+ 9.581	+ 9.549	+ 9.502	+ 9.431	+ 9.346	+ 9.242
	$\pm v$	0.000	- 0.321	- 0.637	- 0.947	- 1.251	- 1.562	- 1.866
	$\pm w$	0.000	- 0.410	- 0.813	- 1.208	- 1.593	- 1.962	- 2.314
$10^\circ$	$u$	+13.833	+13.816	+13.775	+12.716	+13.626	+13.520	+13.385
	$+v$	0.000	- 0.279	- 0.558	- 0.822	- 1.092	- 1.361	- 1.632
	$\pm w$	0.000	- 0.524	- 1.041	- 1.547	- 2.044	- 2.530	- 2.989

\* + or - according as point is west or east of origin.

origin for Lower Burma and the tables already calculated can be made use of. The difference from going *via* Chittagong caused by this is less than 0.01 second in latitude and longitude and 0.1 second of azimuth as may be seen by considering the closing errors shown in tables II, III, IV. Corresponding to case I,  $\delta a = 1 \text{ km}$  at Kaliānpur, cases I, III, IV arise at the second origin: but the latter two cases are readily eliminated by tables XVII to XX, remembering that the values of  $L$  in these tables are now naturally to be increased by  $14^\circ$ . In this way tables, under heading  $u_g, v_g$  or  $w_g$  as the case may be (suffix  $g$  indicating that the route followed is a geodesic) giving the change of coordinates due to change of axes and origin for the four cases, have been prepared for the whole of India and Burma and are now given (tables XXIX-XXXIV). Those for cases III and IV have already been given for India up to longitude  $94^\circ$  and now only the additions necessary for Burma are given (*vide* tables XXXV, XXXVI).

THE EARTH'S AXES AND TRIANGULATION.

Case I.— $\delta a = 1$  km.

Values of  $u_y$   
in seconds.

TABLE XXIX.

Long. Lat.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long. Lat.	
40°													1.392	1.370	1.351	1.335	1.321							40°
39													1.514	1.401	1.472	1.456	1.443							39
38													1.606	1.582	1.563	1.547	1.535							38
37													1.669	1.644	1.625	1.609	1.598							37
36								1.846	1.803	1.764	1.731	1.701	1.677	1.658	1.642	1.631	1.624	1.623	1.627	1.638	1.656			36
35								1.840	1.806	1.768	1.735	1.706	1.682	1.663	1.647	1.636	1.630	1.629	1.633	1.644	1.661			35
34								1.824	1.782	1.744	1.711	1.683	1.659	1.640	1.624	1.614	1.609	1.608	1.612	1.622	1.638			34
33								1.775	1.733	1.695	1.662	1.634	1.610	1.591	1.575	1.565	1.561	1.560	1.564	1.574	1.589			33
32	2.197	2.119	2.046	1.977	1.912	1.852	1.796	1.746	1.700	1.658	1.620	1.587	1.550	1.535	1.516	1.500	1.490	1.485	1.484	1.488	1.498	1.514		32
31	2.089	2.012	1.939	1.871	1.807	1.747	1.692	1.642	1.596	1.555	1.518	1.485	1.457	1.433	1.414	1.398	1.387	1.382	1.381	1.385	1.395	1.411		31
30	1.954	1.877	1.805	1.737	1.674	1.616	1.562	1.512	1.467	1.426	1.390	1.357	1.329	1.305	1.286	1.270	1.259	1.253	1.253	1.256	1.267	1.282		30
29	1.795	1.720	1.647	1.580	1.517	1.450	1.405	1.356	1.312	1.271	1.235	1.203	1.175	1.151	1.132	1.116	1.105	1.100	1.099	1.103	1.112	1.127		29
28	1.608	1.533	1.463	1.397	1.334	1.277	1.223	1.175	1.132	1.090	1.054	1.022	0.994	0.971	0.953	0.937	0.926	0.921	0.920	0.924	0.933	0.948		28
27	1.398	1.324	1.255	1.189	1.127	1.070	1.017	0.969	0.924	0.885	0.849	0.816	0.787	0.765	0.748	0.733	0.723	0.718	0.717	0.721	0.730	0.744		27
26	1.164	1.091	1.022	0.956	0.898	0.841	0.787	0.739	0.695	0.656	0.620	0.589	0.562	0.538	0.520	0.506	0.496	0.491	0.490	0.494	0.502	0.516		26
25	0.910	0.838	0.769	0.704	0.644	0.587	0.534	0.487	0.444	0.405	0.369	0.339	0.312	0.280	0.270	0.256	0.246	0.241	0.240	0.244	0.253	0.266		25
24								0.213	0.171	0.132	0.096	0.066	0.039	0.018	0.002	0.016	0.026	0.033	0.034	0.028	0.019	0.007		24
23								0.084	0.126	0.165	0.200	0.229	0.255	0.278	0.296	0.310	0.321	0.328	0.329	0.324	0.313	0.302		23
22								0.401	0.443	0.482	0.516	0.546	0.572	0.593	0.611	0.626	0.635	0.645	0.646	0.640	0.630	0.617		22
21								0.710	0.782	0.821	0.854	0.884	0.909	0.931	0.950	0.964	0.975	0.983	0.984	0.980	0.969	0.956		21
20													1.267	1.288	1.307	1.322	1.333	1.340	1.341	1.337	1.327	1.312		20
19													1.643	1.664	1.682	1.697	1.708	1.716	1.717	1.713	1.702	1.688		19
18													2.037	2.058	2.076	2.091	2.103	2.110	2.111	2.106	2.095	2.083		18
17													2.440	2.470	2.489	2.505	2.516	2.522	2.523	2.519	2.510	2.495		17
16													2.877	2.898	2.916	2.931	2.943	2.949	2.950	2.946	2.937	2.923		16
15													3.321	3.342	3.360	3.374	3.385	3.391	3.392	3.388	3.380	3.367		15
14													3.770	3.800	3.817	3.831	3.842	3.849	3.850	3.846	3.837	3.825		14
13													4.252	4.273	4.291	4.305	4.315	4.321	4.322	4.318	4.310	4.299		13
12													4.737	4.758	4.776	4.791	4.800	4.807	4.808	4.804	4.796	4.783		12
11													5.235	5.256	5.274	5.288	5.298	5.303	5.304	5.300	5.292	5.281		11
10													5.742	5.763	5.781	5.795	5.805	5.811	5.812	5.808	5.800	5.788		10
9													6.260	6.281	6.299	6.313	6.323	6.329	6.330	6.326	6.318	6.306		9
8													6.787	6.808	6.826	6.840	6.850	6.857	6.858	6.854	6.846	6.833		8

The sign is + above the horizontal dividing line and - below it.

The sign is + above the horizontal dividing line and - below it.

Case I.—δa = 1 km.

Values of v<sub>y</sub>  
in seconds.

TABLE XXX.

Table with 23 columns for longitude (60° to 102°) and 23 rows for latitude (40° to 8°). The table is divided into 'Positive' and 'Negative' regions. The 'Positive' region covers longitudes 60°-81° and latitudes 40°-21°. The 'Negative' region covers longitudes 81°-102° and latitudes 40°-21°. Data values are provided in seconds for each grid intersection.









in seconds.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.	
Lat.	P o s i t i v e																				N e g a t i v e			Lat.
40°													0.216	0.177	0.130	0.102	0.066						40°	
39													0.194	0.159	0.125	0.091	0.058						39	
38													0.171	0.140	0.110	0.080	0.050						38	
37													0.147	0.121	0.094	0.068	0.043						37	
36								0.214	0.100	0.167	0.114	0.122	0.100	0.078	0.056	0.035	0.014	0.007	0.029	0.049	0.072		36	
35								0.171	0.151	0.133	0.115	0.097	0.079	0.062	0.041	0.027	0.010	0.008	0.022	0.030	0.054		35	
34								0.127	0.112	0.097	0.084	0.071	0.058	0.045	0.032	0.020	0.008	0.004	0.016	0.028	0.040		34	
33								0.081	0.071	0.061	0.052	0.044	0.036	0.028	0.020	0.012	0.005	0.002	0.010	0.018	0.026		33	
32	0.091	0.082	0.074	0.067	0.061	0.055	0.049	0.042	0.036	0.030	0.025	0.021	0.017	0.014	0.011	0.007	0.004	0.001	0.001	0.003	0.006	0.009	32	
31	0.004	0.000	0.004	0.006	0.007	0.008	0.009	0.010	0.011	0.013	0.013	0.012	0.010	0.008	0.006	0.005	0.003	0.001	0.001	0.003	0.004	0.006	31	
30	0.084	0.084	0.082	0.080	0.076	0.071	0.067	0.064	0.061	0.057	0.052	0.046	0.040	0.032	0.025	0.019	0.013	0.005	0.003	0.016	0.017	0.023	30	
29	0.173	0.167	0.160	0.152	0.144	0.135	0.126	0.118	0.110	0.101	0.091	0.080	0.069	0.057	0.044	0.033	0.021	0.000	0.004	0.017	0.029	0.041	29	
28	0.203	0.252	0.241	0.229	0.215	0.201	0.187	0.173	0.159	0.145	0.130	0.114	0.099	0.081	0.064	0.047	0.030	0.012	0.006	0.024	0.042	0.059	28	
27	0.354	0.338	0.321	0.305	0.287	0.268	0.249	0.230	0.211	0.191	0.170	0.150	0.128	0.106	0.084	0.062	0.039	0.016	0.008	0.032	0.055	0.078	27	
26	0.448	0.426	0.404	0.382	0.360	0.337	0.313	0.288	0.263	0.237	0.211	0.185	0.159	0.132	0.104	0.076	0.048	0.019	0.010	0.039	0.068	0.096	26	
25	0.546	0.510	0.491	0.463	0.435	0.407	0.378	0.347	0.316	0.285	0.253	0.222	0.190	0.158	0.126	0.092	0.058	0.023	0.012	0.047	0.082	0.115	25	
24								0.408	0.371	0.334	0.296	0.259	0.222	0.184	0.146	0.107	0.067	0.027	0.014	0.055	0.094	0.133	24	
23								0.469	0.426	0.383	0.341	0.298	0.255	0.212	0.168	0.123	0.077	0.031	0.016	0.063	0.100	0.153	23	
22								0.530	0.483	0.433	0.384	0.336	0.288	0.239	0.188	0.137	0.086	0.034	0.018	0.070	0.122	0.173	22	
21								0.504	0.540	0.485	0.430	0.376	0.322	0.267	0.210	0.153	0.096	0.038	0.020	0.078	0.136	0.193	21	
20													0.356	0.294	0.232	0.169	0.106	0.042	0.022	0.086	0.150	0.212	20	
19													0.391	0.322	0.254	0.185	0.116	0.046	0.024	0.094	0.164	0.232	19	
18													0.426	0.351	0.276	0.201	0.126	0.050	0.026	0.102	0.178	0.253	18	
17													0.462	0.381	0.300	0.219	0.137	0.054	0.023	0.111	0.193	0.275	17	
16													0.500	0.412	0.325	0.237	0.148	0.059	0.031	0.121	0.209	0.297	16	
15													0.539	0.444	0.350	0.265	0.169	0.063	0.033	0.129	0.225	0.320	15	
14													0.578	0.477	0.375	0.273	0.170	0.067	0.035	0.138	0.241	0.343	14	
13													0.618	0.510	0.400	0.291	0.182	0.072	0.038	0.148	0.257	0.366	13	
12													0.658	0.543	0.426	0.310	0.194	0.077	0.040	0.157	0.274	0.390	12	
11													0.698	0.578	0.453	0.329	0.206	0.082	0.043	0.167	0.291	0.414	11	
10													0.739	0.609	0.479	0.349	0.218	0.087	0.046	0.177	0.308	0.436	10	
9													0.780	0.642	0.505	0.360	0.230	0.092	0.049	0.187	0.325	0.464	9	
8													0.821	0.675	0.531	0.389	0.242	0.097	0.052	0.197	0.342	0.489	8	

The signs below the horizontal dividing line are opposite to those above it.

Long.	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°	Long.
Lat.	P o s i t i v e																				Lat.		
30°	0.023	0.030	0.037	0.044	0.050	0.055	0.059	0.063	0.066	0.070	0.074	0.078	0.081	0.083	0.084	0.084	0.083	0.081	0.079	0.074	0.069	0.063	30°
29	0.041	0.052	0.065	0.077	0.088	0.098	0.107	0.116	0.124	0.132	0.141	0.150	0.158	0.165	0.171	0.176	0.180	0.184	0.186	0.187	0.186	0.185	29
28	0.059	0.076	0.093	0.110	0.125	0.140	0.155	0.169	0.183	0.197	0.211	0.224	0.237	0.249	0.260	0.270	0.279	0.288	0.295	0.301	0.305	0.309	28
27	0.078	0.100	0.122	0.143	0.164	0.184	0.205	0.225	0.244	0.263	0.281	0.299	0.317	0.334	0.350	0.365	0.380	0.394	0.406	0.417	0.427	0.436	27
26	0.096	0.123	0.150	0.177	0.209	0.229	0.255	0.281	0.305	0.330	0.353	0.376	0.398	0.420	0.442	0.463	0.483	0.502	0.520	0.536	0.551	0.565	26
25	0.115	0.148	0.180	0.212	0.244	0.275	0.307	0.338	0.368	0.398	0.427	0.455	0.482	0.510	0.537	0.564	0.590	0.614	0.636	0.657	0.677	0.697	25
24	0.133	0.172	0.210	0.248	0.285	0.322	0.360	0.396	0.432	0.468	0.502	0.528	0.568	0.605	0.643	0.680	0.717	0.753	0.788	0.823	0.858	0.891	24
23	0.153	0.198	0.242	0.285	0.327	0.370	0.413	0.457	0.499	0.539	0.579	0.527	0.572	0.616	0.659	0.703	0.745	0.786	0.826	0.866	0.907	0.947	23
22	0.173	0.223	0.273	0.321	0.369	0.418	0.467	0.515	0.563	0.611	0.658	0.524	0.575	0.625	0.675	0.724	0.772	0.818	0.864	0.911	0.957	1.003	22
21	0.193	0.249	0.305	0.359	0.413	0.468	0.523	0.577			0.737	0.823	0.880	0.936	0.991	1.047	1.101	1.154	1.206	1.257	1.307	1.356	21
20	0.212	0.274	0.336	0.398	0.459	0.520	0.581	0.642			0.818	0.922	0.985	1.048	1.108	1.169	1.228	1.285	1.341	1.395	1.448	1.500	20
19	0.232	0.300	0.360	0.438	0.505						0.900	1.020	1.088	1.156	1.222	1.286	1.348	1.408	1.466	1.522	1.576	1.629	19
18	0.253	0.328	0.403	0.478	0.552						0.980	1.110	1.180	1.248	1.314	1.378	1.439	1.498	1.555	1.610	1.663	1.715	18
17	0.275	0.356	0.437	0.518	0.600						1.060	1.200	1.270	1.336	1.398	1.456	1.511	1.563	1.612	1.659	1.704	1.747	17
16	0.297	0.385	0.472	0.560	0.648						1.140	1.290	1.360	1.424	1.483	1.537	1.586	1.631	1.672	1.710	1.746	1.780	16
15											1.220	1.380	1.450	1.512	1.568	1.618	1.663	1.704	1.741	1.774	1.803	1.829	15
14											1.300	1.470	1.540	1.600	1.654	1.702	1.744	1.781	1.813	1.841	1.865	1.885	14
13											1.380	1.560	1.630	1.688	1.740	1.786	1.826	1.860	1.889	1.913	1.933	1.949	13
12											1.460	1.650	1.720	1.776	1.826	1.869	1.905	1.934	1.958	1.977	1.992	1.999	12
11											1.540	1.750	1.820	1.874	1.922	1.961	1.991	2.011	2.021	2.029	2.029	2.029	11
10											1.620	1.850	1.920	1.972	2.018	2.056	2.085	2.104	2.113	2.113	2.113	2.113	10
9											1.700	1.950	2.020	2.070	2.114	2.151	2.179	2.198	2.207	2.207	2.207	2.207	9
8											1.780	2.050	2.120	2.168	2.205	2.232	2.249	2.256	2.256	2.256	2.256	2.256	8

Case III.— $u_0=1''$ .

Values of  $u, v, w$   
in seconds.

TABLE XXXV\*.

Lat.	Long.	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°
Values of $u$ (positive).														
29°-8°		0.977	0.973	0.969	0.964	0.959	0.954	0.949	0.944	0.938	0.932	0.926	0.919	0.912
Values of $v$ .														
29°	Positive	0.119	0.128	0.137	0.147	0.156	0.165	0.175	0.184	0.193	0.202	0.211	0.219	0.228
28		0.114	0.123	0.132	0.141	0.150	0.159	0.167	0.176	0.185	0.194	0.202	0.211	0.219
27		0.109	0.118	0.126	0.135	0.144	0.152	0.160	0.169	0.177	0.185	0.195	0.202	0.210
26		0.104	0.113	0.121	0.129	0.137	0.145	0.153	0.161	0.169	0.177	0.185	0.193	0.201
25		0.100	0.108	0.116	0.124	0.131	0.138	0.146	0.153	0.161	0.169	0.176	0.185	0.192
24		0.095	0.103	0.110	0.118	0.125	0.133	0.140	0.147	0.155	0.162	0.169	0.176	0.183
23		0.091	0.098	0.105	0.112	0.119	0.126	0.134	0.141	0.147	0.154	0.161	0.168	0.175
22		0.087	0.093	0.100	0.107	0.114	0.120	0.127	0.134	0.140	0.147	0.153	0.160	0.166
21		0.082	0.089	0.095	0.102	0.108	0.114	0.120	0.127	0.133	0.140	0.146	0.152	0.158
20		0.078	0.084	0.090	0.096	0.102	0.108	0.114	0.120	0.126	0.132	0.138	0.144	0.150
19		0.074	0.079	0.085	0.091	0.097	0.102	0.108	0.114	0.120	0.125	0.131	0.136	0.142
18		0.069	0.075	0.080	0.086	0.091	0.097	0.102	0.107	0.113	0.118	0.123	0.129	0.134
17		0.065	0.071	0.076	0.081	0.086	0.091	0.096	0.101	0.106	0.111	0.116	0.121	0.126
16		0.061	0.066	0.071	0.076	0.081	0.085	0.090	0.095	0.100	0.104	0.109	0.114	0.118
15		0.057	0.062	0.066	0.071	0.075	0.080	0.084	0.089	0.093	0.098	0.102	0.106	0.110
14		0.053	0.058	0.062	0.066	0.070	0.074	0.079	0.083	0.087	0.091	0.095	0.099	0.103
13	0.049	0.053	0.057	0.061	0.065	0.069	0.073	0.076	0.080	0.084	0.088	0.091	0.095	
12	0.045	0.049	0.053	0.056	0.060	0.063	0.067	0.070	0.074	0.077	0.081	0.084	0.088	
11	0.041	0.045	0.048	0.051	0.055	0.058	0.061	0.064	0.068	0.071	0.074	0.077	0.080	
10	0.038	0.041	0.044	0.047	0.050	0.053	0.055	0.058	0.061	0.064	0.067	0.070	0.073	
9	0.034	0.037	0.040	0.043	0.046	0.048	0.050	0.052	0.055	0.057	0.060	0.063	0.065	
8	0.030	0.033	0.036	0.039	0.042	0.044	0.045	0.047	0.049	0.051	0.053	0.056	0.058	
Values of $w$ .														
29°	Positive	0.245	0.264	0.283	0.303	0.322	0.341	0.360	0.379	0.398	0.416	0.435	0.453	0.471
28		0.242	0.261	0.281	0.300	0.319	0.338	0.357	0.375	0.394	0.412	0.431	0.449	0.467
27		0.240	0.259	0.278	0.297	0.316	0.334	0.353	0.372	0.390	0.408	0.427	0.444	0.462
26		0.238	0.257	0.276	0.294	0.313	0.332	0.350	0.369	0.387	0.405	0.423	0.441	0.458
25		0.236	0.255	0.273	0.292	0.311	0.329	0.347	0.365	0.383	0.401	0.419	0.437	0.454
24		0.234	0.253	0.271	0.290	0.308	0.326	0.344	0.363	0.380	0.398	0.416	0.433	0.451
23		0.232	0.251	0.269	0.287	0.306	0.324	0.342	0.360	0.378	0.395	0.413	0.430	0.448
22		0.231	0.249	0.267	0.285	0.304	0.322	0.339	0.357	0.375	0.392	0.410	0.427	0.444
21		0.229	0.247	0.265	0.284	0.302	0.319	0.337	0.355	0.372	0.390	0.407	0.424	0.441
20		0.228	0.246	0.264	0.282	0.300	0.317	0.335	0.352	0.370	0.387	0.404	0.421	0.438
19		0.227	0.244	0.262	0.280	0.298	0.315	0.333	0.351	0.368	0.385	0.402	0.419	0.436
18		0.225	0.242	0.261	0.278	0.296	0.313	0.331	0.348	0.366	0.383	0.400	0.416	0.433
17		0.224	0.241	0.259	0.277	0.295	0.312	0.329	0.347	0.364	0.381	0.398	0.414	0.431
16		0.222	0.240	0.258	0.275	0.293	0.310	0.327	0.345	0.362	0.379	0.395	0.412	0.429
15		0.221	0.239	0.256	0.274	0.291	0.309	0.326	0.343	0.360	0.377	0.393	0.410	0.427
14		0.220	0.238	0.255	0.273	0.290	0.307	0.324	0.341	0.358	0.375	0.392	0.408	0.425
13	0.219	0.237	0.254	0.271	0.289	0.306	0.323	0.340	0.357	0.373	0.390	0.406	0.423	
12	0.218	0.236	0.253	0.270	0.288	0.305	0.322	0.339	0.355	0.372	0.388	0.405	0.421	
11	0.218	0.235	0.252	0.270	0.287	0.304	0.321	0.338	0.354	0.371	0.387	0.403	0.420	
10	0.217	0.234	0.251	0.269	0.286	0.303	0.320	0.337	0.353	0.369	0.386	0.402	0.418	
9	0.216	0.233	0.250	0.268	0.285	0.302	0.319	0.336	0.352	0.368	0.385	0.401	0.417	
8	0.215	0.232	0.249	0.267	0.284	0.301	0.318	0.335	0.351	0.367	0.384	0.400	0.416	

\* Extension of tables XVII, XVIII for Burma and Assam.

Case IV.  $-w_0=1''$ . Values of  $u, v, w$  in seconds. TABLE XXXVI\*.

Lat.	Long.	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°
Values of $u$ (negative).														
29° - 8'		0.196	0.212	0.227	0.243	0.258	0.273	0.288	0.303	0.318	0.333	0.348	0.363	0.377
Values of $v$ .														
29°	Positive	0.085	0.082	0.081	0.078	0.076	0.074	0.071	0.068	0.065	0.062	0.059	0.056	0.052
28°		0.065	0.063	0.061	0.059	0.057	0.055	0.052	0.049	0.046	0.043	0.040	0.037	0.033
27°		0.045	0.043	0.041	0.039	0.037	0.035	0.032	0.030	0.027	0.024	0.021	0.018	0.014
26°		0.026	0.024	0.022	0.020	0.018	0.016	0.013	0.011	0.008	0.005	0.003	0.000	0.003
25°	Positive	0.007	0.005	0.004	0.002	0.000	0.002	0.004	0.006	0.009	0.012	0.014	0.018	0.020
24°		0.012	0.013	0.015	0.017	0.019	0.021	0.023	0.025	0.028	0.031	0.033	0.036	0.039
23°		0.030	0.032	0.033	0.035	0.037	0.039	0.041	0.043	0.046	0.048	0.051	0.054	0.057
22°		0.048	0.050	0.051	0.053	0.055	0.057	0.059	0.061	0.063	0.065	0.068	0.071	0.074
21°		0.066	0.068	0.069	0.071	0.072	0.074	0.076	0.078	0.080	0.082	0.085	0.088	0.090
20°		0.084	0.085	0.086	0.088	0.090	0.091	0.093	0.095	0.097	0.099	0.102	0.104	0.106
19°		0.100	0.102	0.104	0.105	0.107	0.108	0.110	0.112	0.114	0.116	0.118	0.120	0.122
18°		0.118	0.119	0.121	0.122	0.124	0.125	0.127	0.129	0.130	0.132	0.134	0.136	0.138
17°		0.135	0.136	0.138	0.139	0.140	0.142	0.143	0.145	0.146	0.148	0.150	0.152	0.154
16°		0.151	0.152	0.154	0.155	0.157	0.158	0.159	0.161	0.163	0.164	0.166	0.168	0.170
15°		0.167	0.169	0.170	0.172	0.173	0.174	0.175	0.177	0.178	0.179	0.181	0.183	0.185
14°		0.184	0.185	0.187	0.188	0.189	0.190	0.192	0.193	0.194	0.196	0.197	0.199	0.200
13°		0.201	0.202	0.204	0.205	0.206	0.207	0.208	0.209	0.210	0.212	0.213	0.214	0.215
12°		0.216	0.217	0.219	0.220	0.221	0.222	0.223	0.224	0.225	0.227	0.228	0.229	0.230
11°		0.232	0.233	0.235	0.236	0.237	0.238	0.239	0.240	0.241	0.242	0.243	0.244	0.245
10°		0.248	0.249	0.251	0.252	0.252	0.253	0.254	0.255	0.256	0.257	0.258	0.259	0.260
9°		0.264	0.265	0.266	0.267	0.267	0.268	0.269	0.270	0.271	0.272	0.273	0.274	0.275
8°		0.279	0.280	0.281	0.282	0.282	0.283	0.284	0.285	0.286	0.287	0.288	0.289	0.290
Values of $w$ .														
29°	Positive	1.019	1.015	1.011	1.006	1.001	0.996	0.990	0.984	0.978	0.972	0.965	0.958	0.950
28°		1.009	1.005	1.001	0.997	0.992	0.987	0.982	0.976	0.970	0.963	0.956	0.949	0.942
27°		1.000	0.997	0.992	0.988	0.983	0.978	0.972	0.966	0.960	0.954	0.947	0.940	0.933
26°		0.992	0.988	0.984	0.979	0.974	0.969	0.963	0.958	0.952	0.945	0.939	0.932	0.925
25°		0.984	0.980	0.975	0.971	0.966	0.961	0.956	0.950	0.944	0.938	0.931	0.925	0.918
24°		0.976	0.972	0.968	0.963	0.958	0.954	0.948	0.943	0.937	0.931	0.924	0.917	0.910
23°		0.969	0.965	0.961	0.956	0.951	0.947	0.942	0.936	0.930	0.924	0.918	0.911	0.904
22°		0.962	0.958	0.954	0.949	0.944	0.939	0.934	0.928	0.923	0.917	0.910	0.904	0.897
21°		0.955	0.951	0.947	0.943	0.938	0.934	0.928	0.923	0.917	0.911	0.905	0.898	0.891
20°		0.949	0.945	0.941	0.937	0.932	0.927	0.922	0.916	0.911	0.904	0.898	0.892	0.885
19°		0.943	0.939	0.935	0.931	0.926	0.921	0.916	0.911	0.905	0.899	0.893	0.886	0.879
18°		0.938	0.934	0.930	0.926	0.921	0.916	0.911	0.906	0.900	0.894	0.888	0.882	0.875
17°		0.933	0.929	0.925	0.921	0.916	0.912	0.907	0.901	0.896	0.890	0.883	0.877	0.870
16°		0.928	0.924	0.920	0.916	0.911	0.907	0.902	0.896	0.891	0.885	0.879	0.872	0.865
15°		0.924	0.920	0.916	0.911	0.907	0.902	0.897	0.892	0.886	0.880	0.874	0.868	0.861
14°		0.920	0.916	0.912	0.908	0.903	0.898	0.893	0.888	0.883	0.877	0.870	0.864	0.857
13°		0.916	0.912	0.908	0.904	0.899	0.894	0.889	0.884	0.879	0.873	0.867	0.861	0.854
12°		0.913	0.909	0.905	0.901	0.896	0.892	0.887	0.881	0.876	0.870	0.864	0.858	0.851
11°		0.909	0.905	0.901	0.897	0.892	0.888	0.883	0.878	0.872	0.866	0.860	0.854	0.847
10°		0.906	0.902	0.898	0.894	0.890	0.885	0.880	0.875	0.869	0.863	0.857	0.851	0.844
9°		0.903	0.899	0.895	0.891	0.887	0.882	0.877	0.872	0.866	0.860	0.854	0.848	0.841
8°		0.900	0.896	0.892	0.888	0.884	0.879	0.874	0.869	0.863	0.857	0.851	0.845	0.838

\* Extension of tables XIX, XX for Burma and Assam.

It will be noticed that in tables XXXI, XXXIV discontinuities in the values of  $w_g$  in the neighbourhood of lat.  $20^\circ$ , long.  $91^\circ$ ,  $92^\circ$  are easily apparent. More careful examination of tables XXIX, XXX, XXXII, XXXIII reveals similar but much less marked discontinuities in the values of  $n_g$  and  $r_g$ . These are inevitable in view of the method by which the quantities have been found, and the differences are in agreement with those which may be computed by equations (42)—(47) of Chapter I following the two paths (viz. by direct geodesic and by two geodesics through the second origin) to such a point as  $L = 91\frac{1}{2}^\circ$ ,  $\lambda = 20^\circ$ . The amounts are not however sufficiently large to be of practical importance: and moreover they do not actually occur to the same extent in the actual triangulation of India as in the tables, for the tables have been extended somewhat beyond the triangulation limits for facility of subsequent interpolation. It may be mentioned in passing that the azimuth of a ray of length 40 miles is altered by an amount of order  $0''\cdot 1$  when its terminal latitude or longitude is altered by an amount of order  $0''\cdot 001$ : so that the taking out of azimuth to more than one place of decimals is not really defensible when the coordinates are given to only three places. In the present instance the ordinary procedure of the department has been followed and three places of decimals have been kept, with the idea that at any time the latter two of these may be disregarded.

## CHAPTER IV.

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### Geometrical change from one Spheroid of Reference to another.

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1. In selecting a spheroid of reference for the geoid there is no doubt as to the direction of the polar axis ; for this is the axis about which heavenly bodies appear to rotate. Hence all possible spheroids of reference are defined by the size of their axes and the position of their centres.

Consider two such spheroids. Let the semi-axes of one be  $a, b$  and of the other  $a' = a + da, b' = b + db$ . Select the origin of coordinates at the centre of the first spheroid and let the coordinates of the centre of the second be  $aa, a\beta, b\gamma$  where  $a, \beta, \gamma$  are small quantities.

In relating a point on a geoid to the spheroid the natural course seems to be to draw the normal through the point to the spheroid and to find out the coordinates of the point where this normal meets the spheroid. So long as the spheroid and geoid are not widely different this normal may, without appreciable error be considered as the vertical to the geoid and also as a straight line. For supposing there is a plumb-line deflection of 1 minute and a separation of the geoid and spheroid by 300 feet, the divergence of the normal from the vertical only amounts to about one inch which only affects coordinates by 0.001 of a second. It is accordingly satisfactory to relate a point on the geoid to one on the spheroid by merely producing the vertical of the geoid until it meets the spheroid. Considering then the relation between the points thus obtained on two reference spheroids corresponding to a point on the geoid, it is clear that all these points may with sufficient accuracy be regarded as being on a straight line, this straight line being normal to one of the three surfaces, whichever is most convenient.

To any triangle formed by three points on the geoid there is a corresponding triangle on any reference spheroid. The angles of these triangles are not identical. Those on the spheroid have different spheroidal excesses. The angles of a triangle observed on the geoid accordingly require correcting before they can be properly applied to a spheroid of reference. If this is properly done then this point relationship given above will hold. It is a fault in reduction of most, if not all, survey observations that geoidal and spheroidal angles have been treated as identical.

2. The coordinates of a point  $P$  on the first spheroid may be represented by

$$a \cos \phi \cos L, \quad a \cos \phi \sin L, \quad b \sin \phi$$

while those of a related point  $P'$  on the second spheroid may be represented by

$$aa + a' \cos \phi' \cos L', \quad a\beta + a' \cos \phi' \sin L', \quad b\gamma + b' \sin \phi'$$

where  $\phi' = \phi + d\phi$ ,  $L' = L + dL$ . It is necessary to find expressions for  $d\phi$  and  $dL$ . It is customary to decide on a point on a spheroid of reference as origin. All spheroids of reference are supposed to pass through this. Suppose that the origin lies in the plane of  $xz$ , so that  $L$  vanishes at the origin. In notation of previous chapters  $dL = v$ ,  $d\phi = u_1$ . At the origin these quantities reduce to  $v_0$  and  ${}_0u_1$ . The value of  $v_0$  is not obtained directly, but is derivable after azimuth has been decided on by the relation

$$w_0 = v_0 \sin \lambda_0 \dots \dots \dots (1)$$

This of course does not show the error of longitude of the origin: it merely shows by how much it will be changed if the azimuth is changed on the supposition of a plumb-line deflection in prime vertical.

Since the origin on either spheroid is identical in position the expression for its coordinates may be equated. Hence putting  $\phi = \phi_0$  and  $L = 0$

$$\left. \begin{aligned} a \cos \phi_0 &= aa + a' \cos \phi_0' \cos v_0 \\ 0 &= a\beta + a' \cos \phi_0' \sin v_0 \\ b \sin \phi_0 &= b\gamma + b' \sin \phi_0' \end{aligned} \right\} \dots \dots \dots (2)$$

Neglecting second order quantities and substituting from (1) for  $v_0$  (2) may be written

$$\left. \begin{aligned} a + \frac{da}{a} \cos \phi_0 - {}_0u_1 \sin \phi_0 &= 0 \\ \beta + w_0 \frac{\cos \phi_0}{\sin \lambda_0} &= 0 \\ \gamma + \frac{db}{b} \sin \phi_0 + {}_0u_1 \cos \phi_0 &= 0 \end{aligned} \right\} \dots \dots \dots (3)$$

These equations serve to determine  $a, \beta, \gamma$  in terms of the axes changes and changes at the origin: if the quantities  $\frac{da}{a}, \frac{db}{b}$  are multiplied by cosec  $1''$  the results are expressed in seconds.

3. Two further conditions are obtained by expressing the fact that the normal to the first spheroid at any point passes through the related point on the other spheroid.

The normal at  $P$  is

$$\frac{x - a \cos \phi \cos L}{\cos \phi \cos L} = \frac{y - a \cos \phi \sin L}{\cos \phi \sin L} = \frac{z - b \sin \phi}{\sin \phi}$$

and the conditions that  $P'$  should lie on this are

$$\frac{aa + d(a \cos \phi \cos L)}{\cos \phi \cos L} = \frac{a\beta + d(a \cos \phi \sin L)}{\cos \phi \sin L} = \frac{b\gamma + d(b \sin \phi)}{\sin \phi}$$

whence

$$\frac{a}{\cos \phi \cos L} + \frac{da}{a} - \tan \phi \cdot u_1 - \tan L \cdot v = \frac{\beta}{\cos \phi \sin L} + \frac{d\beta}{\beta} - \tan \phi \cdot u_1 + \cot L \cdot v$$

$$= (1 - e^2) \left\{ \frac{\gamma}{\sin \phi} + \frac{d\gamma}{\gamma} + \cot \phi \cdot u_1 \right\} \dots \dots (4)$$

From the first of equations (4)

$$v (\tan L + \cot L) = \sec \phi \left( \frac{a}{\cos L} - \frac{\beta}{\sin L} \right)$$

or

$$v = \sec \phi (a \sin L - \beta \cos L) \dots \dots \dots (5)$$

Eliminating  $v$  from (4) it follows that

$$\begin{aligned} \left( \frac{a}{\cos \phi \cos L} + \frac{da}{a} - u_1 \tan \phi \right) \cot L + \left( \frac{\beta}{\cos \phi \sin L} + \frac{db}{a} - u_1 \tan \phi \right) \tan L \\ = (1 - e^2) (\tan L + \cot L) \left( \frac{\gamma}{\sin \phi} + \frac{db}{b} + u_1 \cot \phi \right) \end{aligned}$$

whence

$$(a \cos L + \beta \sin L) \sec \phi + \frac{da}{a} - u_1 \tan \phi = (1 - e^2) \left( \frac{\gamma}{\sin \phi} + \frac{db}{b} + u_1 \cot \phi \right)$$

and expressing the results in seconds this may be written

$$u_1(1 - e^2 \cos^2 \phi) = (a \cos L + \beta \sin L) \sin \phi - (1 - e^2) \gamma \cos \phi + \sin \phi \cos \phi \left\{ \frac{da}{a} - (1 - e^2) \frac{db}{b} \right\} \operatorname{cosec} 1'' \dots (6)$$

The relation between  $u_1$  and  $u$  is given by (16) of Chap. III, and in terms of  $da$  and  $db$  is

$$u = \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 + \frac{1}{2} \left( \frac{da}{a} - \frac{db}{b} \right) \sin 2\lambda \operatorname{cosec} 1'' \dots \dots \dots (7)$$

The quantities  $a, \beta, \gamma, \frac{da}{a}, \frac{db}{b}$  all enter linearly into the equations. Their several effects can accordingly be computed separately and combined afterwards in any desired way. Cases corresponding to each of the four quantities  $\frac{da}{a}, \frac{db}{b}, u_1$  and  $w_0$  will now be considered.

Case (i)  $da = 1 \text{ km.}$   $u_1 = 0,$   $(u_0 = 12'' \cdot 063)$

From (3)

$$\begin{aligned} a + A \cos \phi_0 = 0 \quad \text{where } A = \frac{da}{a} \operatorname{cosec} 1'' = 32'' \cdot 3437 \\ \beta = \gamma = 0 \\ \therefore a = -29'' \cdot 536 \end{aligned}$$

From (5) and (6)

$$\left. \begin{aligned} v &= a \sec \phi \sin L \\ u_1 (1 - e^2 \cos^2 \phi) &= u_1 \cdot \frac{\sin^2 \phi}{\sin^2 \lambda} = a \sin \phi \cos L + \frac{1}{2} A \sin 2\phi \\ u &= \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 + \frac{1}{2} A \sin 2\lambda \\ &= \sin 2\lambda \left( \frac{u_1}{\sin 2\phi} + \frac{1}{2} A \right) \end{aligned} \right\} \dots \dots \dots (8)$$

Case (ii)  $db = 1 \text{ km.}$   $u_1 = 0$   $(u_0 = -12'' \cdot 1034)$

From (3)

$$\begin{aligned} a = \beta = 0 \\ \gamma + B \sin \phi_0 = 0 \quad \text{where } B = \frac{db}{b} \operatorname{cosec} 1'' = 32'' \cdot 4516 \\ \therefore \gamma = -13'' \cdot 2244 \end{aligned}$$

From (5) and (6)

$$\left. \begin{aligned} v &= 0 \\ u_1 (1 - e^2 \cos^2 \phi) &= - (1 - e^2) (\gamma + B \sin \phi) \cos \phi \\ u &= \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 - \frac{1}{2} B \sin 2\lambda \end{aligned} \right\} \dots \dots \dots (9)$$

Case (iii)  ${}_0u_1 = 9'' \cdot 978$   $(u_0 = 10'')$

From (3)  $a - {}_0u_1 \sin \phi_0 = 0$   
 $\beta = 0$   
 $\gamma + {}_0u_1 \cos \phi_0 = 0$   
 $\therefore a = 4'' \cdot 0659$   $\gamma = - 9'' \cdot 1118$

From (5) and (6)

$$\left. \begin{aligned} v &= a \sec \phi \sin L \\ u_1 (1 - e^2 \cos^2 \phi) &= a \sin \phi \cos L - (1 - e^2) \gamma \cos \phi \\ u &= \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 \end{aligned} \right\} \dots \dots \dots (10)$$

Case (iv)  $w_0 = 1''$   $v_0 = w_0 \operatorname{cosec} \lambda_0 = 2'' \cdot 447$

$a = \gamma = 0$   
 From (3)  $\beta + \frac{\cos \phi_0}{\sin \lambda_0} = 0$   $\therefore \beta = -2'' \cdot 2347$

From (5) and (6)

$$\left. \begin{aligned} v &= -\beta \sec \phi \cos L \\ u_1 (1 - e^2 \cos^2 \phi) &= \beta \sin L \sin \phi \\ u &= \left( 1 + \frac{e^2}{2} \cos 2\phi \right) u_1 \end{aligned} \right\} \dots \dots \dots (11)$$

To find the azimuth change  $w$ , the following equation holds for all cases

$$v - v_0 = w \operatorname{cosec} \lambda - w_0 \operatorname{cosec} \lambda_0 \dots \dots \dots (12)$$

in which  $v_0$  includes the entire origin change of longitude and is not restricted to that due to plumb-line deflection only. The equation follows from the fact that either side of it gives the difference between spheroidal and geoidal longitude. It is proved otherwise in the following chapter (*vide* equation (4)). With reference to the case IV (or (iv)) it will be noticed that the value of  $u \propto \sin L \sin \phi$ . The value found by the method of Chapter I was independent of  $\phi$ . The two cases however are not geometrically similar. In the case of Chapter I an azimuth change of origin involves a twist about the normal at the origin. In the present case the fixed axis is the polar axis and any twist introduced to give any desired azimuth change is only a component of a twist round an axis parallel to the polar axis. This makes it clear why the effect on latitude of this azimuth change is zero at the equator, the equator being at right angles to the axis of twist.

The values of  $u, v, w$  have been computed for the four cases by means of equations (8) to (12). The values of  $\lambda$  and  $L$  are the same as those of tables XXVII, XXVIII, and hence it is easy to make a comparison between the values of  $u, v, w$  found by the method of the present chapter which may be denoted by  $u_r, v_r, w_r$ , (related points on two spheroids) and  $u_g, v_g, w_g$  (found by following a geodesic). For this purpose values of  $u_r - u_g$  &c. are exhibited in tables XXXVII—XL.



TABLE XXXVII.  
Case I.— $\delta a = 1$  km.

TABLE XXXVIII.  
Case II.— $\delta b = 1$  km.

L'—L	0°	4°	8°	12°	16°	20°	24°	0°	4°	8°	12°	16°	20°	24°	L'—L
$\lambda$	Values of $(u_r - u_g)$ in seconds.							Values of $(u_r - u_g)$ in seconds.							$\lambda$
38°	0.000	+0.010	+0.049	+0.117	+0.201	+0.315	+0.442	+0.078	+0.073	+0.043	-0.005	-0.062	-0.148	-0.248	38°
34	+0.011	+0.021	+0.053	+0.099	+0.165	+0.249	+0.347	+0.017	+0.008	-0.012	-0.044	-0.093	-0.156	-0.234	34
30	+0.005	+0.014	+0.031	+0.061	+0.105	+0.154	+0.216	+0.008	+0.001	-0.011	-0.032	-0.064	-0.105	-0.159	30
26	+0.005	+0.008	+0.013	+0.022	+0.038	+0.058	+0.076	-0.004	-0.006	-0.010	-0.020	-0.032	-0.052	-0.075	26
22	-0.003	-0.010	-0.017	-0.028	-0.043	-0.061	-0.090	+0.003	+0.008	+0.012	+0.017	+0.024	+0.029	+0.035	22
18	-0.004	-0.017	-0.038	-0.070	-0.117	-0.180	-0.253	-0.001	+0.011	+0.025	+0.040	+0.069	+0.098	+0.139	18
14	+0.005	-0.012	-0.048	-0.103	-0.186	-0.290	-0.418	-0.031	-0.022	+0.001	+0.034	+0.084	+0.144	+0.216	14
10	+0.035	+0.013	-0.040	-0.120	-0.239	-0.386	-0.573	-0.114	-0.100	-0.065	-0.017	+0.060	+0.147	+0.259	10
	Values of $\pm (v_r - v_g)^*$ in seconds.							Values of $\pm (v_r - v_g)^*$ in seconds.							
38	0.000	+0.038	+0.069	+0.080	+0.067	-0.005	-0.116	0.000	-0.068	-0.133	-0.201	-0.266	-0.317	-0.371	38
34	0.000	+0.016	+0.026	+0.020	-0.019	-0.100	-0.234	0.000	-0.030	-0.064	-0.101	-0.127	-0.152	-0.168	34
30	0.000	+0.003	+0.003	-0.015	-0.071	-0.162	-0.312	0.000	-0.008	-0.024	-0.057	-0.043	-0.044	-0.043	30
26	0.000	+0.001	-0.005	-0.026	-0.000	-0.193	-0.351	0.000	-0.003	-0.009	-0.015	-0.005	-0.004	+0.015	26
22	0.000	-0.002	-0.003	-0.031	-0.091	-0.185	-0.316	0.000	-0.001	-0.011	-0.011	-0.008	-0.007	+0.011	22
18	0.000	+0.002	+0.004	-0.012	-0.061	-0.153	-0.293	0.000	-0.008	-0.025	-0.043	-0.052	-0.059	-0.056	18
14	0.000	+0.021	+0.033	+0.034	-0.001	-0.079	-0.204	0.000	-0.033	-0.069	-0.107	-0.147	-0.173	-0.196	14
10	0.000	+0.047	+0.082	+0.108	+0.100	+0.044	-0.068	0.000	-0.070	-0.137	-0.217	-0.285	-0.348	-0.400	10
	Values of $\pm (w_r - w_g)^*$ in seconds.							Values of $\pm (w_r - w_g)^*$ in seconds.							
38	0.000	+0.371	+0.735	+1.093	+1.436	+1.756	+2.058	0.000	-0.388	-0.777	-1.170	-1.566	-1.960	-2.363	38
34	0.000	+0.273	+0.542	+0.806	+1.058	+1.292	+1.509	0.000	-0.280	-0.564	-0.853	-1.143	-1.438	-1.740	34
30	0.000	+0.168	+0.329	+0.494	+0.641	+0.782	+0.913	0.000	-0.171	-0.341	-0.523	-0.700	-0.889	-1.090	30
26	0.000	+0.052	+0.107	+0.160	+0.205	+0.238	+0.265	0.000	-0.053	-0.114	-0.177	-0.251	-0.325	-0.416	26
22	0.000	-0.068	-0.130	-0.197	-0.267	-0.338	-0.411	0.000	+0.067	+0.124	+0.182	+0.228	+0.268	+0.288	22
18	0.000	-0.191	-0.384	-0.573	-0.760	-0.950	-1.136	0.000	+0.100	+0.379	+0.558	+0.728	+0.885	+1.030	18
14	0.000	-0.328	-0.648	-0.969	-1.285	-1.598	-1.903	0.000	+0.324	+0.642	+0.953	+1.255	+1.542	+1.815	14
10	0.000	-0.466	-0.928	-1.380	-1.830	-2.278	-2.707	0.000	+0.463	+0.920	+1.366	+1.805	+2.233	+2.636	10

TABLE XXXIX.  
Case III.— $u_0 = 1''$

TABLE XL.  
Case IV.— $w_0 = 1''$

L'—L	0°	4°	8°	12°	16°	20°	24°	0°	4°	8°	12°	16°	20°	24°	L'—L
$\lambda$	Values of $(u_r - u_g)$ in seconds.							Values of $\pm (u_r - u_g)^*$ in seconds.							$\lambda$
38°	-0.031	-0.029	-0.024	-0.015	-0.002	+0.014	+0.033	0.000	+0.031	+0.064	+0.095	+0.128	+0.158	+0.188	38°
34	-0.016	-0.014	-0.009	-0.001	+0.014	+0.030	+0.050	0.000	+0.023	+0.047	+0.069	+0.093	+0.116	+0.138	34
30	-0.006	-0.004	+0.002	+0.011	+0.025	+0.042	+0.062	0.000	+0.013	+0.028	+0.042	+0.056	+0.071	+0.084	30
26	-0.001	+0.002	+0.007	+0.017	+0.031	+0.048	+0.070	0.000	+0.004	+0.009	+0.013	+0.019	+0.024	+0.028	26
22	0.000	+0.002	+0.009	+0.018	+0.033	+0.051	+0.072	0.000	-0.006	-0.011	-0.016	-0.021	-0.025	-0.030	22
18	-0.005	-0.002	+0.004	+0.014	+0.029	+0.047	+0.070	0.000	-0.016	-0.031	-0.047	-0.061	-0.076	-0.090	18
14	-0.014	-0.012	-0.005	+0.005	+0.021	+0.040	+0.063	0.000	-0.027	-0.053	-0.078	-0.103	-0.127	-0.151	14
10	-0.029	-0.026	-0.020	-0.008	+0.007	+0.027	+0.051	0.000	-0.038	-0.074	-0.110	-0.145	-0.180	-0.214	10
	Values of $\pm (v_r - v_g)^*$ in seconds.							Values of $(v_r - v_g)$ in seconds.							
38	0.000	+0.019	+0.037	+0.055	+0.073	+0.091	+0.108	+0.083	+0.078	+0.062	+0.037	+0.001	-0.045	-0.100	38
34	0.000	+0.013	+0.026	+0.038	+0.051	+0.064	+0.075	+0.040	+0.035	+0.020	-0.005	-0.038	-0.085	-0.140	34
30	0.000	+0.008	+0.010	+0.023	+0.030	+0.037	+0.044	+0.013	+0.009	0.003	-0.031	-0.066	-0.110	-0.164	30
26	0.000	+0.002	+0.005	+0.007	+0.009	+0.011	+0.014	+0.001	-0.004	-0.019	-0.044	-0.078	-0.122	-0.176	26
22	0.000	-0.002	-0.004	-0.006	-0.009	-0.012	-0.014	+0.002	-0.003	-0.017	-0.043	-0.077	-0.121	-0.184	22
18	0.000	-0.007	-0.014	-0.022	-0.029	-0.035	-0.042	+0.014	+0.008	0.006	-0.031	-0.066	-0.110	-0.164	18
14	0.000	-0.011	-0.023	-0.035	-0.046	-0.057	-0.068	+0.036	+0.031	+0.016	-0.010	-0.044	-0.089	-0.143	14
10	0.000	-0.017	-0.032	-0.049	-0.065	-0.081	-0.096	+0.068	+0.064	+0.047	+0.020	-0.011	-0.059	-0.114	10
	Values of $\pm (w_r - w_g)^*$ in seconds.							Values of $(w_r - w_g)$ in seconds.							
38	0.000	+0.067	+0.133	+0.198	+0.262	+0.326	+0.387	+0.586	+0.585	+0.580	+0.574	+0.564	+0.550	+0.536	38
34	0.000	+0.065	+0.130	+0.194	+0.257	+0.319	+0.380	+0.407	+0.406	+0.403	+0.398	+0.393	+0.382	+0.371	34
30	0.000	+0.065	+0.128	+0.191	+0.253	+0.315	+0.375	+0.235	+0.235	+0.235	+0.231	+0.226	+0.221	+0.214	30
26	0.000	+0.064	+0.127	+0.191	+0.252	+0.313	+0.372	+0.075	+0.074	+0.073	+0.073	+0.072	+0.070	+0.068	26
22	0.000	+0.065	+0.128	+0.191	+0.253	+0.313	+0.371	-0.082	-0.082	-0.082	-0.081	-0.079	-0.078	-0.070	22
18	0.000	+0.065	+0.129	+0.192	+0.254	+0.315	+0.373	-0.234	-0.234	-0.232	-0.229	-0.225	-0.220	-0.214	18
14	0.000	+0.065	+0.129	+0.193	+0.256	+0.317	+0.378	-0.384	-0.383	-0.380	-0.376	-0.370	-0.362	-0.350	14
10	0.000	+0.066	+0.131	+0.196	+0.260	+0.323	+0.383	-0.533	-0.532	-0.528	-0.522	-0.512	-0.501	-0.486	10

\* + or - according as point is west or east of origin.

4. The differences  $w_r - w_g$  and  $v_r - v_g$  are never sufficiently great to have any important effect on geodetic results. In the case of  $w_r - w_g$  larger values are met with. The deflections of plumb-line in the prime vertical are affected by the amount  $(w_r - w_g) \cot \lambda$ , a quantity which may be as much as several seconds. In the practical case however where, according to the most recent determinations  $\delta a = 0.924 \text{ km}$  and  $\delta b = 0.743 \text{ km}$ : the combined effect of Cases I and II in these proportions are not large, the two cases tending to cancel each other; as may be seen from table XLI below, in which the values of  $(w_r - w_g) \cot \lambda$  are also given.

TABLE XLI.

$\lambda$	$L' - L$	$\text{Cot } \lambda$	$0.924 \times \text{Case I}$	$0.743 \times \text{Case II}$	Combined effect	Discrepancy in plumb-line deflection
38°	24°	1.380	1.2902	-1.256	0.034	0.187
34	24	1.483	1.394	-1.293	0.101	0.150
30	24	1.732	0.844	-0.810	0.034	0.059
26	24	2.050	0.245	-0.309	-0.064	-0.131
22	24	2.475	-0.380	0.214	-0.166	-0.411
18	8	3.078	-0.355	0.282	-0.073	-0.225
14	4	4.011	-0.303	0.241	-0.062	-0.249
10	4	5.671	-0.431	0.344	-0.087	-0.493

The only case in which the discrepancy in prime vertical deflection would be considerable occurs in low latitudes and this case does not concern Indian Triangulation as in these latitudes there are no great longitude differences: for the case of Burma special treatment, as given in Chapter III, is in any case necessary.

The conclusion is that either this method or that of the preceding chapter could be used with practically satisfactory results. The discrepancies however indicate how far theoretical accuracy has been departed from in failing to project geoidal angles on to the spheroid of reference before introducing them in the computations.

The figures in tables XXXVII—XL may be noticed to be a little irregular. This is doubtless due to the fact that the computations of Chapter III were not made with sufficient accuracy to ensure the last figure always being correct. This was not considered to be sufficiently important to justify the extra labour which would have been necessary. The results are fully accurate enough for all practical purposes to which they can be put.

5. Three methods of finding the change in coordinates due to any proposed changes of the axes of the spheroid and the latitude and azimuth at the origin have now been given. That of Chapter I gives a means of computing these along any path defined by a relation between  $\lambda$  and  $L$ . Chapter III gives the results for the special case when the path selected is the geodesic through the origin and the point at which the changes are required: and in the present chapter the geometrical relation between corresponding points on two spheroids is worked out. All these methods give somewhat different results. The latter two have the advantage over the former of being free from any ambiguity due to multiple values and inconsistency: and the differences met with between them are not of amount sufficiently large to be troublesome. The reason for their discrepancy is examined in the following chapter: and the conclusion is arrived at, from theoretical considerations, in view of the methods by which the observations of the triangulation of India have been reduced, that **the method of calculation along geodesics as set forth in Chapter III is the correct one to use.** The detailed tables XXIX—XXXIV permit this to be done readily.

## CHAPTER V.

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Laplace's Equation and the Choice of a Spheroid of Reference.

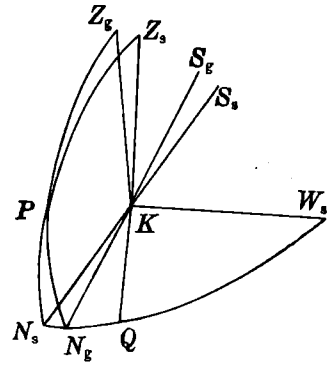
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1. When a large survey is begun one of the first essentials is the selection of a point as origin. The coordinates of this point have to be decided on. The longitude is of little consequence and the meridian through the point may be taken as that from which all deduced longitudes are measured. The latitude and azimuth can be observed astronomically: but their geodetic values depend on the plumb-line deflection existing at the point. Plumb-line deflection is of course merely the deviation of the vertical from the normal to some assumed figure of reference. In choosing the origin it must be eventually decided whether to consider the deflection there as nil—in other words, choosing a spheroid of reference parallel to the geoid at the origin—or whether considerations of topographical features and irregularities of density justify the adoption of certain values of the deflection in the two components. The question of height above geoid of the point selected as origin also arises. If an error of 10 feet is made in this it is practically equivalent to assuming a spheroid with axes 10 feet different from those actually selected. In the case of the origin of the Indian Survey there is no reason to suppose an error of nearly so much, and so no further consideration will be given to this point here.

Having decided on the origin  $O$ , it is next necessary to decide on a figure of reference. This will generally be referred to as the "spheroid" in opposition to the "geoid" or sea level equipotential surface of the earth. It is not implied by this that the figure of reference must be a spheroid, though the almost universal practice is to take a spheroid as a reference figure.

2. Chains of triangulation may now be computed *rigidly* if proper corrections given below in § 16 are applied: and the coordinates, latitude, longitude, height and azimuth at a distant point  $K$  may be deduced. Suppose that astronomical azimuth and latitude are observed at  $K$  and also that the arc  $OK$  is observed as a telegraphic longitude arc. Let  $\lambda$ ,  $L$  be the geodetic latitude and longitude of  $K$  and let  $A$  be the geodetic azimuth at  $K$  of some reference mark: these are quantities brought up by triangulation. Let  $\eta$ ,  $\xi$  be the plumb-line deflection at  $K$  in meridian and prime vertical (positive for southerly and westerly deflections of the plumb bob) referred to the selected spheroid of reference. These quantities obviously differ for different spheroids of reference.

In the figure suffixes *s, g* refer to spheroidal and geoidal points respectively: or in other words points derived from triangulation and star observations respectively. *P* is the pole and *Z* the zenith. The astronomic azimuth of a point *Q* is clearly  $A + \xi \tan \lambda$  being  $S_g K S_s$  greater than the geodetic azimuth. Let suffix zero denote quantities appertaining to the origin of the survey: so that the values of  $\eta_0, \xi_0$  have been decided in some way or other.



3. First consider the longitude observations. They depend on the interval of time between the meridians of *O* and *K* as shown by star transits. If the zenith at a station is displaced in the prime vertical, the meridian is also displaced as a result, the direction of the pole being fixed. Time stars are observed when they transit the plane  $Z_g P$  instead of the plane  $Z_s P$ . With a westerly deflection the zenith is moved towards the east and the result is that stars are observed too soon by the angle  $Z_g P Z_s = \xi \sec \lambda$ . If  $\xi$  is expressed in seconds of arc, the star time as given by the local meridian is early by  $\xi \sec \lambda / 15$  seconds of time. Now of the two stations *O, K*, if *O* is the more westerly and *T* is the time interval between the transits of a star at the two meridians, then the time interval between the two spheroidal meridians is

$$T - \frac{\xi \sec \lambda}{15} + \frac{\xi_0 \sec \lambda_0}{15}$$

and this should be the same as  $\frac{L - L_0}{15}$ , or the difference of longitude in time, on that spheroid on which  $\xi, \xi_0$  represent the plumb-line deflections in prime vertical at *K* and *O*.

The quantity *T* is an observation quantity; suppose its error is  $\delta T$ : also suppose the error in longitude generated in the triangulation is  $\delta L$ . It follows that

$$\xi_0 \sec \lambda_0 - \xi \sec \lambda = L - L_0 - \delta L - 15 (T - \delta T) \quad \dots \dots \dots (1)$$

Now consider the azimuth observations. Let *A'* be the astronomically observed azimuth which has an error  $\delta A'$ : *A* and  $\delta A$  being the geodetic azimuth computed on the spheroid and its error. Then

$$\left. \begin{aligned} A' - \delta A - A + \delta A &= \xi \tan \lambda \\ A'_0 - \delta A'_0 - A_0 &= \xi_0 \tan \lambda_0 \end{aligned} \right\} \dots \dots \dots (2)$$

Eliminating  $\xi, \xi_0$  between (1) and (2) it follows that

$$(A'_0 - \delta A'_0 - A_0) \operatorname{cosec} \lambda_0 - (A' - \delta A' - A + \delta A) \operatorname{cosec} \lambda = L - L_0 - \delta L - 15 (T - \delta T) \quad \dots \dots (3)$$

which is an elaborated form of Laplace's equation.

4. Suppose now that the computations had been carried out on a slightly different spheroid. If this had been done vigorously the quantity  $\delta L$ , being itself a small quantity will not be changed appreciably, while *A',*  $\delta A',$  *A'\_0,*  $\delta A'_0,$  *T,*  $\delta T$  are all quantities which are not affected by the change in spheroid. The only quantities in (3) which change appreciably are *A, L - L\_0* and  $\lambda$ . The  $\lambda$  terms are multiplied by small coefficients and their variations can be neglected. Hence differentiating (3) for change of spheroid it follows that in the notation of Chapter I where *u v w* represent changes in latitude, longitude and azimuth

$$w \operatorname{cosec} \lambda - v_0 \operatorname{cosec} \lambda_0 = v - v_0 \quad \dots \dots \dots (4)$$

which is equation (12) of Chapter IV.

It might be expected that this equation would be in accordance with those found in Chapter I. As was noticed there, however, the quantities  $u, v, w$  are many-valued, a separate set of values appertaining to each route along which the integration is performed. Equation (4) on the other hand is free from any ambiguity and accordingly cannot be in accord with the equations of Chapter I. If numerical quantities are substituted it is at once clear that the relation (4) is not satisfied. Consider the values of

$$v = v_x - f(v_x - v_y) \text{ and } w = w_x - f'(w_x - w_y)$$

where  $f$  and  $f'$  are fractional quantities. These expressions are the values of  $v, w$  computed along routes intermediate to those of  $v_x$  and  $v_y$ . Taking case where  $\delta a = 1 \text{ km.}, \lambda = 30^\circ, L = 66^\circ, v_0 = w_0 = 0$  from the tables VII—X it follows that

$$v \sin 30^\circ = 4.027 - .013 f$$

and

$$w = 3.784 - .504 f$$

which cannot be made equal by any positive fractional values of  $f$  and  $f'$ . In the same way the tables of Chapter III show that the relation (4) is not satisfied along a geodesic.

5. It has generally been considered that azimuth and longitude observations both give the same information, namely deflection of the plumb-line in prime vertical, and nothing more: and in so far as the results differ by the two methods the reason is that the observations are burdened by errors. Clarke states\* that "the observations of the difference of longitude gives "us no information that is not also given by the observation for azimuth". With this principle Colonel Sir Sidney Burrard† has used the longitude observations of India to correct azimuth observations for the accumulation of error due to triangulation, considering the differences of the resulting plumb-line deflection found by the two observations to be entirely accounted for by observation error in the triangulation.

6. The explanation of these apparent inconsistencies was not discovered for some time. Equation (4) is perfectly correct if the triangulation is properly computed. The ordinary process of computation is not quite correct. Angles are measured by means of a theodolite and reduced to the horizontal plane of the geoid. This is not quite the same thing in general as the horizontal plane of the spheroid. If the computation is to be effected on the spheroid (on which all the various formulæ are based) the observed angles should be projected on to the selected spheroid of reference, and so will differ according to what spheroid is selected. The actual amount by which the geoidal angle must be altered to get the spheroidal angle depends on two things (vide § 16 below)

(1) the deflection of the plumb-line or inclination of the geoidal (astronomic) vertical to the spheroidal (triangulated) vertical.

(2) the inclination to the horizontal of the rays between which the geoidal angle is measured.

The first of these quantities varies appreciably with change of spheroid and accordingly the correction to the geoidal angle varies according to the spheroid used. The actual case under consideration is represented in symbols by supposing  $\delta L$  in (1) to contain not only the error due to faulty observations but also the error due to failure to correct the geoidal angles to spheroidal angles. This is purely a computation error. The actual "grinding" process has treated these errors as errors of observation.

This perhaps explains why Laplace's equation is in general not satisfied so well as the probable errors of the several observations on which its formation depends would cause to be expected.

\* "Geodesy" by Col. A. R. Clarke, p. 291.

† Appendix No. 5 of G.T. Volume XVIII. "On the azimuth observations of the G.T.S. of India".

7. It also explains how it is that the different values of  $u, v, w$  arise as noticed in Chapter I. In this case the fact that the closing errors of circuits will differ from one spheroid to another *unless all the geoidal angles are reduced to spheroidal angles* makes the distribution of closing errors have a different effect according as the spheroid is altered.

Equation (4) may be re-written to meet the actual case as follows :

$$[w \operatorname{cosec} \lambda - v] = \Delta L \dots \dots \dots (5)$$

where  $\Delta L$  is the change in computation error of longitude difference due to the treatment of spheroidal and geoidal angles as identical. The fact that  $\Delta L$  is not zero would be of more serious importance in the question of change of spheroid had equation (3) been used in all possible cases as a condition for the series of triangulation to satisfy. When the main Indian triangulation was adjusted the longitude arcs either were not available or else were ignored, so that Laplace's condition was not imposed on the triangulation. In correcting the azimuth observations, Colonel Sir Sidney Burrard introduced the condition for the first time in India.

8. Clarke's statement quoted in § 5 was deduced from an equation which arises in his work : but it may be seen to be true without any analysis. The longitude observation fixes the meridian plane at a point, that is the plane through the zenith and the pole, by taking the time of stars transiting this plane. It obviously does no more than fix this plane with relation to another. The azimuth observation practically draws the great circle through the pole and zenith and locates where this cuts the horizon, by means of a horizontal angle measured from a fixed point. The position of the pole and the place of observation being already given the fixing of one other point suffices to fix the meridian plane. Thus longitude and azimuth observations both merely fix the position of the meridian plane and nothing more. The inclination of this plane to the meridian plane deducible from triangulation is the deflection in prime vertical.

9. None the less equation (3) does definitely give some information as to the error of computation generated in the triangulation, and to this extent Clarke's statement needs modification. When the practical case is considered from equation (3) it may be seen that the identity of plumb-line deflection, whether derived from longitude observations or azimuth observations, affords some information concerning the slightly faulty method of computing from geoidal angles instead of from spheroidal angles. For split up the error  $\delta L$  into  $\delta_1 L$  due to faulty observation and  $\delta_2 L$  due to faulty computation. Suppose next that the spheroid of reference is changed so that it is necessary to substitute  $A+w$  for  $A$  and  $L+v$  for  $L$ . The quantity  $\delta_1 L$  remains unaltered, but  $\delta_2 L$  obviously is a variable according to the spheroid used and from (3) it follows that

$$(A'_0 - A_0 - w_0) \operatorname{cosec} \lambda_0 - (A' - A - w) \operatorname{cosec} \lambda = L - L_0 + v - v_0 - 15 T - \delta_2 L + \Delta E \dots \dots (6)$$

in which the only variables are  $w, w_0, v - v_0$  and  $\delta_2 L$ , and  $\Delta E$  is the combined and fixed effect of observation errors. It is possible to form sixteen equations of the form (6) from the longitude and azimuth observations of India. Expressing  $w, w_0, v - v_0$  in terms of  $\delta a, \delta b, u_0$  and  $w_0$  it is possible to solve these equations for  $\delta a, \delta b, u_0, w_0$  so as to make  $\Sigma (\Delta E - \delta_2 L)^2$  a minimum : *i.e.* since  $\Delta E$  is equally likely to be positive or negative  $\Sigma \overline{\Delta E}^2 + \Sigma \overline{\delta_2 L}^2$  is a minimum. But as the value of  $\Delta E$  is not being varied, this implies that  $\Sigma \overline{\delta_2 L}^2$  is a minimum. Now when the spheroid differs widely from the geoid it is clear that the computation errors increase : and conversely when the spheroid approximates more closely to the geoid these errors diminish. The fact that  $\Sigma \overline{\delta_2 L}^2$  is made a minimum affords one criterion for the spheroid being in close agreement with the geoid for the *area* over which the triangulation of India extends. It is of course possible to consider what spheroid suits the actual deflections best : but this is an entirely different point of view from that indicated above, and deals only with the actual localities in which the deflections are measured : and moreover is burdened by the errors of computation involved in treating geoidal and spheroidal angles as identical.

10. The interest in the method is chiefly theoretical. The quantities to be dealt with are very small: and in most cases the effects of observation error may well mask those due to the computation error. Sixteen equations of the form (6) are given below. These can be solved for  $\delta a, \delta b, u_0, w_0$  or, treating  $\delta a, \delta b$  as known, for  $u_0, w_0$  only. It was not anticipated that the former course would give reliable values of  $\delta a, \delta b$  but the solution was none the less made. Afterwards the solution of  $u_0, w_0$  only taking the latest values of  $\delta a, \delta b$  was performed. Referring to these two solutions as *A* and *B*, one difficulty of the application of (6) arises in *A*, but to a very much less extent in *B*. This difficulty is the selection of the route along which  $u, v, w$  in terms of  $\delta a, \delta b$  shall be determined. The actual courses of the triangulation series are numerous, and the case seems to be best met by taking the geodesic solution of Chapter III, for this in general leads to a medial path through the triangulation. In solution *B* it so happens that the  $\delta a$  and  $\delta b$  terms very nearly cancel one another. The form of the equations is as follows

$$(v_1 \sin \lambda - w_1) \delta a + (v_2 \sin \lambda - w_2) \delta b + (v_3 \sin \lambda - w_3) u_0 + (v_4 \sin \lambda - w_4) w_0 + A' - A - (A'_0 - A_0) \sin \lambda \operatorname{cosec} \lambda_0 - \sin \lambda (15T - L + L_0) = 0 \quad (7)$$

in which  $A'_0 - A_0 = 1'' \cdot 29 - w_0$  (vide Chapter I §4).

The sixteen arcs from Kalianpur give rise to sixteen equations which are exhibited in the table.

TABLE XLII.

Azimuth station	Longitude station	Coordinates of Azimuth station		Coefficients of				$A' - A$	$15T - L + L_0$	Absolute term $A - 1 \cdot 29 \operatorname{cosec} \lambda_0 \sin \lambda + A' - A - (15T - L + L_0) \sin \lambda$	Residual <i>A</i>	Absolute term <i>B</i>	Residual <i>B</i>	
		$\lambda$	$L$	$\delta a$	$\delta b$	$u_0$	$w_0$							
Karachi Observatory	Karachi T. O.	24 49 50	87 1 35	+0.080	-0.050	+0.168	+0.042	-1.4	+0.5	-2.9	-1.6	-2.9	-3.0	
Dehra Dun Observatory (old)	Dehra Dun Longitude Station	30 19 57	78 3 35	-0.017	+0.018	-0.005	+0.242	-11.0	-25.7	-0.5	-2.5	-0.5	-2.3	
Quetta T. O.	Quetta T. O.	30 11 57	67 0 32	+0.457	-0.450	+0.160	+0.251	-4.4	+2.4	-7.2	-3.4	-7.1	-8.8	
Calcutta Base-line, S. end	Calcutta	22 36 56	88 22 54	+0.118	-0.118	-0.172	-0.043	-8.0	-11.0	-5.9	-5.4	-5.9	-5.6	
Orejhar	Pyzabad T. O.	26 46 56	82 12 8	-0.085	+0.085	-0.071	+0.107	-4.1	-0.5	-5.3	-7.4	-5.3	-6.2	
Jalpaiguri	Jalpaiguri	26 31 17	86 44 13	-0.199	+0.196	-0.172	+0.109	-4.7	-20.4	+3.0	-1.0	+3.0	+2.0	
Nagarkhana	Chittagong T. O.	22 22 58	91 48 30	+0.226	-0.051	-0.227	-0.043	-8.7	-11.7	-5.4	-0.1	-5.2	-5.1	
Bolarum P.W.D. Office	Bolarum	17 30 19	78 31 11	+0.046	-0.045	-0.014	-0.257	-1.1	-3.5	-1.0	+1.4	-1.0	+0.0	
Vizagapatam Base-line, N. end	Waltair	18 1 3	83 13 43	+0.268	-0.267	-0.092	-0.235	-1.4	-3.3	-1.4	+2.7	-1.4	+0.2	
Karaundi	Jubbulpore T. O.	23 10 40	79 50 43	+0.017	-0.021	-0.058	-0.086	-4.0	-10.2	-1.2	-1.1	-1.2	-1.0	
Colaba Observatory	Bombay	18 53 49	72 48 40	-0.198	+0.199	+0.080	-0.199	+1.0	-6.8	+2.2	+2.2	+2.2	+3.7	
Deesa T. O.	Deesa T. O.	24 15 30	72 11 6	+0.003	-0.003	+0.086	+0.009	-4.6	+3.6	-7.4	-6.0	-7.4	-7.4	
Mangalore	Mangalore	12 52 14	74 50 43	-0.281	+0.261	+0.048	-0.434	-2.8	-2.0	-3.1	-2.2	-3.1	+0.1	
Bangalore Base-line, S.W. end	Bangalore	13 0 41	77 35 0	-0.007	+0.007	+0.003	-0.430	-5.3	+2.0	-6.7	-3.4	-6.7	-3.6	
St. Thomas' Mount Trestle	Madras	13 0 15	80 11 41	+0.233	-0.233	-0.044	-0.429	-4.0	-7.2	-3.1	+2.4	-3.1	0.0	
Kudankulam Observatory	Nagarkoil	8 10 22	77 41 27	+0.005	-0.007	0.000	-0.615	-7.7	+1.6	-8.4	-3.6	-8.4	-3.0	
Sum of squares										...	360.23	203.83	350.88	292.30
Square root of mean square										...	4.74	3.56	4.74	4.27

The solution *A*, i.e. the most probable value of  $\delta a$ ,  $\delta b$ ,  $u_0$ ,  $w_0$ , is

$$\left. \begin{aligned} \delta a &= 33\cdot08 \text{ km} \\ \delta b &= 22\cdot63 \text{ km} \\ u_0 &= +6''\cdot10 \\ w_0 &= -7''\cdot71 \end{aligned} \right\} \dots \dots \dots A$$

to which correspond the residuals under "Residual *A*" in the table. If 0·92363 km and 0·74273 km are substituted for  $\delta a$  and  $\delta b$  (*vide* Chapter I §3) the following most probable values of  $u_0$  and  $w_0$  are arrived at (solution *B*) :—

$$\left. \begin{aligned} u_0 &= +1''\cdot01 \\ w_0 &= -7''\cdot28 \end{aligned} \right\} \dots \dots \dots B$$

and the corresponding residuals are shown in the table under heading "Residual *B*".

12. Solution *A* is obviously of no practical use as the values of  $\delta a$  and  $\delta b$  are much larger than it is possible could be correct. Solution *B* is not unreasonable. A southerly deflection at Kaliaupur has previously been inferred, the estimated amount being 4". The value of  $w_0$  indicates an easterly deflection of 16"·2. The value computed from the topography, but taking no account of compensation is 10"·7 (*vide* Prof. Paper 13, p. 116). The solutions however have been given more as illustrations of a principal than for their numerical values. The residuals show that the solution is not highly successful in satisfying the equations: yet the values of  $u_0$ ,  $w_0$  derived from *B* are reasonable and the residuals might fairly be attributed to observation errors.

13. *The choice of a figure of reference for the geoid.* In surveying a surface such as the geoid, in the first place of unknown form, it is necessary at the outset to decide on some figure of reference to which measurements may be referred. This figure of reference may be of any form whatever—a particular case would be any set of three orthogonal planes. In traverse operations a single plane is chosen, on the assumption that for a limited area the geoid is not much different from a plane: and it would be possible to extend the application of the plane of reference by the introduction of a third coordinate, namely that at right angles to the plane. If at each triangulation station the direction of the geoidal vertical is determined by means of latitude and longitude observations the data is sufficient to enable the position of any point to be expressed by means of its three coordinates, quite independently of the shape of the geoid. The statement of the position of numerous points on the geoid in fact determines the shape of the geoid. If however the position of the several points are referred to a figure which approximates in shape and position to the geoid, the actual shape of the geoid is much more readily grasped by the small deviations it exhibits from the well known reference figure. The choice of such a figure is extremely useful and greatly decreases the labour of calculation of the positions of points on the geoid. The closer the approximation between it and the geoid the smaller are the quantities which express the difference of the two surfaces: and, as a result, when these quantities appear in formulæ as square and power terms they may be neglected in many of the computations which arise. It is important none the less to recognise that the two surfaces cannot always be treated as identical, and to examine each case thoroughly. Moreover it is to be borne in mind that the complexity of computation will be much increased if a very complex figure of reference is selected. A balance must be struck between the two considerations, and it has been customary to adopt a spheroid. This is at the same time a comparatively simple figure for computation and also a fairly close approximation to the geoid. This choice need not at all imply that no other geometrical figure can be found which approximates more closely to the geoid. Suppose even that the geoid was in actual fact an ellipsoid (not of revolution) not very different from a spheroid. It would still be strictly accurate to refer it to a spheroid of reference:



and it is probable that this would be the simplest course to follow in dealing with the results of any one survey, for instance the Indian Survey. Or the geoid might be referred with strict accuracy to a sphere: but in this case the residuals in a vertical direction might be inconveniently large.

14. This does not appear to have been quite the point of view usually taken, seeing that much energy has been devoted to finding the spheroid which best fits the whole earth. The origin of this research was doubtless the desire to uphold the Newtonian theory that the earth, being a revolving gravitating mass, should approximate in form to an oblate spheroid: rather than to the prolate spheroid which early measurements led the French school to believe in. This question was finally settled in favour of the Newtonian theory by the measurements of the arcs in Peru and Lapland: and the matters now to be investigated are the relatively minor deviations of the geoid from the oblate spheroid. Given a ready means of converting coordinates from one spheroid to another, each survey may properly select the spheroid most suitable to its own requirements. In any case the several large surveys of the world are expressed in terms of different spheroids, and for purposes of intercomparison it is necessary to develop a method of changing from one spheroid to another. An interesting question is to consider how closely the several spheroids, which best fit the respective surveys, agree *inter se*: to account for any differences: and to see whether a theory of density distribution can be found which will bring all these spheroids into agreement. The same question may be considered by taking the surveys on the spheroids they happen to have been reduced on and afterwards expressing the results in terms of a single spheroid and the local differences of the geoid from this general spheroid. Even if this general spheroid is so selected as to make the differences from the geoid a minimum it still remains only a convenient figure of reference and a more or less close approximation to the geoid.

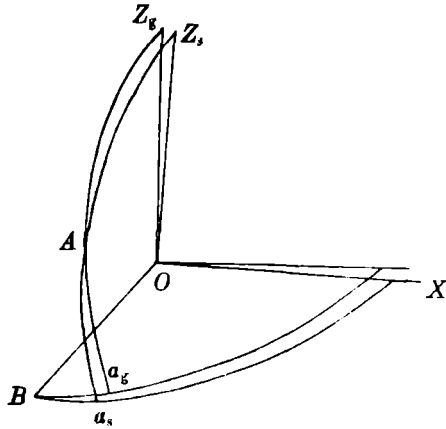
15. In the case of triangulation the usual procedure is as follows: horizontal angles are measured on the geoid, that is to say a theodolite is set up and levelled so that its horizontal circle is tangential to a level surface of the geoid. Spheroidal excess, calculated from the assumed spheroid, is applied to these angles. Further computations of the latitude and longitude of the points of triangulation are then carried out as though the spheroid and geoid were identical.

Now in certain disturbed districts the geoid is of considerably different curvature from the adopted spheroid: and the excess over  $180^\circ$  of the sum of the three angles of a triangle observed on the geoid is not the same as that computed from the spheroid. On account of the relative smallness of this excess in triangles of the size which occur in triangulation, this difference is not of great importance, though it gives rise to the two entirely different methods of Chapters III and IV. But if the rays observed have a considerable elevation, such as  $5^\circ$ , a very appreciable error is introduced, as will shortly be explained. It is necessary to be more precise. The most natural way of relating a point on the geoid to the spheroid is by giving the coordinates (latitude and longitude) of that point of the spheroid the normal—or more strictly for large distances the orthogonal confocal hyperbola—at which passes through the point on the geoid; and by stating the height of the geoid above the spheroid measured along this normal as well as the angle between this normal and the normal to the geoid (deflection of the plumb-line) and the azimuth of the plane containing the two normals.

Defining the position of a geoidal point in this way for the present, the separation of the geoid and spheroid need not be considered. To each point on the geoid there is a corresponding point on the spheroid: and consequently to each geoidal triangle a spheroidal triangle corresponds. It is with such spheroidal triangles that computations of latitude and longitude etc. really deal, the formulæ being deduced from properties of the spheroid. Consider then the relation between the angles of a geoidal triangle and the corresponding spheroidal triangle.

16. Suppose a theodolite is set up at a point  $O$  and levelled in the ordinary way: at this point two zeniths may be distinguished,  $Z_g$  that of the geoid and  $Z_s$  that of the spheroid, the former being indicated by the direction of the theodolite when the altitude is set to  $90^\circ$ .

The plane  $Z_g O Z_s$  is the plane of deflection and the line  $OB$  at right angles to this plane is parallel to both spheroid and geoid and is chosen as axis of  $Y$ : so that  $OB$  may be regarded indifferently as belonging to the spheroid or the geoid. Consider another point  $A$  and draw the great circles  $Z_s A a_s$  and  $Z_g A a_g$ . Then  $O a_s$  and  $O a_g$  are the traces of the ray  $OA$  on the spheroid and geoid respectively. Suppose that the horizontal angle between  $OA$  and  $OB$  is required. Observation by the theodolite gives the angle  $a_g OB$ : but the angle required for computation on the spheroid is  $a_s OB$ . Denote by  $a$  the geoidal angle of elevation of  $A$  and by  $z$  the azimuthal angle  $X_g O a_g$ , corresponding quantities for the spheroid being  $a + \delta a$ ,  $z + \delta z$ . From triangles  $Z_g BA$  and  $Z_s BA$



$$\cos AB = \sin a \cos Z_g B - \cos a \sin Z_g B \cos(z_g - 90^\circ) = \sin(a + \delta a) \cos Z_s B - \cos(a + \delta a) \sin Z_s B \cos(z_s - 90^\circ)$$

But  $Z_g B = Z_s B = 90^\circ$ ; hence

$$\cos a \sin z = \cos(a + \delta a) \sin(z + \delta z) \quad \dots \dots \dots (1)$$

*i.e.*  $\tan a \cdot \delta a = \cot z \cdot \delta z \quad \dots \dots \dots (2)$

neglecting second order terms.

Now  $\delta z$ ,  $\delta a$  are the corrections which should be applied to geoidal quantities to correct them into spheroidal quantities and make them suitable for spheroidal formulæ.  $\delta a$  is  $A a_s - A a_g$  and is approximately  $\epsilon \cos z$ , where  $\epsilon$  is the total plumb-line deflection which is in the plane  $O X_g Z_g$ : so that (2) may be written

$$\delta z = \epsilon \tan a \sin z \quad \dots \dots \dots (3)$$

The correction  $\delta z$  accordingly is greatest when  $z = 90^\circ$ , that is when the observed object is in the plane of no deflection, and its magnitude in this case is  $\epsilon \tan a$ . Now values of  $\epsilon$  up to one minute have been observed: and if at the same time a ray of elevation of  $4^\circ$  is observed, the horizontal angle may need a correction of 4 seconds—a very appreciable quantity in geodetic triangulation. The figures here given are roughly applicable to a ray through Jharipani (Dehra Dun district) where the deflection exceeds one minute and considerable angles of depression occur. In the case of a triangle at two of whose corners there is no deflection while a considerable deflection occurs at the third, a large triangular error will be apparent. More usually however the deflection is not so widely different at the three corners and the angular errors partially compensate one another in the sum, thus masking the error, but leaving the triangle distorted.

17. It is of interest to note that deflections have a corresponding effect on the measurements of base lines. Suppose that an element  $ds$  is measured along a line inclined at an angle  $a$  to the geoidal horizontal and  $a + \delta a$  to the spheroidal vertical. Its reduced length is generally taken as  $ds \cos a$ , whereas reduced to the spheroid it is  $ds \cos(a + \delta a)$ , so that a correction

of  $-ds \sin a \cdot \delta a$  is required. The error on the whole line is  $\int ds \cdot \sin a \delta a$ . This is equal to  $s \sin a_m \delta a_m$  where  $a_m$  and  $\delta a_m$  are values which occur at some part of the line. If  $a$  is fairly constant  $s \sin a$  is approximately the difference of level of the two ends of the base and the error is approximately  $(h_2 - h_1) \delta a_m$ . Owing to  $h_2 - h_1$  being small compared with the length this is only liable to affect the length by a quantity of as much as 1 in  $10^6$  in extreme cases.

18. Deflections are usually stated in terms of their westerly and southerly components,  $\xi, \eta$ . It is clear that the effect of either component on a ray can be computed independently and then the two results combined. In the case of a ray of azimuth  $A$  it follows from (3) that a correction to the geoidal azimuth of amount  $\delta A$  is required where

$$\delta A = (-\xi \cos A + \eta \sin A) \tan a \dots \dots \dots (4)$$

Consider now the case of a traverse. Denote the successive points by 1, 2, 3 . . .  $n$ : let  $a_n$  be the angle of elevation of  $n+1$  from  $n$  and let  $\beta_n$  be the elevation of  $n-1$  from  $n$ . Also let  $A_n$  be the azimuth of  $n, n+1$  and  $B_n$  that of  $n, n-1$  and  $c_n$  be the arc subtended at centre of earth by  $n, n+1$

$$\left. \begin{array}{l} \text{Then} \quad a_n + \beta_{n+1} = -c_n \\ \text{and} \quad A_n = \beta_{n+1} + 180^\circ - K_n \end{array} \right\} \dots \dots \dots (5)$$

where  $K_n$  is the convergence.

This traverse may be regarded as the flank of a series of triangulation: and in proceeding along it, the accumulation of azimuth error will be estimated. Now the flank of a triangulation series may, without much loss of generality be considered to proceed along a great circle of the earth (or a geodesic to be more precise). The great circles on the earth which are most conveniently considered are the meridians: but it is clear that by changing the system of coordinates to which points are referred any great circle may be regarded as a meridian of a different system of coordinates. It will accordingly be sufficient to consider the case of a meridian (not necessarily one of the system with the axes of rotation as pole). Along such a meridian the azimuthal angle  $A$  is zero or  $180^\circ$ . Suppose then that the traverse 1 2 3 . . .  $n$  lies on this meridian and that  $\mu_n$  is the component of the plumb-line deflection at  $n$  in a direction perpendicular to this meridian (but not necessarily east and west as the meridian may be any great circle).

The correction to the angle at  $n$  will now be

$$\delta C_n = \mu_n (\tan a_n + \tan \beta_n) \dots \dots \dots (6)$$

which may be written sufficiently accurately for the present purpose

$$\delta C_n = \mu_n (a_n + \beta_n) \dots \dots \dots (7)$$

since  $a_n$  and  $\beta_n$  seldom if ever are so large as  $5^\circ$  in triangulation of a geodetic kind.

Let  $\delta h_n$  be the height of  $n+1$  above  $n$ : then very approximately, if  $R$  is the radius of the earth

$$\delta h_n = R c_n (a_n + \frac{1}{2} c_n) = -R c_n (\beta_{n+1} + \frac{1}{2} c_n) \dots \dots \dots (8)$$

and

$$\delta C_n = \frac{\mu_n}{R} \left( \frac{\delta h_n}{c_n} - \frac{\delta h_{n-1}}{c_{n-1}} \right) - \frac{1}{2} \mu_n (c_{n-1} + c_n) \dots \dots \dots (9)$$

The accumulated azimuth error of the side  $n, n+1$  is accordingly  $C_n$  where

$$\begin{aligned} C_n &= \frac{1}{R} \sum \frac{\delta h_n}{c_n} (\mu_n - \mu_{n+1}) - \frac{1}{2} \sum \mu_n (c_{n-1} + c_n) \dots \dots \dots (10) \\ &= {}_1C_n + {}_2C_n. \end{aligned}$$

Some attention to detail of the limits of this summation is necessary to obtain the precise value in a particular case: but the present object is to discuss the accumulation of the error and so this detail need not be considered now. It is clear that the first expression on the right hand side of (10) is not liable to great increase: for  $\delta h_n$  is equally likely to be positive or negative

as also is  $\mu_n - \mu_{n+1}$  (considering that  $\mu_n$  is the deflection at right angles to the line  $n, n+1$ ). The most probable value of the expression  ${}_1C_n$  is

$$\frac{\sqrt{n}}{R} \cdot \frac{\delta h}{c} (\mu - \mu')$$

where  $\frac{\delta h}{c}$  and  $\mu - \mu'$  are values intermediate to the extreme values met with. To get an idea of the magnitude which this might reasonably reach after 25 sides put  $\frac{\delta h}{cR} = \frac{1}{50}$  corresponding to an angular elevation of more than  $1^\circ$  and  $\mu - \mu' = 10''$ . The value is then  $5 \times \frac{1}{50} \times 10'' = 1''$ . Now angular elevations of  $1^\circ$  are average, but changes of deflection of as much as  $10''$  in the distance between two stations are not usual, though occasionally much bigger changes occur. It is felt then that the estimate of  $1''$  for 25 rays is fair and that the danger of accumulation of error from this term is not considerable. The second term of (10) remains and its magnitude is liable to be somewhat greater. It may be written

$${}_2C_n = -\mu_m \Sigma c$$

where  $\mu_m$  is some value intermediate to the extreme values of  $\mu$  met with:  $\Sigma c$  is merely the whole arc subtended by the terminal stations at the earth's centre. Taking  $\Sigma c = \frac{1}{10}$  radian which corresponds to a series about 400 miles long we get

$${}_2C_n = -\frac{1}{10} \mu_m$$

In a series along the first range of the Himalayas deflections at right angles to this range of as much as  $40''$  are of common occurrence. If  $\mu_m = 40''$

$${}_2C_n = -4''$$

Now this is an error of magnitude about what might possibly occur in a single angle at which  $\mu = 60''$  and  $\tan a = \frac{1}{15}$ : so that the conclusion may be drawn that the danger of failing to correct observed angles for deflection of the plumb-line is almost confined to the angles themselves and is not liable to produce a cumulative error of azimuth, if the angles were utilised as in a traverse. The effect however may be felt in a way different from that considered above owing to the distribution of triangular error. Each triangle which contains a station where the deflection differs considerably from those at the other stations is liable to be deformed when the angles are adjusted to equal two right angles plus the spherical excess calculated on the spheroid. The amount of this deformation and the effect on computed coordinates of the stations of the triangulation do not appear to be such as can be estimated for a general case. Its effect in the actual triangulation of India is mixed up with the effect of error of observation and its amount is in general considerably less, as appears from the solution of the modified Laplace equations given above in §§ 10—12.

19. In observations for azimuth the result of using a point at considerable angular elevation about the station of observation as a reference point seems to have always been ignored. Yet the same geometrical fact which causes the horizontal trace of the ray through the pole to be displaced in azimuth also gives rise to an azimuthal deflection of the horizontal trace of the ray through the reference point, unless it happens that the ray is in the azimuth at right angles to that of geoidal deflection. Suppose that at any point the southerly and westerly deflections are  $\eta, \xi$  respectively. The horizontal (spheroidal) trace of the ray through the pole and geoidal zenith will be deflected in azimuth by  $+\xi \tan \lambda$ ,  $\lambda$  being the altitude of the pole. The horizontal trace of a ray in azimuth  $A$  and angular elevation  $a$  will be deflected by

$$-\tan a (\xi \cos A - \eta \sin A) \quad \text{vide (4)}$$

The difference between astronomic and geodetic azimuth is accordingly

$$\xi \tan \lambda + (\xi \cos A - \eta \sin A) \tan a$$

instead of the simpler expression  $\xi \tan \lambda$  usually taken, which is more closely approximated to as  $a$  decreases.

20. The advantages and disadvantages of the methods of correction worked out in Chapters I, III, IV may now be considered. The method of Chapter I in which  $u_N$  and  $v_N$  are taken as the changes of latitude and longitude has the justification and weaknesses referred to on pages 10-12. In a triangulation system where the bulk of the triangulation is along parallels and meridians this solution would be satisfactory were it not for the azimuths. The azimuth computed by the corrections at the ends of a ray of triangulation along a parallel differs by an appreciable amount from those for a ray along a meridian, and at first sight it appears that the difference is the necessary correction to the angle contained by these two rays. This however is not satisfactory as it is clear that the longitudinal and meridional series must be a little bent and that the whole error should not be forced into the junction angles. It would be equivalent to putting all the angular closing errors of a traverse which followed approximately the sides of a rectangle into the four angles at the corners of the rectangle. Moreover the final azimuth will not agree with the longitude as laid down in Laplace's equation. Laplace's equation might be adopted as a mode of determining the azimuth changes: but obviously the result would be inconsistent with the latitude and longitude changes found viz.  $u_N, v_N$ . In fact it appears that this method could rightly be applied merely to the junction points of the triangulation series. After changes for these points had been found, the corresponding changes along the series might be adjusted as is done in closing a traverse. This would involve a consideration in detail of all the series and has the disadvantage of being most laborious, and when done it is inconvenient in that the solution for the case of a further change of axes would have to be taken up right from the beginning. It might be supposed to be advantageous in that it takes cognisance of the actual form of the triangulation: but seeing that it is based on a method which is not entirely justifiable there seems to be little advantage in this partial approach to accuracy in the final stages of the reduction.

21. The method of computation along the geodesics at once gets rid of the difficulty of dual values of the changes. It is obvious however that the values obtained for the changes vary according as one origin or another is selected: for the closing errors in a triangle formed by joining any point and two selected origins exist just as much here as in the first method. This closing error however does not occur all at one point as in that method, but is satisfactorily distributed. There is some trouble in computing the geodesics: but this is of minor importance seeing that it has been done once for all and correction tables have been made out from which the coordinates of any point may be deduced by interpolation. These tables permit of the changes due to any desired changes in the elements  $a, b$  and latitude and azimuth at the origin being made immediately and admit of further changes being subsequently made when this becomes desirable.

Both this and the first method are based on the idea of the accuracy of the ratios of the sides in the triangulation: this is almost independent of small changes in the spheroid and the consequent minute changes in the spherical excess of any observed triangle. It may be noted here that in the case of an equilateral triangle with observed angles of equal weight the ratios are unaffected by the amount of spherical excess as this would be distributed equally. But the angles of the geoid have been used in place of spheroidal angles and from this some disturbance must have arisen.

While then the ratios of the sides may be regarded as practically perfect so far as corrections due to size of spheroid are concerned, it must at the same time be remembered that the observation errors have a cumulative effect on the ratio of a side to the original base as the side considered is separated more widely from the base: and the treatment of the observed angles as applicable to the spheroid without correction will aggravate this. The magnitude of the errors

so developed is indicated where closure has been made on additional base lines. It is clear that in a network of triangulation these additional bases may be reached from the origin by various routes and the length of these routes must accordingly be duly considered.

The following figures are taken from the circuits and base-lines of the N.W. Quadrilateral\* in which there are 5 circuits and 4 measured bases.

TABLE XLIII.

(1) Number of equation	(2) Logarithmic closing error $\times 10^6$	(3) Number of triangles	(2) <sup>2</sup> + (3)
1	4.40	51	0.380
2	6.82	96	0.465
3	7.19	36	1.438
4	7.96	95	0.666
5	16.38	123	2.180
6	12.46	185	0.841
7	15.09	88	2.586
8	0.53	138	0.002
			Sum = 8.558
			Mean = 1.070
			Square root of Mean = 1.035

It appears that the mean error per triangle in side ratio is 1.035 in the 6th place of logs which corresponds to an error of one part in 420,000 showing that a high order of accuracy has been attained.

22. In the third method, that of geometrical transformation, the idea of the spheroid as merely a figure of reference is used as a basis for the argument. It is free from the difficulties of multiple values, one for each route traversed, and gives a definite set of values for the changes at any point. Being geometrically correct it naturally satisfies the Laplace condition: but it does not keep the constancy of side ratios, though the departure from constancy is not serious. No attempt is made to correct for the distance between geoid and spheroid which in conjunction with large deflections such as have so far been discovered would make very small changes in the coordinates. It is to be remembered however that the original geoidal triangles have been applied without angular correction to the old spheroid of reference, although considerable corrections must have been necessary in some cases. To put this matter right now, deflections at many stations would need to be observed: and to make use of the information that might be gained by observation, it would be necessary to re-grind the whole triangulation of India. It is the object of the present investigation to avoid this immense piece of work: but as has just been pointed out, it could not be undertaken until many deflection observations had been made. Had the corresponding corrections been made in the first instance, which would have been possible if comparatively rough latitudes and azimuths had been observed at each triangulation station, the method of change of coordinates explained in Chapter IV would have been absolutely correct. The fact that this was not done is the source of the present difficulty and

\* Vide G. T. Volume II of the Survey of India, pages 303, 304.

practically disposes of the usefulness of this method. Further the Laplace equations should by right have been applied in the original grinding: they were not. This omission also makes an objection to any method of computing short of regrinding. And so the geometrical accuracy of the method of Chapter IV is vitiated. It is useless to insist on a method which strictly accords with Laplace's equations when the original quantities which are to be corrected fail to satisfy those equations.

23. As remarked above in §7 the reason of the multiple values of  $u$ ,  $v$ ,  $w$  according to traverse route followed is that in the computations no attempt has been made to correct observed geoidal angles to angles on the particular spheroid which is selected as a reference figure. Had these corrections been applied the method of geometrical change explained in Chapter IV would have given the changes  $u_r$ ,  $v_r$ ,  $w_r$  which would then have been applicable on changing from one spheroid of reference to another. But seeing that no such corrections were applied, and that the closing errors of circuits were dispersed and treated as errors of observations it is clear that this method is not strictly applicable. The portion of the closing errors due to this lack of correction to the angles is small compared with those due to errors of observation: so in the main no great fault was committed. A greater fault was the neglect of closing on the longitude arcs, or in other words applying Laplace's longitude equation. What is at first sight naturally regarded as a defect of the methods of Chapters I and III is that equation (4) of this chapter is not satisfied. But when it is considered that Laplace's condition, of which equation (4) is an immediate consequence, was not enforced on the computation of the original triangulation, it is clear that there is nothing to be gained by now enforcing equation (4) on to the small changes to be applied an account of change of spheroid. A preferable course is to make these changes and then apply Laplace's condition to the final result as Sir Sidney Burrard has done in his discussion of the Indian azimuth observations. It is concluded then that the two objections to the methods of Chapters I, III cited above have little weight in view of the slight inaccuracies of method by which the Indian triangulation has been reduced. It remains then to decide merely on what route should be followed in deducing the changes of coordinates by the method of Chapter I. This method is applicable to any route if the "closing errors" are applied as explained at the end of that chapter: and in Chapter III although the results are obtained in a special way, yet these results might have been obtained by the method of Chapter I. It is clear that no route can be laid down as rigorously correct and that the best that can be done is to select a route which appears to be the best. Suppose there were four triangulation series all of equal merit forming a square. Then the route which should be followed from one corner to another is the diagonal, and this produces a result intermediate to those which would be found by following either pair of sides. If one pair of sides was distinctly better triangulation than the other, the best route would doubtless be one closer to the good sides than the bad sides. But in the case of a great network of triangulation it is too complicated to go into such detail and so the diagonal would be selected. Now the geodesic corresponds fairly closely in the case of a spheroid to the diagonal of any square or rectangle, and it gives a satisfactory medial path among the triangulation and medial results for the changes deduced. This choice is slightly arbitrary, but seems the best that can be made. Another arbitrary choice is that of the origin from which changes are computed on the point from which all the selected geodesics radiate. It is apparent from the theory of the "closing errors" that different values would be deduced for the change according as this central point is selected; but the differences are not really appreciable in comparison with the errors due to faulty observation: Kalianpur is very centrally situated as regards India, and as it is the origin of the triangulation it appears that it would merely be an unnecessary complication to select a slightly different point for the point from which the geodesics radiate. It would be useless to go into any great refinement as to

the theoretically best centre for this purpose because it would be constantly disturbed as new triangulation is added to the Indian system.

It appears then that, without there being any rigidly accurate reason for adopting the method and results of Chapter III, yet that it meets the present case quite satisfactorily and has no defect of appreciable magnitude: and that any defect of theoretical precision would be present in any alternative method which might be proposed. The conclusion quoted at the end of Chapter IV is accordingly reiterated, namely that **the method of calculation along geodesics through Kalianpur as set forth in Chapter III is the correct one to use.**



## CHAPTER VI.

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### Strength and Adjustment of Triangulation. Mechanical Analogy.

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#### A criterion of strength of triangulation series.

1. If a mechanical network, which is analogous to a triangulation series in the sense explained in § 10 below, replaces each series in a system of triangulation a mechanical framework is formed. Each mesh in this framework corresponds to a circuit in the triangulation and needs straining to close in a way exactly analogous to the need of adjustment of triangulation circuits. Now in general in Indian triangulation, series follow approximately straight lines and these are generally more or less along meridians or parallels. Consider four series which form a circuit  $A B C D$ , and their mechanical analogues. To effect closing at  $A$  in the mechanical framework strains must be applied: and it seems fairly clear that the strains which will be caused in the side  $A B$  will be of the same nature all along this side, but will differ essentially from those caused in  $B C$ . An example would be that the strains in  $A B$  would be such as to increase the length  $A B$  while those in  $B C$  would be to slightly curve  $B C$ . On this account it appears desirable to consider strains of a particular type as existing throughout  $A B$ : but not existing to the same extent in  $B C$ . The side  $A B$  is thought of as having uniform strength which differs from the uniform strength of  $B C$ . Reverting to the triangulation series it may be remarked that in one series the same strength is aimed at throughout, angles being observed with similar precision and figures of the same type selected as far as possible. When topographical conditions change entirely, as must essentially occur on the passage from plain to hilly country, the series should be considered in sections.

In considering one series of a circuit, it is only necessary to think of a "route" formed by those sides which persist in the general direction of the series, but bearing in mind that the length of each side is expressed in terms of the previous side, and that in any adjustment its relative length to this previous side is the quantity which is to be slightly varied; and similarly its azimuth is relative to the previous side. It will not be very far from the truth in effect if these several sides are for simplicity regarded as of equal length  $l$  and practically in the same direction. Suppose the angle between the  $r^{\text{th}}$  and  $(r+1)^{\text{th}}$  side, originally practically  $180^\circ$ , is changed by the small angle  $\eta_r$ : and that the ratio of the length of the  $(r+1)^{\text{th}}$  side to the  $r^{\text{th}}$ , originally unity, is changed to  $1 + \epsilon_r$ . Expressed in terms of the base, the  $r^{\text{th}}$  side changes in length in the ratio  $\prod_1^r (1 + \epsilon) - 1 \doteq \sum_1^r \epsilon$  and in direction by  $\sum_1^r \eta$ . These quantities  $\epsilon$  and  $\eta$  may be chosen to give any possible small changes in the length and azimuth of the terminal side of the series. Consider the displacements in the terminal

point  $r + 1$ . That in the direction of the series is clearly

$$l\epsilon_1 + l(\epsilon_1 + \epsilon_2) + \dots + l(\epsilon_1 + \dots + \epsilon_r) = \frac{l}{r} (r\epsilon_1 + \overline{r-1}\epsilon_2 + \dots)$$

where  $s$  is the length of the series and accordingly  $s = rl$ .

The most probable value of this is

$$\frac{s\epsilon}{r} \left( r^2 + \overline{r-1}^2 + \dots \right)^{\frac{1}{2}} = \frac{s\epsilon}{r} \sqrt{\frac{r(r+1)(2r+1)}{6}} \doteq s\epsilon \sqrt{\frac{r}{3}} \dots (1)$$

if  $r$  is large,  $\epsilon$  being the most probable value of any of the quantities  $\epsilon_r$ . Similarly the displacement of the terminal point  $r$  in the direction at right angles to the series is

$$l\eta_1 + l(\eta_1 + \eta_2) + \dots \doteq s\eta \sqrt{\frac{r}{3}} \dots (2)$$

Both the quantities  $\epsilon$  and  $\eta$  depend on the probable error of an angle (adjusted by the triangular conditions) in the series. General Ferrero introduced the quantity "m" as a criterion of triangulation,

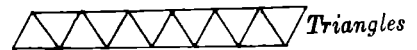
where 
$$m = \sqrt{\frac{\sum \Delta^2}{3n}}$$

$\Delta$  being the triangular error of any triangle and  $n$  the number of triangles considered. This quantity "m" is accordingly the error of mean square of one angle of a triangle. The probable error of an observed angle is  $\cdot 6745 m$ . The probable error of an angle adjusted to satisfy the triangular condition is

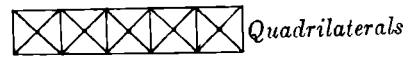
$$a = \sqrt{\frac{2}{3}} \times \cdot 6745 m = \cdot 551 m \dots (3)$$

The formulæ (1) and (2) may accordingly be expressed by saying that the probable displacement of the terminal point of a series of given length in any direction varies as  $m/\sqrt{l}$ .

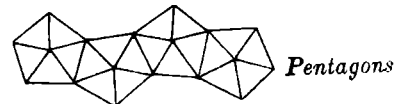
They accordingly show the advantages of figures with long sides. It remains to consider the effect of various types of figures in the series, simple triangles, braced quadrilaterals, central pentagons, hexagons etc. Only regular figures are considered and the sides of each are taken equal to  $l$ . To complete a series of each type of given length  $s$  suppose there are  $n_3, n_4, n_5$  etc. figures, simple triangles, braced quadrilaterals, pentagons, etc. It is clear that the gains in distance in the required direction are



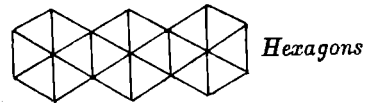
Triangles



Quadrilaterals



Pentagons



Hexagons

- $n_3 l \cos 60^\circ$  in the case of simple triangles
- $n_4 l$  . . . . . quadrilaterals
- $n_5 l (\cos 18^\circ + \frac{1}{2} \cos 54^\circ)$  . . . . . pentagons †
- $2n_6 l \cos 30^\circ$  . . . . . hexagons

and as these must all be equal to  $s$

$$\therefore \frac{1}{2} n_3 = n_4 = 1.245 n_5 = 1.732 n_6$$

Now the probable errors in the determination of a terminal side after a given number of figures of each of the kind mentioned are in the ratio

$$0.82 : 1 : 1.21 : 1.29 *$$

so that at the end of each series of the same length the errors are in the ratio

$$0.82 \sqrt{n_3} : \sqrt{n_4} : 1.21 \sqrt{n_5} : 1.29 \sqrt{n_6}$$

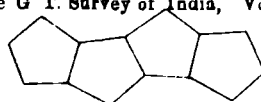
which reduce to

$$1.17 : 1 : 1.08 : 0.98$$

for the cases of triangles, quadrilaterals, pentagons and hexagons respectively.

\* Vide Account of the Operations of the G. T. Survey of India, Vol. II, p. 199.

† An alternative arrangement of the pentagon



gives a result practically the same.

Suppose there is a series composed of  $\alpha$  simple triangles,  $\beta$  braced quadrilaterals,  $\gamma$  pentagons and  $\delta$  hexagons, then the ratio of its terminal probable errors to those of a series of the same length composed of quadrilaterals is

$$\sqrt{\frac{1.17^2 \alpha + \beta + 1.08^2 \gamma + .98^2 \delta}{\alpha + \beta + \gamma + \delta}} : 1$$

which may be approximately written  $1 + f : 1$

where 
$$f = \frac{1}{12} \cdot \frac{2\alpha + \gamma}{\alpha + \beta + \gamma + \delta} \dots \dots \dots (3)$$

Heptagons, nonagons etc. occur rarely and may be treated as pentagons. Octagons, decagons etc. may be treated as hexagons. Combining this result with (1) and (2) the quantity

$$M = (1 + f)^m \sqrt{\frac{18}{l}} \dots \dots \dots (4)$$

is formed in which 18, the average length in miles of sides in the Indian triangulation, is introduced, and " $m$ " is General Ferrero's expression for error of mean square of an angle and  $l$  is the average length of side expressed in miles in the series under consideration.

2. This quantity  $M$  takes cognizance not only of the probable error of the angles in the triangulation but also of the length of side and type of figure. For a given length of triangulation it gives a relative idea of the errors likely to occur in series of different precision and type: for example if there are several series of the same length, say 300 miles each, for which values  $M_1, M_2, M_3 \dots$  have been found by (4), then the probable errors of northing or easting of the terminal point are approximately in ratio  $M_1 : M_2 : M_3 \dots$  and the same is true of the probable errors of length or azimuth of the terminal side. " $M$ " gives a criterion of the value of triangulation considering in proper proportion the excellence of observation and the success in choosing well-proportioned figures which has been attained: " $m$ " only gauges the excellence of observation.

The deduction of the quantity  $M$  is confessedly based on approximations and simplifications. It would not be expected to be very accurate if applied to badly conditioned figures, and it is not intended that this should be done. In geodetic triangulation such figures are exceptional and figures approximately symmetrical largely predominate: and in these cases  $M$  is a practically useful criterion of the excellence or strength of the series.

3. All the triangulation of India has been classified according to values of  $M$  (*vide* table XLIV) and the order of merit of the several series deduced. The series are arranged in chronological order and designated by a serial number. Reference to any series can generally be made more conveniently by use of its serial number than by the rather long and frequently artificial names which have been applied. A consideration of the list shows that the principal and secondary triangulation ranges fairly continuously from very high class work in the best of which No. 76 North Baluchistan Series  $m = 0.221$  and  $M = 0.17$ ; to the least successful secondary triangulation No. 65 Siam Branch in which  $m = 3.711$  and  $M = 4.34$ . The mean square (*vide* note at foot of table) value of  $M$  for the triangulation which was utilised in the grinding of the Indian network is 1.04: that for the whole triangulation 1.51. In some cases so called secondary triangulation proves better than poor principal triangulation: in general there is no marked gap between the two classes. This classification of triangulation into principal and secondary is accordingly dropped after the completion of the series and both are classed as "geodetic" triangulation and placed according to the values of  $M$  yielded by them. The further distinction in Indian triangulation is between "geodetic" and "minor" triangulation. The former is always rigorously computed taking account of spherical excess. The latter, which is generally very much rougher, disregards spherical excess.

TABLE XLIV.

Values of "m" and "M" for all Geodetic Series of the Indian Triangulation.

No	Name of Series	Seasons	± m	l	Number of Independent Figures												f	± M	Order of Merit
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
1	South Pārasnāth Mer.	1831-39	3·308	24·0	6	...	...	...	...	...	...	...	...	...	2	·139	3·26	92	
2	Budhon Meridional ...	1833-43	2·242	19·2	25	1	2	4	...	...	...	...	...	...	...	·135	2·46	86	
3	Amūa Meridional ...	1834-38	1·647	18·9	34	...	...	...	...	...	...	...	...	...	...	·167	1·88	77	
4	Rangīr Meridional ...	1834-64	1·643	20·6	32*	...	...	...	...	...	...	...	...	...	...	·167	1·79	72	
5	Calcutta Longitudinal	1834-69	0·369	26·6	...	...	4	2	2	2	1	1	1	...	...	·051	0·32	8t	
6	Great Arc Meridional, Section 24°-30° ...	1835-66	0·708	22·2	10	...	2	2	...	...	...	...	1	4	·109	0·71	36b		
7	Bombay Longitudinal	1837-63	0·844	27·6	1	1	2	2	...	...	...	...	...	...	2	·077	0·74	38	
8	Great Arc Meridional, Section 18°-24° ...	1838-41	0·567	21·3	17	3	4	...	1	...	...	...	...	...	3	·118	0·59	28b	
9	Great Arc Meridional, Section 8°-15° ...	1840-74	0·390	23·7	1	4	2	5	3	1	...	...	...	...	3	·054	0·36	13b	
10	Singī Meridional ...	1842-62	1·187	24·9	18	3	...	...	...	...	...	...	...	...	2	·131	1·14	51	
11	South Konkan Coast	1842-67	2·176	29·6	16	3	...	...	...	...	...	...	...	...	...	·140	1·93	79	
12	Karāra Meridional ...	1843-45	1·507	16·4	21	1,	...	...	...	...	...	...	1	...	...	·146	1·81	73	
13	North Malūncha Mer.	1844-46	1·266	18·4	6	...	1	2	...	...	...	...	...	...	...	·130	1·42	60	
14	Chendwār Meridional	1844-69	0·841	15·1	17	1	1	1	...	...	...	...	...	...	...	·146	1·06	45	
15	Gora Meridional ...	1845-47	0·973	15·6	23	...	...	...	...	...	...	...	...	1	...	·161	1·21	52b	
16	Calcutta Meridional...	1845-48	1·173	8·5	45	...	...	...	...	...	...	...	...	...	...	·167	1·99	82	
17	South Malūncha Mer.	1845-53	1·606	15·7	11	2	...	...	...	...	...	...	...	...	...	·141	1·97	81	
18	Khānpisura Meridional	1845-62	1·227	27·3	...	4	1	1	2	...	...	...	...	...	1	·071	1·07	46	
19	Gurcāni Meridional...	1846-47	1·165	13·8	32	...	...	...	...	...	...	...	...	...	...	·167	1·55	65	
20	North-East Lon. ...	1846-55	0·446	11·1	96	2	1	3	1	...	...	...	...	...	...	·156	0·65	31b	
21	Hurilōng Meridional	1848-52	1·502	14·3	20	...	...	1	...	...	...	...	...	...	1	·142	1·92	78	
22	North-West Himalaya	1848-53	0·641	25·3	...	6	1	2	...	...	...	...	...	...	3	·021	0·55	26	
23	Gurhāgarh Meridional	1848-62	0·914	13·6	70	...	1	2	1	...	...	...	...	...	1	·157	1·21	52b	
24	East Coast ...	1848-63	0·608	16·9	22	6	3	1	...	...	...	...	...	...	2	·116	0·70	31b	
25	Karūchi Longitudinal	1849-53	0·558	15·8	...	10	2	10	2	2	...	...	...	...	1	·015	0·60	30	
26	Abū Meridional ...	1851-52	0·617	15·9	1	...	3	...	...	...	...	...	...	...	...	·042	0·68	33	
27	North Pārasnāth Mer.	1851-52	0·895	12·6	20	...	...	...	...	...	...	...	...	...	...	·167	1·25	54	
28	Kāthiāwār Meridional	1852-56	0·990	17·4	7	3	1	1	...	...	...	...	...	...	3	·101	1·11	49	
29	Gujarāt Longitudinal	1852-62	0·859	14·2	31	...	2	...	...	...	...	...	...	...	...	·157	1·12	50	
30	Kāthiāwār Minor Lon.	1853	1·481	23·1	1	5	...	...	...	...	...	...	...	...	...	·028	1·34	58	
31	Sābarmati Secondary	1853-54	1·318	5·3	15	2	...	...	...	...	...	...	...	...	...	·147	2·84	88	
32	Great Indus ...	1853-61	0·359	12·7	...	9	1	2	2	2	3	1	...	...	·008	0·13	20b		
33	Rahūn Meridional ...	1853-63	0·327	14·1	...	2	15	2	...	...	...	...	...	...	...	·009	0·37	15t	

Mer. = Meridional

Lon. = Longitudinal.

\* One centred Δ.

|| Centred Quadrilateral.

TABLE XLIV.—(Contd.)

Values of "m" and "M" for all Geodetic Series of the Indian Triangulation.

No.	Name of Series	Seasons	±m	l	Number of Independent Figures											f	±M	Order of Merit	
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
34	Assam Longitudinal...	1854-60	0.579	12.5...		2	3	4	...	1							0.025	0.71	36b
35	Cutch Coast ...	1855-58	0.986	12.5...		5	6	...	1								1.074	1.27	55b
36	Kashmīr Principal ...	1855-60	0.884	19.2...		18	1	1	...								0.004	0.86	40
37	Jogī-Tila Meridional	1855-63	0.481	12.3...				7	1	1							0.009	0.59	28b
38	Sambalpur Lon. ...	1856-57	0.806	19.3 7		2		1	...								0.117	0.87	41
39	(Cutch) Coast Line Sec.	1856-60	0.975	10.8 33													0.167	1.47	61
40	Kāthiāwār Minor Meridional No. 1 ...	1858-59	0.930	9.1 13		1											0.155	1.51	64
41	Kāthiāwār Minor Meridional No. 2 ...	1859-60	1.247	12.1 14		1											1.145	1.75	70
42	Kāthiāwār Minor Meridional No. 3 ...	1859-60	0.969	10.0 17		3	1										0.139	1.48	62
43	Bider Longitudinal ...	1859-72	0.311	22.0 ...		1,  1	2	1	...	3							1.064	0.30	52
44	Eastern Frontier or Shillong Meridional	1860-64	0.409	13.2 ...		6	2	1	1	...	1						0.028	0.49	23b
45	Sutlej Meridional ...	1861-63	0.346	10.6 50													0.167	0.53	25
46	Madras Mer. and Coast	1861-68	0.426	21.6 ...		3	2	6	3	2							0.026	0.40	19
47	Kāthiāwār Minor Meridional No. 4 ...	1863-64	1.154	10.8 14													1.157	1.73	69
48	East Calcutta Lon. ...	1863-69	0.379	10.7 32				2									0.157	0.57	27
49	Mangalore Meridional	1863-73	0.440	20.7 1		1	1	4	2								2.096	0.45	22
50	Kumaun and Garhwāl	1864-65	1.742	26.7 2		4		1									0.048	1.50	63
51	Nāsik Secondary ...	1864-65	2.033	10.5 26													0.167	3.12	91
52	Burma Coast ...	1864-82	0.380	19.8 14		18	5	5	1	2							4.078	0.39	18
53	Jabalpur Meridional	1865-67	0.340	22.4 ...		2		7	1								0.008	0.31	7
54	Madras Longitudinal	1865-80	0.384	21.4 1		1	3	6									0.038	0.37	15f
55	Assam Valley Triangulation	1867-78	1.690	9.5 46		5,  3	1										0.141	2.65	87
56	Brahmaputra Mer. ...	1868-74	0.564	12.0 ...				1	6	...	1						1.009	0.70	31b
57	Coimbatore Minor No. 1	1869-71	1.547	13.6 19				1									0.163	2.07	83
58	Rīlāspur Meridional ...	1869-73	0.302	15.7 ...		6	4	6									0.021	0.33	11
59	Cuddapah Minor ...	1871-72	0.826	17.6 8		1											0.148	0.96	43b
60	Hyderabad Minor ...	1871-72	1.405	19.9 9													0.167	1.56	66
61	Malabar Coast ...	1871, 74, 80	1.532	17.4 12													0.167	1.82	74
62	Jodhpore Meridional	1873-76	0.291	15.6 ...		3	1	7	1	...	1						0.019	0.32	87
63	South East Coast ...	1875-79	0.522	11.7 ...		8		11	2								1.007	0.65	31b

Mer. = Meridional.

Lon. = Longitudinal.

Sec. = Secondary.

Centred Quadrilateral.

T A B L E X L I V.—(Contd.)

Values of “m” and “M” for all Geodetic Series of the Indian Triangulation.

No.	Name of Series	Seasons	± m	l	Number of Independent Figures												f	± M	Order of Merit.
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
64	Eastern Sindh Mer.	1876-81	0.244	13.0	...	3	2	5	2	...	...	...	...	...	...	0.28	0.30	5b	
65	Siam Branch Triangulation	1878-81	3.711	16.1	7	4	...	...	...	...	...	...	...	...	1	1.07	1.31	9a	
66	Mandalay Meridional	1889-95	0.418	27.0	...	13,  3	...	...	...	...	...	...	...	...	1	0.09	0.35	12	
67	Mong Hsat Secondary	1891-93	3.054	24.0	9	1	...	...	...	...	...	...	...	...	1	1.39	3.01	89b	
68	Manipur Longitudinal	1894-99	0.453	28.4	...	5,  2	...	...	...	...	...	...	...	...	...	0.00	0.36	13b	
69	Makran Longitudinal	1895-97	0.285	23.9	2	2,  1	1	...	...	...	...	...	...	...	...	0.69	0.26	4	
70	Mandalay Lon.	1899-1909	1.696	17.2	8	2	1	...	...	...	...	...	...	...	...	1.29	1.96	80	
71	Manipur Minor Mer.	1899-1902 1915-16	0.750	20.9	31	1	...	...	...	...	...	...	...	...	...	1.61	0.81	39	
72	Great Salween	1900-11	0.404	32.6	...	2,  4	2	...	...	...	...	...	...	...	1	0.56	0.32	8t	
73	Kidarkanta Secondary	1902-03	1.323	15.2	3	1	...	...	...	...	...	...	...	...	...	1.25	1.62	67	
74	Kalāt Longitudinal	1904-08	0.365	39.7	...	6,  5	1	...	...	...	...	...	...	...	...	0.00	0.25	3	
75	Baluchistān Triangulation	1908-09	1.348	33.2	...	1	...	...	...	...	...	...	...	...	1	0.83	1.08	47b	
76	North Baluchistān	1908-10	0.221	32.7	...	5,  3	1	...	...	...	...	...	...	...	...	0.09	0.17	1	
77	Gilgit	1909-11	0.443	31.6	5	3,  1	...	...	...	...	...	...	...	...	...	0.93	0.37	15t	
78	Khāsi Hills Secondary	1909-11	2.038	10.7	14	3	...	...	...	...	...	...	...	...	...	1.37	3.01	89b	
79	Mawkmai Secondary	1909-11	1.575	10.9	41	1	...	...	...	...	...	...	...	...	...	1.63	2.35	85	
80	Upper Irrawaddy	1909-11	0.596	30.6	4	5	...	...	...	...	...	...	...	...	...	0.74	0.49	23b	
81	Jaintia Hills Sec.	1910-11	0.986	6.9	23	...	...	...	...	...	...	...	...	...	...	1.67	1.86	76	
82	Bhīr Secondary	1911-12	0.794	17.4	24	...	...	...	...	...	...	...	...	...	...	1.67	0.94	42	
83	Ranchi Secondary	1911-12	1.840	15.2	13	...	...	...	...	...	...	...	...	...	...	1.67	2.34	84	
84	Villupurām Secondary	1911-12	1.184	10.9	18	...	...	...	...	...	...	...	...	...	...	1.67	1.78	71	
85	Sambalpur Meridional	1911-14	0.250	25.7	...	6,  4	1	1	...	...	...	...	...	...	...	0.07	0.21	2	
86	Indo-Russian Connection	1912-13	2.790	10.9	11	7,  2	...	...	...	...	...	...	...	...	...	0.92	3.92	93	
87	Khandwa Secondary	1912-13	0.999	15.2	22	...	...	...	...	...	...	...	...	...	...	1.67	1.27	55b	
88	Ashta Secondary	1913-15	1.048	15.3	21	...	...	...	...	...	...	...	...	...	...	1.67	1.33	57	
89	Buldāna Secondary	1913-14	0.304	12.3	18	...	...	...	...	...	...	...	...	...	...	1.67	0.43	20b	
90	Naldrug Secondary	1913-14	1.465	15.2	27	1	...	...	...	...	...	...	...	...	...	1.61	1.85	75	
91	Nāga Hills Secondary	1913-14	0.913	21.3	7	1	1	...	...	...	...	...	...	...	...	1.39	0.96	43b	
92	Middle Godaverī Sec.	1914-15	0.913	17.1	14	1	...	...	...	...	...	...	...	...	...	1.56	1.08	47b	
93	Kohima Secondary	1914-15	1.094	15.0	13	...	...	...	...	...	...	...	...	...	...	1.67	1.39	59	
94	Cachar	1914-15	1.077	10.5	10	...	...	...	...	...	...	...	...	...	...	1.67	1.65	68	

Mer. = Meridional. Lon. = Longitudinal. Sec. = Secondary. || Centred quadrilateral.

For 42 Series entering the Simultaneous Grinding (shown in italics above)

$$\Sigma M^2 = 45.3591 \quad \text{Mean Square } M = \pm \sqrt{\frac{45.3591}{42}} = \pm 1.04$$

4. Replace each triangulation series by one of its flanks. The network is then nearly similar to a traverse network, with the addition that closure of the length of the last side is necessary as well as its azimuth and the position of its terminal point. The flank of any series is usually not far from straight: or else consists of two or more portions with approximately straight flanks. Consider each such portion separately and denote it by the name "*triangulation line*."

Each triangulation line is liable to be slightly bent and to have its length slightly altered in the course of adjustment. This is effected by the angle at each station, and by the ratio of successive sides (between stations) of the triangulation line being slightly changed. Triangulation lines may be of different strengths according to the series from which they are derived: but it will be assumed that the strength of any one triangulation line is uniform. In other words, if it is necessary to adjust the azimuth at the end of a triangulation line this would be done correctly by giving the angles at all its stations an equal change: and to adjust the length of terminal side it would be correct to change the ratios of successive sides each by equal percentages.

When several triangulation lines are concerned the angular adjustment at any station of one will in general be different from that at any station of any other on account of both the different strengths of the several triangulation lines as well as their directions. The question of strengths has been considered in some detail above (*vide* § 1) and can be taken into account by means of M.

Adjustments of latitude and longitude at the end of a triangulation line may also be effected by a combination of small changes of the angles at the stations and the ratios of sides between stations: but in these cases the most probable adjustment would not be that of changing all the angles and the successive side ratios by equal amounts. The actual difference of this latter course from the most probable one is not very great in triangulation lines of moderate length, and it may be deemed justifiable on the ground of simplicity to make the adjustment by adopting the latter. This would bring the four types of adjustment into one simple scheme: but the more general case will now be explained and the simple case can easily be deduced from this if desired by omission of certain terms.

5. Consider any triangulation line and let the successive stations along the line be denoted by the numbers 0, 1, 2, . . . n, there being altogether n sides in the line. Let  $A_r$  be the azimuth at r of r+1 and  $\lambda_r, L_r$  the latitude and longitude of r. Denote by  $c_r$  the length of the  $r^{\text{th}}$  side and by  $\Delta\lambda_r, \Delta L_r, \Delta A_r$  the increments of latitude, longitude and azimuth along this side (r-1, r). Suppose that the angle at the station r is changed by  $\eta_r$  radians and the ratio of the  $(r+1)^{\text{th}}$  to the  $r^{\text{th}}$  side to  $\frac{c_{r+1}}{c_r} (1 + \epsilon_r)$  and consider what changes will be caused thereby.

The following expressions hold approximately

$$\left. \begin{aligned} \Delta\lambda_r &= -\frac{c_r}{a} \cos A_{r-1} \\ \Delta L_r &= -\frac{c_r}{a} \sin A_{r-1} \sec \lambda_{r-1} \\ \Delta A_r &= -\frac{c_r}{a} \sin A_{r-1} \tan \lambda_{r-1} \end{aligned} \right\} \dots \dots \dots (5)$$

in which  $\Delta\lambda_r, \Delta L_r, \Delta A_r$  are expressed in radians. The differences between  $\rho, \nu$  the principal radii of curvature and a the mean radius are neglected as only approximate equations are required in what follows.

Differentiate (5) with regard to  $c_r, A_{r-1}, \lambda_{r-1}$ . Denoting the changes in latitude, longitude, back azimuth and forward azimuth at station r by  $u_r, v_r, w_r, w_r + \eta_r$  it follows that the change in  $\Delta\lambda_r$  is  $u_r - u_{r-1}$  etc., so that

$$\left. \begin{aligned} u_r - u_{r-1} &= \Delta \lambda_r \left\{ \frac{\delta c_r}{c_r} - \tan A_{r-1} (w_{r-1} + \eta_{r-1}) \right\} \\ v_r - v_{r-1} &= \Delta L_r \left\{ \frac{\delta c_r}{c_r} + \cot A_{r-1} (w_{r-1} + \eta_{r-1}) + \tan \lambda_{r-1} u_{r-1} \right\} \\ w_r - w_{r-1} - \eta_{r-1} &= \Delta A_r \left\{ \frac{\delta c_r}{c_r} + \cot A_{r-1} (w_{r-1} + \eta_{r-1}) + \sec \lambda_{r-1} \operatorname{cosec} \lambda_{r-1} u_{r-1} \right\} \end{aligned} \right\} \dots (6)$$

In these equations  $\cot A_{r-1}$  or  $\tan A_{r-1}$  is liable to be inconveniently large: but this is always accompanied by either  $\Delta \lambda_r$  or  $\Delta L_r$  being correspondingly small. It is convenient to eliminate  $A_{r-1}$  by means of (5) and to write (6)

$$\left. \begin{aligned} u_r - u_{r-1} &= \Delta \lambda_r \frac{\delta c_r}{c_r} - \Delta L_r \cos \lambda_{r-1} (w_{r-1} + \eta_{r-1}) \\ v_r - v_{r-1} &= \Delta L_r \frac{\delta c_r}{c_r} + \Delta \lambda_r \sec \lambda_{r-1} (w_{r-1} + \eta_{r-1}) + \Delta L_r \tan \lambda_{r-1} u_{r-1} \\ w_r - w_{r-1} &= \Delta L_r \sin \lambda_{r-1} \frac{\delta c_r}{c_r} + \Delta \lambda_r \tan \lambda_{r-1} (w_{r-1} + \eta_{r-1}) + \Delta L_r \sec \lambda_{r-1} u_{r-1} \end{aligned} \right\} \dots (7)$$

In accordance with notation explained above, since  $\frac{c_r}{c_0} = \frac{c_r}{c_{r-1}} \cdot \frac{c_{r-1}}{c_{r-2}} \dots \frac{c_1}{c_0}$  and  $\frac{c_r}{c_{r-1}}$  is changed into  $\frac{c_r}{c_{r-1}} (1 + \epsilon_{r-1})$ : then  $\frac{c_r}{c_0}$  is changed into  $\frac{c_r}{c_0} \prod_0^{r-1} (1 + \epsilon_i)$ . Hence  $\frac{\delta c_r}{c_r} = E + \sum_0^{r-1} \epsilon$  where  $c_0$ , which may be regarded as the last side previous to the side 01, is supposed to change to  $c_0 (1 + E_0)$ .

Now suppose that the changes in successive side ratios and angles are in arithmetic progression: *i.e.*

$$\begin{aligned} \epsilon_0 &= \epsilon_1 - \epsilon' = \epsilon_2 - 2\epsilon' = \dots = \epsilon_r - r\epsilon' = \epsilon \\ \eta_0 &= \eta_1 - \eta' = \eta_2 - 2\eta' = \dots = \eta_r - r\eta' = \eta \end{aligned}$$

and

Then 
$$\frac{\delta c_r}{c_r} = E + r\epsilon + \frac{r(r-1)}{2} \epsilon'$$

E and H being the quantities relating to the side from which the triangulation line emanates.

Equations (7) may now be written

$$\left. \begin{aligned} u''_r - u''_{r-1} &= \left( E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} 1'' \Delta \lambda_r - (w''_{r-1} + \eta + (r-1)\eta') \operatorname{cosec} \lambda_{r-1} \Delta L_r \\ v''_r - v''_{r-1} &= \left( E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} 1'' \Delta L_r + (w''_{r-1} + \eta + (r-1)\eta') \sec \lambda_{r-1} \Delta \lambda_r \\ &\quad + w''_{r-1} \tan \lambda_{r-1} \Delta L_r \dots \\ w''_r - w''_{r-1} - \eta - \overline{r-1} \eta' &= \left( E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} 1'' \sin \lambda_{r-1} \Delta L_r \\ &\quad + (w''_{r-1} + \eta + (r-1)\eta') \tan \lambda_{r-1} \Delta \lambda_r + w''_{r-1} \sec \lambda_{r-1} \Delta L_r \dots \end{aligned} \right\} \dots (8)$$

in which  $u, v, w, \eta$  are now expressed in seconds and  $\Delta \lambda, \Delta L$  are in radians, and  $w_0 = H$ .

To apply these equations it is necessary to know the values of  $E_0, \epsilon, \eta', u_0'', w_0''$  and  $v_0''$ . The last quantity is simply additive to all values of  $v$ . The remaining five quantities give rise to five cases: for they can be considered separately and the results combined afterwards, since second order quantities are being neglected. By successive application of (8) the solution may be obtained in the form

$$\left. \begin{aligned} u_r &= (A_r E + B_r \epsilon + G_r \epsilon') \operatorname{cosec} 1'' + C_r \eta + F_r \eta' + D_r u_0 + K_r w_0 \\ v_r - v_0 &= (A_r E + B_r \epsilon + G_r \epsilon') \operatorname{cosec} 1'' + C_r \eta + F_r \eta' + D_r u_0 + K_r w_0 \\ w_r &= (A_r E + B_r \epsilon + G_r \epsilon') \operatorname{cosec} 1'' + C_r \eta + F_r \eta' + D_r u_0 + K_r w_0 \end{aligned} \right\} \dots (9)$$



The coefficients  $A, B$  etc . . . have to be determined for each triangulation line and then the latitude, longitude and azimuth changes are expressed by (9). The  $r^{\text{th}}$  side is changed in the ratio  $1 + E + r\epsilon + \frac{r(r-1)}{2}\epsilon'$  : 1 ; so that the solution is complete.

It is convenient to denote the changes at the end of any, the  $m^{\text{th}}$ , triangulation line in latitude, longitude, azimuth and side by  ${}_mU \quad {}_mV \quad {}_mW \quad {}_mE$  : then taking account of different values of  $\epsilon, \eta$  which occur in different triangulation lines the change in latitude, longitude and azimuth of the  $r^{\text{th}}$  station of the  $m^{\text{th}}$  line may be written

$$\left. \begin{aligned} {}_mU_r &= ({}_m A_u \quad {}_{m-1} E + {}_m B_u \epsilon_m + {}_m G_u \epsilon_m') \operatorname{cosec} 1'' + {}_m C_u \eta_m + {}_m D_u u_0 + {}_m K_u w_0 \\ {}_mV_r - {}_{m-1} V &= ({}_m A_v \quad {}_{m-1} E + {}_m B_v \epsilon_m + {}_m G_v \epsilon_m') \operatorname{cosec} 1'' + {}_m C_v \eta_m + {}_m D_v u_0 + {}_m K_v w_0 \\ {}_mW_r &= ({}_m A_w \quad {}_{m-1} E + {}_m B_w \epsilon_m + {}_m G_w \epsilon_m') \operatorname{cosec} 1'' + {}_m C_w \eta_m + {}_m D_w u_0 + {}_m K_w w_0 \\ {}_mE_r &= {}_{m-1} E + r\epsilon_m + \frac{r(r-1)}{2}\epsilon'_m \end{aligned} \right\} \quad (10)$$

Equations (10) give in the most general form the changes that are effected by introducing the alterations  $\eta, \eta', \epsilon, \epsilon'$  different for each triangulation line.

6. We may now consider the probable relative values of  $\epsilon$  and  $\eta$  (the latter expressed in radians) for this purpose treating  $\epsilon'$  and  $\eta'$  as zero.. Take the case of a series of simple equilateral triangles and let  $p-2, p-1, p$  be three successive stations on a flank: also let  $x_p, y_p, z_p$  etc. be changes which are applied to the several angles, as indicated in the figure. The ratio of the successive flank sides  $Q R/P Q$  being denoted by  $r$ , then

$$\begin{aligned} d \log r &= \cot 60^\circ (x_{p-1} + x'_{p-1} + y_p - y'_{p-1} - z'_{p-1} - z_p) \\ \text{Hence} \quad \Sigma d \log r &= d \log \Pi r \\ &= \cot 60^\circ \left\{ \sum_0^{p-1} (x_r + x'_r - z_{r-1} - z'_{r-1}) + y_p - y'_0 \right\} \\ &\doteq p\epsilon. \end{aligned}$$

Hence the most probable way of getting a particular value for the change in logarithm of the side is by making all the  $x^s$  equal to each other and of opposite sign to all the  $z^s$  which will also all be equal. The  $y^s$  do not come into the case at all except at the two ends, unless an azimuth change is also required.

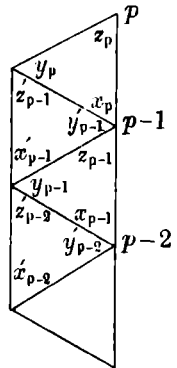
For the azimuth it is clear that

$$\Sigma \eta_{p-1} = \Sigma (x_p + y'_{p-1} + z_{p-1}) = p\eta$$

is the azimuth change, since the  $\eta^s$  are all to be equal.

Now in the most probable distribution of changes obviously all the  $x^s$  are equal as are the  $y^s$  and  $z^s$ . Hence

$$\left. \begin{aligned} p\epsilon &= \frac{p}{\sqrt{3}} \left\{ x + x' - z - z' \right\} + \frac{1}{\sqrt{3}} (y - y') \\ \text{and} \quad p\eta &= p (x + y' + z). \\ \text{Now} \quad x + z &= -y \\ \therefore \quad \eta &= y' - y \\ \text{and} \quad \epsilon &\doteq \frac{1}{\sqrt{3}} (x + x' - z - z') \end{aligned} \right\} \dots \dots \dots (11)$$



The probable values of  $\eta$  and  $\epsilon$  are accordingly in the ratio of the most probable values of  $\frac{y\sqrt{3}}{x-z}$  in which the quantities are subject to the relation

$$x + y + z = 0$$

Following the usual plan of independent multipliers explained below (Chapter VII § 4) we have for the two cases (a) of  $y$  and (b) of  $x-z$

$$\begin{aligned} a_1 &= a_2 = a_3 = 1 \\ b_1 &= b_2 = b_3 = 0 \\ &\text{etc.} \\ u_1 &= u_2 = u_3 = 1 \end{aligned}$$

and

$$3k_1 = 1 \text{ in case (a); } 3k_1 = 0 \text{ in case (b)}$$

$\therefore$

$$u_F = 1 - \frac{1}{3} = \frac{2}{3} \text{ in case (a); } u_F = 2 - 0 = 2 \text{ in case (b)}$$

Hence the ratio of probable values of  $\eta$  and  $\epsilon$  is unity, *i.e.*,  $\eta$  and  $\epsilon$  have the same weight.

7. Now it is clear that in the adjustment of azimuth and side the most probable solution is obtained by introducing equal values of  $\eta$ ,  $\epsilon$  and by putting  $\epsilon' = \eta = 0$ , all round a circuit composed of triangulation lines of equal strength: and if the triangulation lines are of different strengths the value of  $\eta$ ,  $\epsilon$  in any one is inversely proportional to the strength. With latitude and longitude it is clear that the directions of the traverse lines are of essential importance. Thus in the case of a triangulation line along a meridian a change in  $\epsilon$  will give changes in latitude but no change in longitude. Moreover to get any desired change the most probable values of changes in angle and side ratio would not be equal at all the stations along a triangulation line: though it is probable that they would not alter much on one triangulation line if the circuit was composed of several such lines. Taking the case of a triangulation line along a meridian it is clear that to obtain a given change in longitude the change in the angles at the starting end of the line are more effective than equal changes in the angles near the closing point. If the sides were of equal length the most probable changes would be in arithmetical progression. The complexity of different changes at the various stations of a triangulation line on account of this has been deemed in general to overbalance the slight gain in theoretical accuracy of adjustment and at least in some preliminary adjustments which are about to be made (*e.g.* the incomplete Burma triangulation) the plan will be followed of making  $\eta$  and  $\epsilon$  uniform along one triangulation line and putting  $\eta' = \epsilon' = 0$ . When the probable errors of position etc. are considered (*vide* Chapter VII.) it is believed that this plan will be considered fully satisfactory for many cases of geodetic application.

The equations (10) however show how the variation in changes along a triangulation line may be taken into account: and in closing a single triangulation line between two previously fixed sides they give the necessary number of quantities (four) at choice to satisfy the closing condition. It is clear from the theorem stated at the end of § 13 that for the general case of closure the changes at the several stations of a triangulation line should be in arithmetical progression this being a combination of the most probable adjustment firstly for log. side and azimuth and secondly for latitude and longitude. When the number of  $\eta^s$  and  $\epsilon^s$  at choice is in considerable excess of the number of conditions to be satisfied it is believed that little gain in accuracy is obtained by introducing  $\eta^s$  and  $\epsilon^s$ : and certainly this doubles the number of unknowns and greatly increases the labour of formation and solution of the normal equations. It also increases the complexity of the solution, and its subsequent application.

8. In the case of a network of circuits including Laplace stations and extra base lines, equations of form of (10) may be formed; and by equating the right hand sides to the several closing

errors which arise, four equations are formed for each circuit together with one extra for each extra base line or Laplace station. These may be used to determine the most probable values of  $\eta$ ,  $\epsilon$  (and  $\eta'$ ,  $\epsilon'$  if it is thought desirable not to put these equal to zero for each triangulation line), due regard being paid to the strength of each triangulation line.

If  $m = \sqrt{\frac{\sum \Delta^2}{3n}}$  where  $\Delta$  is the triangular error of any one of the triangles of a series, then

the probable value of  $\eta$  in the case of a series of simple triangles in which the triangular error has been dispersed is, by (11),  $\sqrt{\frac{2}{3}} m \sqrt{2} \times .6745 = .779m$ . Account may be taken of the series comprising quadrilaterals and other figures by the introduction of a factor  $\frac{1+f}{1+\frac{1}{6}}$  (see § 1 above) so that the probable values of  $\eta$  and  $\epsilon$  (which are equal) are both equal to

$$.779m \frac{1+f}{1+\frac{1}{6}} = 0.668m(1+f) = a \dots \dots \dots (12)$$

in which  $m$  is supposed to be expressed in radians.

The equations for  $\eta$  and  $\epsilon$  accordingly have to be solved subject to the condition that the sum of the squares of the corrections multiplied by their weights is a minimum. In one triangulation line the sum of the squares of the corrections is

$$\begin{aligned} \Sigma(\eta^2 + \epsilon^2) + \Sigma(\eta'^2 + \epsilon'^2)r^2 &= r(\eta^2 + \epsilon^2) + \frac{r(r+1)(2r+1)}{6}(\eta'^2 + \epsilon'^2) \\ &= r(\eta^2 + \epsilon^2) + \frac{r^3}{3}(\eta'^2 + \epsilon'^2) \end{aligned}$$

when  $r$  is the number of sides in a triangulation line.

Hence the condition to be satisfied is that

$$\Sigma \left[ \left\{ r(\eta^2 + \epsilon^2) + \frac{r^3}{3}(\eta'^2 + \epsilon'^2) \right\} \div m^2(1+f)^2 \right] = \text{a minimum} \dots \dots \dots (13)$$

the summation extending over all the triangulation lines.

### Mechanical Analogy

9. Suppose that there is a network of trilateration, in which the lengths of all sides have been determined by direct measurement. Let  $l$  be the measured length of any side,  $l + \delta l$  the adjusted value, and  $w$  the weight of the determination  $l$ . In making any adjustment of the network, for example to bring a terminal side into agreement with a line slightly differently placed, the principle of least squares demands that  $\Sigma w \delta l^2$  shall be a minimum, all imposed conditions being satisfied.

Now consider a similar framework formed by rods of material obeying Hooke's law of extension and compression, and suppose these rods are freely jointed at their junctions. If this framework is in equilibrium without any strains in action, and, the first side being held fast, the last side is brought into a slightly different position and its length slightly changed, then the several rods will undergo compression or extension and their lengths will be slightly altered. If the unstrained length of any rod is  $l$  and its strained length is  $l + \delta l$ ; and the force in action in it causing this strain is  $F$ : then the work done on the rod is  $\frac{1}{2} F \delta l$ . Now by Hooke's law

$$\frac{\delta l}{l} = \frac{F}{aE}$$

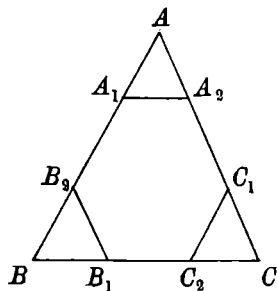
where  $E$  is Young's modulus and  $a$  is the cross section.

Hence

$$\frac{1}{2} F \delta l = \frac{aE}{2l} \cdot \delta l^2$$

and this represents the work done on the rod. The quantity  $\frac{aE}{2l}$  may be varied by suitably choosing  $a$ : suppose it is made equal to  $w$ . Then the work done on the rod is  $w \delta l^2$ . The principle of least work immediately shows that for the set of strains applied to bring the last side into the desired position the total work done must be a minimum: whence  $\Sigma w \delta l^2$  is a minimum. The solution accordingly is the same as that of the most probable adjustment of the similar trilateration network.

10. Now trilateration has never been carried out on a large scale for one obvious reason that no large tract of country is suitable for its execution by ordinary methods. Triangulation on the other hand extends over vast tracts and it is its adjustment which is of importance to geodesists. A mechanical analogy can be supposed for triangulation also. It is less simple than that just described for trilateration and to give it practical shape would be a matter of greater difficulty. Imagine a set of rods, the medial parts of which are laterally rigid, but which are longitudinally extensible without the application of (appreciable) force\*. Let these rods be freely jointed to form a framework similar to a network of triangulation, the angles of the triangulation being maintained by rigid pieces. For example in the  $\Delta ABC$  the rod  $AB$  is freely extensible between  $A_1$  and  $B_2$ : a cross piece  $A_1 A_2$  maintains the angle  $A$  at its proper value and so on. It will be seen then that such a triangle may be enlarged to any size but that it will always remain similar, unless the pieces  $A_1 A_2, B_1 B_2, C_1 C_2$  are changed in length. If these pieces  $A_1 A_2$  etc. are made of material obeying Hooke's law, by properly choosing their cross section it is possible to make them represent the "strength" of the angles  $A, B, C$ . When the system is deformed the work done on  $A_1 A_2$  will be proportional to the square of the change in the angle  $A$  and accordingly a condition of form  $\Sigma w \delta \theta^2 = \text{minimum}$  must be satisfied so that the total work done shall be a minimum. All geometrical conditions such as triangular conditions, central station conditions and side ratio conditions obviously cannot be avoided in the mechanical analogy, so that the solution of the mechanical problem is precisely the same as that of the triangulation adjustment according to the method of least squares.



11. It is clear that if the change in any observed quantity can be made to correspond to the extension of a rod which obeys Hooke's law; and if a system of such rods are linked up in such a way as to represent the geometrical conditions controlling the observations; then a mechanical analogy for the set of observations is obtainable. The governing fact is that the work done on any rod is proportional to the square of its extension: so that by substituting extension for error of observation the equation of minimum squares is transformed into the principle of least work.

12. Consider a framework representing a network of triangulation and suppose that it is held fast at one or more points. It may be necessary to bring a terminal side into agreement with a predetermined value and position. To do this four conditions have to be satisfied:

- (1) The terminal side must be adjusted to the correct length.
- (2) The terminal side must be adjusted to the correct azimuth.
- (3) One extremity of the terminal side must be moved to the correct latitude.
- (4) This extremity of the terminal side must be moved to the correct longitude.

Now these adjustments may be considered one by one. First strain the terminal side to the correct length and hold it so: then change its azimuth, etc. The adjustments may also be performed

\* Approximations to these can readily be conceived, e.g. a rod sliding in a tube.

in any order and the final result is the same. This is immediately obvious mechanically, for it is clear that the final configuration due to small strains has nothing to do with the order in which they are applied.

13. The analogy thus proves an important theorem† in the adjustment of observations, namely that *provided all imposed conditions are maintained, the adjustment conditions may be introduced separately in any order, the previous adjustment conditions in each case being maintained; and the most probable complete adjustment is obtained after the last adjustment condition has been applied.* This enables the circuit adjustments of triangulation to be applied after the figural adjustments, as has been done in the Survey of India. A further theorem is also easily deducible. In the case of the closing of a simple triangulation circuit in which there are four closing conditions to satisfy (*i.e.* the case in which there are no additional base lines or independent azimuth determinations) denote the four closing quantities by  $X, Y, Z, U$ . To effect the closing  $X$  alone in the most-probable manner, changes are concurrently introduced which affect the other quantities by amounts  $y_x, z_x, u_x$ ; and similarly for the other quantities. From the mechanical analogy it is clear that adjustments as follows should be made:—

- (1)  $x \quad y_x \quad z_x \quad u_x$
- (2)  $x_y \quad y \quad z_y \quad u_y$
- (3)  $x_z \quad y_z \quad z \quad u_z$
- (4)  $x_u \quad y_u \quad z_u \quad u$

in which

$$\left. \begin{aligned} x + x_x + x_z + x_u &= X \\ y_x + y + y_z + y_u &= Y \\ z_x + x_y + z + z_u &= Z \\ u_x + u_y + u_z + u &= U \end{aligned} \right\} \dots \dots \dots (14)$$

and the quantities with suffixes are geometrically related to the suffix quantity, *i.e.*

$$\left. \begin{aligned} \frac{y_x}{B_x} = \frac{z_x}{C_x} = \frac{u_x}{D_x} = x \\ \frac{x_y}{A_y} = \frac{z_y}{C_y} = \frac{u_y}{D_y} = y \\ \text{etc.} \end{aligned} \right\} \dots \dots \dots (15)$$

where the coefficients  $A B C D$  depend only on the form of the triangulation and are independent of the closing errors. It accordingly follows that

$$\left. \begin{aligned} x + A_y y + A_z z + A_u u &= X \\ B_x x + y + B_z z + B_u u &= Y \\ C_x x + C_y y + z + C_u u &= Z \\ D_x x + D_y y + D_z z + u &= U \end{aligned} \right\} \dots \dots \dots (16)$$

and by solving these the following equations may be obtained

$$\left. \begin{aligned} x &= a_x X + a_y Y + a_z Z + a_u U \\ y &= b_x X + b_y Y + b_z Z + b_u U \\ z &= c_x X + c_y Y + c_z Z + c_u U \\ u &= d_x X + d_y Y + d_z Z + d_u U \end{aligned} \right\} \dots \dots \dots (17)$$

If the adjustments  $x, y, z, u$  are applied independently and then combined the total effect will be the same as that of the single adjustment  $X, Y, Z, U$  taken simultaneously. By use of the quantities  $x, y, z, u$  in place of the related quantities  $X, Y, Z, U$  it is accordingly possible to treat each closure entirely independently of the remaining three.

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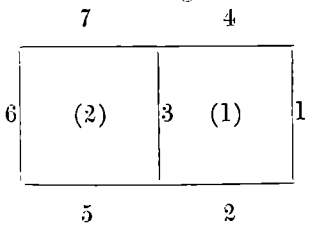
† This theorem was proved analytically in "Account of the Operations of the G.T. Survey of India, Vol. II," Appendix No. 8, pp. 151-153.

If in addition to the 4 ordinary closing quantities additional conditions such as extra bases, or fixings of latitude, longitude or azimuth are introduced, this only makes the relation more involved: it is still possible to express  $x, y, z, u$  in terms of the closing errors and proceed as though each of these quantities was independent of the other.

The principle is perfectly general and is applicable to a whole network of triangulation. As this becomes more complex the determination of the coefficients in the equations which correspond to (17) would become more difficult: but this is not necessary at least in some applications of the theorem. It is an important fact that an undetermined portion of each of the closures may be regarded quite independently of the others. The theorem may be stated as follows:—*It is possible to find quantities related to the several closing errors such that each type may be adjusted separately and independently of the others, and such that the combined effects of these several adjustments will be the same as the most probable simultaneous adjustment.*

These related quantities may be called the “independent errors” of each type. If each is adjusted independently of errors of another type, the other type adjustments being allowed to come in just as they will naturally do while the adjustment of the first type is made independently in the most probable manner, then the combined result of the adjustment of the four types will be the most probable adjustment of the closing errors which can be made.

14. Consider now in more detail the case of more than one circuit forming a network which has to be simultaneously adjusted. The case of two circuits which have a common portion is illustrative of this and will be seen to lead to a result generally applicable. Further these circuits may be considered as formed each for four series and no loss of generality occurs in taking these circuits to be of the same strength. The several series may be characterised by the numbers 1—7 and the circuits by (1) and (2).



The adjustments along any side may be made by the introduction of  $\epsilon$  and  $\eta$  changes: and any one of the four types of closures the  $\epsilon^s$  and  $\eta^s$  along a triangulation line will be in arithmetical progression (*vide* § 16). For the  $r^{th}$  side of the  $k^{th}$  line their values may be represented by

$$\epsilon_k + \overline{r-1} | \epsilon'_k \quad , \quad \eta_k + \overline{r-1} | \eta'_k$$

Consider first the X closure and suppose that  ${}_1x \quad {}_2x$  are the “independent errors” of the two circuits: these quantities have to be determined. Then equations of the following form may be formed:

$$\left. \begin{aligned} \sum_1^4 (A_k \epsilon_k + B_k \epsilon'_k + C_k \eta_k + D_k \eta'_k) &= {}_1x \\ \sum_{3,5,6,7} (A_k \epsilon_k + B_k \epsilon_k + C_k \eta_k + D_k \eta'_k) &= {}_2x \end{aligned} \right\} \dots \dots \dots (18)$$

in which the coefficients  $A \ B \ C \ D$  are independent of the closing errors. The prefixes to  $x$  are the circuit numbers.

To obtain the most probable solution the quantities  $\epsilon, \eta$  must be chosen so as to make

$$\sum_{k=1}^7 \left\{ \sum_{r=1}^{r=n} (\epsilon_k + \overline{r-1} | \epsilon'_k)^2 + \sum_{r=1}^{r=n} (\eta_k + \overline{r-1} | \eta'_k)^2 \right\} = \text{minimum.}$$

The solution of this accordingly gives definite value for all the  $\epsilon^s$  and  $\eta^s$  as linear functions of  ${}_1x$  and  ${}_2x$ . The associated quantities are given by equations of form

$${}_1y_x = {}_1\beta_x {}_1x + {}_1\beta'_x {}_2x$$

$${}_1z_x = {}_1\gamma_x {}_1x + {}_1\gamma'_x {}_2x$$

$${}_1u_x = {}_1\delta_x {}_1x + {}_1\delta'_x {}_2x$$

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$${}_2y_x = {}_2\beta_x {}_1x + {}_2\beta'_x {}_2x \text{ etc.}$$

Similarly for the other closure

$${}_1x_y = {}_1a_y {}_1y + {}_1a'_y {}_2y \text{ etc.}$$

Equations similar to (14) can now be formed for each circuit

$${}_1x + ({}_1a_y {}_1y + {}_1a'_y {}_2y) + ({}_1a_z {}_1z + {}_1a'_z {}_2z) + ({}_1a_u {}_1u + {}_1a'_u {}_2u) = {}_1X \text{ etc. . . . . (19)}$$

From (19) it is clear that the determination of the "independent errors"  ${}_1x$ ,  ${}_2x$  etc. from the known closing conditions can be effected by the solution of simultaneous linear equations of the same number as there are conditions. These equations are soluble without the introduction of the principles of least squares. Their solution however will entail much the same labour as the solution of the normal equations which would arise in the simultaneous adjustment of all the four types of closure. It appears at present to be only of theoretical interest that these closures or rather the adjustment of the related independent errors may be effected, each type independently of the others: for no material reduction in computation would be effected. However it is believed that the mechanical analogy throws some light on the question, and that developments are likely to result from its consideration.

15. The idea of a triangulation line has been introduced in § 4. It is the flank of a nearly straight series of triangulation. For the present, disregard the  $\epsilon$  changes and fix attention on the  $\eta$  changes. Suppose a set of rigid rods are placed similarly to the several rays of the flank and that these are freely jointed at their ends, which accordingly correspond to the stations on the triangulation line. Introduce constraints at each junction which tend to maintain the angles between successive rods equal to the observed values and such that the force necessary to alter any one of these angles by a stated amount is inversely proportional to the probable error of the angle or directly proportional to the *strength* of the angle. Then it is clear that the work done in varying the angles in any way is proportional to the sum of the weighted squares of the changes of these angles. It is clear then that to bring the system into any given displaced formation the angular changes introduced are such as to make either the total work done, or the sum of the weighted squares of the adjustment in the angles, a minimum. That is to say the mechanical deformation is the same as the most probable adjustment. It is clear from this why when the angular strength is given the strength of the triangulation to resist either angular deflection at the end, or linear deflection, or a combination of the two, is the greater according as the number of stations in the line is less, or, in other words, according as the length of side is greater. The quantity M already introduced takes account of this (see Chapter VII § 1) and for certain considerations of probable error makes it unnecessary to consider the triangulation line in any detail.

The  $\epsilon$  changes in a triangulation line (due to the extension of its sides) are precisely analogous to the  $\eta$  changes, if for angular deflection, change in ratio of final to initial side is taken, and for linear deflection at right angles to the line, linear deflection in the direction of the line is substituted. In the case of a single triangulation line (which is a straight line) the  $\eta$  changes are

independent of the  $\epsilon$  changes and the corresponding adjustments can be performed independently: so both could be separately considered by means of this partial mechanical analogy. When however there is a set of triangulation lines in various directions the  $\eta$  and  $\epsilon$  changes (northing and easting) of the several lines get intermixed and for this a complete analogy is required.

It is clear that for the case of trilateration it would only be necessary to arrange a set of *extensible* rods, jointed as described above, and representing a flank of the trilateration to obtain a complete analogy. But in the case of triangulation it is necessary to arrange that if any element changes length the successive elements change length in the same proportion without any force being involved. A simple mechanical analogy completely representing a triangulation line has not yet been discovered: and for this case it at present appears necessary to consider the analogy of the complete series instead of only one of its flanks (vide § 10). For purposes of probable errors it would be possible to replace the series by a simple series composed of equilateral triangles, in the same way that in the triangulation line it may be a convenient simplification to take all the sides of equal length and persisting in the same direction: and it appears from the mechanical analogy that little accuracy would be lost by so doing.

16. It is clear from the partial analogy given in the preceding section that the best adjustment to make in a triangulation line to obtain a given azimuth change is to change all the angles by amounts proportional to their probable errors: for this corresponds to the mechanical set of rods of which the first is held fast and to the last of which a suitable couple is applied. To obtain a given deflection the most probable adjustment is that the angular changes divided by their probable errors should be in arithmetical progression; as would be the case in the mechanical system, when held at one end and subjected to a simple force at right angles to the line at the other end. The exact analogy between the  $\epsilon^s$  and  $\eta^s$  shows that similar conditions hold for the  $\epsilon^s$ . The above statements can be simplified for the case of a triangulation line if the weights of the several angles are considered equal. It appears that a determination of the weights of the angles may best be obtained by a consideration of all the triangular errors, which leads to one value for the probable value of any angle of the series. Any determination of probable error of each angle separately is very much vitiated by the graduation error of the instrument, which is unknown. The simplified statements for a triangulation line with all sides of the same length are:—

(a) for a given angular deflection or a given change of ratio of final side to initial side, that the  $\eta^s$  or  $\epsilon^s$  are all equal.

(b) for a given linear deflection at right angles to or along the line, that the  $\eta^s$  or  $\epsilon^s$  are in arithmetical progression.

On this account the cases of the  $\eta^s$  and  $\epsilon^s$  changing in arithmetical progression (which includes (a) as a special case, the constant difference then being zero) have been considered in equations (8) to (10). Some departure from rigid accuracy occurs when the several sides of the triangulation line are of unequal lengths.



## CHAPTER VII.

### Probable errors of triangulation before and after adjustment.

1. Expressions will now be formed for the probable errors of points and sides generated in one or more series of triangulation, in which only figural conditions have been adjusted.

(1) *Probable errors in logarithm and azimuth of terminal side.*

Equation (12) of chapter VI gives  $a$  the probable value of either  $\eta$  or  $\epsilon$ . The probable error in the logarithm of the terminal side after  $n$  sides of a triangulation line is clearly  $\sqrt{n} \log_{10}(1 + \epsilon) = .4343a \sqrt{n}$ ; and in azimuth is  $a \sqrt{n}$ . Considering the triangulation line as practically straight and the distances between stations as equal then, if  $l$  is the average length of side and  $s$  the length of the line, so that  $nl = s$ ,

$$\begin{aligned} a\sqrt{n} &= .668m(1+f)\sqrt{n} = m(1+f) \sqrt{\frac{18}{l}} \cdot \sqrt{\frac{nl}{18}} \times .668 \\ &= .1575 M \sqrt{s} \dots \dots \dots (1) \end{aligned}$$

Hence if  $M$  is expressed (as is always done) in seconds of arc

Probable error in azimuth at end of a triangulation line =  $0''.1575 M \sqrt{s}$

Probable error in log. side at end of a triangulation line =  $.4343 \sin 1' \times 0.1575 M \sqrt{s}$   
 =  $3.32 \times 10^7 M \sqrt{s}$

For the case of a number of triangulation lines it is necessary to substitute  $\sqrt{\Sigma M^2 s}$  for  $M \sqrt{s}$ . It is convenient to measure lengths of triangulation lines in units of 100 miles: so replace  $s$  by  $100 S$  where the unit of  $S$  is 100 miles and then

Probable error in azimuth of the terminal side of a series of triangulation lines	= $1''.575 \sqrt{\Sigma M^2 S}$	} \dots \dots \dots (2)
Probable error in seventh place of logarithm of the terminal side of a series of triangulation lines	= $33.2 \sqrt{\Sigma M^2 S}$	

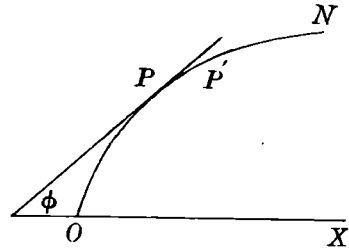
In the above  $S$  is measured along the triangulation, and it is immaterial whether this is straight or not: but if the elements of the summation indicated by  $\Sigma$  are straight, then  $S$  may be replaced by  $L$  the length of any triangulation line, bringing formulæ (2) into similar terms to those of (5) below.

(2) *Probable errors in easting and northing of terminal points.*

Consider any curve defined by  $s$  the distance measured along it and  $\phi$  the angle the tangent makes with  $OX$ . Let this curve be divided into elements of length  $l$ . Suppose that for purposes of adjustment, or on account of errors, the ratio of the  $(m+1)^{th}$  element to the  $m^{th}$  is changed by factor  $1 + \epsilon_m$  and the angle at the junction of these two elements by  $\eta_m$ . Then if  $P P'$  is the  $(m+1)^{th}$  element the relative shift of  $P'$  to  $P$  is

$$l \cos \phi_m \sum_1^{m+1} \epsilon - l \sin \phi_m \sum_1^{m+1} \eta \quad \text{in easting}$$

and  $l \sin \phi_m \sum_1^{m+1} \epsilon + l \cos \phi_m \sum_1^{m+1} \eta \quad \text{in northing}$



and the total change relative to  $O$  of  $N$  is given by

$$\Delta x = l \sum_1^n \left( \cos \phi_m \sum_1^{m+1} \epsilon \right) - l \sum_1^n \left( \sin \phi_m \sum_1^{m+1} \eta \right)$$

$$\Delta y = l \sum_1^n \left( \sin \phi_m \sum_1^{m+1} \epsilon \right) + l \sum_1^n \left( \cos \phi_m \sum_1^{m+1} \eta \right)$$

Hence

$$\begin{aligned} \Delta x &= l \epsilon_1 \sum_1^n \cos \phi + l \epsilon_2 \sum_2^n \cos \phi + l \epsilon_3 \sum_3^n \cos \phi + \dots - l \eta_1 \sum_1^n \sin \phi - \dots \\ &= \epsilon_1 (x_n - x_0) + \epsilon_2 (x_n - x_1) + \epsilon_3 (x_n - x_2) \dots - \eta_1 (y_n - y_0) - \dots \end{aligned}$$

since  $\cos \phi = \frac{x' - x}{l}$ , and  $\sin \phi = \frac{y' - y}{l}$

The most probable value of  $\Delta x$  is accordingly

$$\left[ \epsilon^2 \left\{ (x_n - x_0)^2 + (x_n - x_1)^2 + \dots \right\} + \eta^2 \left\{ (y_n - y_0)^2 + (y_n - y_1)^2 + \dots \right\} \right]^{\frac{1}{2}}$$

and since the probable values of  $\epsilon$  and  $\eta$  are both equal to  $a$  this reduces to  $a \sqrt{\Sigma r^2}$  where  $r$  is the radius vector measured from the point  $N$ . The probable value of  $\Delta y$  is the same. When the elements are increased in number  $\Sigma r^2$  may be replaced by  $\frac{1}{l} \int r^2 ds$ : also by (12) of chapter VI and (1)

$$a = .1575 M \sin 1'' \sqrt{\frac{s}{n}} = .1575 M \sin 1'' \sqrt{l}$$

and  $a \sqrt{\Sigma r^2}$  becomes equal to  $.1575 M \sin 1'' \sqrt{\int r^2 ds}$

Hence the probable value of  $\Delta x$  or  $\Delta y = .1575 M \sin 1'' \sqrt{\int r^2 ds} \dots \dots \dots (3)$

So far only a single triangulation line has been considered. If there are several it is clear that it is necessary to alter this expression to  $.1575 \sin 1'' \sqrt{\Sigma M^2 \int r^2 ds}$ . This is expressed in units of a mile. To express it in feet multiply by 5280. It is convenient to measure  $r$  and  $s$  in units of 100 miles: denote their values in units of 100 miles by  $R$  and  $S$ . Finally if  $P$  (feet) is the probable error at a point  $N$  in casting or northing, and  $R$  is the radius vector measured from  $N$ , and  $S$  is the distance measured along the triangulation, both  $R$  and  $S$  being expressed in units of 100 miles, then

$$P = .1575 \times 5280 \times 10^5 \sin 1'' \sqrt{\sum M^2 \int R^2 dS}$$

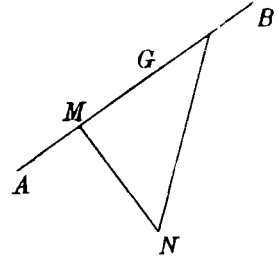
$$= 4.03 \sqrt{\sum M^2 \int R^2 dS} \dots \dots \dots (4)$$

The integral  $\int R^2 dS$  may be taken out for each triangulation line (assumed to be a straight line). If  $AB$  be one of these and  $S$  is measured from  $M$  the foot of the perpendicular from  $N$  on  $AB$  and  $NM = p$ , then  $R^2 = S^2 + p^2$

$$\therefore \int R^2 dS = \left[ \frac{1}{3} S^3 + S p^2 \right]_{MA}^{MB} = AB \left( \frac{1}{3} AB^2 + p^2 + MA \cdot MB \right)$$

Now if  $G$  is the middle point of  $AB$

$$MA \cdot MB = - \left( \frac{L}{2} + MG \right) \left( \frac{L}{2} - MG \right) = - \frac{L^2}{4} + MG^2$$



$$\text{Hence } \int R^2 dS = L \left( \frac{1}{3} L^2 + p^2 - \frac{1}{4} L^2 + MG^2 \right) = L \left( R_0^2 + \frac{L^2}{12} \right)$$

where  $R_0$  is the radius vector to the middle point of  $AB$  and  $L$  is the length  $AB$ . Therefore for a series of triangulation lines

$$P = 4.03 \sqrt{\sum M^2 L \left( R_0^2 + \frac{L^2}{12} \right)} \dots \dots \dots (5)$$

in which the quantities  $L, R_0$  may be measured off a chart in units of 100 miles. From either (4) or (5) it is clear that the probable closing error in northing or easting is different according as different points of a circuit are selected on which to close and from which to start.

2. Formulæ (2), (5) show that the probable errors in azimuth and logarithm of terminal side increase as the square root of the length of the several triangulation lines involved, while those of easting and northing increase at a much more rapid rate namely as the three halves power of the lengths, the triangulation lines remaining similar and similarly situated. On account of this latter fact it is desirable to have more frequent checks to prevent accumulation of errors than would be necessary if only length and azimuth of side were required. These checks may be obtained by measurement of bases and by forming Laplace stations, that is stations whose longitudes are observed telegraphically and at which astronomical azimuths are also observed. These two checks are of precisely equal importance; and applying only one of them does not serve a very useful purpose. In the Indian triangulation eight base lines have been measured to date (1916), excluding the short Mergui base in Burma, and these have been made use of in the adjustment of the triangulation. The longitude arcs were not available when (previous to 1879) the main adjustment was carried out. They do not all admit of the formation of Laplace equations, as the longitude stations are not coincident with the triangulation stations, nor can they be connected with satisfactory accuracy (as regards azimuth) in all cases. Only latterly\* (1906) have they been applied to control the azimuth observations with a view to determining corrected plumb-line deflections in the prime vertical. This application does not improve the probable error in easting and northing of the points concerned or any other points of the triangulation.

3. It is necessary for the full consideration of the problem to find expressions for the probable errors after certain further adjustments, *viz.* closing on extra base lines, closing of circuits or (what has not been done in India) closing on Laplace stations, have been effected. Before treating

\* Account of the Operations of the G. T. Survey of India, Vol. XVIII, Appendix 5.

this question quite generally it may be of interest to consider a special case. Suppose a series of triangulation lines closes between two bases. The probable error in easting or northing at any point of it after the adjustment has been performed is required. It is to be noticed in (4) that each portion of the triangulation is fully taken account of by the portion of the integral  $\int M^2 R^2 dS$  which applies to it: that is to say, this portion of the integral represents the errors of displacement generated in the corresponding part of the triangulation, as if carried on to the closing point by perfect triangulation.

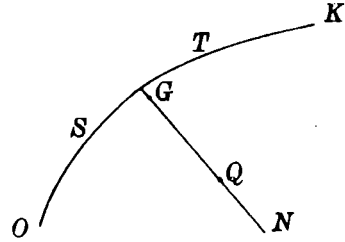
Let  $O, K$  be the two bases and suppose  $ST$  is the portion whose error as generated at  $N$  is required.

As in § 1 the expression for  $\Delta x$  for  $ST$  may be written

$$\Delta x = \sum_s^t (\epsilon_r x_r - \eta_r y_r)$$

The adjustment under consideration does not affect the  $\eta$  terms. The condition of closing gives (assuming uniform strength along  $OK$ )

$$\sum_1^k \epsilon_r = \text{a known quantity}$$



and in adjusting for this  $\epsilon_r$  is replaced by  $\epsilon_r - \frac{1}{k} \sum_1^k \epsilon_r$ .

Denote the adjusted value of  $\Delta x$  by  ${}_a\Delta x$ . Then

$$\begin{aligned} {}_a\Delta x &= \sum_s^t \left\{ x_r \left( \epsilon_r - \frac{1}{k} \sum_1^k \epsilon_r \right) - \eta_r y_r \right\} \\ &= \sum_s^t \epsilon_r x_r - \frac{t-s}{k} \bar{X} \sum_1^k \epsilon_r - \sum_s^t \eta_r y_r \end{aligned}$$

where  $\bar{X}$  is the  $x$ -coordinate of the centre of gravity of  $ST$ . Hence

$${}_a\Delta x = - \frac{t-s}{k} \bar{X} \left( \sum_1^{s-1} \epsilon + \sum_{t+1}^k \epsilon \right) + \sum_s^t \epsilon_r \left( x_r - \frac{t-s}{k} \bar{X} \right) - \sum_s^t \eta_r y_r$$

Hence the probable error of displacement at  $N$ , after  $OK$  has been adjusted, due to the portion  $ST$  is the square root of

$$({}_a\Delta x)^2 + ({}_a\Delta y)^2 = (k-t+s) \left( \frac{t-s}{k} \right)^2 \epsilon^2 NG^2 + \epsilon^2 \sum_s^t R_Q^2 + \eta^2 \sum_s^t R_N^2$$

where  $G$  is the centre of gravity of  $ST$  and  $Q$  lies on  $NG$  and  $QN/GN = (t-s)/k$ : also  $R_Q$  and  $R_N$  are radius vectors measured from  $Q$  and  $N$  respectively. Denoting the resultant probable displacement by  $D$ , this relation may be written, putting  $OK = S_0$ ,  $ST = S$  (measured along the curve)

$$\begin{aligned} D^2 &= k-t+s \left( \frac{S}{S_0} \right)^2 \cdot NG^2 \cdot \epsilon^2 + \epsilon^2 \sum_s^t R_Q^2 + \eta^2 \sum_s^t R_N^2 \\ &= \epsilon^2 \left\{ n N Q^2 + \sum_s^t R_Q^2 \right\} + \eta^2 \sum_s^t R_N^2 \dots \dots \dots (6) \end{aligned}$$

where  $n = k - l + s$  is the number of sides in the line  $ST$ . Hence

$$D^2 = a^2 n \left\{ NQ^2 + \frac{1}{S} \int R_Q^2 dS + \frac{1}{S} \int R_N^2 dS \right\}$$

If the portion  $ST$  is a single triangulation line this becomes

$$D^2 = a^2 n \left\{ NQ^2 + \rho_0^2 + \frac{L^2}{12} + R_0^2 + \frac{L^2}{12} \right\}$$

where  $\rho_0, R_0$  are the distances of the centre of gravity of  $ST$  from  $N$  and  $Q$  respectively and  $S$  now becomes equal to  $L$ . Putting in numerical quantities and expressing the results in feet it follows, as in (4), that

$$D = 4.03 M \sqrt{L \left( R_0^2 + \rho_0^2 + \frac{L^2}{6} + NQ^2 \right)}$$

Finally if  $ST$  is composed of a number of triangulation lines, the probable displacement in any direction

$$P = \frac{D}{\sqrt{2}} = 2.85 \sqrt{\sum M^2 L \left( R_0^2 + \rho_0^2 + \frac{L^2}{6} + NQ^2 \right)} \dots \dots \dots (7)$$

where the points  $Q$  differ for each triangulation line and are found for each as has been done above for  $ST$  alone.

If closure had also been effected on Laplace station as well as on base lines at  $O$  and  $K$  this clearly would become

$$P = 4.03 \sqrt{\sum M^2 L \left( \rho_0^2 + \frac{L^2}{12} + NQ^2 \right)} \dots \dots \dots (8)$$

Equations (7) and (8) illustrate the statement made in §2 that closure only on a base and not on a Laplace station does not improve the results nearly so much as closure of both kinds simultaneously will probably do.

Probable errors after adjustment.

4. Consider any function  $F$  of quantities  $x_1, x_2, x_3 \dots$  which have been found by measurement. If the true values of these quantities are  $x_1 + v_1, x_2 + v_2, \dots$  and the true value of the function is  $F + dF$  then

$$dF = F(x_1 + v_1, x_2 + v_2, \dots) - F(x_1, x_2, \dots) \\ \doteq f_1 v_1 + f_2 v_2 + \dots \dots \dots (9)$$

where  $f_1 = \frac{dF}{dx_1}, f_2 = \frac{dF}{dx_2}, \text{etc.} \dots \dots$

Any conditions which may be imposed on  $x_1, x_2, \dots$  result in equations of form

$$\left. \begin{aligned} a_1 v_1 + a_2 v_2 + \dots + a_n v_n - l_1 &= 0 \\ b_1 v_1 + b_2 v_2 + \dots + b_n v_n - l_2 &= 0 \\ \text{etc.} \end{aligned} \right\} \dots \dots \dots (10)$$

Then  $dF$  may be written

$$dF = f_1 v_1 + f_2 v_2 \dots - k_1 (a_1 v_1 + a_2 v_2 + \dots - l_1) - k_2 (b_1 v_1 + b_2 v_2 + \dots - l_2) \dots$$

in which any values may be assigned to  $k_1, k_2 \dots$ . Hence

$$dF = (f_1 - k_1 a_1 - k_2 b_1 - \dots) v_1 + (f_2 - k_1 a_2 - k_2 b_2 - \dots) v_2 + \dots + k_1 l_1 + k_2 l_2 + \dots \dots \dots (11)$$

Let the reciprocal weights of  $x_1, x_2 \dots$  be  $u_1, u_2 \dots$  and let the reciprocal weight of  $F$  be  $u_F$ . Then it follows that

$$u_F = (f_1 - k_1 a_1 - k_2 b_1 - \dots)^2 u_1 + (f_2 - k_1 a_2 - k_2 b_2 - \dots)^2 u_2 + \dots \dots \dots (12)$$

If  $v_1, v_2 \dots$  are the most probable values of the corrections, then  $u_F$  must be a minimum, and accordingly  $k_1, k_2, \dots$  are to be determined from the following equations

$$\left. \begin{aligned} [u a a] k_1 + [u a b] k_2 + \dots &= [u a f] \\ [u b a] k_1 + [u b b] k_2 + \dots &= [u b f] \end{aligned} \right\} \dots \dots \dots (13)$$

in which  $[u a b]$  is written for  $u_1 a_1 b_1 + u_2 a_2 b_2 + \dots$ , etc.

Reverting to (12) and developing the squares it is clear that

$$\begin{aligned} u_F &= [u f f] - 2 [u a f] k_1 - 2 [u b f] k_2 - \dots \\ &\quad + k_1 ([u a a] k_1 + [u a b] k_2 + \dots) + k_2 ([u a b] k_1 + [u b b] k_2 + \dots) + \dots \\ &= [u f f] - [u a f] k_1 - [u b f] k_2 - \dots \dots \dots (14) \end{aligned}$$

by using the equations (13)\*.

5. In considering the probable errors of triangulation it will be permissible to treat the quantities  $\epsilon$  and  $\eta$  as errors in observed quantities, and as being independent except in so far as the several closing conditions relate them. Distinguish the several triangulation lines which make up a network of triangulation by the prefixes 1, 2, 3, etc., and the several stations on any triangulation by suffixes 1, 2, 3, etc. It will also be permissible in an investigation into the probable errors after adjustment to replace the triangulation by its projection on a plane, the projection of the general map of India naturally being selected for this purpose. This enables the closing conditions to be written in the following form, either for closure round a circuit or for closure between base lines or Laplace stations :

$$\left. \begin{aligned} \sum \sum_r \epsilon_s &= 0 && \text{side closure} \\ \sum \sum_r \eta_s &= 0 && \text{azimuth closure} \\ \sum \sum_r (x\epsilon - y\eta)_s &= 0 && \text{easting closure} \\ \sum \sum_r (y\epsilon + x\eta)_s &= 0 && \text{northing closure} \end{aligned} \right\} \dots \dots \dots (15)$$

the first summation for  $S$  corresponding to the number of sides in a triangulation line and the second for  $r$  corresponding to the number of triangulation lines: and  $x, y$  being the coordinates of any station referred to the closing point of the particular circuit as origin.

\* This deduction is taken from p. 229 of "A Treatise on the Adjustment of Observations" by T. W. Wright, New York, 1884.

\* There may be any number of each type of closures, and not necessarily the same number for each type. The triangulation lines over which the summations are taken differ in part for each closure of any type. Thus in the case supposed in Chapter VI §14 the line 3 occurs in both circuits. The quantities whose probable errors are required are of the same form as the left hand sides of equations (15) : but the limits are different.

It is clear that the coefficients in (13) may be written, agreeably with the notation already adopted

$$\left. \begin{aligned} [u a a] &= \sum_r [u a a] \\ [u a b] &= \sum_r [u a b] \\ \text{etc.} \end{aligned} \right\} \dots \dots \dots (16)$$

and accordingly the portion relating to each triangulation line may be separately computed. [For the side and azimuth closures the coefficients  $a, b$  are all unity : for the other two closures they are  $x$  or  $y$ . Consider the case of a triangulation line which forms part of a closed circuit, so that there are closures of each type. It may also form part of a closure between base lines and Laplace stations. If so the coefficients of the several  $\epsilon^s$  and  $\eta^s$  remain the same as in corresponding type of closure in the circuit, and the coefficients reduce to type  $[u a a]$ . For the case of clearness take the specific case shown in the diagram when there are base lines and Laplace stations at  $A$  and  $B$ .

Consider the line 2. It enters into the following relations :

	A	1	2	3	B
$a$	-----				
$b$		6	4		
$c$		5			
$d$					

$$\begin{aligned} e \quad & \sum_2 (x\epsilon - y\eta)_s + \sum_4 (x\epsilon - y\eta)_s + \sum_5 (x\epsilon - y\eta)_s + \sum_6 (x\epsilon - y\eta)_s = 0 \\ f \quad & \sum_2 (y\epsilon + x\eta)_s + \sum_4 (y\epsilon + x\eta)_s + \sum_5 (y\epsilon + x\eta)_s + \sum_6 (y\epsilon + x\eta)_s = 0 \end{aligned}$$

The symbolic coefficients  $a b c d e f$  are each written opposite one of these equations.

Along any triangulation line for  $u$  write  $a^2$ ,  $a$  being the probable value of  $\epsilon$  or  $\eta$ . Suppose that  $n$  is the number of sides in the triangulation line, all considered of equal length. Then omitting the prefix 2 for simplicity and considering only the line 2

$$\begin{aligned} [u a a] &= na^2 \\ [u a b] &= ua^2 \\ [u a c] &= 0 \\ [u a d] &= 0 \\ [u a e] &= a^2 \sum x = na^2 X \\ [u a f] &= a^2 \sum y = na^2 Y \end{aligned}$$

where  $X, Y$  are the coordinates of the centre of gravity of the line 2. These are typical of all the combinations of  $a b c d$  with  $a b c d e f$ . The remaining typical coefficients are represented by

$$[u e e] = [u f f] = a^2 \sum (x^2 + y^2)$$

If  $x_1, x_2$  are the limits of  $x$  and  $x_1 = X - x_0$  and  $x_2 = X + x_0$

$$\begin{aligned} \Sigma x^2 &\doteq \frac{n}{x_2 - x_1} \int x^2 dx \\ &= \frac{n}{3} (x_2^3 + x_2 x_1 + x_1^3) \\ &= \frac{n}{3} (\overline{X + x_0}^3 + (X + x_0)(X - x_0) + \overline{X - x_0}^3) \\ &= n(X^2 + \frac{1}{3}x_0^2) \dots \dots \dots (17) \end{aligned}$$

Similarly  $\Sigma y^2 = n(Y^2 + \frac{1}{3}y_0^2)$

Hence

$$\Sigma (x^2 + y^2) = n(R^2 + \frac{1}{3}r_0^2) = n(R^2 + \frac{l^2}{12}) \dots \dots \dots (18)$$

Finally  $[uee] = na^2 (R^2 + \frac{l^2}{12})$

Also  $[uef] = a^2 \Sigma (xy - yx) = 0$

The complete result is given in tabular form :—

*Values of  ${}_2[uaa] \doteq na^2$ , etc.*

	<i>a</i>	<i>b</i>	<i>c</i>	<i>d</i>	<i>e</i>	<i>f</i>	
Side	1	1	0	0	X	Y	Base
Side	1	1	0	0	X	Y	
Azimuth	0	0	1	1	-Y	X	Laplace
Azimuth	0	0	1	1	-Y	X	
Easting	X	X	-Y	-Y	$R^2 + \frac{l^2}{12}$	0	
Northing	Y	Y	X	X	0	$R^2 + \frac{l^2}{12}$	

It is clear from this that there are really only four types of quantities, and that a closure on an outer circuit or base gives rise to the same coefficient as the corresponding closure within the circuit. All the quantities of the form  ${}_r[ua b]$  are given by the following schemes, in which the four closures, side azimuth, easting, northing are represented by S, A, E, N.

Values of  ${}_r[ua b] \doteq na^2$

	S	A	E	N	
S	1	0	X	Y	
A	0	1	-Y	X	
E	X	-Y	$R^2 + \frac{l^2}{12}$	0	
N	Y	X	0	$R^2 + \frac{l^2}{12}$	(19)



The value of  $na^2$  is given by (1) in terms of  $M$  and  $S$ , or  $M$  and  $L$  if straight portions are considered separately.

The above scheme deals with the cases in which the coordinates  $x, y$  which occur all refer to the same origin. It only remains to take the case where two origins occur and to form the typical expression  $[uaf]$ . The coefficients  $a$ , etc. are either  $0, 1, x$  or  $\pm y$  in which  $x, y$  are the coordinates of any point on the line referred to the point of closure for which the corresponding condition of closure was formed. When it is desired to find the probable error with regard to any point  $O$  (for example Kalianpur, the origin of the survey), then the coefficients  $f$ , etc. are either  $0, 1, x_0$  or  $\pm y_0$  where  $x_0, y_0$  are coordinates referred to this point. If the circuit closing point is  $P$  let  $x_p, y_p$  be the coordinates of  $P$  with regard to origin  $O$  then

$$x_0 = x_p + x \quad y_0 = y_p + y.$$

The values of  $[uaf] \div na^2$  are now indicated in tabular form similar to (19). As only the ratios of the coefficients in (13) are required  $na^2$  may be replaced by  $M^2L$  by means of (1).

	$S_f$	$A_f$	$E_f$	$N_f$	
S	1	0	$X + x_p = X_0$	$Y + y_p = Y_0$	
A	0	1	$-Y - y_p = -Y_0$	$X + x_p = X_0$	. . . . . (20)
E	$X$	$-Y$	$R^2 + \frac{L^2}{12} + x_p X + y_p Y$	$-x_p Y + y_p X$	
N	$Y$	$X$	$x_p Y - y_p X$	$R^2 + \frac{L^2}{12} + x_p X + y_p Y$	

All the quantities  $X, Y, R$  are measured from  $P$  while  $x_p, y_p$  are coordinates of  $P$  and  $X_0, Y_0$  are the coordinates of the mid point of the triangulation line relative to  $O$ .

6. The method explained in §§ 4,5 will now be applied to the determination of probable errors in the N.W. Quadrilateral of the Indian triangulation, after all adjustments have been carried out. The probable errors most generally required will be those with reference to the origin of the survey at Kalianpur (Sironj Base). In putting down the conditions of closure any closing point may be chosen: but when probable errors with regard to Kalianpur are desired, Kalianpur will naturally be selected as the closing point and origin for  $X$  and  $Y$ . Chart I shows all the triangulation of India: in charts II . . . V it is represented diagrammatically, each series being replaced by one or more straight lines, which may be regarded as the equivalent triangulation lines. The Indian triangulation was divided for purposes of adjustment into five portions, *viz.* the N.W. Quadrilateral, S.E. Quadrilateral, N.E. Quadrilateral, Southern Trigon, S.W. Quadrilateral. These were adjusted in the order stated, so that the first two were quite independently adjusted while the third was adjusted on the first two: the fourth was adjusted on the second, and the fifth was adjusted on the first, second and fourth. The Burma quadrilateral (chart IV) has just\* been adjusted by the methods of Chapter VI and is being adjusted on the eastern series (Shillong Meridional, No. 44) of the N.E. Quadrilateral.

In chart II the series which were taken account of simultaneously in each quadrilateral or trigon are shown in full lines, while some additional series afterwards adjusted on these series are shown in broken lines. The series which are common to adjacent quadrilaterals or trigon are distinguished by heavy lines. The numbers written by the side of each triangulation line in the chart are those which have been applied in Table XLIV to the several series of the triangulation. The eight base lines of the triangulation of India and the Mergui base in Burma are shown: also the points at which it has been possible to form Laplace equations. The circuits are indicated by roman numerals and the points of closure by small arcs at the closing angle, *e. g.* at  $D$ . The junction points of the triangulation lines are distinguished by letters  $A, B . . . Z, a, b . . .$  with suffixes 1, 2, 3, 4, 5, 6 corresponding to the N.W. Quadrilateral, S.E. Quadrilateral, N.E. Quadrilateral, S. Trigon, S.W. Quadrilateral, Burma Quadrilateral.

\* 1917.

The first step is to find  $M^2L$  for each triangulation line,  $L$  being the length in units of 100 miles. Chart II is on the scale of 100 miles = 1 inch, so that  $L$  is the length of each line on the chart in inches. As an example take the line between the Sironj base and the Dehra base. This is composed of two triangulation lines representing a portion of the Great Arc Series, No. 6. From table XLIV the value of  $M$  is 0.71 and by measurement on the chart the values of  $L$  for the two parts are found to be 2.15 and 2.22 inches. Hence the values of  $M^2L$  are 1.083 and 1.109 respectively.

The values of  $M^2L$  for all the triangulation lines are exhibited in table XLV together with certain related quantities to which reference will be made later. To proceed with the formation of the equations of form (13) which are necessary to determine the several probable errors, all quantities of the types indicated in (20) have to be formed. To any coefficient  $[u g h]$  several component terms, each corresponding to a particular triangulation line, may contribute. Some of these are exhibited in table XLV, while the remainder are found in table XLVI. To go more into detail, the N.W. Quadrilateral (*vide* chart II) is divided into five circuits I, II, III, IV, V. In each circuit there are four types of closure—side, azimuth, easting, northing—which may be characterised by suffixes  $s, a, e, n$ . These give rise to twenty conditions  $I_s, I_a, I_e, I_n, II_s, \dots, V_n$ . In addition there are three extra base lines giving conditions  $VI_s, VII_s, VIII_s$ . To investigate the additional value of having as many Laplace stations as there are base lines, the cases have also been worked out for three Laplace closures at the same points as the base line closures, giving rise to conditions  $VI_a, VII_a, VIII_a$ . The 26 conditions which result make it necessary to determine 26 multipliers  $k_1 \dots k_{26}$  by equations of form (13). In table XLVII the coefficients of the left hand sides of these equations are given. The method by which these coefficients are derived will now be given in detail for a few of them. The letters used correspond to those shown marginally in table XLVII.

By (20)  ${}_r[uaa] \div M^2L = 1$ : hence  $[uaa] = \Sigma M^2L$  round circuit I = 3.66 from table XLV. Also  $[uab] = 0$ ,  $[uac] = \Sigma X M^2L = 2.20$  from table XLV. In the formation of any coefficients involving any pair of  $a, b, c$  or  $d$  it is clear that the summation extends around the circuit I, for the corresponding conditions relate to this complete circuit. The case is different when a coefficient involving one of the quantities,  $a, b, c, d$  and one other quantity, say  $e$ , are considered. In this case both circuits I and II are involved and it is only the part common to the two circuits that has to be considered. Moreover in this case two origins are introduced. The portion common to circuits I and II is the line  $D_1T_1$ . It will be seen by (20) that

$$[uae] = M^2L \text{ for } D_1T_1 = .70 \quad \text{from table XLV.}$$

$$[uaf] = 0$$

$$[uag] = X_0 M^2L \text{ for } D_1T_1, \quad X_0 \text{ referring to the circuit to which } g \text{ relates, } i. e. \text{ circuit II} \\ = +.41$$

$$[uah] = Y_0 M^2L \text{ for } D_1T_1 \text{ from circuit II} = -2.30$$

This deals with all the conditions which relate to a portion which occurs also in condition  $a$ , except the base line conditions  $VI$ , and  $VII_s$ . The most complex case is the coefficient corresponding to northing or easting relations in two circuits which have a portion in common. An example of this is  $[udg]$ . The circuits involved I, II have the portion  $T_1D_1$  only in common. Hence from (20) as may be seen opposite the entry I, II in table XLVI,  $[udg] = M^2L (x_p Y - y_p X) = -1.14$ . Similarly  $[udh] = M^2L (R^2 + L^2/12 + x_p X + y_p Y) = 7.35$ . The quantities in tables XLV, XLVI depend on measurements of  $L_1, R_1, X, Y$  taken from a chart. These are not quite precise, and so quantities deduced from them are not quite precise. All the quantities which are known by symmetry to be equal in pairs have been separately determined by way of a check. They differ slightly as will be seen in table XLVI, the worst case being the coefficient which occurs in line II, III and also in line III, II of which the values 29.56 and 29.68 are obtained. The quantity made use of is the mean 29.62 shown in block type.

TABLE XLV.  
N.W. Quadrilateral.

Base line closures	Circuit	Line	Series	M	L	A = M·L	Closing point referred to Kalianpur	R <sup>2</sup>	L <sup>2</sup> /12	R <sup>2</sup> +L <sup>2</sup> /12	X	Y	A × (R <sup>2</sup> +L <sup>2</sup> /12)	ΔX	ΔY	Base line closures	Circuit	Line	A = M·L	ΔX	ΔY
4 1 3	I	D <sub>1</sub> T <sub>1</sub>	33	0-37	5-10	0-608	D <sub>1</sub> x <sub>p</sub> = -0-01 y <sub>p</sub> = +5-18	6-45	2-17	8-62	-0-04	-2-54	6-02	-0-03	-1-77	VI 1	I	A <sub>1</sub> B <sub>1</sub>	1-083	+0-88	-4-46
		T <sub>1</sub> A <sub>1</sub>	25	0-60	0-99	0-353		26-50	0-08	26-68	+0-42	-5-13	9-38	+0-15	-1-81			B <sub>1</sub> C <sub>1</sub>	1-109	+0-08	-2-16
		A <sub>1</sub> H <sub>1</sub>	8	0-71	2-15	1-083		17-63	6-39	18-02	+0-81	-4-12	19-50	+0-88	-4-46			+2-102	+1-88	-6-62	
		B <sub>1</sub> C <sub>1</sub>	6	0-71	2-22	1-100		4-67	0-41	4-08	+0-88	-1-95	5-52	+0-98	-2-16						
		C <sub>1</sub> D <sub>1</sub>	22	0-55	1-38	0-418		0-46	0-18	0-62	+0-53	-0-42	0-26	+0-23	-0-18						
													3-601								
	II	E <sub>1</sub> U <sub>1</sub>	23	1-21	1-36	1-090	E <sub>1</sub> x <sub>p</sub> = -1-54 y <sub>p</sub> = +5-94	0-48	0-15	0-63	-0-04	-0-60	1-26	-0-08	-1-37	VII 3	I II IV V	C <sub>1</sub> D <sub>1</sub>	0-418	+0-23	-0-18
		U <sub>1</sub> X <sub>1</sub>	23	1-21	1-85	2-706		5-16	0-20	5-45	-0-11	-2-27	14-75	-0-30	-6-14			D <sub>1</sub> E <sub>1</sub>	0-306	+0-10	-0-12
		X <sub>1</sub> Y <sub>1</sub>	23	1-21	1-13	1-654		14-19	0-11	14-40	-0-27	-3-77	23-80	-0-45	-8-23			E <sub>1</sub> F <sub>1</sub>	0-338	+0-15	-0-10
		Y <sub>1</sub> S <sub>1</sub>	23	1-21	1-44	2-107		25-37	0-17	25-54	-0-27	-5-03	53-80	-0-67	-10-60			F <sub>1</sub> G <sub>1</sub>	0-163	+0-02	-0-02
		S <sub>1</sub> T <sub>1</sub>	25	0-60	0-70	0-252		33-53	0-04	33-57	+0-20	-5-79	8-45	+0-05	-1-40			+1-223	+0-40	-0-42	
		T <sub>1</sub> D <sub>1</sub>	33	0-37	6-10	0-698		11-24	2-17	13-41	+0-59	-3-30	9-36	+0-41	-2-30						
D <sub>1</sub> E <sub>1</sub>	22	0-55	1-01	0-396	0-25	0-08	0-33	+0-32	-0-39	0-10	+0-10	-0-12									
											9-713										
5 4 4 4	III	J <sub>1</sub> L <sub>1</sub>	32	0-43	1-99	0-368	J <sub>1</sub> x <sub>p</sub> = -4-26 y <sub>p</sub> = +3-36	1-00	0-33	1-33	-0-90	-0-44	0-49	-0-33	-0-18	VIII 5	III V	J <sub>1</sub> L <sub>1</sub>	0-368	-0-33	-0-18
		L <sub>1</sub> N <sub>1</sub>	32	0-43	1-67	0-309		6-71	0-23	6-94	-1-98	-1-67	2-14	-0-61	-0-52			L <sub>1</sub> N <sub>1</sub>	0-309	-0-61	-0-52
		N <sub>1</sub> P <sub>1</sub>	25	0-60	0-28	0-822		7-81	0-43	8-24	-1-01	-2-81	6-77	-0-82	-2-15			+1-373	-0-18	-0-24	
		P <sub>1</sub> Q <sub>1</sub>	25	0-60	1-02	0-368		8-53	0-00	8-62	+0-61	-3-85	9-17	+0-24	-1-05						
		Q <sub>1</sub> S <sub>1</sub>	25	0-40	1-46	0-526		12-91	0-13	13-09	+1-85	-3-08	6-88	+0-07	-1-62						
		S <sub>1</sub> Y <sub>1</sub>	23	1-21	1-44	2-106		12-00	0-17	12-17	+3-45	-2-45	25-61	+5-16	-5-18						
	Y <sub>1</sub> X <sub>1</sub>	23	1-21	1-13	1-653	7-42	0-11	7-53	+2-46	-1-19	12-46	+4-05	-1-97	+1-373	-0-45	-0-93					
	X <sub>1</sub> U <sub>1</sub>	23	1-21	1-85	2-707	6-00	0-29	7-19	+2-61	+0-30	10-46	+7-06	+0-81								
	U <sub>1</sub> V <sub>1</sub>	46	0-53	0-98	0-275	5-00	0-08	5-98	+2-21	+1-01	1-65	+0-61	+0-28								
	V <sub>1</sub> J <sub>1</sub>	45	0-53	1-04	0-545	0-92	0-31	1-23	+0-88	+0-36	0-67	+0-48	+0-21								
												9-679									
	IV	F <sub>1</sub> V <sub>1</sub>	37	0-59	2-37	0-825	F <sub>1</sub> x <sub>p</sub> = -2-50 y <sub>p</sub> = +6-55	1-44	0-47	1-91	+0-01	-1-20	1-58	+0-01	-0-99	+1-373	-0-45	-0-93			
V <sub>1</sub> U <sub>1</sub>		45	0-53	0-98	0-275	4-05		0-03	5-03	+0-45	-2-18	1-83	+0-12	-0-60							
U <sub>1</sub> E <sub>1</sub>		23	1-21	1-36	1-090	2-55		0-15	2-70	+0-93	-1-30	5-37	+1-85	-2-50							
E <sub>1</sub> F <sub>1</sub>		22	0-55	1-11	0-336	0-30		0-10	0-40	+0-46	-0-30	0-13	+0-15	-0-10							
											3-426										
											8-46	+2-13	-4-28								
5 6 3	V	G <sub>1</sub> H <sub>1</sub>	33	0-43	1-73	0-320	G <sub>1</sub> x <sub>p</sub> = -2-68 y <sub>p</sub> = +6-80	0-87	0-26	1-12	-0-56	-0-75	0-36	-0-18	-0-24	+1-373	-0-45	-0-93			
		H <sub>1</sub> J <sub>1</sub>	32	0-43	2-03	0-376		7-59	0-34	7-93	-1-20	-2-48	2-93	-0-45	-0-03						
		J <sub>1</sub> V <sub>1</sub>	46	0-53	1-04	0-545		0-65	0-31	0-86	-0-50	-3-05	5-37	-0-27	-1-66						
		V <sub>1</sub> F <sub>1</sub>	37	0-59	2-37	0-825		2-26	0-47	2-72	+0-39	-1-45	2-24	+0-32	-1-20						
		F <sub>1</sub> G <sub>1</sub>	22	0-55	0-54	0-163		0-04	0-02	0-06	+0-13	-0-15	0-01	+0-02	-0-02						
													2-229								
											10-96	-0-56	-4-05								

TABLE XLVI.

N.W. Quadrilateral. Common points of adjacent circuits.

Circuits	Line	First circuit referred to second		For first circuit			A X r <sub>p</sub>	A Y y <sub>p</sub>	Sum of last 3 columns	A F x <sup>p</sup>	A X y <sup>p</sup>	Sum of last 2 columns	
		x <sub>p</sub>	y <sub>p</sub>	ΔX	ΔY	A(R <sup>2</sup> +L <sup>2</sup> /12)							
I II	II I	D <sub>1</sub> T <sub>1</sub>	+ .63	- .76	- .03	- 1.77	+ 6.02	- 0.02	+ 1.24	+ 7.34	+ 7.35	- 1.11	- 0.02
		T <sub>1</sub> D <sub>1</sub>	- .63	+ .76	+ .41	- 2.30	+ 9.36	- 0.26	- 1.74	+ 7.96			
II III	III II	U <sub>1</sub> X <sub>1</sub>	+ 2.72	+ 2.58	- .30	- 6.14	+ 14.75	- 3.50	- 59.20	+ 29.56	- 63.45	+ 3.40	- 59.06
		X <sub>1</sub> U <sub>1</sub>			- .45	- 6.23	+ 23.80						
III II	II III	Y <sub>1</sub> S <sub>1</sub>	- 2.72	- 2.58	- .67	- 10.60	+ 53.80	- 44.15	+ 16.30	+ 29.68	+ 17.19	+ 41.95	+ 59.14
		S <sub>1</sub> Y <sub>1</sub>			+ 5.16	- 5.16	+ 25.61						
IV II	II IV	E <sub>1</sub> U <sub>1</sub>	+ 0.96	- 0.61	+ 4.05	- 1.97	+ 12.46	- 1.77	- 1.58	+ 2.02	+ 2.20	+ 2.48	- 1.13
		U <sub>1</sub> E <sub>1</sub>	- 0.96	+ 0.61	- 1.35	- 2.39	+ 5.37						
IV III	III IV	V <sub>1</sub> V <sub>1</sub>	- 1.76	- 3.10	+ 7.00	+ 0.81	+ 19.48	- 1.07	- 1.01	- 0.33	- 0.32	- 0.69	+ 1.94
		V <sub>1</sub> V <sub>1</sub>	+ 1.76	+ 3.10	+ .61	+ .28	+ 1.65						
III V	V III	V <sub>1</sub> J <sub>1</sub>	- 1.83	- 3.44	+ 4.05	+ 0.81	+ 19.48	- 0.72	- 0.71	- 0.71	- 0.71	+ 1.65	- 1.36
		J <sub>1</sub> V <sub>1</sub>	+ 1.83	+ 3.44	- .27	- 1.60	+ 5.37						
IV V	V IV	F <sub>1</sub> V <sub>1</sub>	+ 0.38	- 0.25	+ 0.01	- .99	+ 1.58	- 0.12	+ 0.25	+ 1.83	+ 1.83	+ 0.38	- 0.08
		V <sub>1</sub> F <sub>1</sub>	- 0.38	+ 0.25	+ .32	- 1.20	+ 2.24						



7. To find the probable errors at any point of side azimuth easting or northing it is necessary to form the right hand sides of the equations of which the left hand side coefficients are exhibited in table XLVII, i.e. to complete the formation of equations (13). The quantities [u a f] etc. are of the same nature as those already formed: but differ in the lines for which they have to be

TABLE XLVIII.

Line	Circuit	M	L	A = M <sup>2</sup> L	R <sup>2</sup>	$\frac{L^2}{12}$	X	Y	R <sup>2</sup> + $\frac{L^2}{12}$	A (R <sup>2</sup> + $\frac{L^2}{12}$ )	AX	AY	X <sub>0</sub>	Y <sub>0</sub>
K <sub>1</sub> L <sub>1</sub>	III	0.43	1.55	0.286	1.03	0.20	-0.91	-0.45	1.23	0.35	-0.26	-0.13	-5.17	+2.91
M <sub>1</sub> N <sub>1</sub>	III	0.43	0.76	0.144	8.70	0.05	-2.06	-2.11	8.75	1.26	-0.30	-0.30	-6.32	+1.27
O <sub>1</sub> P <sub>1</sub>	III	0.80	0.57	0.205	7.21	0.03	-0.14	-2.68	7.24	1.48	-0.03	-0.55	-4.40	+0.68
R <sub>1</sub> S <sub>1</sub>	III	0.80	0.53	0.191	15.27	0.02	+2.31	-3.15	15.29	2.02	+0.44	-0.60	-1.95	+0.21
V <sub>1</sub> W <sub>1</sub>	III	0.63	0.74	0.208	2.45	0.05	+1.43	+0.64	2.50	0.52	+0.30	+0.13	-2.83	+4.00
V <sub>1</sub> W <sub>1</sub>	V	0.53	0.74	0.208	7.85	0.05	+0.05	-2.80	7.00	1.64	+0.01	-0.58	-2.83	+4.00
W <sub>1</sub> J <sub>1</sub>	III	0.53	1.20	0.337	0.37	0.12	+0.55	+0.25	0.49	0.16	+0.19	+0.08	-3.71	+3.61
W <sub>1</sub> J <sub>1</sub>	V	0.53	1.20	0.337	10.87	0.12	-0.83	-3.19	10.99	3.70	-0.28	-1.08	-3.71	+3.61

TABLE XLIX.

Circuit, Base and Laplace closures.

Line	Circuit	Closing point of circuit referred to Kalianpur		(1)	(2)	(3)	(4)	(5)	(6) =	(7) =	(8)	(9)	(10)	(11) =	(12)	(13)	(14) =	(15)	(16) =
		x <sub>p</sub>	y <sub>p</sub>	$\frac{A}{M^2L}$	$\Delta x_p$	$\Delta y_p$	$\Delta X$	$\Delta Y$	$(2)+(4) = \Delta X_0$	$(3)+(5) = \Delta Y_0$	$\sum_{i=1}^n \frac{L_i^2}{12}$	$A \sum x_p$	$A \sum y_p$	$(6)+(9) = A \sum y_p$	$A \sum x_p$	$\Delta X y_p$	$\Delta Y x_p$	$(12)+(13)$	$\sum_{i=1}^n L_i^2$
A <sub>1</sub> B <sub>1</sub>	I, VI	-0.91	+5.18	1.083	-0.90	+5.61	+0.88	-4.46	-0.11	+1.15	+19.50	-0.80	-23.10	-4.40	+4.06	-4.56	-0.50	1.52	1.65
B <sub>1</sub> C <sub>1</sub>	I, VI	-0.91	+5.18	1.109	-1.01	+5.74	+0.98	-2.16	-0.03	+3.58	+5.52	-0.89	-11.10	-6.56	+1.97	-5.08	-3.11	10.84	12.02
C <sub>1</sub> D <sub>1</sub>	I, VII	-0.91	+5.18	0.418	-0.38	+2.17	+0.22	-0.18	-0.16	+1.90	+0.26	-0.20	-0.93	-0.87	+0.16	-1.14	-0.98	22.96	9.69
D <sub>1</sub> E <sub>1</sub>	II, VII	-1.54	+5.94	0.306	-0.47	+1.82	+0.10	-0.12	-0.37	+1.70	+0.10	-0.15	-0.71	-0.76	+0.18	-0.59	-0.41	32.38	9.91
E <sub>1</sub> F <sub>1</sub>	IV, VII	-2.50	+6.55	0.336	-0.84	+2.20	+0.15	-0.10	-0.69	+2.10	+0.13	-0.37	-0.66	-0.90	+0.28	-0.98	-0.73	43.33	14.66
F <sub>1</sub> G <sub>1</sub>	V, VII	-2.88	+6.80	0.163	-0.47	+1.11	+0.02	-0.02	-0.45	+1.09	+0.01	-0.06	-0.14	-0.19	+0.06	-0.14	-0.08	51.81	8.46
G <sub>1</sub> H <sub>1</sub>	V, VIII	-2.88	+6.80	0.320	-0.92	+2.18	-0.18	-0.24	-1.10	+1.94	+0.36	+9.52	-1.63	-0.75	+0.69	+1.22	+1.91	48.60	15.58
H <sub>1</sub> J <sub>1</sub>	V, VIII	-2.88	+6.80	0.376	-1.08	+2.56	-0.45	-0.93	-1.53	+1.63	+2.98	+1.30	-0.32	-2.04	+2.68	+3.06	+5.74	35.65	13.40
J <sub>1</sub> L <sub>1</sub>	III, VIII	-4.26	+3.36	0.286	-1.22	+0.96	-0.26	-0.13	-1.48	+0.83	+0.35	+1.11	-0.44	+1.02	+0.55	+0.87	+1.42	35.40	10.12
L <sub>1</sub> N <sub>1</sub>	III, VIII	-4.26	+3.36	0.368	-1.57	+1.24	-0.33	-0.16	-1.90	+1.08	+0.49	+1.41	-0.54	+1.36	+0.68	+1.11	+1.79	35.48	13.66
M <sub>1</sub> N <sub>1</sub>	III, VIII	-4.26	+3.36	0.141	-0.61	+0.48	-0.30	-0.30	-0.91	+0.62	+1.26	+1.28	-1.01	+1.63	+1.28	+1.01	+2.29	41.59	5.90
N <sub>1</sub> P <sub>1</sub>	III, VIII	-4.26	+3.36	0.309	-1.32	+1.04	-0.61	-0.52	-1.93	+0.82	+2.14	+2.60	-1.75	+2.09	+2.22	+2.05	+4.27	42.03	12.99
O <sub>1</sub> P <sub>1</sub>	III	-4.26	+3.36	0.205	-0.87	+0.69	-0.09	-0.65	-0.90	+0.14	+1.48	+0.13	-1.85	-0.24	+2.34	+0.10	+2.44	19.85	4.07
N <sub>1</sub> P <sub>1</sub>	III	-4.26	+3.36	0.822	-3.50	+2.78	-0.82	-2.15	-4.32	+0.61	+8.77	+3.49	-7.22	+3.04	+0.16	+2.76	+11.92	28.66	23.55
P <sub>1</sub> Q <sub>1</sub>	III	-4.26	+3.36	0.368	-1.57	+1.24	+0.24	-1.05	-1.33	+0.19	+3.17	-1.02	-3.53	-1.38	+4.47	-0.81	+3.66	13.45	4.95
R <sub>1</sub> S <sub>1</sub>	III	-4.26	+3.36	0.191	-0.81	+0.64	+0.44	-0.60	-0.37	+0.04	+2.92	-1.67	-2.02	-0.97	+3.56	-1.48	+1.08	3.87	0.74
Q <sub>1</sub> S <sub>1</sub>	III	-4.26	+3.36	0.526	-2.24	+1.67	+0.97	-1.62	-1.27	+0.15	+6.88	+4.13	-5.44	-2.69	+6.90	+3.26	+3.64	6.06	3.19
S <sub>1</sub> T <sub>1</sub>	II	-1.54	+5.94	0.252	-0.39	+1.50	+0.05	-1.46	-0.34	+0.04	+8.45	-0.08	-8.67	-0.30	+2.25	-0.30	+1.95	1.81	0.47
T <sub>1</sub> A <sub>1</sub>	I	-0.91	+5.18	0.353	-0.32	+1.83	+0.15	-1.81	-0.17	+0.02	+9.38	-0.14	-9.37	-0.13	+1.65	-0.78	+0.87	0.32	0.11
T <sub>1</sub> D <sub>1</sub>	I	-0.91	+5.18	0.668	-0.64	+3.62	-0.03	-1.77	-0.67	+1.85	+6.02	+0.03	-9.17	-3.12	+1.61	+0.16	+1.77	10.04	7.01
T <sub>1</sub> D <sub>1</sub>	II	-1.54	+5.94	0.698	-1.07	+4.15	+0.41	-2.30	-0.66	+1.85	+9.36	-0.63	-13.66	-4.93	+3.54	-2.44	+1.10		
S <sub>1</sub> Y <sub>1</sub>	II	-1.54	+5.94	2.107	-3.24	+12.51	-0.57	-10.60	-3.81	+1.91	+53.80	+0.88	-62.96	-8.28	+16.32	+3.39	+19.71	4.28	0.02
S <sub>1</sub> Y <sub>1</sub>	III	-4.26	+3.36	2.106	-8.97	+7.08	+5.16	-5.16	-3.81	+1.92	+25.61	-21.98	-17.34	-13.71	+21.98	-17.34	+4.64		
Y <sub>1</sub> X <sub>1</sub>	II	-1.54	+5.94	1.654	-2.55	+9.82	-0.45	-6.24	-3.00	+3.58	+23.80	+0.69	-37.07	-12.58	+0.61	+2.67	+12.28	8.00	13.38
Y <sub>1</sub> X <sub>1</sub>	III	-4.26	+3.36	1.653	-7.04	+5.55	+4.05	-1.07	-2.99	+3.58	+12.46	-17.26	-6.62	-11.41	+8.39	-13.61	-5.22		
X <sub>1</sub> U <sub>1</sub>	II	-1.54	+5.94	2.706	-4.17	+16.07	-0.30	-6.14	-4.47	+0.93	+14.75	+0.46	-36.47	-21.26	+0.46	+1.78	+11.24	16.48	44.59
X <sub>1</sub> U <sub>1</sub>	III	-4.26	+3.36	2.707	-11.53	+9.10	+7.07	+0.81	-4.46	+9.91	+19.46	-30.12	+2.72	-7.94	+3.45	-23.76	-27.21		
U <sub>1</sub> E <sub>1</sub>	II	-1.54	+5.94	1.990	-3.08	+11.82	-0.08	-1.37	-3.14	+10.45	+1.25	+0.12	-8.14	-0.77	+2.11	+0.48	+2.59	30.21	60.12
N <sub>1</sub> P <sub>1</sub>	IV	-2.50	+6.55	1.690	-4.98	+13.03	+1.85	-2.59	-3.13	+10.44	+6.37	+4.63	-16.96	-16.22	+6.48	-12.12	-5.64		
U <sub>1</sub> V <sub>1</sub>	III	-4.26	+3.36	0.275	-1.17	+0.92	+0.61	+0.28	-0.56	+1.20	+1.65	-2.60	+0.94	-0.01	-1.19	-2.05	-3.24	23.38	6.43
U <sub>1</sub> Y <sub>1</sub>	IV	-2.50	+6.55	0.275	-0.69	+1.80	+0.12	-0.60	-0.57	+1.20	+1.38	-0.30	-3.93	-2.85	+1.50	-0.79	+0.71		
V <sub>1</sub> W <sub>1</sub>	III	-4.26	+3.36	0.208	-0.80	+0.70	+0.30	-0.13	-0.59	+0.83	+0.82	-1.28	+0.44	-0.32	-0.35	-1.01	+1.56	24.05	5.00
V <sub>1</sub> W <sub>1</sub>	V	-2.88	+6.80	0.208	-0.60	+1.41	+0.01	-0.58	-0.59	+0.83	+1.64	-0.03	-3.94	-2.33	+1.67	-0.07	+1.60		
W <sub>1</sub> J <sub>1</sub>	III	-4.26	+3.36	0.337	-1.44	+1.13	+0.19	+0.08	-1.25	+1.21	+0.18	-0.81	+0.27	-0.38	-0.34	-0.64	-0.98		
W <sub>1</sub> J <sub>1</sub>	IV	-2.88	+6.80	0.337	-0.97	+2.29	-0.23	-1.08	-1.25	+1.21	+3.70	+0.81	-7.32	-2.81	+3.11	+1.60	+5.01		
V <sub>1</sub> F <sub>1</sub>	V	-2.50	+6.55	0.925	-2.06	+5.40	+0.01	-0.90	-2.05	+4.41	+1.58	-0.03	-6.18	-4.93	+2.48	-0.07	+2.41	35.29	29.11
V <sub>1</sub> F <sub>1</sub>	V	-2.88	+6.80	0.825	-2.38	+5.61	+0.32	-1.20	-2.06	+4.41	+2.24	-0.92	-8.16	-6.81	+3.46	-2.18	+1.28		

computed. The first step is to compute the necessary quantities for each line; it will only remain then to combine by simple addition the several lines which go to form the route between the point of reference and the point whose relative probable errors are sought. The reference point will be in general  $A_1$  (Kalianpur), though any other point could also be used. The quantities of scheme (20) have to be formed as was done in tables XLV and XLVI. This has to be done for each section. Typical cases are  $S_1T_1$ ,  $T_1D_1$ ,  $D_1E_1$ . The first  $S_1T_1$  enters only in closing condition of circuit II;  $T_1D_1$  enters in closing conditions of circuits I and II; and  $D_1E_1$  enters into closing condition of circuit II as well as base and Laplace closure between Dehra and Chach, VI. For most of the lines values of  $A$ ,  $AX$ ,  $AY$ ,  $A(R^2 + L^2/12)$  are given in table XLV. For the others which are required the values are now exhibited in table XLVIII. In table XLIX all the quantities necessary for forming  $[uaf] \dots [u/f] \dots$  are given. In the column headed "circuit" the base line closures are indicated, and these require only  $A$ ,  $AX$ ,  $AY$ ,  $AX_0$ ,  $AY_0$ , which are the same quantities as for the circuits. To form  $[u/f]$ ,  $A(R_0^2 + L^2/12)$  is required, where  $R_0$  refers to Kalianpur. This quantity is also given for all lines. It is deduced from values of  $X_0$ ,  $Y_0$ .

As an example consider the probable errors at  $U_1$ , selecting the route  $A_1 B_1 C_1 D_1 E_1 U_1$ . This gives a good example of the method of forming the right hand sides of the equations.

The sections  $A_1B_1$ ,  $B_1C_1$  enter into circuits I and VI  
 ,,  $C_1D_1$  I ,, VII  
 ,,  $D_1E_1$  II ,, VII  
 ,,  $E_1U_1$  II ,, IV

TABLE L. (formed by (21))

Equations	Side closure						Azimuth closure	Easting closure						Northing closure.
	$A_1 B_1$	$B_1 C_1$	$C_1 D_1$	$D_1 E_1$	$E_1 U_1$	Total		$A_1 B_1$	$B_1 C_1$	$C_1 D_1$	$D_1 E_1$	$E_1 U_1$	Total	
I 1	+1.08	+1.11	+0.42			+2.61	0	-0.11	-0.03	-0.16			-0.30	+6.72
I 2	0	0	0			0	+2.61	-1.15	-3.58	-1.99			-6.72	-0.30
I 3	+0.88	+0.98	+0.22			+2.08	+6.80	-4.40	-6.56	-0.87			-11.83	+4.59
I 4	-4.46	-2.16	-0.18			-6.80	+2.08	-0.50	-3.11	-0.98			-4.59	-11.83
II 5				+0.31	+1.99	+2.30	0				-0.37	-3.14	-3.51	+12.15
II 6				0	0	0	+2.30				-1.70	-10.45	-12.15	-3.51
II 7				+0.10	-0.08	+0.02	+1.49				-0.76	-6.77	-7.53	-2.18
II 8				-0.12	-1.37	-1.49	+0.02				-0.41	+2.59	+2.18	-7.53
III 9						0	0						0	0
III 10						0	0						0	0
III 11						0	0						0	0
III 12						0	0						0	0
IV 13					+1.99	+1.99	0						-3.13	+10.44
IV 14					0	0	+1.99						-10.44	-3.13
IV 15					+1.85	+1.85	+2.59						-16.22	+5.64
IV 16					-2.59	-2.59	+1.85						-5.64	-16.22
V 17						0	0						0	0
V 18						0	0						0	0
V 19						0	0						0	0
V 20						0	0						0	0
VI 21	+1.08	+1.11				+2.19	0	-0.11	-0.03				-0.14	+4.73
VII 22			+0.42	+0.31		+0.73	0			-0.16	-0.37		-0.53	+3.69
VIII 23						0	0						0	0
VI 24						0	+2.19	-1.15	-3.58				-4.73	-0.14
VII 25						0	+0.73			-1.99	-1.70		-3.69	-0.53
VIII 26						0	0						0	0

There are four cases according as probable errors of side, azimuth, easting or northing are sought. The scheme (20) may be rewritten, entering the numbers of the column in table XLIX in place of the actual symbolical quantities.

	$S_f$	$A_f$	$E_f$	$N_f$	
$S$	(1)	zero	(6)	(7)	
$A$	zero	(1)	- (7)	(6)	. . . . . (21).
$E$	(4)	- (6)	(11)	- (14)	
$N$	(5)	(4)	(14)	(11)	

For side error  $A_1$   $B_1$  contributes (1), 0, (4) and (5) to equations 1, 2, 3, 4, of circuit I and (1) to equation 21 of VI: to all other equations nothing. The numerical quantities taken from table XLIX are (1) = +1.08, (4) = +0.88, (5) = -4.46. The complete process is shown in table L.

The azimuth closure can at once be written down from the side closure by rearrangement of terms and changing of certain signs in accordance with (21).

8. In a similar way all the quantities occurring on the right hand sides of equations (13) can be formed as required, the necessary data being taken from Table XLIX. The solution of the equations which arise for the N. W. Quadrilateral is effected in the latter portion of the next chapter (VIII). It remains only to refer to the quantities  $[u, ff]$  occurring in equation (14). From (19) it is clear that the necessary quantities for determining these are  $A = M^2L, AX_0, AY_0, A (R_0^2 + L^2/12)$ , the suffix zero indicating  $A_1$ , or Kalianpur as origin. All these quantities are given in table XLIX in columns (1), (6), (7), (15) for each section of line: and the corresponding quantities for a set of lines are obtained by summation of the sectional quantities.

From (2) and (5) the probable errors before adjustment are proportional to  $\sqrt{[u, ff]}$  the multiplying factors being  $1'' \cdot 575$  for azimuth,  $33 \cdot 2$  for 7th place of logarithm of side and  $4 \cdot 03$  for easting and northing in feet: for it is clear that  $[u, ff] = \Sigma M^2L$  in the cases of azimuth and side and  $[u, ff] = M^2L (R_0^2 + L^2/12)$  in the case of easting or northing. By (14) the probable errors after adjustment are proportional to  $[u, ff] - [u, af] k_1 - [u, bf] k_2 \dots$  the multiplying factors being as just given for the several cases. The ratio of probable error after adjustment to probable error before adjustment is  $K$  where

$$K = \sqrt{1 - \frac{[u, af]}{[u, ff]} k_1 - \frac{[u, bf]}{[u, ff]} k_2 - \dots} \quad (22)$$

Numerical values of  $K$  will be given in the next chapter.

## CHAPTER VIII.

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**Numerical values of the probable and actual errors in the Indian triangulation. Note on the solution of linear equations.**


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1. The formulæ (2) and (5) of chapter VII will now be applied to the actual circuits and closures of the Indian triangulation and the numerical results compared with the actual closing errors which have been found in the several circuits. The question is a little complicated by the fact that the triangulation has been adjusted in six portions and so, in the case of some circuits, the closing errors are those due to several series some of which have been adjusted in a neighbouring quadrilateral. In practice this is of little account, as the best quadrilaterals were first adjusted and those adjusted later are of considerably lower quality, so that the probable errors brought in by the adjusted quadrilaterals form only a small part of the total probable error and it is of little account whether the probable errors before or after adjustment are employed. In consideration of this question the fact that the flanking quadrilaterals were previously adjusted will be ignored. What has been said regarding the relative excellence of the several quadrilaterals does not apply to the Burma quadrilateral, which had only been begun when the Indian quadrilaterals were adjusted.

It will be seen from the equations that the quantities required for each line are  $M^2L$  and  $M^2L\left(R^2 + \frac{L^2}{12}\right)$  whence  $L$  is the length of line and  $R$  the distance of its mid point from the closing point of the circuit concerned, both expressed in 100 miles. These quantities have already been taken out (table XLV) for the N.W. quadrilateral. It remains to obtain these for the rest of the triangulation. For this purpose charts III, IV, V are given. These as well as chart II are on the scale 100 miles to an inch: so that  $L$  and  $R$  are the lengths on the charts in inches. See also § 6, chapter VII. The actual measurements and necessary deductions are now shown in table LI. For the Base-line and Laplace closures it is only necessary to compute  $\Sigma M^2L$  along the route. This has already been done for each element of the route which enters into one of the circuits: and only a few remain to be formed for the Laplace closures. The values for the elements are combined to form the necessary values of  $\Sigma M^2L$  for the complete routes. The results are shown in table LII.

Having thus found the probable values of closing errors in all forms of closure, the next step is to compare the results with the closing errors which have actually been found. This is done for the Base-line and Laplace closure in table LII and for the circuit closures in table LIII. In each case the actual error is given and then this is divided by the theoretical error, giving a quantity  $f$ , the ratio of the actual error to the probable error. The actual errors of 7th place of logarithm of side, of azimuth in seconds, of easting and northing in feet are denoted by  $\Delta S$ ,  $\Delta A$ ,  $\Delta E$  and  $\Delta N$  respectively.







The values of  $\Delta A$  in table LII are taken from the second table of chapter IX in which the accumulated errors of azimuth at Laplace stations are found. These are the errors *after* the adjustment of the triangulation has been performed.

TABLE LIII.

	Circuit	$B^2 = \Sigma A$	$33 \cdot 2 B$	$\Delta S$	$1 \cdot 575 B$	$\Delta A$	$D = \sqrt{AC}$	$4 \cdot 03 D$	$\Delta E$	$\Delta N$	$f_s$	$f_a$	$f_c$	$f_n$	Reference number
		$\frac{\Delta S}{33 \cdot 2 B}$	$\frac{\Delta S}{1 \cdot 575 B}$	$\frac{\Delta E}{4 \cdot 03 D}$	$\frac{\Delta N}{4 \cdot 03 D}$										
N.W. Quad.	I	3.661	63.512	+ 68.2	3.013	+ 5.508	40.68	25.703	+ 14.613	+ 57.509	1.074	1.990	0.560	2.237	1
	II	9.713	103.484	- 124.8	4.909	+ 1.560	111.51	42.537	+ 18.296	- 30.413	1.205	0.315	0.430	0.026	2
	III	9.679	103.285	- 79.6	4.900	- 3.254	79.30	35.887	+ 26.065	+ 39.087	0.761	0.664	0.726	1.089	3
	IV	3.426	61.453	+ 150.0	2.915	- 4.232	8.46	11.723	- 21.477	+ 3.638	2.456	1.452	2.088	0.310	4
	V	2.220	49.568	- 5.3	2.351	- 3.000	10.06	13.343	- 24.688	- 0.505	0.107	1.276	1.843	0.038	5
S.E. Quad.	I	2.120	48.339	- 54.9	2.293	+ 0.212	24.23	19.836	- 20.676	+ 5.043	1.664	0.092	1.042	0.254	6
	II	1.112	34.993	+ 31.9	1.660	- 4.968	7.78	11.256	+ 19.623	- 21.785	0.961	2.993	1.726	1.936	7
	III	3.306	60.368	- 17.6	2.863	- 3.888	27.04	20.956	+ 23.073	- 14.322	0.627	1.393	1.101	0.663	8
N.E. Quad.	I	13.202	120.640	+ 287.1	5.724	+ 11.598	123.78	44.834	+ 81.933	+ 94.856	2.360	2.026	1.628	2.118	9
	II	21.630	154.413	- 7.3	7.255	- 14.317	68.10	11.961	- 96.904	- 35.551	0.047	1.056	8.102	2.972	10
	III	19.677	147.275	+ 482.3	6.987	- 6.798	62.63	31.893	- 16.380	+ 93.415	3.275	0.973	0.513	2.929	11
	IV	14.681	127.222	- 641.0	6.035	+ 10.177	36.52	24.363	- 5.388	- 122.087	5.038	1.688	0.221	5.013	12
	V	9.601	102.654	+ 169.4	4.679	- 0.791	21.05	18.490	- 21.320	+ 38.675	1.647	0.162	1.154	2.092	13
	VI	12.660	118.192	+ 82.1	5.607	+ 4.843	24.84	20.086	+ 9.377	+ 21.306	0.695	0.864	0.467	1.061	14
	VII	11.539	112.780	- 132.0	5.350	- 5.739	23.13	19.380	- 2.710	- 36.047	1.178	1.073	0.139	1.860	15
	VIII	5.931	80.875	+ 196.7	3.637	+ 4.441	20.44	18.220	- 2.449	+ 30.193	2.432	1.157	0.131	1.679	16
	IX	7.495	90.902	- 284.6	4.312	- 2.279	11.96	13.936	+ 3.269	+ 31.703	3.130	0.528	0.234	2.275	17
	X	15.323	120.678	+ 190.9	6.168	+ 0.848	37.00	24.800	- 10.284	- 21.600	1.468	0.105	0.414	0.871	18
	XI	11.923	114.640	- 193.7	5.438	+ 2.403	29.81	22.004	- 13.044	+ 13.023	1.690	0.443	0.592	0.592	19
	XII	2.362	51.028	+ 102.0	2.421	- 13.140	4.72	6.737	- 57.230	+ 17.465	2.011	5.429	6.542	1.995	20
S. Trigon.	I	0.900	104.447	+ 136.6	4.955	- 0.251	18.00	17.090	+ 7.624	+ 2.218	1.308	0.051	0.446	0.130	21
	II	2.591	53.462	- 22.7	2.538	- 3.680	24.68	20.102	+ 28.479	- 0.101	0.425	1.451	1.416	0.005	22
	III	1.332	38.313	+ 39.9	1.818	+ 4.303	6.27	10.091	- 17.218	0.000	1.041	2.367	1.706	0.000	23
	IV	1.091	34.661	- 0.4	1.644	- 9.040	3.39	7.411	+ 42.500	- 27.515	0.012	5.500	5.794	3.712	24
	V	0.769	28.917	+ 11.7	1.372	- 0.319	1.35	4.683	- 0.973	- 4.233	0.404	0.233	0.208	0.904	25
S.W. Quad.	I	7.677	91.997	+ 189.8	4.364	- 7.806	48.44	23.049	- 54.726	- 64.255	2.063	1.789	1.952	2.291	26
	II	3.989	66.300	- 212.5	3.145	- 7.113	24.8	6.218	+ 38.058	+ 6.862	3.207	2.260	6.121	1.104	27
	III	9.162	100.496	+ 27.4	4.708	- 6.719	35.86	24.132	+ 6.238	- 43.878	0.261	1.408	0.258	1.818	28
	IV	2.834	56.872	- 185.9	2.698	+ 2.057	1.74	5.320	- 1.915	+ 15.238	3.269	1.006	0.228	2.668	29
	V	4.027	66.842	- 62.2	3.101	+ 3.048	3.30	7.323	- 6.712	+ 11.028	6.934	0.964	0.916	1.016	30
	VI	6.536	84.892	+ 257.1	4.027	- 3.912	18.03	17.536	- 23.569	+ 6.057	3.030	0.972	1.344	0.346	31
Burma Quad.	I	2.032	47.343	+ 45.0	2.246	- 2.267	19.69	17.881	- 47.521	+ 41.790	0.951	1.009	2.658	2.337	32
	II	3.405	61.254	+ 47.0	2.906	- 6.002	14.27	12.543	- 17.550	- 17.563	0.767	2.065	1.399	1.399	33
	III	6.103	82.037	- 189.0	3.892	+ 3.606	4.24	6.836	+ 2.914	+ 1.427	2.304	0.927	0.426	0.269	34

2. It is not to be expected that the actual errors will be the same as the probable errors: but in a considerable number of cases, values of the ratio of actual to probable errors, that is  $f$ , should be distributed according to the laws of probability. In a given number of cases the probability is that those values of  $f$  which are comprised within certain limits will form a certain percentage of the total cases. The probability integral, between the proper ordinates, represents this distribution of errors. It is tabulated in most books on minimum squares\*. By means of this it is seen that 10 per cent of the errors will most probably fall in each of the regions  $A, B, \dots, J$  of table LIV defined by limiting values of  $f$ .

The actual values of  $f$  found in 167 cases of closures are classified in these columns. Each value of  $f$  is followed by a number in brackets which refers to the corresponding closing condition in tables LII or LIII.

\* Vide Wright's "Adjustment of Observations", § 213.

TABLE LIV.

Values of <i>f</i> from to	A 0 ·185	B ·185 ·375	C ·375 ·572	D ·572 ·777	E ·777 1·000	F 1·000 1·249	G 1·240 1·539	H 1·539 1·900	I 1·900 2·438	J 2·438 ∞
<b>Side</b>										
(i) Circuit ...	·012 (24) ·047 (10) ·107 (5)	·261 (28)	·425 (22) ·404 (25) ·527 (8)	·695 (14) ·761 (3) ·767 (33)	·934 (30) ·951 (32) ·961 (7)	1·041 (23) 1·074 (1) 1·178 (16) 1·205 (2)	1·308 (21) 1·468 (18)	1·647 (13) 1·654 (6) 1·690 (19)	2·011 (20) 2·063 (26) 2·304 (34) 2·380 (9) 2·432 (16)	2·458 (4) 3·080 (31) 3·130 (17) 3·207 (27) 3·269 (29) 3·275 (11) 5·038 (12) 7
(ii) Base ...	3 ·017 (56) ·051 (54) ·137 (43) 3	1 ·347 (46)	3 ·408 (44)	3 ·614 (47)	3 ·893 (35) ·013 (49) ·953 (41) 3	4 1·146 (5)	2 1·283 (52)	3 1·574 (37)	5 1·080 (36)	1 4·210 (40)
<b>Azimuth</b>										
(i) Circuit ...	·051 (21) ·092 (6) ·104 (18) ·162 (13)	·233 (25) ·315 (2)	·442 (10) ·528 (17)	·684 (3)	·864 (14) ·927 (34) ·964 (30) ·973 (31) ·973 (11)	1·009 (32) 1·073 (15) 1·096 (20) 1·157 (16)	1·276 (5) 1·393 (8) 1·408 (23) 1·451 (22) 1·452 (4)	1·686 (12) 1·789 (26)	1·956 (10) 1·960 (1) 2·026 (9) 2·065 (33) 2·260 (27) 2·367 (23)	2·903 (7) 5·428 (20) 5·500 (24) 2·629 (38) 2·730 (41) 2·923 (54) 3·259 (51) 6·125 (59)
(ii) Laplace ...	4	2 ·214 (35) ·223 (59)	2	1 ·733 (57)	5 ·792 (55)	4 1·106 (30)	5 1·258 (60) 1·447 (48)	2 1·774 (46) 1·872 (43)	6 1·917 (42) 1·970 (53) 2·179 (60)	3 2·629 (38) 2·730 (41) 2·923 (54) 3·259 (51) 6·125 (59)
<b>Easting</b>	·134 (16) ·139 (15)	·208 (25) ·221 (12) ·229 (20) ·234 (17) ·258 (28)	·414 (18) ·426 (34) ·430 (2)	·592 (19) ·728 (3)	·918 (30)	1·042 (6) 1·101 (8) 1·154 (13)	1·344 (31) 1·399 (33) 1·416 (22)	1·706 (23) 1·726 (7) 1·823 (9) 1·843 (5)	1·952 (26) 2·098 (4)	2·653 (32) 5·734 (24) 6·121 (27) 6·542 (20) 8·102 (10)
<b>Northing</b>	2 ·000 (33) ·005 (22) ·038 (5) ·130 (21) 4	5 ·209 (34) ·254 (6) ·310 (4) ·346 (31) 4	7 ·446 (21) ·467 (14) ·513 (11) ·569 (1)	2 ·592 (19) ·683 (8)	1 ·871 (18) ·904 (25) ·926 (2)	3 1·061 (14) 1·089 (3) 1·104 (27)	3 1·399 (33)	4 1·679 (16) 1·818 (28) 1·860 (15)	2 1·916 (30) 1·936 (7) 1·995 (20) 2·092 (13) 2·118 (9) 2·237 (1) 2·275 (17) 2·291 (26) 2·337 (32) 9	5 2·865 (29) 2·929 (11) 2·972 (10) 3·712 (24) 5·013 (12)
<b>Side and Azimuth excluding Laplace</b>	10	4	6	5	11	9	8	6	12	11
<b>Northing and Easting</b>	8	9	7	4	4	6	4	7	11	10
<b>All except Laplace</b>	18	13	13	9	15	15	12	13	23	21
			66					84		

3. On examination of table LIV it is immediately noticeable that the cases in which the actual error is greater than the computed probable error are more numerous than the cases when it is less. Considering all the 167 cases, *f* is less than unity for 70 cases and greater than unity for the remaining 97 cases. This inequality is largely attributable to the Laplace closures,

although in this case values of azimuth, adjusted for all circuit conditions, have been used. In the Laplace closures there are 13 cases of actual error greater than computed probable error and only 4 cases of actual error less than computed error. It is believed that the explanation of this is that given in § 6, Chapter V. The error due to acceptance of geoidal angles uncorrected, instead of spheroidal angles, to some extent magnifies the triangular error and so increases the value of  $M$ , and as a result the closures of azimuth in circuits and of the deduced quantities side, northing and easting are in better agreement with the formulæ than the closures on Laplace points; since the former do not depend on absolute errors while the latter do. If the Laplace closures are ignored the number of cases, less than and greater than the formulæ give, are 66 and 84 respectively. If the formulæ values were increased in the ratio 1.1 to 1.0\* the figures would become 74 and 76 respectively. It appears then that the formulæ give values of the probable error which are some ten per cent below what the 150 cases would lead to. This is not a serious deviation from the facts and, apart from mere chance, may be attributed partly to

- (1) the use of geoidal instead of spheroidal angles.
- (2) the fact that  $M$  is based on certain simplifying assumptions regarding the regularity of the triangles and polygons in the series of triangulation\*.

The total number of cases falling in each class A, ... J is shown at the bottom of the table and from this it is seen that the errors are fairly distributed in the various classes, except that in classes I, J a considerable excess of cases occur. The excess of large errors over the number which is given by the formulæ, viz.  $45 - 30 = 15$  or  $50\%$ , in classes I, J is to be attributed to the neglect of certain sources of error. One such source of error is that, already mentioned, of treating geoidal and spheroidal angles as identical: and it may be that other undetected sources also exist. However the formulæ give practically a satisfactory indication of the probable accuracy of side, azimuth, easting or northing. They should be a useful guide to the care which ought to be expended on observing and selecting a series in order that a result of any stated precision may be arrived at. As work on such a series progresses, the value of  $M$  may be taken out and observations increased in number, or rays increased in length, until the value of  $M$  is reduced to a quantity sufficiently small to give the proper precision.

4. It has just (June 1917) been noticed that the question of probable errors of side, azimuth, easting and northing generated in a chain of triangles were considered by General Walker and Mr. W. H. Cole in 1882†. The deduction is based on the equations by which the simultaneous reduction of the triangulation of India had been effected, and the equations obtained—*vide* xxviii, xxix, xxx *ibid*—are somewhat complex. These equations are comparable with (2) and (5) of Chapter VII of this work. Dealing with the case of a simple chain of equilateral triangles (on p. 104) with sides of 15 miles and chain of length  $8^\circ$  of arc, it is found in the Appendix that the

$$\begin{aligned} \text{e. m. s. (i. e. } \frac{\text{probable error}}{\cdot 6745} \text{) in azimuth} &= 6'' \cdot 93\epsilon, \text{ average value} \\ \text{latitude} &= 0'' \cdot 55\epsilon \\ \text{longitude} &= 0'' \cdot 59\epsilon \end{aligned}$$

the first quantity being somewhat dependent on the direction of the chain. It appears that  $\epsilon$  is the quantity now denoted by  $m$ . To obtain results by the method of present work put  $M =$

$$\frac{7}{8} m \sqrt{\frac{18}{15}} = 1 \cdot 278m. \text{ Then taking } 8^\circ \text{ as equivalent to 550 miles,}$$

$$\text{Probable error in azimuth} = 1 \cdot 575 \times 1 \cdot 278\epsilon \sqrt{5 \cdot 5} = 4'' \cdot 72\epsilon$$

$$\text{Mean error in azimuth} = \frac{4 \cdot 72}{\cdot 6745} \epsilon = 6'' \cdot 99\epsilon$$

$$\text{Probable error in easting or northing} = 4 \cdot 03 \times 1 \cdot 278\epsilon \sqrt{5 \cdot 5} \times \frac{5 \cdot 5}{\sqrt{3}} \text{ feet} = 38 \cdot 3\epsilon \text{ feet.}$$

\* See also § 4 below.

† *Vide* G.T.S. Vol. VII, Appendix No. 3.

$$\text{Mean error in easting or northing} = \frac{38 \cdot 3\epsilon}{\cdot 6745} = 56 \cdot 8\epsilon \text{ feet.}$$

These results are in accord with those found by the old formulæ. On pp. 105, 106 of the Appendix\* additional quantities were introduced to take account of geometrical irregularity, double instead of single chains, length of side. The latter two considerations have been dealt with in the present work by use of the quantity *M*. In the appendix under reference geometrical irregularity of the magnitude which might occur in Survey of India work is represented by an augmenting factor  $\kappa$  and it is stated that "we may as a rule put  $\kappa = 1 \cdot 4$  in hilly country and  $\kappa = 1 \cdot 1$  in the plains". The introduction of this factor would increase most of the probable errors computed in the present work in ratio 1·1, an amount which would practically equalise the number of cases of errors exceeding and falling short of the probable error as already deduced in § 3. The independent opinion of the author before seeing this appendix was that it was better to leave this out of account: but it is a question as to whether the factor  $\kappa$  might not with advantage be incorporated in *M* in some cases.

The formulæ developed in the Appendix\* do not appear to have been put to much use: and as far as can be seen were lost sight of. They are only applicable to straight (or approximately straight) chains of triangles, and not to circuits of all forms.

**Note on the solution of equations.**

5. In the adjustment of triangulation, and the calculation of its probable errors, groups of linear equations involving a large number of unknowns frequently arise. Although the solution is not necessarily required to a high order of accuracy, yet the work of elimination has generally of necessity been performed using a large number of significant figures to safeguard the solution against accumulation of computation inaccuracy. Some of the multipliers in the process of elimination become very large owing to the fact that the denominators consist of terms of which the positive and negative portions are not very different in amount. Taking the denominator to be of the form  $\Sigma a - \Sigma \beta$  where *a*,  $\beta$  represent the positive and negative terms respectively, while  $\Sigma a$  and  $\Sigma \beta$  may each be formed to a fairly high percentage accuracy, the quantity  $\Sigma a - \Sigma \beta$  may be inaccurate by a considerable percentage.

If the ordinary Gaussian method of arranging the elimination is followed, it is to be noted that most of the solution is independent of the R. H. S.† of the equations, and is in fact just a process of elimination of the several unknowns. In some cases, *e. g.* that occurring in Chapter VII, solutions are required for a number of sets of values of the R. H. S.† It is accordingly in this case desirable to retain the R. H. S.† in symbolical form. But this has an advantage of an entirely different kind, as it permits of any number of successive approximations in the solution being made, without repeating the eliminating process, which accordingly need not be performed with such exactness as would otherwise be necessary.

First consider the L. H. S.‡ of the equations. Following Gauss's method of arrangement denote the equations by

$$\left. \begin{array}{l} a_1 x_1 + a_2 x_2 + \dots + a_n x_n = A \\ b_1 x_1 + b_2 x_2 + \dots \phantom{+ a_n x_n} = B \\ \dots \phantom{+ a_1 x_1 + a_2 x_2 + \dots} \phantom{+ a_n x_n} \\ n_1 x_1 + n_2 x_2 + \dots \phantom{+ a_n x_n} = N \end{array} \right\} \dots \dots \dots (1)$$

If the first equation is multiplied by  $-\frac{b_1}{a_1}$  and added to the second,  $x_1$  is eliminated. Similarly

\* *Vide* O.T.S. Vol. VII, Appendix No. 3. † R.H.S. = right hand side. ‡ L.H.S. = left hand side.

if it is multiplied by  $-\frac{r_1}{a_1}$  and added to the  $r$ th equation  $x_1$  is eliminated. The following equations are formed

$$\left. \begin{aligned} \left( b_2 - \frac{b_1}{a_1} a_2 \right) x_2 + \left( b_3 - \frac{b_1}{a_1} a_3 \right) x_3 + \dots &= B - \frac{b_1}{a_1} A \\ \left( r_2 - \frac{r_1}{a_1} a_2 \right) x_2 + \left( r_3 - \frac{r_1}{a_1} a_3 \right) x_3 + \dots &= R - \frac{r_1}{a_1} A \end{aligned} \right\} \dots (2)$$

To eliminate  $x_2$  the same process is applied, the multipliers in this case all having the denominator  $b_2 - \frac{b_1}{a_1} a_2$ . It is to be observed that the denominator of the multiplying factors is always the first coefficient of the first equation of the set being operated on. The successive denominators are  $a_1, b_2 - \frac{b_1}{a_1} a_2, c_3 - \frac{c_1}{a_1} a_3 - \frac{c_2 - \frac{c_1}{a_1} a_2}{b_2 - \frac{b_1}{a_1} a_2} \left( b_3 - \frac{b_1}{a_1} a_3 \right)$  etc.

In the solution of normal equations the diagonal coefficients are generally larger than the others. Being of the forms  $\sum u a^2$  and  $\sum u a b$  respectively the component parts of the first form are all positive while those of the second form are equally likely to be positive and negative, and accordingly tend to cancel. Accordingly in more cases than not  $a_1 > a_r$  ( $r \neq 1$ ),  $b_2 > b_r$  ( $r \neq 2$ ) . . . so that  $b_2 - \frac{b_1}{a_1} a_2$  is not likely to be small compared with  $b_2$ . But as the denominators become more complex there is more possibility of their becoming small. To avoid this, as well as may be foreseen before the actual computations are carried out, it is accordingly convenient to rearrange the equations in such order that the diagonal coefficients are of increasing magnitude. Before doing this however it is desirable that all these diagonal coefficients should be brought up to as near as may be the same order of magnitude. It is inconvenient to have quantities entering the computation, some with many figures preceding the decimal point while in others there are no figures before the decimal point and a number of zeros following the decimal point.

Any coefficient say  $f_r$  can be changed to  $10^a f_r$  or  $\frac{1}{10^a} f_r$  if at the same time for  $x_r$  is written  $\frac{1}{10^a} x_r$  or  $10 x_r$ , this being done in all the equations; and solution subsequently being performed for  $\frac{1}{10^a} x_r$  or  $10 x_r$  as the case may be. It is convenient to use as multiplier a power of ten, as this involves no loss of precision in the coefficient and no labour in transforming it. If the process thus suggested is carried out in the case of symmetrical equations—such as normal equations—the symmetry is destroyed. As the symmetry is advantageous it is desirable to avoid destroying it, and this may be arranged for as follows. When any column is multiplied by a power of ten the corresponding row should be multiplied by the same quantity. Then the equations (1) may be written

$$\left. \begin{aligned} a_1 x_1 + a_2 x_2 + \dots + 10^a a_r X_r + \dots &= A \\ 10^a r_1 x_1 + 10^a r_2 x_2 + \dots + 10^{2a} r_r X_r &= R \cdot 10 \end{aligned} \right\} \dots (3)$$

in which  $X_r = \frac{x_r}{10^a}$  which are quite symmetrical.

In the solution of (1) the quantities which should be brought up to about the same order are the diagonal coefficients, as these enter more than the others into the computations. It may be seen from (3) that the diagonal coefficients can only be conveniently changed by powers (positive or

negative) of  $10^2$ . The first step then is to apply this process of dealing with the diagonal coefficients; the second is to rearrange the equations in such order that the modified diagonal coefficients occur in increasing order of magnitude. It can always be arranged that the largest diagonal coefficient is not so much as one hundred times the smallest.

6. Suppose now that (1) represents a set of equations dealt with and arranged as explained above. They are now in an order as favourable for solution as can be arranged for by mere inspection.

In proceeding with the elimination as explained in §5, notice that the R. H. S. quantities  $A, B, C$  do not in any way affect the elimination. The work on the left hand side of the equation is entirely independent of the values  $A, B, C$ . . . . . Suppose then that a solution has been taken out, which is essentially only approximate: the degree of approximation depends on the number of figures retained in the arithmetical processes. Instead of the correct quantities  $x_1, x_2$ . . . this solution will determine slightly different quantities  $x_1 - \delta x_1, x_2 - \delta x_2$ . On substituting these in the L.H.S of (1) the values found will be  $A - \delta A, B - \delta B$ . . . . . Hence

$$\left. \begin{aligned} a_1 \delta x_1 + a_2 \delta x_2 + \dots &= \delta A \\ b_2 \delta x_1 + \dots &= \delta B \end{aligned} \right\} \dots \dots \dots (4)$$

These equations have the same coefficients on the L.H.S. as those of the original equations (1). Hence the process of elimination in order to determine  $\delta x_1, \delta x_2$ . . . is the same as that already performed for  $x_1, x_2$ . . . : and it is only necessary to change the portion of the computation involving  $A, B, C$ . In this way a second approximation is easily arrived at. Clearly this result may again be treated in the same way, and so successive sets of higher approximation may be obtained, without any necessity of increasing the accuracy of the elimination process. The gain in accuracy is arrived at by means of accurate substitution, and this substitution may be made absolutely perfect by keeping all the figures resulting from the substitution. As an example, if the original solution consists of numbers of 4 significant figures, and the coefficients are given to 4 figures the products will consist of 7 or 8 significant figures. It may be pointed out here that if the coefficients are given to much higher accuracy it is not necessary to use the full amount of figures for the elimination process: but this must be done in the substitution.

The upshot of this is that the various multiplications and divisions which arise in the process of elimination may all be performed by slide rule, which greatly facilitates the work. The substitutions must be carried out with higher, or even absolute, accuracy, as can most conveniently be done by arithmometer.

7. As remarked in §5 it is sometimes convenient to have the solution in terms of symbolic values of the R.H.S. This gives rise to another general method of procedure as regards higher approximation to any desired degree, which will now be described. Other attendant advantages will be seen to arise in this method.

Suppose the solution of (1) is expressed in the form

$$\left. \begin{aligned} x_1 &= a_1 x_1 A + b_1 x_1 B + c_1 x_1 C + \dots \\ &\dots \dots \dots \\ x_r &= a_r x_r A + b_r x_r B + c_r x_r C + \dots \end{aligned} \right\} \dots \dots \dots (5)$$

Then  $a_1 x_1, a_2 x_2$ . . . are the solutions of

$$\left. \begin{aligned} a_1 x_1 + a_2 x_2 + \dots &= 1 \\ b_1 x_1 + b_2 x_2 + \dots &= 0 \\ r_1 x_1 + r_2 x_2 + \dots &= 0 \end{aligned} \right\} .$$



and  $r_1x_1, r_2x_2 \dots$  are the solutions of

$$\left. \begin{aligned} a_1x_1 + a_2x_2 & \dots = 0 \\ \dots & \dots = 0 \\ r_1x_1 + r_2x_2 & \dots = 1 \\ \dots & \dots = 0 \end{aligned} \right\} \dots \dots \dots (6)$$

In (6) all the quantities on the R.H.S. are zero, except the  $r$ th, which is unity. The elimination processes are identical for all values of  $r$ , but the work on the R.H.S. differs for each value of  $r$ . The accurate solution of all the sets of form (6) gives values of all the quantities  $r_x$ : but any actual solution will give quantities slightly different, viz.  $r_x - \delta_r x_x$ . On substituting these, instead of getting the quantities unity and zero as values of the R.H.S. slightly different quantities are obtained, as indicated.

*Values of R.H.S. of equations (6) resulting from an approximate solution.*

Values of R.H.S.				Approximate solutions.			
	1	2	3 . . . . .	1	2	3	
1	$1 + a_1$	$a_2$	$a_3 \dots$	$a''_1$	$a''_2$	$a''_3$	
2	$\beta_1$	$1 + \beta_2$	$\beta_3 \dots$	$b''_1$	$b''_2$	$b''_3$	
3	$\gamma_1$	$\gamma_2$	$1 + \gamma_3 \dots$	$c''_1$	$c''_2$	$c''_3$	
	$\dots$	$\dots$	$\dots$	$\dots$	$\dots$	$\dots$	$\dots$

If the original equations are symmetrical, so also are the quantities of the approximate solution.

It is clear that if the approximate calculation has been properly carried out all the quantities  $a, \beta, \gamma \dots$  are small compared with unity.

8. Suppose the solutions 1, 2, 3 . . . are combined in any way, taking for  $x_1$  the value  $Aa''_1 + Bb''_2 + \dots$  and similarly for the other  $x$ 's. Then it is clear that the corresponding values of the R. H. S. will be

$$\begin{aligned} A(1 + a_1) + Ba_2 + Ca_3 & \dots, \\ A\beta_1 + B(1 + \beta_2) + C\beta_3 & \dots, \\ A\gamma_1 + B\gamma_2 + C(1 + \gamma_3) & + \dots, \end{aligned}$$

Putting  $A = 1 - a_1 \dots B = -\beta_1 \dots C = -\gamma_1 \dots$  these become

$$\begin{aligned} 1 - a_1^2 - \beta_1 a_2 - \gamma_1 a_3 & \dots, \\ - a_1 \beta_1 - \beta_1 \beta_2 - \gamma_1 \beta_3 & \dots, \end{aligned}$$

Only products of the small quantities  $a, \beta \dots$  now occur except in the quantity unity in the first line: so that a higher approximation is readily found in this way. The process can obviously be repeated as often as is desirable.

9. The question can also be considered otherwise. Suppose the true value of any quantity  $x$  is  $u-v$ ,  $u$  being a value obtained by solution and  $v$  a small correction. Then the solutions of equations

$$\begin{array}{l}
 a_1 x_1 + a_2 x_2 + \dots \\
 b_1 x_1 + b_2 x_2 + \\
 \dots
 \end{array}
 =
 \begin{array}{|c|c|c|}
 \hline
 A & B & C \\
 \hline
 1 & 0 & 0 \\
 \hline
 0 & 1 & 0 \\
 \hline
 \dots & \dots & \dots \\
 \hline
 \end{array}
 \text{ are }
 \begin{array}{|c|c|c|}
 \hline
 A & B & C \\
 \hline
 a x_1 & b x_1 & c x_1 \\
 \hline
 a x_2 & b x_2 & c x_2 \\
 \hline
 \dots & \dots & \dots \\
 \hline
 \end{array}
 \quad (7)$$

and as above

$$\begin{array}{l}
 a_1 v_1 + a_2 v_2 + \dots \\
 b_1 v_1 + b_2 v_2 + \dots \\
 \dots
 \end{array}
 =
 \begin{array}{|c|c|c|}
 \hline
 A & B & C \\
 \hline
 a_1 & a_2 & a_3 \\
 \hline
 \beta_1 & \beta_2 & \beta_3 \\
 \hline
 \dots & \dots & \dots \\
 \hline
 \end{array}$$

Whence by means of (7) the solutions are represented by

$$\left. \begin{array}{l}
 {}_s v_1 = a_s a x_1 + \beta_s b x_1 + \gamma_s c x_1 + \dots \\
 {}_s v_2 = a_s a x_2 + \beta_s b x_2 + \gamma_s c x_2 + \dots \\
 {}_s v_3 = a_s a x_3 + \beta_s b x_3 + \gamma_s c x_3 + \dots
 \end{array} \right\} \dots \quad (8)$$

Hence

$${}_s v_r = a_s a x_r + \beta_s b x_r + \gamma_s c x_r + \dots \quad (9)$$

The quantities  ${}_s x_r$  etc. may as a first approximation be replaced by the determined quantities  ${}_a u_r$  etc. The solution then takes the form

$${}_s u'_r = {}_s u_r - a_s a u_r - \beta_s b u_r \dots \quad (10)$$

and if further approximation is required this value may be substituted in (8) for  ${}_a x_r$  giving the next approximation to  ${}_s v_r$ . Any number of successive approximations may be made in this manner. The *R. H. S.* corresponding to (9) may be written down. They are

$$\begin{array}{l}
 1 + a_1 - a_1(1 + a_1) - a_2 \beta_1 - a_3 \gamma_1 \dots, \quad a_2 - a_2(1 + a_1) - a_2 \beta_2 - a_3 \gamma_2 \dots, \\
 \beta_1 - \beta_1(1 + a_1) - \beta_2 \beta_1 - \beta_3 \gamma_1 \dots, \quad 1 + \beta_2 - a_2 \beta_1 - \beta_2(1 + \beta_2) - \gamma_2 \beta_3 \dots, \\
 \dots
 \end{array}$$

or

$$\begin{array}{l}
 1 - a_1^2 - a_2 \beta_1 - a_3 \gamma_1 \dots, \quad -a_1 a_2 - a_2 \beta_2 - a_3 \gamma_2 \dots, \\
 -\beta_1 a_1 - \beta_2 \beta_1 - \beta_3 \gamma_1 \dots, \quad 1 - \beta_1 a_2 - \beta_2^2 - \beta_3 \gamma_2 \dots,
 \end{array}$$

which may be written with abbreviated form

$$\begin{array}{l}
 1 + a'_1 \quad a'_2 \quad a'_3 \dots \\
 \beta'_1 \quad 1 + \beta'_2 \quad \beta'_3 \dots \\
 \dots
 \end{array}$$

where

$$a'_1 = -a_1^2 - a_2 \beta_1 - a_3 \gamma_1, \dots \text{ etc.}$$

These residuals  $\alpha', \beta'$  etc., are composed of binary products of  $\alpha, \beta$  etc. and accordingly are of a smaller order than  $\alpha, \beta \dots$ . Starting with the second approximate values  $\alpha', \beta'$  and these residuals  $\alpha', \beta'$  another approximation may be made: and so on as far as is necessary to attain the accuracy of solution desired.

10. It may at times be convenient to split up a set of equations and apply the above process to any portion consisting of an equal number of rows and columns. This may be done with advantage when a considerable number of the coefficients are zero. It is to be remembered that the actual numerical labour of solving  $n$  equations varies as the cube of  $n$ , so that a group of say 30 equations presents a formidable piece of computation. By the method now proposed perhaps this labour may be considerably reduced: but one certain advantage is a substitution check at a comparatively early stage of the computation. This is a check against actual computation blunders, as well as against accumulation of error due to lack of absolute exactness in the calculation on account of the necessity of limiting the number of figures employed. In this way, as has been shown above, much of the work can be performed readily with a slide rule, greatly accelerating the work.

In a large class of equations, *e.g.* the normal equations which occur in the method of least squares, there is complete symmetry about a diagonal. This reduces the work of elimination to about one-half. It is important then that in dealing with equations of this class that the symmetry should be preserved. Gauss's arrangement secures this for his method of solution, and it will now be shown that symmetry is maintained when the equations are split up as just suggested.

Denote the equations by

$$\left. \begin{aligned} (1,1) k_1 + (1,2) k_2 + \dots + (1,n) k_n &= (i) \\ (2,1) k_1 + (2,2) k_2 + \dots + (2,n) k_n &= (ii) \\ \dots &\dots \\ (n,1) k_1 + (n,2) k_2 + \dots + (n,n) k_n &= (n) \end{aligned} \right\} \dots \dots \dots (11)$$

Suppose the solution of

$$\left. \begin{aligned} (1,1) k_1 + (1,2) k_2 \dots + (1,r) k_r &= [i] \\ (2,1) k_1 + (2,2) k_2 \dots + (2,r) k_r &= [ii] \\ \dots &\dots \\ (r,1) k_1 + (r,2) k_2 \dots + (r,r) k_r &= [r] \end{aligned} \right\} \dots \dots \dots (12)$$

where  $r$  is less than  $n$ , is

$$\left. \begin{aligned} k_1 &= {}_1k_1 [i] + {}_2k_1 [ii] + \dots + {}_rk_1 [r] \\ k_2 &= {}_1k_2 [i] + \dots \\ \dots &\dots \\ k_r &= {}_1k_r [i] + \dots + {}_rk_r [r] \end{aligned} \right\}^* \dots \dots \dots (13)$$

Then from the first  $r$  equations of (11) it is seen that

$$\left. \begin{aligned} [i] &= (i) - (1,r+1) k_{r+1} - (1,r+2) k_{r+2} - \dots - (1,n) k_n \\ [ii] &= (ii) - (2,r+1) k_{r+1} - (2,r+2) k_{r+2} - \dots - (2,n) k_n \\ \dots &\dots \\ [r] &= (r) - (r,r+1) k_{r+1} - (r,r+2) k_{r+2} - \dots - (r,n) k_n \end{aligned} \right\} \dots \dots \dots (14)$$

Substituting from (14) in (13) it follows that

---

\* This may be a first or higher approximation as appears most suitable.

$$\left. \begin{aligned}
 k_1 &= {}_1k_1 (i) + {}_2k_1 (ii) + \dots + {}_rk_1 (r) \\
 &\quad - k_{r+1} \sum {}_sk_1 (s, r+1) - k_{r+2} \sum {}_sk_1 (s, r+2) - \dots - k_n \sum {}_sk_1 (s, n) \\
 k_t &= {}_1k_t (i) + {}_2k_t (ii) + \dots + {}_rk_t (r) \\
 &\quad - k_{r+1} \sum {}_sk_t (s, r+1) - k_{r+2} \sum {}_sk_t (s, r+2) \dots - k_n \sum {}_sk_t (s, n)
 \end{aligned} \right\} \dots (15)$$

the summation indicated by  $\Sigma$  referring to  $s$  to which all values from 1 to  $r$  are to be given.

The values  $k_1, k_2, \dots, k_r$  are now to be substituted in the latter  $n - r$  equations of (11). It is clear that thereby the coefficients of  $k_{r+1}, k_{r+2}, \dots, k_n$  are altered. The original coefficients are clearly symmetrical, and it is only necessary to show that the change in the coefficient of  $k_u$  in the  $r + 1^{th}$  equation—the first of the equations dealt with—is the same as the change in the coefficient of  $k_{r+1}$  in the  $u^{th}$  equation. By (15) and (11) it is seen that the change in coefficient of  $k_u$  in the  $r + 1^{th}$  equation is

$$-(r+1, 1) \sum {}_sk_1 (s, u) - (r+1, 2) \sum {}_sk_2 (s, u) - \dots - (r+1, r) \sum {}_sk_r (s, u) \dots (16)$$

while the change in the coefficient of  $k_{r+1}$  in the  $u^{th}$  equation is

$$-(u, 1) \sum {}_sk_1 (s, r+1) - (u, 2) \sum {}_sk_2 (s, r+1) - \dots - (u, r) \sum {}_sk_r (s, r+1) \dots (17)$$

Remembering that  ${}_sk_t = {}_tk_s$ , notice that the sum of the coefficients of  ${}_sk_t$  and  ${}_tk_s$  in (16) is  $-(r+1, t) (s, u) - (r+1, s) (t, u)$  and the corresponding quantity in (17) is  $-(u, t) (s, r+1) - (u, s) (t, r+1)$ . These quantities are the same, since  $(u, t) = (t, u)$  etc.

The equations resulting from this method of solution are accordingly symmetrical.

11. In some cases after the solution of a set of normal equations has been effected, additional conditions may have to be introduced. The form of the equations, *vide* (13), is only modified thereby by the addition of a number of terms at the end of the original equations and an addition of the same number of equations at the end. If then the original equations have been solved in the manner explained above, it is possible to proceed immediately to derive the solution of the larger number of equations, making use of the solution already obtained. If the ordinary method of solution of the original equations had been followed this would have been of little help in proceeding to the solution of the larger number of equations.

These methods will now be given effect to in the solution of the 26 equations of p. 116. In table XLVII the coefficients are marked off in the stages for which solution will be performed.

In cases where a highly accurate solution is not desired the values of the residuals are not required. It is however of importance to verify that the solution does not contain any blunders, as may easily occur in the numerical work. A check on this is obtained by substituting the values obtained for solution A, B, . . . in the last equation. This equation only enters into the final eliminant from which the value of the last unknown is determined, and not into any of the previous equations used for the actual solution—that is the first of each group of equations formed by successive elimination of the first, second, . . . unknowns. If then this last equation is satisfied with satisfactory precision it is an indication that no blunder has been committed. It is not certain from this that the residuals of the other equations are equally small, and nothing short of substitution in each of these will make this point quite clear: but it is a sure check against any serious blunder.

12. Application of the method suggested above will now be made to the equation whose solution is necessary for the determination of the several probable errors of the N.W. Quadrilateral after adjustment. The L.H.S. of the equations are indicated in table XLVII. Conformably with §10 only a portion of the complete set is dealt with at first. It is at once clear that the first 8 equations and the next 12 equations form convenient groups. The first step is to solve the first 8 equations for the quantities  $k_1 \dots k_8$  ignoring for the present quantities  $k_0 \dots k_{16}$  which occur in the 5th—8th equations. The R.H.S. are taken as zero or unity, *vide* (6). As an example of (3) multiplications by powers of 10 are introduced and the order of equations arranged to make the diagonal quantities in increasing order of magnitude. In this particular case there is little gained by the former procedure, which is introduced merely to illustrate the method. The arrangement of the work is shown in tabular form in table LV, of which detailed explanation is now given.

T A B L E L V.

Equation Number	Right Hand Side	Left Hand Side							
		1 $k_3$	2 $k_4$	3 $k_7$	4 $k_8$	5 $\frac{1}{10}k_1$	6 $\frac{1}{10}k_2$	7 $\frac{1}{10}k_5$	8 $\frac{1}{10}k_6$
1	+1.000	+40.68	0	+7.35	+1.14	+22.0	+103.8	-0.3	+17.7
1(1)	+4.1335		0	-0.296	-0.0001	-1.354	-2.993	-0.0001	+0.247
2	0	+40.68	-1.14	+7.35	-103.8	+22.0	-17.7	-0.3	
3	0	0	+111.51	0	+4.1	+23.0	-8.4	+232.2	
4	-0.1807	0	-1.33	-0.206	-3.975	-18.76	+0.54	-3.2	
5	0	0	+111.51	+111.51	-23.0	+4.1	-282.2	-8.4	
6	-0.280	0	-0.3	-0.3	-6.2	-2.906	+0.084	-4.96	
7	0	+70.0	+366.0	0	+366.0	0	+70.0	0	
8	-0.541	0	-11.9	-11.9	-11.9	-56.16	+0.162	-9.576	
9	0	+70.0	+366.0	0	0	+366.0	0	+70.0	
9(1)	-0.0006	+1.000	+40.68	-1.14	+7.35	-103.8	+22.0	-17.7	-0.3
9(2)	+4.1333			+0.0046	-0.0008	+6.389	-6.344	-0.0073	+0.0004
10	-0.1807	0	+110.18	-0.3	+0.206	-2.906	+4.24	-8.346	+279.0
11	0	+0.280	-0.3	+0.206	-2.906	+0.62	-0.496	-0.1	
12	-0.280	0	+111.48	+111.48	-23.62	+1.194	-282.2	-8.896	
13	0	-0.1807	-1.33	-1.33	+18.76	-3.975	+3.2	+0.54	
14	0	0	+354.1	+354.1	-56.16	-56.16	+70.16	-9.576	
15	-2.552	+2.552	-264.9	-264.9	+56.144	+56.144	+45.17	-45.17	
16	0	-0.541	-11.9	-11.9	+101.1	+101.1	+76.56	+24.83	
17	+0.0074	0	+9.576	+9.576	-11.9	-11.9	+971.0	+0.16	
18	0	+0.435	+0.435	+0.435	+0.435	+0.435	+971.0	+0.131	
19	-0.435	0	-7.7	-7.7	-7.7	-7.7	-7.7	+0.131	
20	0	+0.0074	+0.0074	+0.0074	+0.0074	+0.0074	+0.0074	+963.3	
21	-0.435	+0.0074	-0.0074	-0.0074	-0.0074	-0.0074	-0.0074	-7.7	
16	-0.1807	+0.280	+1.000	+110.15	0	-2.781	+4.86	-8.842	+279.0
16(1)	-4.433					+0.0171	-1.401	+0.0036	+0.3892
16(2)		+0.116				-0.0802	+0.299	+0.0123	+0.1142
16(3)			+3.7709			+0.0005	+0.0043	-0.0027	+2.773
17	-0.280	-1.807	0	+110.15	-4.86	-2.781	-279.0	-8.842	
18	0	0	0	0	0	0	0	0	
19	-0.541	+2.552	0	0	+89.2	-0.016	+24.99	-10.342	
20	-0.0046	+0.007	+0.2525	0	-0.07	+1.227	-0.223	+7.045	
21	-2.552	-0.541	0	0	+89.2	+89.2	+10.342	+24.99	
22	+0.008	+0.012	-0.4412	0	-0.21	-0.21	+0.39	-13.31	
23	+0.0074	+0.435	0	0	0	0	+963.3	0	
24	-0.0145	+0.022	+0.0803	0	0	0	+7.1	+22.40	
25	-0.435	+0.0074	0	0	0	0	0	+963.3	
26	+0.4577	-0.0709	-2.533	0	0	0	0	-7.067	

TABLE LV (Continued).

Equation Number	Right Hand Side				Left Hand Side				
					$k_8$	$\frac{1}{10} k_1$	$\frac{1}{10} k_2$	$\frac{1}{10} k_3$	$\frac{1}{10} k_6$
22	- .028	- .1807	0	+ 1.000	+110.15	- 4.86	- 2.781	-279 0	- 8.842
22(1)	- .0114					+ .0299	+ .0802	- .1144	- .0122
22(2)		- .4436				- .1401	+ .0171	+ .3895	- .0036
22(3)			+ .0002			+ .0008	- .0025	+ .0864	+ .0879
22(4)				+ 3.7709		+ .0043	+ .0005	- 2.773	- .0027
23	- .5456	+ 2.5527	+ .02525	0		+ 89.13	+ .1067	+ 24.77	- 3.297
	- .0012	- .0080	0	+ .04412		- .21	- .1227	- 12.31	- .390
24	- 2.544	- .5422	- .0441	0			+ 88.99	+ 10.732	+ 12.68
	- .001	- .0046	0	+ .02525			- .07	- 7.045	- .22
25	- .0071	+ .4372	+ .0803	0				+ 962.6	+ 22.40
	- .0709	- .4577	0	+ 2.533				- 706.7	- 22.40
26	+ .0227	- .0835	- 2.533	0					+ 256.0
	- .0022	- .0145	0	+ .0803					- .71
27	- .5468	+ 2.5447	+ .02525	+ .04412	+ 1.000	+ 88.92	- .016	+ 12.46	- 3.687
27(1)	- .5473						+ .0005	+ .0051	- .0051
27(2)		+ 2.5635					+ .0001	+ .0174	- .0015
27(3)			- .01525				0	+ .0039	+ .0366
27(4)				- .0787			0	+ .1239	- .0011
27(5)					+ 1.0075		0	- .0069	- .0006
28	- 2.545	- .5468	- .04412	+ .02525	0		+ 88.92	+ 3.687	+ 12.46
	- .0001	+ .0005	0	+ .00001	+ .00018		0	+ .002	- .001
29	- .078	- .0205	+ .0803	+ 2.533	0			+ 255.9	0
	+ .0766	- .3566	- .0035	- .006	- .1401			- 1.75	+ .5165
30	+ .0205	- .078	- 2.533	+ .0803	0				+ 255.9
	- .0227	+ .1055	+ .001	+ .0018	+ .04146				- .15
31	- 2.5451	- .5463	- .04412	+ .02525	+ .00018	+ 1.000	+ 88.92	+ 3.689	+ 12.459
31(1)	- 2.564							+ .0015	+ .0174
31(2)		- .5463						- .0051	+ .0051
31(3)			+ .0786					+ .0011	- .1238
31(4)				- .0153				+ .0367	+ .0039
31(5)					+ .00018			- .0020	+ .0020
31(6)						+ 1.0075		- .0006	- .0069
32	- .0014	- .3771	+ .0768	+ 2.527	- .1401	0		+ 254.15	+ .5165
	+ .1056	+ .0227	+ .0018	- .001	0	- .04149		- .153	+ .5169
33	- .0022	+ .0275	- 2.532	+ .0821	+ .04146	0			+ 255.75
	+ .3566	+ .0766	+ .006	- .0035	0	- .1401			- 1.75
34	+ .1042	- .3544	+ .0786	+ 2.526	- .1401	- .04149	+ 1.000	+ 254.0	- .0004
34(1)	+ .1042								0
34(2)		- .3544							0
34(3)			+ .0786						0
34(4)				+ 2.526					0
34(5)					- .1401				0
34(6)						- .04149			0
34(7)							+ 1.000		0
35	+ .3544	+ .1040	- 2.526	+ .0786	+ .04146	- .1401	+ 0.000		+ 254.0
	0	0	0	0	0	0	+ .000002		0
36	+ .3544	+ .1040	- 2.526	+ .0786	+ .04146	- .1401	+ 0.000	+ 1.000	+ 254.0
Case	I	II	III	IV	V	VI	VII	VIII	
Solution 1st approximation.	Case I	$k_3$	$k_4$	$k_7$	$k_8$	$\frac{1}{10} k_1$	$\frac{1}{10} k_2$	$\frac{1}{10} k_3$	$\frac{1}{10} k_6$
	II	+ .1016	- .00001474	- .004025	- .0001035	- .006155	- .028835	+ .00041	+ .001395
	III		+ .1016	+ .0001053	- .004027	+ .028835	- .006144	- .001396	+ .0001094
	IV			+ .03423	+ .00000182	- .00017	+ .000864	+ .0003095	- .00094
	V				+ .03423	- .000875	- .000172	+ .00094	+ .0003095
	VI					+ .01133	+ .00000202	- .000552	+ .000163
	VII						+ .01133	- .0001635	- .000552
	VIII							+ .003937	0
								+ .003937	

The first, second, fifth and sixth lines and columns of table XLVII are multiplied by 10 and their rows and columns are arranged in order 3, 4, 7, 8, 1, 2, 5, 6: giving rise to equations whose coefficients are written down in column headed "Left Hand Side 1 . . . 8," and in lines whose

equation numbers are 1 to 8. The Right Hand Sides of these equations are  
 It is however only necessary to write down the first column under heading R.H.S.  
 To eliminate the first quantity  $k_3$  from these 8 equations, the first equation is multiplied  
 successively by  $-\frac{0}{40 \cdot 68} = 0$ ,  $\frac{-7 \cdot 35}{40 \cdot 68} = -\cdot 1807$ ,  $\frac{-1 \cdot 14}{40 \cdot 68} = -\cdot 280$  . . . . .

1	0	0
0	1	0
0	0	1

and the results written below the 2nd, 3rd, 4th equations in old face type. These multipliers are also applied to the right hand side of the first equation which is unity, and therefore appear in this column. The work is performed by sliderule, making one setting representing division by 40·68, and reading off opposite the quantities 0, 7·35, 1·14 etc. these quantities being at once entered in R.H.S. column (shown in old face type). In completing the multiplication of the first equation by these several factors, it is noticed that the quantity to be entered is the product of the factor of the particular line by the coefficient of the particular column in equation No. 1, e. g. old face figures  $-3 \cdot 975$  (line 3, column 5) =  $-\cdot 1807 \times 22 \cdot 0$ . The old face figures thus formed for all the equations Nos. 2-8, are added to the corresponding coefficients and give rise to 7 equations (numbered 9-15) from which  $k_3$  has been eliminated. The process as regards the right hand side has only actually been applied for case I in which the right hand side of the first equation is unity and the rest are zero (*vide* § 7): but it will be easily seen that for the other cases all the old face quantities would be zero having zero as a factor. Case II is accordingly brought in conveniently after the elimination of  $k_3$  has been completed.

The necessary multipliers for the next elimination, that of  $k_4$ , are now duly entered of the R.H.S. for case II. They are the old face figures  $+\cdot 0280$ ,  $-\cdot 1807$ ,  $+2 \cdot 552$  . . . and the process of elimination is proceeded with. In this way the groups of equations 16-21, 22-26, 27-30, 31-33, 34-35, 36 are formed successively in the last of which only  $\frac{1}{10} k_6$  occurs. From this values of  $\frac{1}{10} k_6$  are written down at the bottom of the table for each of the eight cases. These values would be substituted in 34, except that the coefficient of  $k_6$  is so small that they are negligible. Thus equations 34 (1) . . . 34 (7) are formed giving values of  $\frac{1}{10} k_5$  for seven cases. Then the values of  $\frac{1}{10} k_6$  and  $\frac{1}{10} k_5$  are substituted in 31, giving 31 (1) . . . 31 (6) from which six values of  $\frac{1}{10} k_3$  are formed. These several substitutions are shown in old face. The results of the terms on the left hand sides are combined and with changed sign applied to the right hand side for the corresponding case, e. g. 31 (3)  $+\cdot 0786 = -\cdot 04412 - \cdot 0011 + \cdot 1238$ . All the values obtained are exhibited at the foot of the table: they are the values of  $x$ , in the notation of § 7. Since the equations are symmetrical the solution is also symmetrical and  $x_s = x_r$ : so that it is only necessary to take out half of the quantities. Had the equations not been symmetrical all work below the diagonal in each group of eliminants would have had to be completed and the full number of cases, eight, would have been necessary in substituting in each of equations 1, 9, 16, 22, 27, 31, 34, 36.

13. An approximate solution of the first eight equations has now been found. As it will be necessary to substitute from the solution in the remaining equations, it is desirable in any case at this stage to check the solution; but this substitution at the same time enables a higher approximation to be reached by (8). This is desirable; although a high order of accuracy of final solution is not desired, yet it is proper to avoid the introduction of computation inaccuracy at an early stage in the work. The first step is to substitute in the equations and so to find the quantities  $a, \beta$  . . . of § 7. In performing this substitution a sliderule cannot be used, as greater precision is desired. If an arithmometer is used, then there is little extra labour in taking out the work to the full number of figures which occur. When this has been done (10) gives a means of an infinite number of successive approximations. As exemplifying this, full accuracy is kept in this substitution, which is reproduced for Case I only in table LVI. It is to be remarked that there is no advantage in keeping the solution of

table LV to as many figures as has been done, as the latter figures cannot be accurate: and their presence adds to the labour of multiplication. However in the present case the substitution had already been carried out by the computer before this simplification could be given effect to.

TABLE LVI.

Equation	1	2	3	4	5	6	7	8
Case I	+4·133,088,00	0	+746,760,00	+115,824,00	+2·235,200	+10·546,080,00	-030,480	+1·798,320,00
	0	-000,509,623,2	+000,016,873,6	-000,108,330	+0·001,630,012	-0·000,324,28	+000,280,808	+0·000,004,422
	-0·020,583,75	+004,689,85	-448,827,75	0	-0·016,602,5	-0·002,575	+033,810	-1·135,855
	-0·000,117,90	-000,760,725	0	-011,641,265	+0·002,380,5	-0·000,424,35	+029,207,7	+0·000,860,4
	-0·135,410	+638,890	-025,235,5	+141,665	-2·252,730	0	-430,850	0
	-2·903,073	-634,370	-683,205	-118,223,5	0	-10·553,010	0	-2·018,450
	-0·000,123	-007,257	-003,444	-115,702	+0·029,700	0	+398,110	0
	+0·024,091,5	-000,418,5	+303,669	-011,718	0	+0·097,650	0	+1·354,545
Sum	+0·999,471,76	+000,072,001,8	-000,268,446,4	+000,095,876	-0·001,421,088	-0·002,203,63	+000,058,608	-0·000,560,178

Table LVII gives the results of the substitution for all cases, that is the quantities  $\alpha \beta \gamma \dots$  in notation of § 7. To avoid constant repetition of zeros, they have all been multiplied by  $10^3$ . It is now perfectly straightforward to substitute in (10) and obtain a second approximation. Since the solution is known to be symmetrical, it is not necessary to perform the substitution on both sides of the diagonal: but as a check it may sometimes be useful to do so. When two determinations of what is known by symmetry to be one quantity differ slightly, the mean can be taken.

TABLE LVII.

r	1	2	3	4	5	6	7	8
$10^3 \alpha_r$	-528,24	+072,001,8	-266,446,4	+095,876	-1·421,988	-3·203,63	+058,598	-566,179
$10^3 \beta_r$	-129,469,2	-285,112	-010,756	-190,739,6	+2·639,45	-464,812	+153,302	+248,962
$10^3 \gamma_r$	-156,075,2	-127,360	-749,292	+077,603,2	+286,00	-436,038	-077,414	-1·448,378
$10^3 \delta_r$	+217,747	-1·210,784,8	+334,403,2	-977,14	+3·003,062	-230,44	+2·077,762	+842,254
$10^3 \epsilon_r$	-112,024	+492,20	+067,01	+1·171,762	-015	+132,62	-3·072,0	+196,4
$10^3 \zeta_r$	-329,39	-100,006	-160,966	-026,95	+251,92	-474,2	-105,0	-618,7
$10^3 \eta_r$	-012,175	-067,41	+022,785	-359,55	+131,75	-122,5	+1·105,4	+075,7
$10^3 \theta_r$	+015,73	+016,317	+370,834	+028,735	-020,22	+014,75	-099,78	+1·087,68

The result of the second approximation, found by means of (8), is given in table LVIII which has been rearranged in the order of the original quantities.

TABLE LVIII.

	$k_1$	$k_2$	$k_3$	$k_4$	$k_5$	$k_6$	$k_7$	$k_8$
1	+1·1325	+000,006	-031,44	+288,135	-055,14	+016,27	-001,702	-008,836
2		+1·132,5	-2·8,135	-081,44	-016,27	-055,14	+006,835	-001,702
3			+101,552	0	+004,079	+013,936	-004,021	-000,109,5
4				+101,552	-013,936	+004,079	+000,108,5	-004,021
5					+303,63	0	+003,002	+090,41
6						+303,63	-009,41	+003,092
7							+034,241	0
8								+034,241

14. This solution corresponds to that indicated in (13). The next step is to form (15) with a view to substitution for  $k_1 \dots k_8$  in the equations Nos. 9—20. For this only values of  $k_5 \dots k_8$  are required since the coefficients of  $k_1 \dots k_4$  are zero in these equations. However values of  $k_1 \dots k_4$  are also required at a later stage, so the complete series of quantities  $k_1$  to  $k_8$  are expressed



according to (15). Denote the quantities given in table LVIII by  ${}_sK_r$ , where  $r$  and  $s$  have all values from 1 to 8. It is necessary to compute all the quantities  $\Sigma {}_sK_r(s, t)$  where  $s$  has all values from 1 to 8 and is the quantity to which  $\Sigma$  refers, for each value of  $r$  from 1 to 8 and each value of  $t$  from 9 to 20:  $(s, t)$  in the notation of (11) indicate the coefficients shown in table XLVII, while  ${}_sK_r$  are the quantities of table LVIII. Owing to zero coefficients it has only to be taken for values 9 to 16 and so the number of quantities  $\Sigma {}_sK_r(s, t)$  actually to be computed is  $8 \times 8 = 64$ ; and among them there is a sort of skew symmetry due to equality in pairs of coefficients in table XLVII, making altogether only 32 independent quantities. Further the summation  $\Sigma$  although relating to 8 values of  $s$ , actually only gives rise to four terms owing to zero coefficients. The details of the computation are given in full in table LIX. The computation is taken out to full accuracy to exhibit the symmetry which afterwards occurs when substituting in equations Nos. 9—20.

TABLE LIX.

Values of  ${}_sK_r(s, t)$  and of  $\Sigma {}_sK_r(s, t)$ , the latter in old face type.

r	s	t = 0	10	11	12	13	14	15	16
1	5	-356,755,80	0	-607,127,80	+348,484,8	-100,728,00	0	-102,000,00	+142,812,60
	6	0	+105,286,00	+102,826,40	+264,712,9	0	+032,377,30	+042,139,30	+030,099,50
	7	+002,246,64	-039,094,94	-050,413,24	-100,588,2	+000,136,18	-002,331,74	-003,438,04	-002,314,72
	8	+203,939,95	+011,682,20	+522,148,50	-201,802,7	+012,103,05	+000,706,80	+012,015,60	-017,846,70
Sum		-151,569,21	+077,834,16	-322,566,14	+250,916,8	-097,488,49	+030,752,36	-051,292,14	+152,759,68
2	5	-106,266,90	0	-294,712,9	+102,826,40	-032,377,30	0	-030,099,50	+042,139,30
	6	0	-356,755,80	-348,484,8	-867,127,80	0	-109,728,60	-142,812,60	-102,009,00
	7	-011,662,20	+202,939,95	+261,092,7	+522,148,50	-000,706,80	+012,103,95	+017,846,70	+012,015,60
	8	+039,094,94	+002,246,64	+100,588,2	-050,413,24	+002,331,74	+000,136,16	+002,314,72	-003,438,04
Sum		-077,834,16	-151,569,21	-250,916,8	-322,566,14	-030,752,36	-097,488,49	-152,759,68	-051,292,14
3	5	+026,981,130	0	+066,365,33	-026,779,28	+008,117,210	0	+007,546,15	-010,584,61
	6	0	+090,165,92	+098,075,52	+226,738,72	0	+027,732,64	+036,001,24	+025,781,60
	7	+005,307,720	-092,362,37	-119,102,02	-237,641,10	+000,321,680	-005,503,77	-008,122,42	-006,468,56
	8	+002,482,245	+000,143,22	+006,412,35	-003,213,77	+000,148,645	+000,008,68	+000,147,56	-000,210,17
Sum		+034,191,095	-002,053,23	+041,751,18	-039,895,43	+008,587,535	+022,232,55	+035,665,53	+009,529,26
4	5	-080,165,92	0	-228,738,72	+088,075,52	-027,732,84	0	-026,781,60	+036,091,24
	6	0	+026,301,130	+025,779,28	+086,365,33	0	+008,117,210	+010,584,61	+007,546,15
	7	-000,143,22	+002,482,245	+003,213,77	+006,412,35	-000,008,68	+000,148,645	+000,219,17	+000,147,56
	8	+032,362,37	+005,307,720	+237,641,10	-119,102,02	+005,503,77	+000,321,680	+005,406,56	-008,122,42
Sum		+002,053,23	+034,191,095	+039,895,43	+041,751,18	-022,232,55	+008,587,535	-009,529,26	+035,665,53
5	5	+2,546,786,10	0	+6,404,360,10	-2,487,741,6	+783,323,70	0	+728,216,50	-1,016,501,70
	6	0	+2,546,786,10	+2,487,741,6	+8,043,360,10	0	+783,323,70	+1,010,501,70	+728,216,50
	7	-004,081,44	+071,023,24	+091,585,04	+182,737,2	-000,247,36	+004,236,04	+006,245,84	+004,205,12
	8	-2,283,447,70	-131,221,20	-5,875,131,00	+2,044,524,2	-136,191,70	-007,052,80	-135,107,80	+200,900,20
Sum		+259,256,96	-060,197,96	+620,814,14	+639,519,8	+646,884,64	-003,716,76	+599,261,74	-614,488,38
6	5	0	0	0	0	0	0	0	0
	6	0	+3,546,786,10	+2,487,741,6	+8,043,360,10	0	+783,323,70	+1,010,501,70	+728,216,50
	7	+131,221,20	-2,283,447,70	-2,044,524,2	-5,875,131,00	+007,052,80	-136,191,70	-200,900,20	-135,107,80
	8	-071,023,24	-004,081,44	-182,737,2	+001,585,04	-004,236,04	-000,247,36	-004,205,12	+006,245,84
Sum		+060,197,96	+259,256,96	-639,519,8	+620,814,14	+003,716,76	+646,884,64	+814,488,38	+599,261,74
7	5	+020,005,24	0	+050,306,84	-019,541,44	+006,153,08	0	+005,720,20	-008,008,28
	6	0	-643,182,70	-628,271,20	-1,617,400,70	0	-197,825,60	-257,471,80	-183,908,50
	7	-045,108,12	+786,515,77	+1,014,218,42	+2,023,643,10	-002,739,28	+046,910,17	+060,166,82	+046,577,76
	8	0	0	0	0	0	0	0	0
Sum		-025,192,88	+143,333,07	+436,254,06	+386,700,96	+003,413,80	-150,915,73	-182,584,88	-145,349,02
8	5	+043,182,70	0	+1,617,400,70	-628,271,20	+197,825,60	0	+183,908,50	-257,471,80
	6	0	+020,005,24	+050,306,84	+019,541,44	0	+006,153,08	+005,720,20	0
	7	0	0	0	0	0	0	0	0
	8	-786,515,77	-045,108,12	-2,023,643,10	+1,014,218,42	-046,910,17	-002,739,28	-046,577,76	+069,166,82
Sum		-143,333,07	-025,192,88	-386,700,96	+436,254,06	+150,915,73	+003,413,80	+145,349,02	-182,584,88

15. Values of  $k_5, k_6, k_7, k_8$  given in (15) contain terms in  $k_1 \dots k_{16}$  of which the coefficients have just been found in table LIX. These are to be substituted in equations 9—16, they do not occur in equations 17 to 20. The formation of the products and the collecting of coefficients is carried out in Table LX. In this the values of  $\Sigma {}_sK_r(s, t)$  are rewritten with sign changed at the top, kept to six places, while the multiplying coefficients are all shown in the first column. The previously existing

coefficients of  $k_9 \dots k_{20}$  are also included. The complete coefficients of  $k_9$  to  $k_{20}$  after including the portions due to substitution of  $k_5$  to  $k_{13}$  are shown in Table LX in old face. Thus twelve symmetrical equations relating  $k_9$  to  $k_{20}$  have been formed,  $k_1$  to  $k_8$  having been eliminated. The solution of these twelve equations is performed in a manner similar to that employed for the solution of the first eight. The equations are first rearranged in increasing order of the diagonal coefficients, the whole process being given in table LXI. It does not seem likely that a second approximation is necessary in this case, so a verification, as described in § 11, is carried out in table LXII showing the degree of precision with which the last equation of the group of 12 is satisfied for each of the 12 cases.

TABLE LX.

Substitution of  $k_5-k_8$  in equations 9 to 20.

r	t	9	10	11	12	13	14	15	16	17	18	19	20	
5	$\Sigma, K_r(s,t)$	-250,257	+060,108	-620,814	-630,520	-646,885	+003,717	-500,264	+814,488					
6		-060,198	-260,257	+630,520	-620,814	-003,717	-646,885	-14,488	-500,264					
7		+025,193	-143,333	-436,254	-386,701	-003,414	+150,916	+182,565	+145,349					
8		+143,333	+025,193	+386,701	-436,254	-150,916	-003,414	-145,349	+182,565					
Equation 9	Multiplier	-1.6774	+ .3895	- 4.0167	- 4.1377	- 4.1853	+ .0240	- 3.8772	+ 5.2607					
	0	0	0	0	0	0	0	0	0					
	- 1.32	- .0333	+ .1892	+ .5750	+ .5104	+ .0045	- .1992	+ .2410	- .1919					
	- 22.97	- 3.2024	- .5787	- 8.8825	+ 10.0298	+ 3.4665	+ .0784	+ 3.3387	- 4.1940					
	(9,t)	+ 9.68	0	+ 16.81	- 11.33	+ 0.28	0	+ 0.12	0.60					
	Sum	+ 4.6770	0	+ 4.4867	- 4.0365	- 0.4343	- 0.0968	- 0.6595	+ 0.2838	+ 0.55	0	- 0.27	- 1.66	
10	0	0	0	0	0	0	0	0	0					
	+ 6.47	- .3895	- 1.6774	+ 4.1377	- 4.0167	- .0240	- 4.1853	- 2.897	- 3.8772					
	+ 22.97	- .5787	- 3.2924	- 10.0208	- 8.8825	- .0784	+ 3.4666	+ 4.1940	+ 3.3387					
	- 1.32	- .1892	- .0333	- .5104	+ .5759	+ .1992	+ .0045	+ .1919	- 2.410					
	(10,t)	0	+ 9.68	+ 11.33	+ 16.81	0	+ .28	+ .60	+ .12					
	Sum	+ 4.6770	+ 4.0365	+ 4.4867	+ 0.0968	- 0.4343	- 0.2838	- 0.6595		0	+ 0.55	+ 1.66	- 0.27	
11	+ 16.27	- 4.2151	- .9794	- 10.1006	- 10.4050	- 10.5248	+ .0605	- 9.7500	+ 13.2517					
	+ 6.32	- .3804	- 1.6355	+ 4.0418	- 3.9235	- .0235	- 4.0883	- 5.1476	- 3.7873					
	+ 29.62	+ .7462	- 4.2455	- 12.0218	- 11.4541	- .1011	+ 4.4701	+ 5.4082	+ 4.3052					
	- 50.10	- 8.4710	- 1.4889	- 22.8540	+ 25.7826	+ 8.9191	+ .2018	+ 8.5501	- 10.7908					
	(11,t)	+ 16.81	+ 11.33	+ 79.30	0	+ .61	- .28	- .32	+ 1.44					
	Sum	+ 4.4867	+ 4.0365	+ 37.4654	0	- 1.1203	+ 0.3040	- 1.2193	+ 1.5388	+ 0.48	- 0.21	- 0.71	- 1.36	
12	- 6.32	+ 1.6385	- .3804	+ 3.0285	+ 4.0118	+ 4.0883	- .0235	+ 3.7873	- 5.1476					
	+ 16.27	- .9794	- 4.2181	+ 10.4050	- 10.1006	- .0605	- 10.5248	- 13.2517	- 9.7500					
	+ 30.10	+ 1.4889	- 8.4710	- 25.7826	- 22.8540	- .2018	+ 8.9191	+ 10.7909	+ 5.5001					
	+ 29.62	+ 4.2455	- .7462	+ 11.4541	- 12.0218	- 4.4701	- .1011	- 4.3052	+ 5.4082					
	(12,t)	- 11.33	+ 16.81	+ 79.30	+ .23	+ .61	- .28	+ 1.44	- .32					
	Sum	- 4.9395	+ 4.4867	0	+ 37.4654	- 0.3040	- 1.1203	- 1.5388	+ 1.2193	+ 0.21	+ 0.48	+ 1.36	- 0.71	
13	+ 1.99	- .5159	+ .1198	- 1.2354	- 1.2726	- 1.2873	+ .0074	- 1.1925	+ 1.6208					
	0	0	0	0	0	0	0	0	0					
	- 0.05	- .0020	+ .0115	+ .0349	+ .0309	+ .0003	- .0121	- .0146	- .0116					
	- 1.37	- .1984	- .0345	- .5298	+ .5977	+ .2068	+ .0047	+ .1991	- .2501					
	(13,t)	+ .28	0	+ .61	+ .28	+ 3.43	0	+ 2.13	+ 4.28					
	Sum	- 0.4343	+ 0.0968	- 1.1203	- 0.3040	+ 2.3497	0	+ 1.1220	- 2.9200	+ 0.83	0	+ 0.32	- 1.20	
14	0	0	0	0	0	0	0	0	0					
	+ 1.99	- .1198	- .5159	+ 1.2726	- 1.2354	- .0074	- 1.2873	- 1.6208	- 1.1925					
	+ 1.37	+ .0345	- .1984	- .5077	- .5298	- .0047	+ 0.2068	+ .2501	+ .1991					
	- 0.08	- .0115	- .0020	- .0309	+ .0349	+ .0121	+ .0003	+ 0.116	- .0146					
	(14,t)	0	+ .28	+ .61	+ .28	+ 3.43	0	+ 4.28	+ 2.13					
	Sum	- 0.0968	- 0.4343	+ 0.3040	- 1.1203	+ 2.3497	+ 2.9200	+ 1.1220	+ 1.6208	0	+ 0.83	+ 1.20	+ 0.32	
15	+ 1.85	- .4796	+ .1114	- 1.1485	- 1.1831	- 1.1967	+ .0069	- 1.1086	+ 1.5068					
	+ 2.50	- .1550	- .6715	+ 1.6584	- 1.6070	- .0066	- 1.6754	- 2.1095	- 1.5521					
	+ 2.02	+ .0509	- .2895	- .8812	- .7811	- .0069	+ .3048	+ .3688	+ .2936					
	- 1.36	- .1049	- .0343	- .5269	+ .5933	+ .2052	+ .0046	+ .1977	- 2.163					
	(15,t)	+ .12	+ .60	+ .32	+ 1.44	+ 2.13	+ 4.28	+ 8.46	0					
	Sum	- 0.5505	- 0.2839	- 1.2192	- 1.5388	+ 1.1220	+ 2.9200	+ 5.8084	0	+ 0.01	+ 0.09	+ 1.83	+ 0.38	
16	- 2.59	+ .8715	- .1550	+ 1.6070	+ 1.6584	+ 1.6754	- .0096	+ 1.5521	- 2.1095					
	+ 1.85	- .1114	- .4796	+ 1.1831	- 1.1485	- .0069	- 1.1967	+ 1.5068	- 1.1086					
	+ 1.36	+ .0343	- .1949	- .5933	- .5259	- .0048	+ .2052	+ .2183	+ .1977					
	+ 2.02	+ .2895	+ .0509	+ .7811	- .8812	- .3040	- .0060	- .2936	+ .3688					
	(16,t)	- .60	+ .12	+ 1.44	+ .32	+ 4.28	+ 2.13	0	+ 8.46					
	Sum	+ 0.2830	- 0.0505	+ 1.5388	- 1.2192	- 2.9210	+ 1.1220	0	+ 5.8084	- 0.99	+ 0.01	- 0.38	+ 1.83	
17	(17,t)	+ 0.55	0	+ 0.48	+ 0.21	+ 0.83	0	+ 0.01	- 0.99	+ 2.23	0	- 0.56	- 4.05	
18	(18,t)	0	+ 0.55	- 0.21	+ 0.48	0	+ 0.83	+ 0.99	+ 0.01	0	+ 2.23	+ 4.05	- 0.56	
19	(19,t)	- 0.27	+ 1.66	- 0.71	+ 1.36	+ 0.32	+ 1.20	+ 1.83	- 0.38	- 0.56	+ 4.05	+ 10.06	0	
20	(20,t)	- 1.66	- 0.27	- 1.36	- 0.71	- 1.20	+ 0.32	+ 0.38	+ 1.83	- 4.05	- 0.56	0	+ 10.06	

TABLE LXI.

Equation Number	Right hand side	Left Hand Side.											
		$k_{17}$	$k_{16}$	$k_{15}$	$k_{14}$	$k_{13}$	$k_{12}$	$k_{11}$	$k_{10}$	$k_9$	$k_8$	$k_7$	$k_6$
1	+1.00 +3.77446	+2.23	0	+ .83 - .4318	0	+ .65 - .01716	0	+ .01 + .00188	-.99 + .08799	-.58 - .00054	- 4.05 - 2.35406	+ .48 + .00226	+ .21 - .003034
2	0	+2.23	0	+ .83	0	+ .65	+ .90	+ .01	+ 4.05	-.56		-.21	+ .48
3	0	0	+2.35	0	-.43	+ .10	+1.12	-2.92	+ .32	- 1.20		- 1.12	-.36
4	0	0	-.309	0	-.2047	-.43	+.0037	+.3695	+ .2084	+ 1.5074		-.1787	-.0783
5	0	0	0	+2.35	-.10	-.43	+2.02	+1.12	+ 1.20	+ .32		+ .36	- 1.12
6	0	0	0	0	+4.68	0	-.66	+ .28	-.27	- 1.66		+ 4.49	- 4.94
7	0	0	0	0	-.136	0	-.0025	+ .2442	+ .1381	+ .9989		+ .1184	-.053
8	0	0	0	0	0	+4.68	-.25	-.66	+ 1.06	-.27		+ 4.04	+ 4.46
9	0	0	0	0	0	0	+5.81	0	+ 1.83	+ .39		- 1.22	- 1.54
10	0	0	0	0	0	0	-.00004	+ .00444	+ .0025	+ .0183		-.00215	-.00044
11	0	0	0	0	0	0	0	0	0	0		0	0
12	0	0	0	0	0	0	0	0	0	0		0	0
13	0	+1.00 +3.77446	+2.23 (1) (2)	0	+ .83 - .2801 - .4318	0	+ .55 - .00941 - .01716	+ .99 + .18603 + .08799	+ .01 - .00089 + .00188	+ 4.05 + .4373 - 2.35406	-.56 - .3255 - .00054	-.21 - .00099 - .003034	+ .48 + .00604 + .00226
14	0	0	+2.041	0	-.6347	+ .10	+1.1163	-2.5515	+ .5281	+ .3074		- 1.2987	-.4392
15	0	0	0	+2.35	-.10	-.43	+2.92	+1.12	+ 1.20	+ .32		+ .36	+ 1.12
16	0	0	0	-.309	-.2047	-.43	+.0037	+.3695	+ .2084	+ 1.5074		-.1787	-.0783
17	0	0	0	0	+4.68	0	-.66	+ .28	-.27	- 1.66		+ 4.94	- 4.49
18	0	0	0	0	-.136	0	-.0025	+ .2442	+ .1381	+ .9989		+ .1184	-.053
19	0	0	0	0	0	+4.68	-.25	-.66	+ 1.06	-.27		+ 4.04	+ 4.46
20	0	0	0	0	0	0	+5.81	0	+ 1.83	+ .39		- 1.22	- 1.54
21	0	0	0	0	0	0	-.00004	+ .00444	+ .0025	+ .0183		-.00215	-.00044
22	0	0	0	0	0	0	0	0	0	0		0	0
23	0	0	0	0	0	0	0	0	0	0		0	0
24	0	+1.00 +3.77446	+2.23 (1) (2) (3)	0	+ .83 - .2801 - .4318	0	+ .55 - .00941 - .01716	+ .99 + .18603 + .08799	+ .01 - .00089 + .00188	+ 4.05 + .4373 - 2.35406	-.56 - .3255 - .00054	-.21 - .00099 - .003034	+ .48 + .00604 + .00226
25	0	0	+2.041	0	-.6347	+ .10	+1.1163	-2.5515	+ .5281	+ .3074		- 1.2987	-.4392
26	0	0	0	+2.35	-.10	-.43	+2.92	+1.12	+ 1.20	+ .32		+ .36	+ 1.12
27	0	0	0	-.309	-.2047	-.43	+.0037	+.3695	+ .2084	+ 1.5074		-.1787	-.0783
28	0	0	0	0	+4.68	0	-.66	+ .28	-.27	- 1.66		+ 4.94	- 4.49
29	0	0	0	0	-.136	0	-.0025	+ .2442	+ .1381	+ .9989		+ .1184	-.053
30	0	0	0	0	0	+4.68	-.25	-.66	+ 1.06	-.27		+ 4.04	+ 4.46
31	0	0	0	0	0	0	+5.81	0	+ 1.83	+ .39		- 1.22	- 1.54
32	0	0	0	0	0	0	-.00004	+ .00444	+ .0025	+ .0183		-.00215	-.00044
33	0	0	0	0	0	0	0	0	0	0		0	0
34	0	+1.00 +3.77446	+2.23 (1) (2) (3) (4)	0	+ .83 - .2801 - .4318	0	+ .55 - .00941 - .01716	+ .99 + .18603 + .08799	+ .01 - .00089 + .00188	+ 4.05 + .4373 - 2.35406	-.56 - .3255 - .00054	-.21 - .00099 - .003034	+ .48 + .00604 + .00226

TABLE LXI.—(Continued).

Table with 12 columns: Equation Number, Right Hand Side, and Left Hand Side (k14 to k12). Rows 35-63 contain numerical data for various equations, showing adjustments and final values.

TABLE LXI.—(Concluded).

Equation number	Right Hand Side.								Left Hand Side.				
	$k_{10}$	$k_{11}$	$k_{20}$	$k_{11}$	$k_{12}$	$k_{10}$	$k_{11}$	$k_{20}$	$k_{11}$	$k_{12}$			
64	- .0985 + .13785	+ .1823 + .29144	+ 1.26325 + 1.35583	- .531 - .52842	+ .0494 + .10925	+ .04385 + .03326	0 o	+ 1.00 + 1.04117	+ 1.551 (1) (2) (3) (4) (5) (6) (7) (8)	+ .2065 + .02232 + .120 + .01123 + .0259 + .01601 + .02233 + .0093 + .00369 + .00914	+ .1039 + .06039 + .01123 - .00805 + .01303 + .00468 - .01543 + .00466 - .00186	+ .3157 + .001487 + .004562 - .003343 + .01666 + .01099 + .01543 + .00825 + .002772	- 1.0181 + .01515 + .004936 + .05529 + .01110 + .05121 + .01649 + .00950 + .0274
65	+ .3345 + .0051	- 1.8296 + .0243	- .2184 + .1681	+ .1985 + .0716	- .00693 - .00657	- .1209 - .00584	- .06703 o	0 - .1331	+ 3.2072 + .0275	+ .01384 + .01384 + .1234 - .0070	- .3986 + .0420 - .1234 + .02116	+ .00506 + .1305 + .0588 + .0702	
66	+ 1.8513 + .0026	+ .3518 + .0122	- .1834 + .0847	- .4221 + .0350	+ .1301 - .00294	+ .1331 + .00294	- .06703 o	0 + .06703					
67	+ .0177 + .0078	+ .6401 + .0371	+ .0502 + .257	- 1.484 + .108	- .8892 + .0101	- 1.2294 + .0086	+ .6754 o	0 + .2034			+ 25.846 - .0642	- 21.32 + 21.32	
68	- .566 + .026	- .133 + .1231	+ .523 + .853	+ .152 - .3588	+ 1.2049 + .0334	+ .0289 + .0296	+ .2034 o	0 + .0754				+ 26.4896 - .7078	
69	+ .3396 + .34375	- 1.8539 - 1.8482	- .2865 - .39877	+ .2681 + .24649	- .0135 - .03583	- 1.2674 - 1.4314	- .08703 - .05683	- 1.331 - .14073	+ 1.00 + 1.00361	+ 3.1797 (1) (2) (3) (4) (5) (6) (7) (8) (9)	0 - .00205 - .00338 + .004045 + .02314 + .01527 + .02143 + .01147 + .003851 + .00235	- .4386 + .00205 + .000181 + .007626 + .001531 + .007663 + .005934 + .001309 + .003779 + .000150	+ .14456 + .00205 + .000181 + .007626 + .001531 + .007663 + .005934 + .001309 + .003779 + .000150
70	+ 1.8539 o	+ .3396 + .612	- .2681 - .2069	- .3865 - 1.376	+ .12674 + .8993	- .0135 - 1.2383	+ .1331 + .6754	- .06703 + .2034	0 o	+ 3.1797 o	- 1.4456 + 25.7818	- .4386 o	
71	+ .0099 + .0468	+ .612 + .2558	- .2069 + .0533	- 1.376 + .037	+ .8993 - .00186	- 1.2383 + .8993	+ .6754 + .00924	- .2034 + .01936	0 + .13795			+ .0005 + .01005	
72	- .012 + .0154	- .0099 + .0842	+ 1.376 + .0176	- .0122 + .0006	+ 1.2383 + .00576	+ .8993 + .00305	+ .00924 + .00005	+ .01936 + .04545	0 + .04545			+ 25.7818 + .0066	
73	+ 1.8539 + 1.8482	+ .3396 + .34375	- .2681 - .24649	- .3865 - .39877	+ .12674 + 1.4314	- .0135 - .03583	+ .1331 + .14073	- .06703 - .05683	0 + 1.00361	+ 3.1797 (1) (2) (3) (4) (5) (6) (7) (8) (9) (10)	- 1.4456 - .000681 + .001089 + .001531 + .007626 + .005934 + .007663 + .003851 + .001200 + .00078 - .000256	+ .4386 + .000138 + .00205 + .002314 + .004645 + .02143 + .01527 + .003851 + .01147 + .00078 + .00235	
74	+ .0869 + .0842	+ .3562 + .0154	- .2601 + .0122	- 1.339 + .0176	- .90116 + .00576	- 1.2558 + .0066	+ .66616 + .00665	- .22276 + .00395	+ .13795 + .04545	0 + .04545	+ 25.7213 - .0000	+ .01995 + .01005	
75	+ .0842 + .2558	+ .0154 + .0743	+ .0122 + 1.3936	- .0176 + .0533	+ .00576 + .0175	- .0066 - .00186	+ .00665 + .01836	+ .00395 + .00924	+ .04545 + .13795			+ 25.7732 - .0695	
76	+ .1211 o	+ .3716 o	- .2723 o	- 1.3566 o	- .8054 o	- 1.2564 o	+ .6722 o	- .2258 o	+ .13795 o	+ .04545 o	+ 1.00 o	+ 25.7147 o	
77	- .3716 o	+ .1211 o	+ 1.3566 o	- .2723 o	+ 1.2564 o	- .8054 o	+ .2258 o	+ .6722 o	- .04545 o	+ .13795 o	0 o	+ 25.7147 o	
78	- .3716 o	+ .1211 o	+ 1.3566 o	- .2723 o	+ 1.2564 o	- .8054 o	+ .2258 o	+ .6722 o	- .04545 o	+ .13795 o	0 o	+ 25.7147 o	
Case I	+ 1.6926	0	- .5202	- .33746	- .0294	- .0171	+ .1879	- .08888	+ .1081	+ .58125	+ .004700	- .01145	
II	+ 1.6926	+ .33746	- .5202	+ .0294	- .0171	+ .0294	+ .1879	- .08888	+ .1081	+ .58125	+ .01445	+ .004700	
III		+ 1.8675		+ .1636	- .02398	- .3407	+ .87417	- .1254	+ .07752	- .01059	+ .052756	+ .052756	
IV			+ 1.8675	+ .1636	- .02398	- .3407	+ .87417	- .1254	+ .07752	- .01059	+ .052756	+ .052756	
V				+ .3308	0	+ .02144	+ .07044	- .01127	+ .04502	- .04882	+ .04882	+ .04882	
VI					0	+ .07044	+ .02144	- .04502	- .01127	- .04882	+ .04882	+ .04882	
VII							+ .6713	0	- .01787	+ .04426	+ .02514	+ .008781	
VIII								+ .6713	- .01787	+ .04426	+ .02514	+ .008781	
IX									0	+ .0058846	+ .0017075	+ .0017075	
X										+ .3153	+ .0017075	+ .0053646	
XI											0	0	
XII												+ .028888	

\* As the coefficient of  $k_{11}$  vanishes no elimination is required, and this equation gives  $k_{11}$  direct.

**TABLE LXII.**  
Verification of solution of 12 equations.

+ .3554460	0	- .1692420	- .0706886	- .0061740	- .0035910	+ .0304500	- .0186648	+ .0227010	+ .1220625	+ .0009889	- .00303450
0	+ .8124480	+ .1619808	- .2496960	+ .0082080	- .0141120	+ .0426624	+ .0901920	- .2790000	+ .0518880	+ .00693600	+ .00228032
+ .1872720	- .1218556	- .6723000	0	- .0589060	- .0086328	+ .1226320	- .3147012	+ .0431440	+ .0279072	+ .00361240	- .01890216
+ .3779562	+ .526240	0	- 2.0016000	+ .0268576	- .1832320	+ .9790704	+ .3815840	- .0689224	+ .1404480	+ .05908672	+ .01186080
+ .1452360	- .0841740	- .8081840	+ .1184612	- 1.6341520	0	- .1059136	- .3479736	+ .0556736	- .2223988	+ .17201090	- .24136440
- .0767790	- .1320060	+ .1076702	+ .7345640	0	+ 1.4852920	+ .3162756	+ .0062656	- .2021398	- .0506023	- .21038140	- .15634180
- .2893960	- .1368752	+ .5246780	+ 1.3462218	- .0330176	+ .1084776	- 1.0398020	0	+ .0275198	- .0681604	- .01025560	- .01352274
+ .1084336	- .2292384	- 1.0664874	+ .4156540	- .0859368	- .0261568	0	- .8180860	+ .0534972	+ .0218014	+ .01071232	- .03189080
+ .1470160	- .7905000	- .1705440	+ .1054272	- .0153272	- .0612272	- .0243032	- .0601936	+ .4288080	0	+ .00729588	- .00240390
- .4126875	- .0767510	+ .6550392	+ .0890340	- .031042	+ .0080017	- .0314248	+ .0126877	0	- .2239630	- .00125493	- .06390897
0	0	0	0	0	0	0	0	0	0	0	0
- .5414415	+ .1764623	+ 1.97676732	- .3068073	+ 1.8307842	- 1.3047054	+ .32002407	+ .9704678	- .06622823	+ .20101158	0	+ 1.45713338
+ .0010848	+ .00018843	- .00062188	+ .0004903	+ .0003820	+ .0001141	+ .00114887	- .0003241	- .00034663	+ .00009416	+ .00015156	+ .99999141

16. The residuals in table LXII are sufficiently small. Accordingly the values of  $k_9 \dots k_{20}$  have been found satisfactorily for the 12 latter cases. It remains to find their values for the first 8 cases, and also values of  $k_1$  to  $k_8$  for all cases. In this the work is much simplified by the known symmetry of the solution. The introduction of cases 1 to 8 — *i.e.* giving the R.H.S. of the first 8 equations values 1, 0 . . . , (case 1) 0, 1, . . . (case 2) etc. causes the R.H.S. of the latter 12 equations to take the values  $-\sum^s (s, t) {}_r K_s$  for case  $r$  and equation  $t, s$  being taken from 1 to 8: and these quantities accordingly have to be found for each of the cases 1 to 8. Values of these quantities with sign reversed have already been given in table LIX. It is necessary then to combine these cases 9 to 20 in such a way as to give the solution for  $k_9$  to  $k_{20}$  for these related cases. For case  $r$  the value of  $k_u$  is  ${}_r k_u = -\sum^t k_u \left\{ \sum^s (s, t) {}_r K_s \right\}$ ,  $t$  being given all values from 9 to 20: but in fact  $\sum^s (s, t) {}_r K_s$  vanishes for values of  $t$  above 16. The process is carried out in table LXIII.

The next step is to find  ${}_u k_u$  for values of  $t$  and  $u$  from 1 to 8. Having found values of  ${}_u k_u$  for all values of  $u$  and values of  $t$  from 9 to 20, it is possible to write down values of  ${}_u k_t$  by symmetry. Equation (15) then enables the remaining quantities to be found as is done in table LXIV. The symmetry occurring largely simplifies the process while still affording a check.

This leads up to the solution of the combined 20 equations for all the 20 fundamental cases. The results of the solution, compiled from tables LXI, LXIII, LXIV, are given in table LXV. The solution of these 20 equations enables the probable errors of the N.W. Quadrilateral after adjustment of circuit conditions only, to be written down. By the incorporation of the next three equations, corresponding results can be given for the case when circuits and base line closures have been made (the actual adjustment carried out). Finally by incorporation of the last three equations the corresponding results obtainable if Laplace closures were introduced at each extra base can be given. Accordingly before giving the application of table LXV, the further solution of 23 and 26 equations will be carried out.

17. In table LXV are also shown certain multipliers. They are the coefficients of  $k_{21}, k_{22}, k_{23}$  in table XLVII. It is necessary to find  $k_u$  in terms of  $k_{21}, k_{22}, k_{23}$ , for all values of  $u$  between 1 and 20 by means of (15), with a view to substituting in equations 21, 22, 23. For this values of  $\sum^s {}_s k_r (s, t)$  are required, where  ${}_s k_r$  are the values given in table LXV and  $t$  has values 21, 22, 23, so that  $(s, t)$  are the multipliers just alluded to. Each of these products  ${}_s k_r (s, t)$  are given in table LXVI, and their sums  $\sum^s {}_s k_r (s, t)$  for values of  $s$  from 1 to 20, the latter in old face type. These are the coefficients of  $-k_u$  in the expressions for  $k_1 \dots k_{20}$  as found from the first 20 equations. These quantities have to be substituted in equations 21, 22, 23, and accordingly multiplied by the respective coefficients. The process is carried out in table LXVII, where the coefficients of the three equations giving  $k_{21}, k_{22}, k_{23}$ , are formed. The process has been carried out with full accuracy to illustrate the complete symmetry of the resulting equations. The solution of these equations is very simple and is given in table LXVIII, with verification at the foot of the table.

TABLE LXIII.

r	t	9	10	11	12	13	14	15	16	17 to 20	18	19	20
1	1	+ .1516	- .0778	+ .2226	- .2509	+ .6975	- .0808	+ .0513	- .1594	Columns			
2	2	+ .0778	+ .1516	- .0418	+ .0326	+ .0268	- .0675	+ .1528	+ .0513	17 to 20			
3	3	+ .0432	+ .0822	- .0389	+ .0486	- .0222	- .0357	- .0005	- .0005	comprise			
4	4	- .0259	- .0839	+ .0418	- .0116	+ .0083	+ .0038	+ .0371	+ .0371	only			
5	5	+ .0682	+ .2502	- .0305	- .0395	- .0169	- .0637	- .0345	- .0345	zeros			
6	6	+ .0252	- .1433	- .0363	- .3847	- .0824	- .0824	+ .1526	+ .1453				
7	7	+ .1433	+ .0252	+ .3867	- .4363	- .1500	- .0034	- .1453	+ .1826				
8	8	0	0	- .0348	+ .0480	+ .1636	- .0240	+ .0214	- .0704	K <sub>r</sub> to K <sub>r</sub>			
9	9	0	- .3008	0	- .0348	- .0240	- .1636	- .0704	+ .0214	are not			
10	10	0	0	- .0480	- .0348	- .0240	- .1636	- .0704	+ .0214	required			
11	11	0	0	- .0348	- .0240	- .1636	- .0704	+ .0214	- .0704				
12	12	0	0	- .0240	- .1636	- .0704	+ .0214	- .0704	+ .0214				
13	13	0	0	- .0704	+ .0214	- .0704	+ .0214	- .0704	+ .0214				
14	14	0	0	- .0214	- .0704	+ .0214	- .0704	+ .0214	- .0704				
15	15	0	0	- .0704	+ .0214	- .0704	+ .0214	- .0704	+ .0214				
16	16	0	0	- .0214	- .0704	+ .0214	- .0704	+ .0214	- .0704				
17	17	0	0	- .0704	+ .0214	- .0704	+ .0214	- .0704	+ .0214				
18	18	0	0	- .0214	- .0704	+ .0214	- .0704	+ .0214	- .0704				
19	19	0	0	- .0704	+ .0214	- .0704	+ .0214	- .0704	+ .0214				
20	20	0	0	- .0214	- .0704	+ .0214	- .0704	+ .0214	- .0704				
1	1	+ .05015	0	- .01123	- .01297	+ .01506	- .00774	+ .00110	- .01276	+ .03269			
2	2	+ .02515	0	- .00573	- .00648	+ .00774	- .00234	+ .00327	+ .00561	+ .04236			
3	3	+ .01133	0	+ .00145	- .00195	- .00141	+ .00153	- .00176	- .00267	+ .01221			
4	4	+ .00668	0	+ .00139	- .00204	- .00384	- .00021	- .00020	- .00251	+ .00021			
5	5	- .06576	0	+ .02160	- .03127	- .10883	- .01009	- .01282	- .05735	- .15651			
6	6	- .01901	0	- .02225	- .03036	- .00661	- .00552	- .01076	- .01214	- .11723			
7	7	+ .06833	0	+ .01518	- .01801	- .00656	- .00382	+ .00301	- .01023	+ .01156			
8	8	+ .04740	0	- .01346	- .02134	- .02449	- .04008	- .00311	+ .01385	- .00227			
9	9	0	0	- .02575	+ .00673	+ .00234	- .00504	- .03681	- .00222	- .04236			
10	10	0	0	- .05015	- .01227	- .04074	- .01586	- .00110	- .00110	- .03349			
11	11	0	0	- .00688	+ .00139	- .00021	- .00304	+ .00251	- .00420	- .00122			
12	12	0	0	- .01133	+ .00195	- .00653	- .01341	- .00067	- .00770	- .01022			
13	13	0	0	- .01891	+ .00306	- .02225	- .00091	+ .01218	- .01744	+ .11723			
14	14	0	0	- .06576	+ .03127	- .04074	- .11683	- .05735	- .01292	- .11684			
15	15	0	0	- .07400	+ .02134	- .00095	- .02469	- .01285	- .00511	- .00227			
16	16	0	0	- .00833	- .01761	- .01518	- .00382	- .00056	- .00056	+ .01356			
17	17	0	0	- .00294	+ .00255	0	- .00109	+ .00134	+ .00134	+ .01436			
18	18	0	0	- .00271	0	- .00093	- .00165	+ .00059	- .00045	+ .01436			
19	19	0	0	- .00116	- .00103	0	- .00117	- .00069	- .00068	+ .00006			
20	20	0	0	- .00007	+ .00155	0	- .00045	- .00025	- .00031	+ .00050			
1	1	+ .06062	- .02294	- .02415	0	+ .00686	- .00020	- .01864	- .00717	- .03422			
2	2	+ .02020	+ .02018	- .01448	0	- .00094	- .03116	- .02127	- .04627	+ .05785			
3	3	+ .00966	+ .07001	- .01697	0	+ .00004	- .00707	+ .00477	+ .00128	- .01528			
4	4	+ .00297	- .00123	+ .01304	0	+ .00160	- .00018	- .00779	- .00161	+ .00529			
5	5	+ .06741	+ .06271	0	- .00278	+ .00516	- .06033	+ .00445	- .00500	+ .00236			
6	6	+ .02721	- .00741	+ .00578	0	- .00033	- .00133	+ .00134	+ .00134	+ .00436			
7	7	- .04077	- .00010	- .00103	0	+ .00069	- .00117	- .00063	- .00023	- .00006			
8	8	- .00910	+ .00119	0	- .00103	+ .00117	- .00029	- .00008	- .00029	- .00019			
9	9	- .01288	- .00269	0	- .02188	- .03416	- .00604	- .00527	- .02227	- .05785			
10	10	- .04584	+ .00602	0	- .02105	- .00250	- .00286	- .00717	- .01059	- .03422			
11	11	+ .00123	+ .00159	0	- .01501	- .00106	- .00169	+ .00101	+ .00101	- .00250			
12	12	+ .00701	- .00088	0	- .01007	- .00707	- .00091	- .00128	- .00091	- .01528			
13	13	+ .02840	- .00197	- .00345	- .01234	+ .18408	- .01234	- .01058	0	+ .02725			
14	14	+ .01274	+ .00301	- .00208	+ .00213	- .00723	0	- .05401	+ .01488	+ .08106			
15	15	+ .00559	+ .00605	+ .00044	+ .00211	- .01006	0	+ .01201	- .00830	+ .01519			
16	16	+ .00701	- .00088	0	- .00642	- .00221	+ .04116	0	- .00324	- .03121			
17	17	0	0	- .00294	- .00345	- .01234	+ .18408	0	- .02714	- .13059			
18	18	0	0	- .00145	- .00208	- .00723	0	0	+ .20134	+ .71901			
19	19	0	0	- .00070	- .00070	- .00070	0	0	+ .27176	- .62591			
20	20	0	0	- .00050	- .00050	- .00050	0	0	- .00221	- .12702			
1	1	+ .02343	+ .00050	- .00110	- .02343	- .29181	0	+ .03160	+ .15663	- .07377			
2	2	- .04241	+ .00145	+ .00650	- .05278	- 1.20049	0	0	0	- .23002			
3	3	- .04495	- .06022	- .00078	- .02579	- .00801	0	0	0	- .23002			
4	4	+ .00412	+ .00044	- .00162	- .01042	- .06625	0	0	0	- .04834			
5	5	+ .02343	+ .00050	- .00110	- .02343	- .29181	0	0	0	- .07377			





TABLE LXVI. Values of \$s\_k (s, t)\$ and \$\Sigma s\_k (s, t)\$, \$s\$ from 1 to 20, the latter in old face type.

Table with 20 columns (t, s) and rows for coefficients of k1 to k20 and their sums. The values are small integers, some in italics, with a 'Sum' row at the bottom.

TABLE LXVII.

TABLE LXVIII.

Large table with columns for Equation 21, Equation 22, Equation 23, and Left Hand Side. Includes coefficients of various terms and solution values.

TABLE LXIX.

18. To pass to the complete solution of the 23 equations the process is precisely similar to that already followed after the solution of 12 equations in §15: the notation given in the corresponding tables concerned—viz tables LXIX, LXX—explains itself and the solution, keeping only 4 decimal places, is exhibited in table LXXI.

In this table are given also the value of  $\Sigma^s k_r$  for all values of  $s$  from 1 to 23. These obviously should correspond to the case where all the R. H. S. of the 23 equations are unity. This solution, which depends on all the fundamental cases, can be verified in the original equations, and affords a complete check of the work. The substitution is carried out in table LXII. It will be seen that the values of the R.H.S. obtained by substitution differ slightly from unity, the greatest difference being .035, showing that no gross error has been committed. It is of interest to consider to what these differences may be attributed. Each value of  $k_r$  is given to four decimal places: and accordingly may be in error by .00005. The sum of 23 such quantities may be wrong by .00115: and in substituting in the original equation such an error is multiplied by coefficients, the largest of which is 111.51, which would admit of the corresponding term containing an error of .1. That this extreme value should be obtained is most unlikely: but it is clear that the actual discrepancies obtained may easily be attributable to this cause. This line of argument shows how many figures it will be necessary to keep to be absolutely sure of all errors being less than a stated amount. For the present object, the solution as found may be considered sufficiently precise.

19. Table LXXI also contains the necessary multipliers to proceed to the solution of the complete 26 equations. The process is exactly similar to that already described in passing from 20 to 23 equations. All results are given in tables LXXIII—LXXIX, and the verification, similar to that of LXXII, in table LXXX. This completes the solution of the equations of table XLVII.

r	t	21	22	23	
1		-.48308	-.30540	+.00933	
2		+.12308	+.04835	-.00880	
3		+.05180	+.00449	+.00278	
4		+.03065	-.00919	+.00261	
5		+.10211	-.06029	-.05482	
6		+.07813	+.01817	+.08229	
7		+.02003	-.01390	-.02325	
8		+.01095	-.01728	-.01521	
9		+.05338	-.01691	-.15830	
10		+.02555	-.01085	-.00918	
11		+.01111	+.00165	+.00849	
12		+.02645	-.35563	+.16911	
13		+.08456	+.08105	+.06785	
14		+.03999	-.03441	-.07504	
15		+.00796	-.00138	-.02435	
16		+.00391	-.10087	+.43200	
17		+.03594	+.06461	-.25455	
18		+.00554	+.03223	+.15147	
19		+.01644	+.06753	-.00328	
20					
	$\Sigma^s k_r$ (s, t)				
	from table LXVI				
	t	21	22	23	Values of $k_u$
21	1	+.137317	+.292777	+.00921	
	2	+.292777	+.125662	+.05870	
	3	+.00320	+.05870	+.12448	
21	1	-.064587	-.115761	+.000030	= -.780318
	2	+.169010	+.014155	-.000028	= +.183137
	3	-.071213	+.001315	+.000009	= -.069689
	4	+.054446	-.029049	+.000008	= +.025414
	5	+.140214	-.019408	-.000175	= +.126631
	6	-.107286	+.014103	+.000203	= -.092020
	7	+.035744	-.094070	-.000074	= +.031600
	8	-.015038	+.005059	-.000109	= -.020144
	9	-.073300	+.003203	-.000506	= -.077009
	10	+.035084	-.003177	-.000317	= +.031590
	11	-.034123	+.001438	+.000200	= -.032485
	12	-.005614	+.000483	+.000027	= -.005134
	13	-.036320	-.095395	+.000349	= -.131306
	14	+.116115	+.023729	+.000313	= +.140157
	15	-.054913	+.010074	-.000240	= -.046079
	16	+.010930	-.068958	-.000080	= -.048108
	17	+.005369	-.047039	-.001350	= -.043020
	18	-.049352	-.018916	-.000750	= -.069018
	19	+.007607	+.000946	+.000430	= +.008983
	20	-.022575	-.019771	-.000203	= -.012549
22	1	-.141695	-.445070	+.000548	= -.586217
	2	+.036034	+.054124	-.000517	= +.089041
	3	-.015183	+.005054	+.000163	= -.009966
	4	+.011608	-.111650	+.000153	= -.098889
	5	+.029895	-.074617	-.003218	= -.017944
	6	-.022874	+.054221	+.001830	= +.036177
	7	+.007621	-.015616	-.001365	= -.009394
	8	-.003206	-.019451	-.000893	= -.025554
	9	-.015628	-.012314	-.000292	= -.037234
	10	+.007480	-.012213	-.005822	= -.010555
	11	-.007275	+.005527	+.003662	= +.001914
	12	-.001203	+.001857	+.000498	= +.001152
	13	-.007744	-.366536	+.006405	= -.367875
	14	+.024757	+.091232	+.005741	= +.121733
	15	-.011708	+.038733	-.001405	= -.022624
	16	+.002330	-.226677	-.001465	= -.225812
	17	+.001145	-.180853	-.024775	= -.204483
	18	-.010522	-.072726	-.013778	= -.079706
	19	+.001822	+.003636	+.007893	= +.013151
	20	-.004813	-.076013	-.003715	= -.084641
23	1	-.001549	-.029210	+.010491	= -.014268
	2	+.000394	+.002838	-.000895	= -.000566
	3	-.000166	+.000264	+.003126	= +.003224
	4	+.000127	-.005822	+.002935	= -.002760
	5	+.000327	-.003891	-.061044	= -.005208
	6	-.000250	+.002828	+.002533	= +.005111
	7	+.000083	-.000816	-.026144	= -.026877
	8	-.000035	-.001614	-.017109	= -.018152
	9	-.000171	-.000642	-.178095	= -.178818
	10	+.000082	-.000637	-.111529	= -.112081
	11	-.000080	+.000288	+.070145	= +.070959
	12	-.000013	+.000097	+.009547	= +.009631
	13	-.000085	-.019114	+.123992	= +.103493
	14	+.000271	+.004768	+.110030	= +.115059
	15	-.000128	+.002020	-.084381	= -.082489
	16	+.000026	-.011821	-.028056	= -.030852
	17	+.000013	-.000431	-.474508	= -.481016
	18	-.000115	-.003793	-.263747	= -.267655
	19	+.000018	+.000100	+.151200	= +.151417
	20	-.000053	-.003904	-.071157	= -.075174

TABLE LXX.

Table with multiple columns for variables t, 21, 22, 23, and rows for r=1, 2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12, 13, 14, 15, 16, 17, 18, 19, 20. Includes sub-tables for r=1, 2, 3, 4, 5, 6, 7, 8, 9, 10 and r=1, 2, 3, 4, 5, 6, 7, 8, 9, 10. Contains numerical values and mathematical expressions like r\_ku Σ r\_k(s,t) from Table LXXVI and LXX.



TABLE LXXI. Values of  $s_r$  for 23 conditions.

Table with 23 columns (1-23) and 23 rows (1-23). Each cell contains a numerical value. The last column is labeled 'Multipliers' with sub-columns K23, K22, K21, K20, K19, K18, K17, K16, K15, K14, K13, K12, K11, K10, K9, K8, K7, K6, K5, K4, K3, K2, K1. The values are arranged in a grid with some bolded cells.

TABLE LXXII. Verification of solution in Table LXXI.

Table with 23 columns (1-23) and 23 rows (1-23). Each cell contains a numerical value. The last column is labeled 'Multipliers' with sub-columns K23, K22, K21, K20, K19, K18, K17, K16, K15, K14, K13, K12, K11, K10, K9, K8, K7, K6, K5, K4, K3, K2, K1. The values are arranged in a grid with some bolded cells.

TABLE LXXIII. Values of  $k_r$  ( $s, t$ ) and  $\Sigma k_r$  ( $s, t$ ),  $s$  from 1 to 23, the latter in old face type.

Table with 24 columns (labeled 1 to 24) and 24 rows (labeled 1 to 24). The table contains numerical values for coefficients k\_r and their sums. The last column (24) contains values in an old face type font.

TABLE LXXIV.

Table with 24 columns (labeled 24 to 24) and 24 rows (labeled 1 to 24). The table is divided into three sections: Equation 24, Equation 25, and Equation 26. It contains numerical values for coefficients k\_s and their sums.

TABLE LXXV.

Table with 6 columns (labeled No., Right Hand Side, k\_24, k\_23, k\_22, k\_21) and 6 rows (labeled 1 to 6). The table contains numerical values for coefficients k\_24, k\_23, k\_22, and k\_21, along with their sums and a 'Solution' row.

TABLE LXXVI.

Table with multiple columns for variables r, t, and u, and rows for values 1 through 26. It includes numerical data and labels like Σk, k, and r\_k.

TABLE LXXVII.

Large table with columns for t, 21, 22, 23 and rows for r=1, u=1, and various numerical entries. It contains references to other tables like LXXIII, LXXVI, and LXXVII.

TABLE LXXVII.—(Continued).

Table with multiple columns and rows containing numerical data. Columns are labeled with 't', '21', '22', '23', and 'rku from LXXI'. Rows are grouped by a large 't' value (5, 6, 7, 8) and a smaller 't' value (12, 13, 14, 15, 16, 17, 18, 19, 20). Each row contains several columns of numbers, some with signs, and some with equals signs, representing data points or calculations.



TABLE LXXVIII. Values of k<sub>r</sub> for 26 conditions.

Table with 26 columns (1-26) and 26 rows (1-26). Each cell contains a numerical value representing k\_r for a specific condition and column.

TABLE LXXIX. Verification of solution in Table LXXVIII.

Table with 26 columns (1-26) and 26 rows (1-26). Each cell contains a complex mathematical expression involving terms from Table LXXVIII, used for verification.

Probable errors after adjustment.

20. Having obtained the solution of the normal equations corresponding to either 20, 23 or 26 conditions it is now possible to determine the probable errors of side, azimuth, easting and northing as explained briefly at the end of Chapter VII. This is first done for the point  $U_1$  for which the necessary quantities forming the R. H. S. of the normal equations have already been found and given in table *L*. The computations are given in full for this point: after which other points are considered in less detail. It is necessary to form the quantities  $[u_f']$  and  $[u_f] - k_1 [uaf] - k_2 [ubf] - \dots$  which may be denoted by  $u_f$  and  $u_F$  respectively [*vide* Chapter VII, equation (14)]. These are the reciprocal weights before and after adjustment and their square roots when multiplied by 33.2, 1.575, 4.03, as the case may be, give the probable errors in 7th place of log side, azimuth (in seconds), easting or northing (in feet) as explained in § 8 of Chapter VII. The factor  $k = \sqrt{\frac{u_F}{u_f}}$  shows the improvement (or otherwise) caused by the adjustment.

Side closure at  $U_1$

$$\begin{aligned}
 [u_f'] &= (u\ 1\ f) + (u\ 2\ f) = 2.61 + 2.30 = 4.91 \\
 [uaf] &= 2.61 \\
 [ubf] &= 0
 \end{aligned}$$

etc. as given in Table *L*.

For 20 conditions

R.H.S. of normal equation (13) are 2.61 . . . . 0

$$\begin{aligned}
 k_r = & 2.61\ 1k_r + 2.08\ 3k_r - 6.80\ 4k_r + 2.30\ 6k_r + .02\ 7k_r - 1.49\ 8k_r + 1.99\ 13k_r \\
 & + 1.85\ 15k_r - 2.59\ 16k_r.
 \end{aligned}$$

and this is required for values of  $r$  1, 3, 4, 5, 7, 8, 13, 15, 16,  ${}_s k_r$  being taken from table *LXV*.

Putting in the coefficients of 2.61 etc. from table *LXV* the computation of  $k_r$  stands as follows:—

TABLE LXXX.

	r	1	3	4	5	7	8	13	15	16
s	[ $u_s r'$ ]	Values of [ $u_s r'$ ] ${}_s k_r$								
1	+ 2.61	+ 3.013	- .165	+ .760	- .347	- .020	- .035	+ .097	+ .112	- .019
3	+ 2.08	- .131	+ .213	0	+ .025	- .013	+ .000	- .032	- .007	- .015
4	- 6.80	- 1.979	0	- .695	+ .169	+ .002	+ .041	- .028	- .050	+ .022
5	+ 2.30	- .306	+ .030	- .057	+ 1.896	+ .048	+ .386	- .828	- .499	- .111
7	+ .02	.000	.000	.000	.000	+ .001	0	+ .001	.000	+ .001
8	- 1.49	+ .020	.000	+ .009	- .250	0	- .085	+ .113	+ .053	+ .018
13	+ 1.99	+ .074	- .030	+ .008	- .716	+ .086	- .151	+ 3.716	- .678	+ 1.740
15	+ 1.85	+ .080	- .006	+ .014	- .401	- .023	- .066	- .630	+ 1.242	0
16	- 2.59	+ .019	+ .019	+ .008	+ .125	- .092	+ .032	- 2.264	0	- 1.739
Sum = $k_r$		+ 0.790	+ 0.061	+ 0.047	+ 0.504	- 0.011	+ 0.122	+ 0.145	+ 0.173	- 0.103
Multiplier = [ $u_r f$ ]		+ 2.61	+ 2.08	- 6.80	+ 2.30	+ .02	- 1.49	+ 1.99	+ 1.85	- 2.59
[ $u_r f'$ ] $k_r$		+ 2.062	+ .127	- .320	+ 1.159	.000	- .182	+ .289	+ .320	+ .267
$u_f = 4.91$		$u_F = u_f - \sum [u_r f'] k_r = 4.91 - 3.72 = 1.19$								
		$K = \sqrt{\frac{u_F}{u_f}} = .49$								

TABLE LXXXI.

Values of  ${}_r k_r$  from Table LXXI.

For 23 conditions

Side closure

r	1	3	4	5	7	8	13	15	16	21	22	
s	[ $urf$ ]	Values of [ $urf$ ] ${}_r k_r$										
1	+2.61	+4.603	- .066	+ .891	- .451	- .050	+ .014	+ .045	+ .144	+ .274	-2.037	-1.530
3	+2.08	- .053	+ .220	- .004	+ .014	- .016	+ .002	- .020	- .002	- .012	- .145	- .021
4	-6.80	-2.164	+ .012	- .770	+ .105	- .013	+ .031	- .242	- .021	- .116	- .173	+ .679
5	+2.30	- .398	+ .015	- .036	+1.940	+ .060	+ .387	- .816	- .502	- .083	+ .277	- .110
7	+0.02	.000	.000	.000	+ .001	+ .001	.000	+ .001	.000	+ .001	+ .001	.000
8	-1.49	- .008	- .002	+ .007	- .251	- .000	- .087	+ .104	+ .051	+ .011	+ .030	+ .035
13	+1.99	+ .492	+ .019	+ .071	- .706	+ .085	- .138	+3.984	- .708	+1.880	- .261	- .732
15	+1.85	+ .102	- .002	+ .006	- .404	- .022	- .063	- .658	+1.258	- .005	- .083	+ .042
16	-2.59	- .272	+ .016	- .044	+ .093	- .099	+ .019	-2.447	+ .068	-1.858	+ .125	+ .585
21	+2.19	-1.709	- .153	+ .056	+ .264	+ .069	- .044	- .288	- .099	- .105	+3.007	+ .641
22	+0.73	- .428	- .007	- .073	- .035	- .007	- .017	- .269	+ .016	- .165	+ .214	+ .822
Sum = $k_r$		+ .165	+ .013	+ .044	+ .570	+ .008	+ .104	- .006	+ .145	- .178	+ .955	+ .411
Multiplier = [ $urf$ ]		+2.61	+2.08	-6.80	+2.30	+0.02	-1.49	+1.99	+1.85	-2.59	+2.19	+0.73
[ $urf$ ] $k_r$		+ .431	+ .027	- .299	+1.311	.000	- .155	- .012	+ .268	+ .461	+2.091	+ .300
$u_f = 4.91$		$u_F = u_f - \sum [urf] k_r = 4.91 - 4.42 = .49$								$K = \sqrt{\frac{u_F}{u_f}} = .32$		

Azimuth closure at  $U_1$

[ $uff$ ] = 2.61 + 2.30 = 4.91

$k_r = 2.61 {}_2 k_r + 6.80 {}_3 k_r + 2.08 {}_4 k_r + 2.30 {}_6 k_r + 1.49 {}_7 k_r + 0.02 {}_8 k_r + 1.99 {}_{13} k_r + 1.85 {}_{15} k_r + 2.59 {}_{16} k_r$

This holds for either 20 or 23 conditions. But the values of  ${}_r k_r$  are different in the two cases. Values of r required are 2, 3, 4, 6 . . . . 16.

For 20 conditions.

Comparing terms in  $k_1$  side closure and  $k_2$  azimuth closure

$2.61 {}_1 k_1 = 2.61 {}_2 k_2$  since  ${}_1 k_1 = {}_2 k_2$   
 $2.08 {}_3 k_1 = 2.08 {}_4 k_2$  „  ${}_3 k_1 = {}_4 k_2$   
 $-6.80 {}_4 k_1 = 6.80 {}_3 k_2$  „  ${}_4 k_1 = - {}_3 k_2$   
 $2.30 {}_5 k_1 = 2.30 {}_6 k_2$  „  ${}_5 k_1 = {}_6 k_2$

etc.

so that  $k_2$  (azimuth) is same as  $k_1$  (side).

As regard's  $k$  (side) and  $k_4$  (azimuth)

$2.61 {}_1 k_3 = 2.61 {}_2 k_4$

etc.

and  $k_1$  (azimuth) =  $k_3$  (side).

For  $k_4$  (side) and  $k_3$  (azimuth)

$2.61 {}_1 k_4 = - 2.61 {}_2 k_3$

so that  $k_3$  (azimuth) =  $k_4$  (side)

etc.

No further computation in this case is necessary; and the multipliers being similarly related numerically and with regard to sign the same value of  $k$  holds for azimuth as for side.

*For 23 conditions.*—The symmetry is lost by the introduction of base-line conditions without corresponding Laplace conditions and the values of  $k_2, k_3, k_4, k_6 \dots k_{16}$  have to be computed.

**TABLE LXXXII.**

*For 23 conditions*

*Azimuth closure*

r	2	3	4	6	7	8	14	15	16	
s	[usf]	Values of [usf] <sub>s</sub> k <sub>r</sub>								
2	+2.61	+3.086	-.784	-.169	-.374	+ .045	-.028	+ .155	+ .009	+ .069
3	+6.80	-2.042	+ .720	-.012	+ .205	-.053	+ .007	-.071	-.007	-.040
4	+2.08	-.135	-.004	+ .236	+ .013	+ .004	-.010	-.045	+ .006	+ .036
6	+2.30	-.330	+ .069	+ .014	+1.940	-.398	+ .046	-.819	+ .106	-.523
7	+1.49	+ .025	-.012	+ .003	-.258	+ .088	.000	+ .112	-.018	+ .057
8	+0.02	.000	.000	.000	.000	.000	+ .001	-.001	-.001	-.001
14	+1.99	+ .118	-.209	-.043	-.708	+ .150	+ .075	+3.790	-1.760	-.732
15	+2.59	+ .009	-.003	+ .008	+ .119	-.031	-.089	-2.290	+1.765	-.008
16	+1.85	+ .049	-.011	+ .032	-.420	+ .071	-.014	-.680	-.005	+1.329
Sum = k <sub>r</sub>		+ .780	-.234	+ .069	+ .517	-.124	-.012	+ .153	+ .095	+ .187
Multiplier = [urf]		+2.61	+6.80	+2.08	+2.30	+1.49	+0.02	+1.99	+2.59	+1.85
[urf]k <sub>r</sub>		+2.039	-1.591	+ .144	+1.189	-.185	.000	+ .305	+ .246	+ .346
$u_f = 4.91 \quad u_F = u_f - \Sigma [urf]k_r = 4.91 - 2.49 = 2.42 \quad K = \sqrt{\frac{u_F}{u_f}} = .70$										

**TABLE LXXXIII.**

Values of  $k_r$  from Table LXXVIII.

*For 26 conditions*

*Side closure*

r	1	3	4	5	7	8	13	15	16	21	22	
s	[usf]	Values of [usf] <sub>s</sub> k <sub>r</sub>										
1	+2.61	+4.688	-.059	+ .861	-.486	-.058	+ .002	+ .712	+ .118	+ .296	-2.032	-1.557
3	+2.08	-.047	+ .244	0	-.002	-.014	-.002	-.032	+ .042	-.002	-.161	-.026
4	-6.80	-2.244	0	-.799	+ .141	-.006	+ .044	-.301	-.007	-.139	-.173	+ .706
5	+2.30	-.428	-.002	-.048	+1.979	+ .057	+ .399	-.808	-.529	-.080	+ .285	-.099
7	+ .02	0	0	0	+ .001	+ .001	0	+ .001	0	+ .001	+ .001	0
8	-1.49	-.001	+ .001	+ .010	-.258	0	-.089	+ .103	+ .055	+ .010	+ .029	+ .033
13	+1.99	+ .543	-.032	+ .088	-.698	+ .074	-.138	+ .059	-.761	+1.905	-.251	-.748
15	+1.85	+ .084	+ .038	+ .002	-.426	-.013	-.068	-.707	+1.346	0	-.110	+ .041
16	-2.59	-.294	+ .003	-.053	+ .090	-.095	+ .018	-2.479	0	-1.884	+ .131	+ .595
21	+2.19	-1.705	-.168	+ .056	+ .272	+ .067	-.042	-.276	-.131	-.111	+3.016	+ .643
22	+0.73	-.436	-.009	-.076	-.031	-.006	-.016	-.274	+ .016	-.168	+ .214	+ .824
Sum = k <sub>r</sub>		+0.160	+ .016	+ .041	+ .582	+ .007	+ .108	-.002	+ .149	-.172	+ .949	+ .412
Multiplier = [urf]k <sub>r</sub>		+2.61	+2.08	-6.80	+2.30	+ .02	-1.49	+1.99	+1.85	-2.59	+2.19	+0.73
[urf]k <sub>r</sub>		+ .418	+ .033	-.279	+1.339	.000	-.101	-.004	+ .276	+ .445	+2.078	+ .301
$u_f = 4.91 \quad u_F = u_f - \Sigma [urf]k_r = 4.91 - 4.45 = .46 \quad K = \sqrt{\frac{u_F}{u_f}} = .31$												

No special explanation is required for tables LXXXIV-LXXXVII, which are now given.

TABLE LXXXIV.

For 20 conditions

At  $U_1$

Easting closure

r	1	2	3	4	5	6	7	8	13	14	15	16	
s	[ $u_{sf}$ ]	Values of [ $u_{sf}$ ] $k_r$											
1	-.30	-.346	0	+.019	-.087	+.040	-.016	+.002	+.004	-.011	+.024	-.013	+.002
2	-6.72	0	-7.758	+1.956	+.425	+.367	+.803	-.090	+.050	-.545	-.250	+.049	+.289
3	-11.83	+.747	+3.443	-1.210	0	-.156	-.204	+.071	-.003	+.180	+.048	+.038	+.084
4	-4.59	-1.336	+.200	0	-.469	+.114	-.061	+.001	+.028	-.019	+.070	-.034	+.015
5	-3.51	+.467	+.191	-.046	+.087	-2.893	0	-.073	-.550	+1.264	-1.084	+.761	+.169
6	-12.15	-.663	+1.615	-.302	-.161	0	-10.016	+2.040	-.254	+3.754	+4.374	-.596	+2.634
7	-7.53	+.056	+.101	+.045	+.002	-.157	+1.264	-.430	0	-.326	-.571	+.093	+.268
8	+2.18	-.029	-.016	+.001	-.013	+.366	+.046	0	+.125	-.165	+.094	-.078	-.027
13	-3.13	-.117	-.254	+.048	-.013	+1.127	+.067	-.136	+.237	-5.845	0	+1.066	-2.736
14	-10.44	+.846	-.389	+.042	+.159	-3.236	+3.750	-7.01	-.452	0	-10.497	+9.126	+3.557
15	-16.22	-.699	+.117	+.053	-.118	+3.616	-.782	+.204	+.578	+5.526	+14.179	-10.889	0
16	-5.64	+.041	-.243	+.041	+.018	+.272	+1.223	-.201	+.070	-4.930	+1.922	0	-3.786
Sum = $k_r$	-1.033	-3.339	+.647	-.170	-.630	-3.017	+.503	-.206	-1.117	-.691	-.564	-.643	
Multiplier = [ $u_{sf}$ ]	-.30	-6.72	-11.83	-4.59	-3.51	-12.15	-7.53	+2.18	-3.13	-10.44	-16.22	-5.64	
[ $u_{sf}$ ] $k_r$	+.310	+22.438	-7.654	+7.90	+2.211	+36.657	-4.465	-.449	+3.496	+7.214	+9.148	+3.627	
$u_f = 93.29$ $u_F = u_f - \sum [u_{sf}]k_r = 93.29 - 73.31 = 19.98$ $K = \sqrt{\frac{u_F}{u_f}} = .46$													

TABLE LXXXV.

For 23 conditions

At  $U_1$

Easting closure

r	1	2	3	4	5	6	7	8	13	14	15	16	21	22	
s	[ $u_{sf}$ ]	Values of [ $u_{sf}$ ] $k_r$													
1	-.30	-.629	+.037	+.008	-.065	+.052	-.026	+.006	-.002	-.074	+.050	-.017	-.032	+.231	+.176
2	-6.72	+.835	-7.039	+2.017	+.436	+.278	+.064	+.115	+.073	-.310	-.399	-.024	-.179	-1.230	-.604
3	-11.83	+.300	+3.551	-1.253	+.021	-.078	-.356	+.092	-.013	+.116	+.124	+.012	+.070	+.827	+.118
4	-4.59	-1.461	+.298	+.008	-.520	+.071	-.028	-.009	+.021	-.163	+.098	-.014	-.078	-.117	+.439
5	-3.51	+.607	+.145	-.023	+.054	-2.960	+.680	-.092	-.501	+1.245	-1.084	+.766	+.126	-.423	+.168
6	-12.15	-1.046	+1.744	-.366	-.075	+.208	-10.221	+2.102	-.241	+3.741	+4.321	-.550	+2.760	+1.130	-.440
7	-7.53	+.145	-.129	+.059	-.014	-.197	+1.303	-.442	-.002	-.321	-.506	+.090	-.289	-.238	+.071
8	+2.18	+.012	-.024	+.002	-.010	+.367	+.043	0	+.127	-.152	+.083	-.075	-.016	-.044	-.051
13	-3.13	-.774	-.145	+.030	-.111	+1.110	+.064	+.133	+.218	-6.266	+.096	+1.114	-2.957	+.411	+1.152
14	-10.44	+2.048	-.620	+.110	+.223	-3.225	+3.712	-.781	-.396	+.322	-19.841	+9.231	+3.831	-1.444	-1.271
15	-16.22	-.895	+.057	+.016	-.050	+3.541	-.746	+.193	+.556	+5.773	+14.342	-11.031	+.047	+7.82	+.367
16	-5.64	-.592	+.150	+.033	-.096	+.203	+1.281	-.217	+.041	-5.325	+2.070	+.046	+4.046	+.271	-1.274
21	-.14	+.100	-.026	+.011	-.004	-.017	+.013	-.004	+.003	+.018	-.020	+.006	+.007	-.192	-.041
22	-.53	+.311	-.048	+.005	+.053	+.625	-.019	+.005	+.013	+.195	-.005	-.012	+.120	-.155	-.597
Sum = $k_r$	-.932	-3.363	+.657	-.188	-.522	-3.056	+.602	-.193	-1.205	-.732	-.497	-.636	-.259	+.047	
Multiplier = [ $u_{sf}$ ]	-.30	-6.72	-11.83	-4.59	-3.51	-12.15	-7.53	+2.18	-3.13	-10.44	-16.22	-5.64	-.14	-.53	
[ $u_{sf}$ ] $k_r$	+.280	+22.599	-7.772	+.863	+1.832	+37.130	-4.533	-.421	+3.772	+8.164	+8.061	+3.587	+.036	-.023	
$u_f = 93.29$ $u_F = u_f - \sum [u_{sf}]k_r = 93.29 - 73.57 = 19.72$ $K = \sqrt{\frac{u_F}{u_f}} = .46$															

TABLE LXXXVI.

For 23 conditions

At  $U_1$

Northing closure

r	1	2	3	4	5	6	7	8	13	14	15	16	21	22	
s	Values of $[u\ s\ f]_s k_r$														
1	+ 6.72	+11.852	- .835	- .171	+ 2.138	- 1.163	+ .579	- .130	+ .037	+ 1.661	- .132	+ .371	+ .071	- 5.244	- 3.939
2	- .30	+ .037	- .364	+ .000	+ .010	+ .012	+ .043	- .005	+ .003	- .014	- .018	- .001	- .008	- .056	- .027
3	+ 4.50	- .117	- 1.378	+ .486	- .008	+ .030	+ .138	- .036	+ .005	- .045	- .048	- .005	- .027	- .321	- .046
4	- 11.83	- 3.764	+ .768	+ .021	- 1.339	+ .183	- .073	- .022	+ .054	- .421	+ .253	- .037	- .202	- .300	+ 1.182
5	+ 12.15	- 2.102	- .503	+ .080	- .188	+ 10.217	- .208	+ .318	+ 2.046	- 4.310	+ 3.753	- 2.652	- .437	+ 1.465	- .582
6	- 3.51	- .302	+ .504	- .106	- .022	+ .060	- 2.963	+ .607	- .069	+ 1.081	+ 1.248	- .161	+ .797	- .326	- .127
7	- 2.18	+ .042	- .037	+ .017	- .004	- .057	+ .377	- .128	- .000	- .003	- .164	+ .026	- .084	- .060	+ .020
8	- 7.53	- .041	+ .082	- .008	+ .035	- 1.208	- .110	- .002	- .437	+ .523	- .285	+ .258	+ .055	+ .151	+ .178
13	+ 10.44	+ 2.581	+ .182	- .101	+ .372	- 3.703	- 3.214	+ .445	- .726	+ 20.901	- .322	- 3.716	+ 0.882	- 1.371	- 3.841
14	- 3.13	+ .613	- .106	+ .034	+ .067	- .967	+ 1.113	- .235	- .119	+ .096	- 5.049	+ 2.768	+ 1.140	- .439	- .391
15	+ 5.64	+ .311	+ .020	- .006	+ .017	- 1.231	+ .259	- .067	- .193	- 2.007	- 4.087	+ 3.836	- .016	- .251	+ .127
16	- 16.22	- 1.703	- .431	+ .096	- .277	+ .584	+ 3.685	- .623	+ .118	- 15.321	+ 5.053	+ .047	- 11.636	+ .780	+ 3.662
21	+ 4.73	- 3.601	+ .866	- .331	+ .120	+ .570	- .439	+ .149	- .095	- .621	+ .663	- .213	- .228	+ 6.405	+ 1.385
22	+ 3.69	- 2.163	+ .332	- .037	- .360	- .177	+ .134	- .035	- .087	- 1.359	+ .449	+ .083	- .833	+ 1.080	+ 1.163
Sum = $k_r$	+ 1.553	- .670	+ .063	+ .561	+ 3.120	- .708	+ .236	+ .537	+ .072	+ .414	+ .604	- 1.537	+ 2.241	+ 1.764	
Multiplier = $[ur\ f]$	+ 6.72	- .30	+ 4.59	- 11.83	+ 12.15	- 3.51	- 2.18	- 7.53	+ 10.44	- 3.13	+ 5.64	- 16.22	+ 4.73	+ 3.69	
$[ur\ f] k_r$	+ 10.436	+ .201	+ .280	- 6.637	+ 37.008	+ 2.485	- .514	- 4.044	+ .752	- 1.296	+ 3.407	+ 24.930	+ 10.611	+ 6.506	
$u_f = 93.29$ $u_F = u_f - \Sigma [ur\ f] k_r = 93.29 - 85.04 = 8.25$ $K = \sqrt{\frac{u_F}{u_f}} = .30$															

TABLE LXXXVII.

For 26 conditions

At  $U_1$

Easting closure

r	1	2	3	4	5	6	7	8	13	14	15	16	21	22	24	25	
s	Values of $[u\ s\ f]_s k_r$																
1	- .30	- .539	0	+ .007	- .099	+ .056	- .020	+ .007	- .000	- .082	+ .052	- .014	- .034	+ .234	+ .179	+ .064	+ .017
2	- 6.72	0	- 12.069	+ 2.218	+ .153	+ .452	+ 1.251	+ .008	+ .151	- 1.161	- 1.834	+ .762	- .304	- 1.424	- .378	+ 5.232	+ 4.009
3	- 11.83	+ .270	+ 3.904	- 1.390	0	+ .012	- .246	+ .077	+ .011	+ .180	+ .523	- .241	+ .012	+ .909	+ .150	- .500	- 1.229
4	- 4.59	- 1.515	+ .105	0	- .533	+ .095	+ .005	- .004	+ .030	- .203	+ .070	- .005	- .094	- .117	+ .477	+ .353	+ .658
5	- 3.51	+ .851	+ .236	+ .004	+ .073	- 3.020	0	- .098	- .600	+ 1.232	- 1.092	+ .808	+ .122	- .436	+ .151	- .310	+ .162
6	- 12.15	- .818	+ 2.262	- .253	+ .012	0	- 10.454	+ 2.107	- .304	+ 3.770	+ 4.265	- .423	+ 2.736	+ 1.073	- .527	- 1.508	+ .524
7	- 7.53	+ .169	+ .007	+ .049	- .007	- .188	+ 1.306	- .448	0	- .279	- .523	+ .053	- .277	- .231	+ .064	- .146	- .165
8	+ 2.18	+ .002	- .049	- .002	- .014	+ .378	+ .055	0	+ .130	- .151	+ .081	- .080	- .015	- .042	- .048	+ .067	- .019
13	- 3.13	- .854	- .541	+ .048	- .138	+ 1.099	+ .973	- .116	+ .217	- 6.384	0	+ 1.197	- 2.999	+ .304	+ 1.176	+ .515	+ .370
14	- 10.44	+ 1.803	- 2.849	+ .461	+ .159	- 3.247	+ 3.064	- .725	- .387	0	- 21.294	+ 9.994	+ 3.892	- 1.717	- 1.263	+ 1.314	+ 3.922
15	- 16.22	- .735	+ 1.839	- .331	- .016	+ 3.732	- .564	+ .114	+ .597	+ 6.201	+ 15.527	- 11.802	0	+ .068	- .362	- .819	- 3.727
16	- 5.64	- .640	- .255	+ .006	- .115	+ .196	+ 1.298	- .208	+ .020	- 5.399	+ 2.157	0	- 4.104	+ .285	+ 1.206	+ .337	- 1.228
21	- .14	+ .109	- .030	+ .018	- .004	- .017	+ .012	- .004	+ .003	+ .018	- .023	+ .006	+ .007	- .193	- .041	- .000	+ .010
22	- .53	+ .316	- .030	+ .007	+ .055	+ .023	- .023	+ .005	+ .012	+ .109	- .064	- .012	- .156	- .599	- .038	- .000	- .000
24	- 4.73	+ 1.002	+ 3.682	- .120	+ .363	- .418	- .587	- .092	- .015	+ .778	+ .596	- .230	+ .282	- .001	- .339	- 6.514	- 1.388
25	- 3.69	+ .266	+ 2.261	- .383	+ .047	+ .160	+ .169	- .061	+ .031	+ .446	+ 1.388	- .648	- .082	+ .262	- .001	- 1.083	- 4.167
Sum = $k_r$	- .570	- 1.587	+ .579	- .070	- .687	- 3.171	+ .550	- .004	- .824	- .173	- .842	- .573	- .192	- .063	- 2.236	- 1.760	
Multiplier = $[ur\ f]$	- .30	- 6.72	- 11.83	- 4.59	- 3.51	- 12.15	- 7.53	+ 2.18	- 3.13	- 10.44	- 16.22	- 5.64	- .14	- .53	- 4.73	- 3.69	
$[ur\ f] k_r$	+ .171	+ 10.665	- 6.860	+ .321	+ 2.411	+ 38.628	- 4.142	- .295	+ 2.579	+ 1.808	+ 13.657	+ 3.232	+ .027	+ .033	+ 10.576	+ 6.468	
$u_f = 93.29$ $u_F = u_f - \Sigma [ur\ f] k_r = 93.29 - 79.27 = 14.02$ $K = \sqrt{\frac{u_F}{u_f}} = .39$																	

To consider the probable errors at some other points the values of the R.H.S. of the corresponding equations have to be found as was done in table L for  $U_1$ . The details are given below in table LXXXVIII.

TABLE LXXXVIII.

Circuits	Equations	At N, computed along route A, T, S, Q, P, N,										At J, along A, to N, as before and L, J,													
		$A_1T_1$	$T_1S_1$	$S_1Q_1$	$Q_1P_1$	$P_1N_1$	Side	Az.	$A_1T_1$	$T_1S_1$	$S_1Q_1$	$Q_1P_1$	$P_1N_1$	East- ing	North- ing	$N_1L_1$	$L_1J_1$	Side	Az.	$N_1L_1$	$L_1J_1$	East- ing	North- ing		
I	1	+ .35						+ .35	0	- .17				- .17	+ .02			+ .35	0			- .17	+ .02		
	2	0						0	+ .35	- .02				- .02	- .17			0	+ .35			- .02	- .17		
	3	+ .15						+ .15	+ 1.81	- .13				- .13	- .87			+ .15	+ 1.81			- .13	- .87		
	4	- 1.81						- 1.81	+ .15	+ .87				+ .87	- .13			- 1.81	+ .15			+ .87	- .13		
II	5		+ .25					+ .25	0	- .34				- .34	+ .04			+ .25	0			- .34	+ .04		
	6		0					0	+ .25	- .04				- .04	- .34			0	+ .25			- .04	- .34		
	7		+ .05					+ .05	+ 1.46	- .30				- .30	- 1.95			+ .05	+ 1.46			- .30	- 1.95		
	8		- 1.46					- 1.46	+ .05	+ 1.95				+ 1.95	- .30			- 1.46	+ .05			+ 1.95	- .30		
III	9		+ .53	+ .37	+ .82	+ 1.72	0			- 1.27	- 1.33	- 4.32	- 6.92	+ .95	+ .31	+ .37	+ 2.40	0	- 1.03	- 1.00	- 10.75	+ 2.55			
	10		0	0	0	0	+ 1.72			- .15	- .19	- .61	- .95	- 6.92	0	0	+ 2.40	- .52	- 1.08	- 1.08	- 2.55	- 10.75			
	11		+ .07	+ .24	- .82	+ .39	+ 4.82			- 2.64	- 1.38	+ 3.04	- 1.03	- 19.22	- .61	- .33	- .55	+ 5.50	+ 2.99	+ 1.36	+ 3.32	- 25.28			
	12		- 1.62	- 1.05	- 2.15	- 4.82	+ .39			+ 3.64	+ 3.66	+ 11.02	+ 19.22	- 1.03	- .52	- .16	- 5.50	- .55	+ 4.27	+ 1.79	+ 25.28	+ 3.32			
VIII	23																	+ .31	+ .37	- .69	0	- 1.93	- 1.00	- 3.83	+ 1.60
	26																	0	+ .68	- .52	- 1.08	- 1.60	- 3.38		

Circuits	Equations	At G, along A, to J, as before and H, G,						At C, along A, B, C,											
		$J_1H_1$	$H_1G_1$	Side	Az.	$J_1H_1$	$H_1G_1$	East- ing	North- ing	A, D,	H, C,	Side	Az.	A, B,	B, C,	East- ing	North- ing		
I	1			+ .36	0			- .17	+ .02	I	1	+ 1.08	+ 1.11	+ 2.19	0	- .11	- .03	- .14	+ 4.73
	2			0	+ .35			- .02	- .17		2	0	0	0	+ 2.19	- 1.15	- 3.58	- 4.73	- .14
	3			+ .15	+ 1.81			- .13	- .87		3	+ .88	+ .98	+ 1.86	+ 6.62	- 4.40	- 0.56	- 10.96	+ 3.61
	4			- 1.81	+ .15			+ .87	- .13		4	- 4.46	- 2.16	- 6.62	+ 1.86	- .50	- 3.11	- 3.61	- 10.96
II	5			+ .25	0			- .34	+ .04	VI	21	+ 1.08	+ 1.11	+ 2.19	0	- .11	- .03	- .14	+ 4.73
	6			0	+ .25			- .04	- .34		24	0	0	0	+ 2.19	- 1.15	- 3.58	- 4.73	- .14
	7			+ .05	+ 1.46			- .30	- 1.95										
	8			- 1.46	+ .05			+ 1.95	- .30										
III	9				0			- 10.75	+ 2.55										
	10				0	+ 2.40		- 2.55	- 10.75										
	11				- .55	+ 5.50		+ 3.32	- 25.28										
	12				- 5.50	- .55		+ 25.28	+ 3.32										
V	17	+ 0.38	+ .32	- .70	0	- 1.53	- 1.10	- 2.63	+ 3.57										
	18	0	0	0	+ .70	- 1.63	- 1.94	- 3.57	2.63										
	19	- .45	- .18	- .63	+ 1.17	- 2.04	- .75	- 2.79	- 7.63										
	20	- .93	- .24	- 1.17	- .63	+ 5.74	+ 1.01	+ 7.66	- 2.79										
VIII	23	+ .38	+ .32	+ 1.38	0	- 1.53	- 1.10	- 6.46	+ 5.17										
	26			0	+ 1.38	- 1.63	- 1.94	- 5.17	- 6.46										

Note.—The values, in columns of Side, Azimuth, Easting and Northing at J, and G, have been brought from the corresponding columns at N, by adding to them the quantities  $N_1L_1$ ,  $L_1J_1$  to find those at J; and similarly the quantities  $J_1H_1$ ,  $H_1G_1$  have been added to the quantities at J, to find those at G.

The computations of k<sub>r</sub> for the following are omitted, and only the values of u<sub>f</sub> and k<sub>r</sub> for Side, Azimuth, Easting and Northing closures are given in the following Table.

TABLE LXXXIX

Table with 5 main columns: C, G, G1, J, N1. Each column contains multiple sub-sections of data for different closure types (e.g., S, A, E, N). Each sub-section lists station numbers (r) and values for k\_r, [urf], and [ur]k\_r. Summary statistics for each section include u\_f and K values.



The results obtained in tables LXXX to LXXXIX are now collected in the following table showing values of  $u_f$  and  $u_F$  for Side, Azimuth, Easting and Northing closures of several points of N.W.Q., the latter for 20, 23 and 26 conditions.

TABLE XC.

	C <sub>1</sub> or Dehra base				U <sub>1</sub>				G <sub>1</sub> or Chach base				J <sub>1</sub>				N <sub>1</sub> or Karachi base			
	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$
Side	2.19	.77	0	0	4.91	1.19	.49	.46	3.70	1.67	0	0	3.00	1.17	1.00	.97	2.32	1.34	0	0
Azimuth	2.19	.77	.76	0	4.91	1.19	2.42	.46	3.70	1.67	1.67	0	3.00	1.17	1.14	.97	2.32	1.34	1.30	0
Easting	13.67	6.44	6.36	2.04	93.29	19.98	19.72	14.02	87.31	33.16	22.50	7.14	58.33	12.43	9.83	6.33	32.28	14.02	12.83	12.32
Northing	13.67	6.44	3.06	2.04	93.29	19.08	8.25	14.02	87.31	33.16	24.18	7.14	58.33	12.43	11.84	9.33	32.28	14.92	14.31	12.32

These values all seem reasonable. Rather unexpected results are 2.42 and 8.25 for 23 conditions for U<sub>1</sub>.

It appears that the greatest probable error of the adjustment of Easting or Northing is  $4\sqrt{33} \doteq 23$  feet: in terms of deflection this is negligible and of the order of probable error of latitude (astronomic) result.

As regards azimuths for 20 or 23 conditions the worst case is  $1.6\sqrt{2.4}$  i.e. probable error of 2".4. Error of 7" is in this case possible and liable to occur.

In N. E. Quadrilateral where triangulation is not so good there will be greater errors. Closed on Laplace stations, however, errors are probably reduced to  $1.6\sqrt{.5}$  and  $1.6\sqrt{1.0}$  i.e. probable error to 1.6 and possible to 5".

In the above the "possible error" is regarded as three times the probable error.

As further discussion of these results is at present impossible, for reasons explained in the preface the chapter is concluded with a tabular statement of the probable errors of log. side to 7th place of decimals, azimuth in seconds and easting and northing in feet; these are obtained as explained in § 20.

TABLE XCI

	C <sub>1</sub> or Dehra base				U <sub>1</sub>				G <sub>1</sub> or Chach base				J <sub>1</sub>				N <sub>1</sub> or Karachi base			
	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$	$u_f$	20 $u_F$	23 $u_F$	26 $u_F$
Side	49	29	0	0	74	36	23	23	64	43	0	0	57	36	34	33	50	39	0	0
Azimuth	2".33	1".39	1".37	0	3".50	1".72	2".46	1".07	3".02	2".03	2".03	0	2".72	1".70	1".69	1".64	2".39	1".33	1".0	0
Easting	feet 14.61	feet 10.24	feet 10.16	feet 5.76	feet 38.93	feet 18.01	feet 17.89	feet 15.07	feet 37.64	feet 23.21	feet 19.10	feet 10.76	feet 30.79	feet 11.23	feet 12.63	feet 12.29	feet 22.89	feet 16.56	feet 14.43	feet 14.16
Northing	feet 14.61	feet 10.24	feet 7.05	feet 5.76	feet 38.93	feet 18.01	feet 11.67	feet 15.07	feet 37.64	feet 23.21	feet 19.83	feet 10.76	feet 30.79	feet 14.23	feet 13.86	feet 12.29	feet 22.89	feet 16.56	feet 15.23	feet 14.15

## CHAPTER IX.

## Deflections of the Plumb-line and values of "g" derived from observations of the Survey of India.

1. The first use of the tables derived in the earlier chapters will now be made use of to display in convenient form all the data of plumb-line deflections available up to the time of writing (May 1917). As regards the deflections in meridian no comment is necessary. The results of observation are immediately available. With the deflections in prime vertical the case is different. It has been stated already that the triangulation of India was not adjusted on the longitude arcs, which have been subsequently observed or reduced. The triangulation accordingly is burdened with an accumulation of error in azimuth which may be largely reduced by adjustment on the longitude arcs. For this purpose it is not essential for the present purpose to reopen the adjustment of the whole triangulation except as regards the azimuth, and the process followed will be substantially that followed by Colonel Sir Sidney Burrard in Appendix 5, G.T.S. Volume XVIII. The numerical results will however be slightly different owing to the improved methods of computing the effect of a change in azimuth at the origin which have been developed in Chapters I—III; the differences will depend mainly on the taking into account of the effect of a change on azimuth at the origin on longitudes of points considerably removed from the origin.

When the triangulation of India was adjusted General Walker decided to adopt a value of the fundamental azimuth (of Surantal from Kalianpur) which differed from the observed value by a small amount (*vide* Chapter I, § 4) and this of course implied a deflection in prime vertical at Kalianpur. This has given rise to a little confusion as regards the longitudes of India. The astronomic longitude of Kalianpur has been determined with reference to Greenwich, but no account of the implied deflection in prime vertical has been hitherto considered.\* Colonel Sir Sidney Burrard in adjusting the azimuth observations eliminated the effect of this oversight by returning to an observed value of the fundamental azimuth. The deflection in meridian remained in terms of the Everest spheroid with Walker's initial azimuth. In dealing with the deflections as a whole it will accordingly be better to keep the azimuths in the same terms as the latitudes, and to recognise that a deflection in prime vertical at Kalianpur is thereby implied. After the adjustments on the longitude arcs have been performed, the results of both azimuth and latitude deflections will be in common terms of Everest spheroid and Walker's origin.

Quantities for correcting all the deflections to refer to any other spheroid and origin are given in table XCV in which all the results are exhibited, as well as the deflections corrected to the special case of Helmert's spheroid and the latest observed value of latitude and azimuth at Kalianpur, as derived from observations at a group of stations surrounding Kalianpur. †

\* *Vide*, p. xv G.T.S. Vol. XVII, Survey of India.

† *Vide*, pp. 7,9 Professional Paper No. 5, Survey of India.

2. As the correction for azimuths has already been treated by Colonel Sir Sidney Burrard *loc. cit.* it will not be necessary to state afresh the various practical difficulties which arose owing to longitude stations not being in general identical with azimuth stations. The observation results exhibited by him will be taken unaltered, and immediately applied to Laplace's equation. This equation has been given in somewhat amplified form in (3) of Chapter V. There is now no occasion to consider observation errors of astronomic azimuths or their determination. The accumulated error of geodetic longitude determination is certainly small compared with that of geodetic azimuth and so will be neglected. The equation may accordingly be written

$$(A - G - \delta G) \operatorname{cosec} \lambda - (A_0 - G_0) \operatorname{cosec} \lambda_0 = A - G \dots \dots \dots (1)$$

where the notation has been changed in conformity with the usual practice and *A, A* and *G, G* signify astronomic and geodetic determinations respectively, roman letters referring to azimuth and italic letters to longitude determinations:  $\delta G$  is the correction necessary to the geodetic value of azimuth: as this is the quantity required it will be convenient to rewrite (1). From Chapter I § 4 it is seen that  $A_0 - G_0 = +1.29$ ,  $\lambda_0 = 24^\circ 7' 12''$ ,  $(A_0 - G_0) \operatorname{cosec} \lambda_0 = 3.16$ . Hence

$$\delta G = A - G - (3''.16 + A - G) \sin \lambda \dots \dots \dots (2)$$

which serves to determine  $\delta G$ . So long as Walker's value of azimuth is adhered to all geodetic longitudes of India require a correction of  $-3''.16$ .

For the Helmert spheroid an additional correction  $\delta_2 G$  is necessary. The corresponding equation to (1) is

$$(A - G - \delta G - \delta_1 G - w) \operatorname{cosec} \lambda = A - G - v \dots \dots \dots (3)$$

since *G, G* are changed by *v* and *w* respectively and the astronomic and geodetic azimuths at the origin have been made identical. Subtracting (3) from (2) it follows that

$$\delta_2 G = 3.1571 \sin \lambda + v \sin \lambda - w \dots \dots \dots (4)$$

The solutions of (2) and (4) and the deduction of  $\delta G$  and  $\delta_2 G$  are now shown in table XCIII.

3 Having obtained the values of  $\delta G$  and  $\delta_2 G$  at all Laplace stations it is next necessary to find values at the intervening azimuth stations. This is done in table XCIV, interpolating according to the number of removes from terminal stations. The Laplace stations are shown in block type. Azimuth stations between which adjustments have been performed are shown in italics.

4. The precision of deflections in prime vertical so far as is due to the astronomical observation, obtained from azimuth observation, is much less than those in meridian.\* As may be seen from results obtained in Chapter VII the probable error of azimuth generated in triangulation is much greater than that in latitude or longitude, all being expressed in seconds. The deflection in prime vertical is derived from the azimuth anomaly by multiplication by  $\cot \lambda$ —a quantity which ranges from 7 to 1.4 in Indian latitudes—and this further increases the lack of precision. Considerable improvement on the other hand should result by the use of Laplace stations, which has been made. A further source of weakness, of varying amount, is the actual azimuth observation itself. The observation is not nearly so satisfactory as that for latitude and involves graduation error of the instrument which, especially in the older observations, introduces a serious uncertainty. It is desirable then to consider the relative degree of reliability of the azimuth observations. On account of the other sources of error, mainly that of accumulation of error in triangulation it is not useful to do this in very great detail, and it is considered sufficient to work out the probable error from the

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\* For probable errors in astronomic latitude vide *G.T.S. Vol. XI, pages 882—982* and *G.T.S. Vol. XI/III App. 7 Table III*. These are seldom so great as  $0''.2$ . The worst case (Gogipatri) is  $\pm 0''.68$ .

results obtained on the several zeros. This takes no account of errors in star places, a defect which was more serious in the early days of the survey than it is at present. Probably graduation error in the instruments has improved at much the same rate as the error of star place, and a fair estimate of probable error will be obtained by consideration of the probable error due to graduation only. The formula used is

$$p. e. \text{ of mean of observations on } n \text{ zeros} = .6745 \sqrt{\frac{\sum \delta^2}{n(n-1)}} \dots \dots (5)$$

where  $\delta$  is the discrepancy of the value derived from any zero from the mean result. The results of the application of this formula are given in table XCIV. Some idea of the probable errors in the geodetic values of latitude, longitude and azimuth is given in Chapter VIII, where the N.W. Quadrilateral is considered in detail.

5. In the table XCV the deflections derived from observations of the Survey of India are given. These have been arranged by degree sheets. On the left hand page of the table the data are expressed in terms of the Everest Spheroid, using the observed value of latitude and the deduced value of azimuth at Kalianpur which General Walker adopted: on the right hand page of the table the quantities are given which must be applied properly to the triangulated values to express in terms of any other spheroid. There are four variables to be considered, giving rise to four cases: these are (1) change of semi-major axis,  $\delta a$ , (2) change of semi-minor axis,  $\delta b$ , (3) change of latitude of origin,  $u_0$  and (4) change of azimuth at origin,  $w_0$ . The cases given correspond to  $\delta a = 1 \text{ km.}$ ,  $\delta b = 1 \text{ km.}$ ,  $u_0 = 1''$ ,  $w_0 = 1''$ , and to obtain the general case these must be combined as follows:—

$$\delta a \times \text{case I} + \delta b \times \text{case II} + u_0 \times \text{case III} + w_0 \times \text{case IV}$$

in which  $\delta a$ ,  $\delta b$  are expressed in kilometres and  $u_0$ ,  $w_0$  are expressed in seconds. Thus in the case of the Helmert Spheroid in which  $a = 6378.2 \text{ km.}$  and  $1/\epsilon = 298.3$  and with revised values of latitude and azimuth at Kalianpur as given on p. 2,  $\delta a = .924$ ,  $\delta b = .743$ ,  $u_0 = .31$ ,  $w_0 = 1.29$ . Deflections in terms of this spheroid are given on the right of the right hand page of the table: but those in terms of any other spheroid may be easily found by making use of different values of  $\delta a$ , etc. It is clear in the notation of this work that the correction to latitude deflection is  $-u$ , to prime vertical deflection  $-v \cos \lambda$  or  $-w \cot \lambda$  according as the deflection is derived from longitude or azimuth observations. Values of these quantities have been taken from tables XVII—XX, XXIX—XXXVI.

It frequently happens that latitude observations have been made on a site not quite identical with the triangulation station, but at some (small) distance from it on the prime vertical through it. Similarly the longitude observations done by wire-telegraphy were made at the telegraph offices and not exactly on the station sites. The coordinates of the triangulation stations are generally the quantities given. But at latitude stations the value of latitude given is that of the latitude station, and in longitude stations the longitude of the longitude station. When a change is made to a different spheroid, since the corrections do not satisfy Laplace equation, slight discrepancies occur between deflection derived from longitude and azimuth observations. This point has been explained in Chapter V and has been taken into account in the case of the Helmert spheroid in table XCIII.

The elevation of the referring mark affects the result of azimuth observations (*vide* §6 Chapter V) and accordingly this has been given in the table except for a few cases where the data could not be found.

Longitude arcs by means of wireless telegraphy were observed in collaboration with the expedition of Cav. de Filippi in 1914 between Dehra Dun and eight stations. Their names and the values of  $A$  are appended. The Dehra Dun observations were made in the transit room adjoining the

Dome Observatory (new) and the transit was 25.2 feet W. of the mark-stone. The longitude (geodetic) of the transit instrument is 7"·46 (equivalent linear measurement being 628.8 + 25.2 = 654.0 feet less than that of the Dehra Dun Haig Observatory when all previous longitude observations were made; the longitude of the transit instrument is accordingly 78° 2' 49"·01 E. of Greenwich. The geodetic values of the eight stations are not yet (1921) available.

TABLE XCII. (Astronomic values).

	Skardu	Kargil	Lamayaru	Leh	Depsang	Suget Karaul	Yarkand	Kashgar
Long.	75° 38' 22"·92	76° 7' 39"·72	76° 46' 33"·86	77° 34' 53"·89	77° 58' 17"·85	78° 1' 36"·09	77° 15' 46"·02	75° 59' 5"·64
Lat.	35 18 18·05 ± 0·17	34 33 38·25 ± 0·51	34 17 4·81 ± 0·23	34 9 54·10 ± 0·19	35 17 26·77 ± 0·18	36 20 54·91 ± 0·16	38 24 22·22 ± 0·31	39 28 19·74 ± 0·31

TABLE XCIII.

Azimuth station Latitude = $\lambda$		Azimuth (Everest)	(1) $A - G \ddagger$	Longitude station $\frac{\text{Longitude from Kailanpur (Everest)}}{\sin \lambda}$	Longitude from Kailanpur (Everest)	$A - G \ddagger$ (2)	(1) - (2) = $\delta G$	(3) $3 \cdot 1571 \times \sin \lambda$	(4) $v \sin \lambda$	(5) $w$	(6) $\frac{(3) + (4) - (5)}{2} = \delta_2 G \S$
Kailanpur 24° 7' 11"	H.S.	A 190° 27' 6"·39 G 190 27 5·10	+ 1·29	Kailanpur ... -4086	A 0 0 0 G 0 0 0	0 + 1"·29	0·0	+ 1·29	0·00	+ 1·29	0·00
Karachi Observatory 24° 49' 50"	S.	A 221 39 9·5 G 221 39 10·9	- 1·4	Karachi T.O. ... -4200	A - 10 38 24·8 G - 10 38 24·3	- 0·6 + 1·1	- 2·5	+ 1·33	+ 2·41	+ 3·62	+ 0·12
Dehra Dun Obs. (old) 30° 19' 57"	S.	A 165 10 58·8 G 166 11 10·7	- 11·9	Dehra Dun -5050	A + 0 23 38·8 G + 0 24 4·6	- 25·7 - 11·4	- 0·5	+ 1·59	- 0·03	+ 1·25	+ 0·31
Quetta T.O. 30° 11' 57"	S.	A 168 31 12·1 G 168 31 17·0	- 4·9	Quetta T.O. ... -5030	A - 10 38 48·3 G - 10 38 45·8	- 2·5 + 0·3	- 5·2	+ 1·59	+ 3·06	+ 4·23	+ 0·42
Calcutta Base S., T.S. 22° 36' 56"	S.	A 177 10 27·3 G 177 10 36·2	- 8·9	Calcutta ... -3846	A + 10 42 0·9 G + 10 42 11·3	- 11·0 - 3·0	- 5·8	+ 1·31	- 2·20	- 0·90	- 0·09
Orejhar 26° 46' 56"	S.	A 308 36 18·8 G 308 36 23·0	- 4·1	Fyznbad T.O. ... -4506	A + 4 28 60·1 G + 4 28 60·6	- 0·5 + 1·2	- 5·3	+ 1·42	- 1·08	+ 0·22	+ 0·12
Jalpaiguri 26° 31' 17"	S.	A 321 33 25·3 G 321 33 30·0	- 4·7	Jalpaiguri ... -4465	A + 11 4 34·8 G + 11 4 56·2	- 20·4 - 7·7	+ 3·0	+ 1·41	- 2·08	- 1·32	+ 0·05
Nagarbhana 23° 22' 56"	H.S.	A 155 47 13·3 G 155 47 23·5	- 10·2	Chittagong T.O. -3808	A + 14 10 47·4 G + 14 10 59·1	- 11·7 - 3·3	- 6·9	+ 1·20	- 2·87	- 1·71	+ 0·04
Dattaung 20° 13' 14"	H.S.	A 171 27 28·3 G 171 27 38·1	- 9·8	Akyab T.O. ... -3456	A + 16 14 21·0 G + 15 14 32·1	- 11·1 - 2·7	- 7·1	+ 1·08	- 3·77	- 1·93	+ 0·25
Kyaunggyi 16° 49' 21"	S.	A 109 26 42·1 G 109 26 58·1	- 16·0	Prome ... -3226	A + 17 33 24·6 G + 17 33 30·9	- 15·3 - 3·9	- 12·1	+ 1·02	- 2·97	- 2·27	+ 0·32
Taungzun 10° 25' 49"	H.S.	A 31 16 18·9 G 31 16 32·7	- 13·8	Moulmein ... -2820	A + 19 58 5·0 G + 19 58 22·5	- 16·6 - 3·8	- 10·0	+ 0·89	- 2·92	- 2·54	+ 0·51
Bolarum P.W.D. 17° 30' 13"	S.	A 25 57 35·8 G 25 57 36·9	- 1·1	Bolarum ... -3008	A + 0 51 50·3 G + 0 51 53·0	- 3·3 0·0	- 1·1	+ 0·95	- 0·18	+ 1·10	- 0·33
Vizaganpatam Base N., S. 18° 1' 3"	S.	A 203 44 24·5 G 203 44 25·9	- 1·4	Waltair ... -3003	A + 5 30 42·6 G + 5 39 45·9	- 3·2 0·0	- 1·4	+ 0·08	- 0·05	+ 0·33	- 0·30
Karundi 23° 10' 40"	H.S.	A 206 22 35·6 G 206 22 30·6	- 4·0	Jabalpur T.O. ... -3936	A + 2 17 34·8 G + 2 17 45·0	- 10·2 - 2·8	- 1·2	+ 1·24	- 0·19	+ 0·79	- 0·04
Colaba Observatory 18° 53' 47"	S.	A 288 5 27·7 G 288 5 26·7	+ 1·0	Bombay ... -3238	A - 4 50 21·8 G - 4 50 28·6	+ 6·6 + 3·3	- 2·2	+ 1·02	+ 0·77	+ 3·05	- 0·26
Deesa T. O. 24° 15' 29"	S.	A 241 16 15·3 G 241 16 19·9	- 4·6	Deesa T.O. ... -4108	A - 5 28 16·4 G - 5 28 12·7	- 3·7 - 0·2	- 4·4	+ 1·30	+ 1·20	+ 2·48	+ 0·04
Mangalore 12° 52' 15"	S.	A 205 52 50·8 G 205 52 53·6	- 2·8	Mangalore ... -2227	A - 2 48 32·0 G - 2 48 35·1	+ 2·2 + 1·3	- 4·0	+ 0·70	+ 0·26	+ 1·55	- 0·50
Bangalore Base S.W., S. 13° 0' 41"	S.	A 224 31 21·7 G 224 31 27·0	- 5·3	Bangalore ... -2252	A - 0 4 20·3 G - 0 4 17·6	- 2·7 + 0·1	- 5·4	+ 0·71	- 0·05	+ 1·23	- 0·56
St. Thomas's Mount S. 13° 0' 15"	S.	A 12 30 5·3 G 12 30 9·3	- 4·0	Madras ... -2250	A + 2 35 20·6 G + 2 35 16·6	- 7·0 - 0·9	- 3·1	+ 0·71	- 0·35	+ 0·59	- 0·53
Kudankulam Obs. 8° 10' 22"	S.	A 185 55 18·8 G 185 55 26·6	- 7·7	Nagarkoil ... -1421	A - 0 13 15·8 G - 0 13 14·3	- 1·6 + 0·2	- 7·9	+ 0·45	- 0·03	+ 1·19	- 0·77

Obs. = observatory, T.O. = Telegraph office. \* *File* G.T.S. Vol. XV, p. (5). † Approximate values. ‡ A, A = Astronomic values; G, G = Geodetic values. § This is the additional correction for Helmert's spheroid. ¶ Derived from unadjusted values of Quetta Secondary Series. This does not enter into the azimuth correction.

TABLE XCIV. (See Index pp. 170-172)

Serial Number	Station	Corrections		Prob- able Error ±	Date of Observation	Serial Number	Station	Corrections		Prob- able Error ±	Date of Observation
		Everest's Spheroid	Helmert's Spheroid*					Everest's Spheroid	Helmert's Spheroid*		
1	Kalianpur H.S.	0.0	0.0	0.31	{ 1836	25	Karachi Obs. S.	-2.5	+0.1	0.19	1855
2	Losalli S.	-0.1	0.0	0.44	{ 1849	39	Karachi Base S., S.	-2.4	+0.1	0.30	1853
3	Salot H.S.	-0.2	0.0	0.27	1849	40	Yusuf S.	-2.0	+0.2	0.53	1858
4	Māta-ka-hūra H.S.	-0.3	0.0	0.34	1849	41	Binnar T.S.	-1.9	+0.2	0.27	1859
5	Gurāria H.S.	-0.4	0.0	0.44	1849	42	Miāni T.S.	-1.8	+0.2	0.32	1859
6	Rāmpūra H.S.	-0.5	0.0	0.49	1849	43	Dājil S.	-1.6	+0.2	0.22	1860
7	Aramlia S.	-0.6	0.0	0.32	1850	44	Dera Din Panāh S.	-1.4	+0.2	0.25	1859
8	Sānd H.S.	-0.7	0.0	0.90	1850	45	Jharkil T.S.	-1.2	+0.2	0.27	1859
9	Tiki H.S.	-0.8	0.0	0.38	1851	46	Umarkhel H.S.	-1.0	+0.3	0.20	1909
10	Kānagar H.S.	-0.9	0.0	0.50	1850	47	Jāoli H.S.	-0.9	+0.3	0.80	1851
11	Gūru Sikkar H.S.	-1.0	0.0	0.45	1850	48	Medwāni H.S.	-0.7	+0.3	0.50	1853
12	Birona S.	-1.1	0.0	0.46	1851	33.	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853
13	Khanbaria S.	-1.2	0.0	0.33	1851	25	Karachi Obs. S.	-2.5	+0.1	0.19	1855
14	Sarla S.	-1.2	+0.1	0.25	1851	49	Andar H.S.	-2.4	+0.1	0.27	1895
15	Didāwa H.S.	-1.3	+0.1	0.21	1851	50	Piaro H.S.	-2.4	+0.1	0.27	1896
16	Vivāria H.S.	-1.4	+0.1	0.19	1851	33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853
17	Lūnki H.S.	-1.5	+0.1	0.19	1851	49	Andar H.S.	-2.4	+0.1	0.27	1895
18	Rojhra H.S.	-1.6	+0.1	0.29	1851	51	Gandpahar H.S.	-2.1	+0.1	0.13	1906
19	Chānga H.S.	-1.7	+0.1	0.28	1852	52	Zawa H.S.	-1.9	+0.1	0.20	1905
20	Mairāb-ka-Shahar T.S.	-1.8	+0.1	0.56	1852	53	Mushelak H.S.	-1.9	+0.2	0.34	1908
21	Khori T.S.	-1.9	+0.1	0.53	1852	54	Gundak H.S.	-1.7	+0.2	0.22	1910
22	Alamkhān T.S.	-2.0	+0.1	0.22	1852	55	Salgiar H.S.	-1.6	+0.2	0.19	1910
23	Chūthi T.S.	-2.1	+0.1	...	1853	56	Tounsa T.S.	-1.5	+0.2	0.31	1910
24	Kārothol H.S.	-2.2	+0.1	0.31	1853	44	Dera Din Panāh S.	-1.4	+0.2	0.25	1859
25	Karachi Obs. S.	-2.5	+0.1	0.19	1855	52	Zawa H.S.	-1.9	+0.1	0.20	1905
26	Pahārgarh H.S.	-0.1	0.0	0.22	1836	57	Kisanen Chappar H.S.	-1.9	+0.1	0.18	1907
27	Kesri H.S.	-0.1	+0.1	0.39	1836	58	Tuzgi H.S.	-1.9	+0.1	0.36	1907
28	Usira H.S.	-0.2	+0.1	0.18	1838	59	Koh-i-Mulik Siah H.S.	-1.9	+0.1	0.30	1907
29	Noh T.S.	-0.3	+0.2	0.93	1837	60	†Quetta T.O. S.	-5.2	+0.4	0.28	1904
30	Datāiri T.S.	-0.3	+0.2	0.40	1836	5	Gurāria H.S.	-0.4	0.0	0.44	1819
31	Kaliāna S.	-0.4	+0.3	0.33	1836	61	Kānkra H.S.	-0.4	0.0	0.34	1862
32	Banog H.S.	-0.5	+0.3	0.28	{ 1836	62	Bānskho H.S.	-0.5	+0.1	0.14	1862
33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	{ 1907	63	Tāsing H.S.	-0.5	+0.1	0.70	1861
33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853	64	Rākhi T.S.	-0.6	+0.2	0.19	1860
34	Rājpur h.s.	-0.5	+0.3	0.26	1914	65	Kheri T.S.	-0.7	+0.2	0.37	1856
35	Spur point h.s.	-0.5	+0.3	0.29	1914	66	Bowra T.S.	-0.7	+0.2	0.52	1853
36	Jhari pani h.s.	-0.5	+0.3	0.25	1914	48	Medwāni H.S.	-0.7	+0.3	0.50	1853
37	Mussooree Dome Obs. H.S.	-0.5	+0.3	0.45	1912	7	Aramlia S.	-0.6	0.0	0.32	1850
38	Nag Tibā H.S.	-0.5	+0.3	0.38	1903	67	Rājgarh H.S.	-0.6	0.0	0.56	1863
32	Banog H.S.	-0.5	+0.3	0.28	{ 1836	68	Gurinda S.	-0.7	+0.1	0.42	1863
					{ 1907	69	Sirsa S.	-0.7	+0.1	0.34	1861
						70	Sangatpur T.S.	-0.8	+0.2	0.54	1860
						47	Jāoli H.S.	-0.9	+0.3	0.80	1851
						47	Jāoli H.S.	-0.9	+0.3	0.80	1851
						71	Murreo h.s.	-0.9	+0.3	0.77	1860
						72	Ganga Choti H.S.	-0.9	+0.3	0.16	1910
						73	Poshkar H.S.	-0.9	+0.3	0.65	1862
						74	Gogipatri H.S.	-0.9	+0.3	0.84	1862
						75	Rustamgarhi h.s.	-0.9	+0.3	0.67	1862
						70	Sangatpur T.S.	-0.8	+0.2	0.54	1860

Note:—In the azimuthal observations, Level corrections were introduced from 1863, vide G.T.S. Vol. II Appendix 9, p. 73 and Diurnal Aberration corrections from 1902 vide Handbook of the Trigonometrical Branch 1902, p. 74.

Obs. = observatory, T.O. = Telegraph Office.

S = station, H.S. = hill station, T.S. = tower station of principal triangulation. The same small letters refer to minor triangulation.

\* Additional correction for Helmert's spheroid.

† Computed from the unadjusted values of longitude and azimuth of Quetta T.O. station: not used for adjusting any azimuth observations.

TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station		Corrections		Probable Error $\pm$	Date of Observation	Serial Number	Station		Corrections		Probable Error $\pm$	Date of Observation
			Everest's Spheroid	Helmert's Spheroid*						Everest's Spheroid	Helmert's Spheroid*		
11	Gūru Sikkar	H.S.	-1.0	0.0	0.45	1850	90	Amūa	H.S.	-1.4	0.0	0.94	1834
76	Thob	H.S.	-1.0	0.0	0.38	1873	111	Nimkār	T.S.	-2.3	+0.1	0.33	1838
77	Jambo	H.S.	-1.0	+0.1	0.21	1874	103	Ramuapur (old)	T.S.	-2.7	+0.2	0.52	1838
78	Mugrala	H.S.	-1.0	+0.1	0.40	1875	91	Karāra	H.S.	-1.8	0.0	0.69	1842
79	Lādimsir	T.S.	-1.0	+0.1	0.45	1862	112	Pabhosa	H.S.	-2.4	+0.1	1.26	1845
80	Mandresa	T.S.	-0.9	+0.2	0.17	1862	113	Sora	T.S.	-2.9	+0.1	0.44	1845
81	Jhambhera	T.S.	-0.9	+0.2	0.63	1862	104	Māsi	T.S.	-3.8	+0.2	0.37	1850
82	Akbar	S.	-0.9	+0.2	0.79	1857							
47	Jāuli	H.S.	-0.9	+0.3	0.80	1851	92	Gurwāni	H.S.	-2.2	0.0	0.61	1845
18	Rojhva	H.S.	-1.6	+0.1	0.29	1851	114	Marār	T.S.	-3.7	+0.1	0.48	1846
83	Malar	H.S.	-1.5	+0.1	0.32	1877	115	Bisaul	T.S.	-5.1	+0.1	0.60	1847
84	Asu	H.S.	-1.4	+0.1	0.49	1880	106	Orejhar	S.	-5.3	+0.1	0.26	1904
85	Vijnot	T.S.	-1.3	+0.1	0.22	1881	93	Gora	H.S.	-2.7	0.0	0.52	1845
86	Dāowāla	T.S.	-1.3	+0.1	0.20	1881	116	Hirdepur	T.S.	-3.0	0.0	0.58	1846
87	Paphra	T.S.	-1.1	+0.1	0.25	1861							
79	Lādimsir	T.S.	-1.0	+0.1	0.45	1862	117	Samenda	T.S.	-3.4	0.0	1.13	1846
1	Kaliānpur	H.S.	0.0	0.0	0.31	1836	118	Rājabāri	T.S.	-3.9	+0.1	1.52	1847
88	Budhon	H.S.	-0.4	0.0	0.36	1898	105	Bāsadela	T.S.	-4.6	+0.1	0.32	1849
89	Rungir (old)	H.S.	-0.8	0.0	0.64	1864	106	Orejhar	S.	-5.3	+0.1	0.26	1904
90	Amūa	H.S.	-1.4	0.0	0.94	1834	119	Naunangurhi	T.S.	-2.3	+0.1	0.34	1852
91	Karāra	H.S.	-1.8	0.0	0.69	1842	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
92	Gurwāni	H.S.	-2.2	0.0	0.64	1845	121	Rāmganj	T.S.	+2.4	+0.1	0.64	1853
93	Gora	H.S.	-2.7	0.0	0.52	1845	122	Jalpaiguri	S.	+3.0	+0.1	0.33	1904
94	Hurilāong	H.S.	-3.1	-0.1	0.44	1849	94	Hurilāong	H.S.	-3.1	-0.1	0.44	1849
95	Chendwār (old)	H.S.	-3.6	-0.1	0.67	1843	123	Mednipur	T.S.	-2.9	0.0	0.34	1850
96	Pārasnūth	H.S.	-3.7	-0.1	0.35	1850	124	Jalālpur	T.S.	-2.6	0.0	0.62	1852
97	Tilabāni	H.S.	-3.9	-0.1	0.76	1845	119	Naunangurhi	T.S.	-2.3	+0.1	0.34	1852
98	Malūncha	H.S.	-4.3	-0.1	0.57	1844	95	Chendwār (old)	H.S.	-3.5	-0.1	0.67	1843
99	Madhpur	T.S.	-4.9	-0.1	0.49	1868	125	Pota	T.S.	-2.8	0.0	0.46	1846
100	Aknāpur	T.S.	-5.2	-0.1	0.83	1869	119	Naunangurhi	T.S.	-2.3	+0.1	0.34	1852
101	Calcutta Base-line S. End	T.S.	-5.9	-0.1	0.37	1845	96	Pārasnūth	H.S.	-3.7	-0.1	0.35	1850
33	Dehra Dun Obs. (old)	S.	-0.5	+0.3	0.34	1853	126	Bichwi	H.S.	-3.2	-0.1	0.30	1851
102	Kaliānpur	T.S.	-1.8	+0.2	0.64	1850	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
103	Ramuapur (old)	T.S.	-2.7	+0.2	0.52	1838	98	Malūncha	H.S.	-4.3	-0.1	0.67	1844
104	Māsi	T.S.	-3.8	+0.2	0.37	1850	127	Sirkānda	T.S.	-1.7	0.0	0.43	1846
105	Hāsadela	T.S.	-4.6	+0.1	0.32	1849	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
106	Orejhar	S.	-5.3	+0.1	0.26	1904	101	Calcutta Base-line S. End	T.S.	-5.9	-0.1	0.37	1845
88	Budhon	H.S.	-0.4	0.0	0.36	1864	128	Anandbās	T.S.	-4.2	-0.1	0.45	1846
107	Gūirmi	T.S.	-0.4	+0.1	0.50	1842	129	Madhpur	T.S.	-2.9	0.0	0.63	1846
108	Sankrāo	T.S.	-0.4	+0.2	0.39	1843	122	Jalpaiguri	S.	+3.0	+0.1	0.33	1904
109	Sira	T.S.	-0.5	+0.2	0.70	1843	101	Calcutta Baseline S. End	T.S.	-5.9	-0.1	0.37	1845
33	Dehra Dun Obs. (old)	S.	-0.5	+0.3	0.34	1853	130	Daulatpur	T.S.	-6.3	-0.1	0.16	1868
89	Rangir (old)	S.	-0.8	0.0	0.64	1834	131	Gangapur	T.S.	-6.5	0.0	0.34	1866
110	Mohammadabad	T.S.	-1.4	+0.1	0.44	1840	132	Lokhinagar	T.S.	-6.5	0.0	0.14	1866
102	Kaliānpur	T.S.	-1.8	+0.2	0.64	1850	133	Semu Tāu	H.S.	-6.8	0.0	0.39	1865
134	Nagarkhana	H.S.	-6.9	0.0	1.29	1905							

\* Additional correction for Helmert's Spheroid.

TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station		Corrections		Probable Error $\pm$	Date of Observation	Serial Number	Station		Corrections		Probable Error $\pm$	Date of Observation
			Everest's Spheroid	Helmer's Spheroid*						Everest's Spheroid	Helmer's Spheroid*		
130	<i>Daulatpur</i>	T.S.	- 6.3	- 0.1	0.15	1868	160	<i>Ubyetaung</i>	H.S.	- 6.8	+ 0.1	0.31	1894
135	<i>Tepri</i>	T.S.	- 4.2	- 0.1	0.30	1869	167	<i>Sinpitaung</i>	H.S.	- 6.8	+ 0.1	0.36	1901
136	<i>Aloakāndi</i>	T.S.	- 2.2	0.0	0.51	1873	168	<i>Loi Hpa Lang</i>	H.S.	- 6.8	+ 0.1	0.16	1903
137	<i>Halkāchar</i>	T.S.	- 1.1	0.0	0.21	1873	169	<i>Loi Hpatan</i>	H.S.	- 6.8	+ 0.1	0.20	1907
138	<i>Alangjāni</i>	T.S.	+ 0.6	0.0	0.79	1874	170	<i>Loi Kiipma</i>	H.S.	- 6.8	+ 0.1	0.17	1908
139	<i>Ataro Bānki</i>	T.S.	+ 1.3	+ 0.1	0.34	1856	171	<i>Loi Hsam Hsum</i>	H.S.	- 6.8	+ 0.1	0.21	1911
122	<i>Jalpaiguri</i>	s.	+ 3.0	+ 0.1	0.33	1904	172	<i>Kumtum Bum</i>	H.S.	- 6.8	+ 0.1	0.22	1910
133	<i>Semu Tān</i>	H.S.	- 6.8	0.0	0.39	1865	173	<i>Kumon Bum</i>	H.S.	- 6.8	+ 0.1	0.17	1911
140	<i>Dawa</i>	H.S.	- 5.3	0.0	0.26	1864	1	<i>Kalianpur</i>	H.S.	0.0	0.0	0.31	{ 1836 1898
141	<i>Rangsanobo</i>	H.S.	- 2.7	0.0	0.32	1861	174	<i>Ahmadpur</i>	H.S.	- 0.1	0.0	0.37	1838
142	<i>Ruikusni</i>	H.S.	- 0.4	0.0	0.29	1858	175	<i>Bhimbhut</i>	H.S.	- 0.2	- 0.1	0.39	1838
138	<i>Alangjāni</i>	T.S.	+ 0.6	0.0	0.79	1874	176	<i>Nilgarh</i>	H.S.	- 0.3	- 0.1	0.23	1839
134	<i>Nagarkhana</i>	H.S.	- 6.9	0.0	1.29	1905	177	<i>Badgaon</i>	H.S.	- 0.5	- 0.1	0.22	1839
143	<i>Fi Tān</i>	H.S.	- 6.9	+ 0.1	0.52	1865	178	<i>Sakri</i>	H.S.	- 0.6	- 0.2	0.30	1838
144	<i>Dattaung</i>	H.S.	- 7.1	+ 0.3	0.35	1866	179	<i>Somtana</i>	H.S.	- 0.7	- 0.2	0.27	1838
144	<i>Dattaung</i>	H.S.	- 7.1	+ 0.3	0.35	1866	180	<i>Dāmargida Obs.</i>	H.S.	- 0.9	- 0.3	0.40	1838
145	<i>Retkumauk</i>	H.S.	- 8.0	+ 0.3	0.41	1916	181	<i>Bolarum P. W. D. Office</i>	s.	- 1.1	- 0.3	0.31	1904
146	<i>Kyaunggyi</i>	s.	- 12.1	+ 0.3	0.34	1904	181	<i>Bolarum P. W. D. Office</i>	s.	- 1.1	- 0.3	0.31	1904
147	<i>Taungzun</i>	H.S.	- 10.0	+ 0.5	0.76	1884	182	<i>Pirmulo</i>	H.S.	- 1.2	- 0.3	0.23	1869
148	<i>Southern Moscos</i>	H.S.	- 10.0	+ 0.5	0.62	1877	183	<i>Vānākonda</i>	H.S.	- 1.2	- 0.3	0.20	1869
149	<i>Mergui Base-line E. End</i>	T.S.	- 10.0	+ 0.5	0.18	1882	184	<i>Singawāram</i>	H.S.	- 1.3	- 0.3	0.19	1871
150	<i>Mergui Base-line W. End</i>	T.S.	- 10.0	+ 0.5	0.32	1882	185	<i>Kālingkonda</i>	H.S.	- 1.3	- 0.3	0.24	1872
151	<i>Natkalintaung</i>	H.S.	- 10.0	+ 0.5	0.20	1881	186	<i>Sānjib</i>	H.S.	- 1.4	- 0.3	0.18	1860
152	<i>Minthangtaung</i>	H.S.	- 10.0	+ 0.5	0.24	1881	187	<i>Vizagapatam Base-line N. End</i>	S.	- 1.4	- 0.3	0.24	1863
146	<i>Kyaunggyi</i>	s.	- 12.1	+ 0.3	0.34	1904	101	<i>Calcutta Base-line S. End</i>	T.S.	- 5.9	- 0.1	0.37	1845
153	<i>Myayabelungkyo</i>	H.S.	- 11.1	+ 0.3	0.45	1889	188	<i>Patna</i>	T.S.	- 4.6	- 0.2	0.19	1852
154	<i>Toungoo</i>	S.	- 10.3	+ 0.2	0.23	1890	189	<i>Chandipur</i>	T.S.	- 4.3	- 0.2	0.30	1854
155	<i>Letpataung</i>	H.S.	- 9.6	+ 0.2	0.27	1891	190	<i>Cuttack</i>	H.S.	- 3.4	- 0.2	0.20	1854
156	<i>Taungpila</i>	H.S.	- 8.8	+ 0.2	0.21	1891	191	<i>Khundābolo</i>	H.S.	- 3.0	- 0.2	0.35	1857
157	<i>Mingun</i>	H.S.	- 8.0	+ 0.2	0.19	1892	192	<i>Rawal</i>	H.S.	- 1.8	- 0.3	0.30	1860
158	<i>Shiennuga</i>	H.S.	- 7.8	+ 0.2	0.36	1892	187	<i>Vizagapatam Base-line N. End</i>	S.	- 1.4	- 0.3	0.24	1863
159	<i>Male</i>	H.S.	- 7.3	+ 0.1	0.19	1892	192	<i>Rawal</i>	H.S.	- 1.8	- 0.3	0.30	1860
160	<i>Ubyetaung</i>	H.S.	- 6.8	+ 0.1	0.31	1894	193	<i>Deodonger</i>	H.S.	- 2.0	- 0.3	0.11	1914
161	<i>Thonbinzin</i>	H.S.	- 6.5	+ 0.1	0.18	1894	194	<i>Sindur</i>	H.S.	- 2.3	- 0.2	0.16	1913
162	<i>Seikpa</i>	H.S.	- 6.3	+ 0.1	0.24	1895	195	<i>Andhari</i>	H.S.	- 2.6	- 0.2	0.20	1913
163	<i>Tamunja</i>	H.S.	- 5.7	+ 0.1	0.30	1896	196	<i>Bhursu</i>	H.S.	- 2.9	- 0.1	0.17	1912
164	<i>Thyoliching</i>	H.S.	- 5.5	+ 0.1	0.19	1898	94	<i>Hurilāong</i>	H.S.	- 3.1	- 0.1	0.44	1849
165	<i>Loijing</i>	H.S.	- 5.0	+ 0.1	0.17	1899	197	<i>Karaundi</i>	H.S.	- 1.2	0.0	0.33	1865
141	<i>Rangsanobo</i>	H.S.	- 2.7	0.0	0.32	1861	198	<i>Sarandi Pat</i>	H.S.	- 1.2	- 0.1	0.33	1865
144	<i>Dattaung</i>	H.S.	- 7.1	+ 0.3	0.35	1866	199	<i>Bhimsain</i>	H.S.	- 1.2	- 0.1	0.18	1866
166	<i>Yeponetaung</i>	H.S.	- 6.9	+ 0.3	0.60	1916	200	<i>Diwai</i>	H.S.	- 1.2	- 0.2	0.16	1867
163	<i>Tamunja</i>	H.S.	- 5.7	+ 0.1	0.30	1896	201	<i>Burgpāili</i>	H.S.	- 1.1	- 0.2	0.18	1867
							181	<i>Bolarum P.W.D. Office</i>	s.	- 1.1	- 0.3	0.31	1904

Additional correction for Helmer's Spheroid.



TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station	Corrections		Prob. able Error ±	Date of Observation	Serial Number	Station	Corrections		Prob. able Error ±	Date of Observation
		Everest's Spheroid	Helmert's Spheroid*					Everest's Spheroid	Helmert's Spheroid*		
91	Karāra H.S.	- 1.8	0.0	0.69	1842	230	Mangalore S.	- 4.0	-0.6	0.58	1873
202	Pathāidi T.S.	- 1.7	-0.1	0.39	1871	231	Nughallibēta H.S.	- 4.7	-0.6	0.23	1871
203	Itamni H.S.	- 1.6	-0.1	0.50	1872	229	Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
204	Karāra H.S.	- 1.5	-0.2	0.20	1873	229	Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
186	Sānjib H.S.	- 1.4	-0.3	0.18	1860	232	Anandalamalai H.S.	- 3.9	-0.5	0.10	1866
97	Tilabani H.S.	- 3.9	-0.1	0.76	1845	233	Injambākum H.S.	- 3.3	-0.5	0.17	1880
205	Kalsibhanga T.S.	- 4.4	-0.2	0.32	1849	234	St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
188	Patna H.S.	- 4.6	-0.2	0.19	1852	234	St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
181	Bolarum P.W.D. Office s.	- 1.1	-0.3	0.31	1904	235	Kistanna H.S.	- 2.7	-0.5	0.37	1864
206	Achola H.S.	- 1.4	-0.3	0.25	1840	236	Dānupa H.S.	- 2.4	-0.4	0.21	1863
207	Nitali H.S.	- 1.6	-0.3	0.27	1840	237	Dhūlipalla S.	- 2.3	-0.4	0.17	1868
208	Kanheri H.S.	- 1.7	-0.3	0.32	1837	238	Parampūdi H.S.	- 1.9	-0.4	0.13	1861
209	Aisunda H.S.	- 1.7	-0.3	0.32	1863	187	Vizagapatam Base-line N. End S.	- 1.4	-0.3	0.24	1863
210	Khānisura H.S.	- 1.8	-0.3	0.42	1846	214	Colaba Obs. S.	- 2.2	-0.3	...	1839
211	Dhauleshvar H.S.	- 1.9	-0.3	0.18	1838	239	Pāchvad H.S.	- 2.7	-0.4	0.16	1865
212	Māndvi H.S.	- 2.1	-0.3	0.32	1841	240	Karabgati H.S.	- 3.0	-0.4	0.22	1865
213	Karānja H.S.	- 2.1	-0.3	...	1839	241	Koramūr H.S.	- 3.6	-0.5	0.21	1873
214	Colaba Obs. S.	- 2.2	-0.3	...	1839	230	Mangalore S.	- 4.0	-0.6	0.58	1873
215	Deesa T.O. s.	- 4.4	0.0	0.26	1904	214	Colaba Obs. S.	- 2.2	-0.3	...	1839
216	Sonāda T.S.	- 3.8	-0.1	0.39	1851	242	Miryā H.S.	- 2.5	-0.3	0.89	1844
217	Putangdi H.S.	- 3.3	-0.2	0.30	1861	243	Chaukola H.S.	- 2.6	-0.3	0.83	1843
218	Sāler H.S.	- 2.7	-0.2	0.60	1845	244	Kumbhari H.S.	- 2.7	-0.4	0.61	1844
219	Pānera H.S.	- 2.6	-0.3	0.47	1843	240	Karabgati H.S.	- 3.0	-0.4	0.22	1865
220	Kalsubai H.S.	- 2.4	-0.3	0.64	1842	229	Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
214	Colaba Obs. S.	- 2.2	-0.3	...	1839	245	Bangalore Base-line N.E. End S.	- 5.5	-0.6	0.23	1870
25	Karachi Obs. s.	- 2.5	+0.1	0.19	1855	246	Kanjumalai H.S.	- 6.3	-0.7	0.21	1869
221	Hāthria H.S.	- 3.1	0.0	0.54	1856	247	Pachupālaiyam s.	- 6.7	-0.7	0.26	1870
222	Dungarpur H.S.	- 3.4	-0.1	0.31	1852	248	Kutipārai S.	- 7.2	-0.7	0.33	1873
223	Ingrodi T.S.	- 3.6	-0.1	0.32	1852	249	Rādhāpuram S.	- 7.8	-0.8	0.14	1869
216	Sonāda T.S.	- 3.8	-0.1	0.39	1851	250	Kudankulam Obs. s.	- 7.9	-0.8	0.21	1869
222	Dungarpur H.S.	- 3.4	-0.1	0.31	1852	234	St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
224	Kunkāvā T.S.	- 3.4	-0.1	0.27	1853	251	Kallapat Trestle S.	- 3.9	-0.5	0.17	1879
7	Arantia S.	- 0.6	0.0	0.32	1850	252	Nayinipiriyān „ S.	- 4.5	-0.6	0.14	1870
225	Indrāwan T.S.	- 1.0	-0.1	0.29	1847	253	Pātharrankota S.	- 5.1	-0.6	0.14	1877
226	Valvādi H.S.	- 1.4	-0.2	0.56	1846	254	Manēgandi S.	- 6.0	-0.7	0.25	1876
210	Khānisura H.S.	- 1.8	-0.3	0.42	1846	255	Rāmnād S.	- 6.4	-0.7	0.39	1875
181	Bolarum P.W.D. office s.	- 1.1	-0.3	0.31	1904	250	Kudankulam Obs. s.	- 7.9	-0.8	0.21	1869
227	Kodungal S.	- 2.2	-0.4	0.18	1872						
228	Darur H.S.	- 2.7	-0.4	0.91	1871						
229	Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870						

\* Additional correction for Helmert's Spheroid.

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Achola H.S.	231	206	7	Birond H.S.	179		20	Dehra Dun Base, E., S.	146		6
Agra-group E. Point	201		6	Bisaul T.S.	318	115	19	Dehra Dun Haig Obs. S.	170		6
Agra-group N. Point	199		6	Bithnok H.S.	71		62	Dehra Dun Obs. (Old) S.	169	33	6
Agra-group S. Point	203		6	Black s.	288		9	" No. VI } between	165		20
Agra-group W. Point	187		6	Bolarum P.W.D. Long. S.	260	181	60	" No. V } Dehra and	166		20
Agra Longitude S.	200		6	Bolikonda H.S.	265		43	" No. IV } Rājpur	167		20
Agra Parade Point	202		6	Bombay Colaba Long. S.	111		7	" No. III }	168		20
Ahmadpur H.S.	214	174	8	Bombay Colaba Obs. S.	112	214	7	Deodongar H.S.	382	193	65
Akampalle h.s.	239		9	Bommasandra s.	267		9	Deo Dongri H.S.	106		18
Akbar S.	63	82	37	Bostān T.S.	155		6	Dera Dun Panāh S.	29	44	32
Aknāpur T.S.	402	100	5	Bowra T.S.	140	66	33	Devanūr s.	238		9
Akyub Longitude S.	419		44	Budhon H.S.	209	88	5	Devaragat h.s.	248		9
Alamkhān T.S.	35	22	25	Bulāwāla h.s.	150		6	Dewarsān T.S.	300		3
Alamvādi H.S.	103		10	Bulbul H.S.	367		6	Dhaiguan S.	127		18
Alangjāni T.S.	395	138	34	Burgpāuli H.S.	253	201	53	Dhamanva T.S.	94		26
Algi H.S.	206		2	Calcutta Base, S., T.S.	403	101	5	Dhānura s.	221		8
Alibagh Observatory S.	115		9	Calcutta Longitude s.	404		5	Dhaulesvar H.S.	122	211	7
Aloākāndi T.S.	397	136	56	Chamardi H.S.	53		30	Dhūlipalla S.	337	237	46
Alsunda H.S.	129	209	7	Chamu H.S.	73		62	Didāwa H.S.	45	15	25
Amritsar Longitude S.	65		23	Chandaos T.S.	156		6	Diwai H.S.	251	200	53
Ameot H.S.	145		6	Chandipur T.S.	380	189	24	Dōddagunta s.	270		9
Amūa H.S.	326	90	5	Chanduria T.S.	392		16	Dotra s.	222		8
Anandalamalai H.S.	274	232	54	Chānga H.S.	38	19	25	Dūbāuli T.S.	361		14
Anandbās T.S.	401	128	16	Chaniāna H.S.	80		26	Dumb h.s.	21		*
Andar H.S.	12	49	32	Charnādhānga T.S.	393		16	Dungarpur H.S.	49	222	23
Andhari H.S.	370	195	85	Chaukola H.S.	131	243	11	Eton T.S.	299		3
Andiāri H.S.	207		2	Chendwār (old) H.S.	371	95	5	Fi Tān H.S.	418	143	52
Ankora H.S.	252		53	Chikalgurki s.	262		9	Fyzabad Longitude S.	317		19
Aranliā S.	89	7	25	Chittagong Longitude S.	412		44	Gandpahar H.S.	11	51	32
Arasākulam S.	281		9	Chūni T.S.	365	120	20	Gang Choti H.S.	54	72	77
Asu H.S.	40	84	64	Chūlli T.S.	36	23	25	Gangapur T.S.	408	131	48
Ataro Bānki T.S.	394	139	34	Colaba Observatory S.	112	214	7	Garinda s.	90	68	23
Badgaon H.S.	220	177	8	Cuttack H.S.	372	190	24	Gattinārāyantippa h.s.	247		9
Bahak H.S.	158		73	Dadnara T.S.	304		20	Gaus T.S.	321		20
Bajamara H.S.	144		73	Daiādhari H.S.	191		6	Ghorārāo H.S.	100		10
Bandūr s.	263		9	Dājil S.	30	43	32	Godhna T.S.	152		6
Bangalore Base, N.E., S.	268	245	9	Dalea H.S.	332		58	Gogipatri H.S.	454	74	22
Bangalore Base, S.W., S.	269	229	9	Dāmargīda Obs. S.	237	180	8	Gora H.S.	325	93	5
Banog H.S.	160	32	6	Dānapa H.S.	271	236	46	Gudali H.S.	347		46
Bānsagopāl T.S.	177		2	Dangarvadi H.S.	51		28	Gundak H.S.	6	54	76
Bānekho H.S.	182	62	33	Dāowāla T.S.	24	86	64	Gurārāo H.S.	184		25
Bāsādela T.S.	315	105	20	Dargawa H.S.	211		2	Gūrmī T.S.	204	107	2
Bellary Longitude s.	258		9	Dariāpur T.S.	378		24	Gūru Sikkar H.S.	76	11	25
Bhānar T.S.	25	41	32	Darur H.S.	245	228	9	Gurwāni H.S.	320	92	5
Bhaorāsa H.S.	208		6	Darutippa S.	272		46	Halda s.	232		8
Bhimbhat H.S.	216	175	8	Dasti S.	9		*	Halkāchar T.S.	396	137	56
Bhimsain H.S.	228	199	53	Datāiri T.S.	154	30	6	Harnāsa T.S.	108		18
Bhursu H.S.	369	196	5	Dattaung H.S.	421	144	52	Harpālsid T.S.	173		2
Bichwi H.S.	364	126	27	Daulatpur T.S.	406	130	48	Hātbenā H.S.	339		58
Bihar H.S.	362		14	Dawa H.S.	409	140	44	Hāthria H.S.	48	221	35
Birona S.	78	12	25	Deesa Telegraph Office s.	79	215	20	h.s.	149		6

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Hirdepur	T.S.	324	116	15	Kheri	T.S.	141	65	33	Majhūr	H.S.	205		2
Hönnavalli	H.S.	137		49	Khimūāna	T.S.	67		23	Mal	H.S.	383		24
Hönnūr	H.S.	264		9	Khirsar	H.S.	62		62	Malar	H.S.	41	83	64
Hurilāong	H.S.	356	94	6	Khojak	H.S.	4		†	Male	H.S.	423	159	66
Imlia	T.S.	306		12	Khori	T.S.	37	21	25	Malūncha	H.S.	375	98	5
Indrāwan	T.S.	107	225	18	Khujnaur	s.	147		20	Mandāla	s.	235		8
Injambākam	H.S.	350	233	54	Khundābolo	H.S.	381	191	24	Mandresa	T.S.	59	80	45
Inrogdi	T.S.	52	223	29	Kidarkanta	H.S.	157		22	Māndvi	H.S.	118	212	7
Isampur	H.S.	139		33	Kisanen Chapper	H.S.	3	57	74	Mangāndi	S.	287	254	63
Jabalpur Longitude	s.	227		53	Kistama	H.S.	273	235	46	Mangalore Longitude	S.	134	230	64
Jalālpur	T.S.	352	124	21	Kodangal	S.	240		9	Manichauk	T.S.	313		20
Jalpaiguri Longitude	s.	389	122	34	Koh-i-Malik Siah	H.S.	1	59	74	Marār	T.S.	319	114	19
Jambo	H.S.	72	77	62	Koramūr	H.S.	133	241	49	Martaban	h.s.	440		52
Jāoli	H.S.	56	47	22	Kudankulam Obs.	S.	284	250	9	Mashelak	H.S.	5	53	76
Jarūra	T.S.	297		3	Kumbhāri	H.S.	132	244	11	Māsi	T.S.	305	104	20
Jetgarh	H.S.	85		23	Kumon Bum	H.S.	430	173	80	Māta-ka-hūra	H.S.	185	4	25
Jhambhera	T.S.	64	81	45	Kumtum Bum	H.S.	431	172	80	Māvinhūnda	H.S.	125		49
Jharipani (IX)	h.s.	162	36	20	Kundgol	H.S.	136		49	Mednipur	T.S.	354	123	21
Jharkil	T.S.	27	45	32	Kunkāvāv	T.S.	50	224	28	Medwāni	H.S.	138	48	22
Kainath	H.S.	93		26	Kurseong	h.s.	387		20	Mergui Base E,	T.S.	446	149	52
Kaliāna	S.	153	31	6	Kutipārai	S.	289	248	9	Mergui Base W.,	T.S.	447	150	52
Kaliānpur	H.S.	195	1	6	Kyaunggyi	S.	428	146	52	Miāni	T.S.	23	42	32
Kaliānpur	T.S.	180	102	20	Lachkuwa	h.s.	171		6	Mingun	H.S.	425	157	66
Kālingkonda	H.S.	341	185	43	Lādi	H.S.	215		8	Minthangtaung	H.S.	448	152	52
Kallapat	Trestle S.	292	251	63	Lādinslar	T.S.	33	79	45	Mira Donger	H.S.	117		11
Kalsibhānga	T.S.	377	205	17	Lakarwas	H.S.	83		25	Mirya	H.S.	120	242	11
Kalsubai	H.S.	116	220	10	Lakhinagar	T.S.	407	132	48	Mohammadabad	T.S.	210	110	4
Kāmkhera	H.S.	213		25	Lambatach	H.S.	143		22	Mooltan Longitude	S.	32		32
Kānākhhera	T.S.	301		3	Letpataung	H.S.	437	155	66	Morali	H.S.	95		26
Kanheri	H.S.	128	208	7	Linganapalle	h.s.	241		9	Moulmein Longitude	S.	441		44
Kanjamalai	H.S.	286	246	9	Lingmāra	H.S.	330		53	Mugrala	H.S.	61	78	62
Kankesvar	H.S.	114		11	Lohārgara	T.S.	391		16	Murree	h.s.	55	71	22
Kānkra	H.S.	183	61	33	Loi Hpa Lang	H.S.	433	168	72	Murree Observatory	s.	451		22
Kānnagar	H.S.	81	10	25	Loi Hpatan	H.S.	434	169	72	Mussooree Dome	H.S.	161	37	6
Karabgati	H.S.	126	240	49	Loi Hsam Hsum	H.S.	436	171	72	Myayabeingkyo	H.S.	439	153	16
Karachi Base S,	S.	14	39	32	Loijing	H.S.	413	165	68	Nagrkhāna	H.S.	411	134	52
Karachi Longitude	S.	16		22	Loi Kiipma	H.S.	435	170	72	Nagarcoil Longitude	S.	283		9
Karachi Observatory	S.	17	25	32	Lora	H.S.	327		5	Nag Tibā	H.S.	159	38	73
Karanja	H.S.	113	213	7	Losalli	S.	197		2	Nāharmāu	H.S.	225		5
Karāra	H.S.	312	91	5	Lūnki	H.S.	42	17	25	Namthabad	s.	261		9
Karāundi	H.S.	226	197	53	Mach	h.s.	8		†	Natkalintaung	H.S.	445	151	52
Kardo	H.S.	92		26	Madhpur	T.S.	376	99	5	Naunangarhi	T.S.	351	119	20
Karia	H.S.	340	204	58	Madhpur	T.S.	400	129	16	Navalūr	H.S.	135		49
Kārothol	H.S.	13	24	25	Madras Observatory	S.	348		54	Nayinipiriyan Trestle	S.	294	252	63
Kūtpālaiyam	s.	277		9	Mahabaleswar	H.S.	119		11	Niālamari	H.S.	257		46
Kuulia	H.S.	357		*	Mahadeo Pokra	H.S.	358		*	Nilgarh	H.S.	217	176	8
Ken	H.S.	130		7	Mahar	H.S.	363		14	Nimbāgal	s.	265		9
Kesri	H.S.	189	27	6	Mahesari	T.S.	174		6	Nimkār	T.S.	298	111	3
Khāmor	H.S.	87		23	Mahwari	H.S.	368		5	Nitali	H.S.	230	207	7
Khankharia	S.	47	13	25	Mairāb-ka-Shāhar	T.S.	39	20	25	Noh	T.S.	186	29	6
Khānpisura	H.S.	121	210	7	Majala	H.S.	124		49	Nojli	T.S.	151		6

Obs. = Observatory. \* Special Triangulation. † Old Baluchistān Triangulation. ‡ Thuyetno and Cape Negrais Series.

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Nuāon	T.S.	353	21	Rām Thal	S.	70	23	Southern Moscos	H.S.	443	148	52
Nughallibēta	H.S.	260	231	Rāmuapur (old)	T.S.	296	103	Spurpoint (VIII)	h.s.	163	35	20
Ongole	H.S.	346	46	Rangir (old)	S.	212	89	St. Thomas's Mount	Tres. S.	349	234	54
Orejhār	S.	316	106	Rāngrai	s.	219	8	Sultan-ka-Got	T.S.	22		32
Oria	h.s.	77	25	Ranganobob	H.S.	399	141	Sūrāntal	H.S.	193		6
Pabhosa	H.S.	311	112	Rānigarh	H.S.	172	20	Takalkhera	s.	218		8
Pachapālaiyam	s.	276	247	Ranjitgarh	T.S.	57	23	Talegaon	s.	236		8
Pāchvad	H.S.	123	239	Rāwal	H.S.	343	192	Tamunja	H.S.	415	163	68
Pahārgarh	H.S.	190	26	Retkamauk	H.S.	427	145	Tunakarukulam	S.	280		9
Palladpur	T.S.	360	14	Rewat	H.S.	84	23	Tarbhān	S.	104		10
Pāldi	H.S.	97	29	Robat	S.	449	†	Tāsing	H.S.	181	63	38
Pandalagudi	s.	290	9	Rojhra	H.S.	43	18	Taungpila	H.S.	426	156	66
Paphra	T.S.	31	87	Rustamgarhi	h.s.	452	75	Taungzun	H.S.	442	147	52
Pamampūdi	H.S.	338	238	Sakri	H.S.	223	178	Telu	H.S.	60		62
Pārasnāth	H.S.	373	96	Sāler	H.S.	105	218	Teona	H.S.	355		21
Parewa	T.S.	308	20	Salighar	H.S.	20	55	Tepri	T.S.	405	135	56
Pariāon	T.S.	310	20	Salimpur	T.S.	198	2	Thikri	H.S.	109		18
Pārnera	H.S.	98	219	Salot	H.S.	192	3	Thob	H.S.	74	76	62
Pātāngdi	H.S.	99	217	Samdari	H.S.	75	62	Thonbinzin	H.S.	417	161	66
Pathāndi	T.S.	333	202	Samenda	T.S.	323	117	Thyoliching	H.S.	414	164	68
Pātharankota	S.	295	253	Sānd	H.S.	88	8	Tiki	H.S.	82	9	25
Pathārdi	T.S.	314	20	Sandawat	H.S.	444	52	Tilabani	H.S.	374	97	5
Patna	T.S.	379	188	Sangatpur	T.S.	66	70	Tinsia	H.S.	196		25
Pāvāgad	H.S.	101	10	Sānjib	H.S.	342	186	Tiruvēndipuram	s.	293		63
Pāvāgāda	H.S.	266	9	Sankrāo	T.S.	178	108	Tonglu	h.s.	385		20
Pavia	H.S.	302	3	Sarandi Pat	H.S.	328	198	Tōnsalgutta	s.	243		9
Pēddapād	s.	244	9	Sarey Khan Latitude	S.	329	53	Toungoo	S.	438	154	66
Peshawar Longitude	S.	18	32	Sarkāra	T.S.	175	2	Tounsa	T.S.	28	56	76
Pballut	h.s.	384	20	Sarla	S.	46	14	Tungat	h.s.	246		9
Pialmudi	s.	242	9	Saugor	H.S.	224	5	Tuzgi	H.S.	2	58	74
Piaro	H.S.	10	50	Sawaiपुर	T.S.	68	23	Ubyetaung	H.S.	422	160	66
Pirmulo	H.S.	249	182	Seikpa	H.S.	416	162	Umakkhel	H.S.	19	46	32
Poshkar	H.S.	453	73	Semu Tān	H.S.	410	133	Usira	H.S.	188	28	6
Pota	T.S.	359	125	Senchal	h.s.	386	20	Utiāmau	T.S.	307		12
Potenda	S.	303	3	Shāhpur	T.S.	58	23	Valvādi	H.S.	110	226	18
Prome Longitude	S.	429	52	Sheimmaga	H.S.	424	158	Vānākonda	H.S.	256	183	43
Punnē Observatory	S.	285	9	Shorpur	H.S.	148	6	Vijayāpati	S.	282		9
Quetta Telegraph Office	s.	7	60	Shūlakarai	s.	278	9	Vijnot	T.S.	26	85	64
Rādhāpuram	S.	279	249	Sidhpur	S.	102	10	Virāria	H.S.	44	16	25
Raikuanj	H.S.	398	142	Siliguri	s.	388	20	Vizagapatām Base, N.	S.	344	187	24
Rājābāri	T.S.	322	118	Sindur	H.S.	335	194	Voi	s.	233		8
Rājgarh	H.S.	86	67	Singāwāram	H.S.	336	184	Waltair Longitude	S.	345		43
Rājpur	h.s.	164	34	Sinpitaung	H.S.	432	167	Yeponetaung	H.S.	420	166	71
Rājoli	H.S.	229	53	Sirkunda	T.S.	366	127	Yē-rugunta	h.s.	259		9
Rākhi	T.S.	142	64	Sironj Base, N.E.	S.	194	6	Yētimalai	S.	275		9
Ramai	H.S.	334	203	Sirsa	S.	69	69	Yūsuf	S.	34	40	32
Rāmāgh Observatory	s.	15	32	Sirsa	T.S.	176	109	Zuwa	H.S.	450	52	74
Rānganj	T.S.	390	121	Sitāpār	H.S.	331	63					
Rāngir	H.S.	254	63	Somtana	H.S.	234	179					
Rānood	s.	291	255	Sonāda	T.S.	96	216					
Rānpura	H.S.	91	6	Sora	T.S.	309	113					

# Deflections of the Plumb-line

in terms of

any Spheroid.

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.											
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
1	30 C	Koh-i-Malik Siah H.S.	5393	G 29 51 31'95	G 60 52 19'71	A 321 52 15'7 G 321 52 0'4	Kachakoh E 0 42	+ 26'7	"	1	
	O	No. 449 <i>ut infra</i>									
2	L	Tuzgi H.S.	3131	G 28 53 14'38	G 62 14 58'97	A 216 16 46'3 G 216 16 37'3	Shuri E 0 40	+ 16'3		2	
3	34 C	Kisanen Chapper H.S.	4362	G 29 3 54'41	G 64 22 24'93	A 166 46 22'6 G 166 46 15'7	Malik Proji E 1 7	+ 12'4		3	
4	J	Khojak H.S.	7851	A 30 51 20'21 G 30 51 24'85	G 66 34 41'08				- 4'6	4	
5	J	Mashelak H.S.	7941	G 30 13 30'77	G 66 44 46'79	A 241 55 31'4 G 241 55 27'8	Takatu E 1 20	+ 6'2		5	
		No. 450 <i>ut infra</i>									
6	M	Gundak H.S.	8163	G 31 9 49'49	G 67 23 28'55	A 276 29 27'1 G 276 29 25'2	Basha E 0 56	+ 3'1		6	
7	N	Quetta Tel Office S.	5500	A 30 11 55'91 G 30 11 57'37	A 67 0 29'27 G 67 0 31'69	A 166 31 12'1 G 166 31 11'8	Takatu E 3 5	+ 0'5	- 1'5	7	
8	O	Mach h.s.	3522	A 29 52 20'46 G 29 52 31'51	G 67 18 39'42				- 11'1	8	
9	O	Dasti S.	316	A 29 0 27'61 G 29 0 29'93	G 67 53 51'59				- 2'3	9	
10	35 J	Piaro H.S.	1438	G 26 3 14'21	G 66 34 7'96	A 159 22 15'3 G 159 22 12'8	Kuliri E 0 16	+ 5'1		10	
11	M	Gandpahar H.S.	723	G 27 25 1'26	G 67 30 43'99	A 192 10 55'2 G 192 10 46'4	Kharko D 0 9	+ 17'0		11	
12	N	Andar H.S.	4047	G 26 1 22'07	G 67 12 10'58	A 181 7 6'5 G 181 7 1'7	Sulemani D 0 24	+ 9'8		12	
13	P	Károthol H.S.	260	A 24 53 44'78 G 24 53 46'69	G 67 53 32'47	A 121 36 58'0 G 121 36 55'3	Kara E 0 38	+ 5'8	- 1'9	13	
14	P	Karachi Base-line S. End S.	46	G 24 52 59'63	G 67 9 24'77	A 205 23 30'5 G 205 23 29'2	Karachi Base-line N. End E 0 10	+ 2'8		14	
15	P	Rāmbāgh Obsy. s.	...	A 24 51 20'58 G 24 51 21'44	G 67 0 55'21				- 0'9	15	
16	P	Karachi Long. S.		G 24 51 2'44	A 67 0 52'88 G 67 0 53'22				- 0'3	16	
17	P	Karachi Obsy. S.	35	A 24 49 50'14 G 24 49 50'25	G 67 1 35'13	A 221 39 9'5 G 221 39 8'4	Mutrani E 0 25	+ 2'4	- 0'1	17	
18	38 N	Peshawar Long. S.	...	A 71 33 14'63 G 71 33 0'27					+ 11'9	18	
19	P	Umārkhel H.S.	3036	G 32 25 31'07	G 71 15 20'79	A 73 9 24'0 G 73 9 15'3	Sistarg E 0 18	+ 13'7		19	
20	39 A	Salighar H.S.	8284	G 31 29 53'82	G 68 25 5'49	A 1 52 50'1 G 1 52 46'2	Tanispa E 0 33	+ 6'4		20	
21	D	Dumb h.s.	183	A 28 15 18'30 G 28 15 21'09	G 68 14 9'96				- 2'8	21	
22	D	Sultan-ka-Got T.S.	213	A 28 4 8'05 G 28 4 9'41	G 68 36 32'23				- 1'4	22	
23	H	Miāni T.S.	300	G 28 34 15'20	G 69 50 46'91	A 188 2 16'6 G 188 2 4'5	Routi D 0 5	+ 22'2		23	
24	H	Dāowāla T.S.	282	G 28 20 12'87	G 69 50 30'68	A 28 49 27'9 G 28 49 21'3	Ghundi D 0 3	+ 12'2		24	
25	H	Bhanar T.S.	256	G 28 8 55'00	G 69 17 11'38	A 197 50 8'8 G 197 50 0'7	Khai D 0 5	+ 15'1		25	
26	H	Vijnot T.S.	276	G 28 2 3'30	G 69 50 32'77	A 159 35 15'6 G 159 35 10'0	Dewari D 0 5	+ 10'5		26	

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

· · XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$ .			Deflection in Prime Vertical	Deflection in Meridian	
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$			
1			+8.80			-0.17			-0.58			+1.77			+10.11	+16.4		1
2			+8.23			-0.30			-0.55			+1.83			+9.56	+6.5		2
3			+7.08			-0.25			-0.48			+1.83			+8.57	+3.6		3
4	+1.64			-5.19			+0.98			+0.17			-1.78				-2.8	4
5			+5.71			-0.09			-0.38			+1.78			+7.38	-1.5		5
6			+5.30			0.00			-0.34			+1.74			+7.04	-4.2		6
7	+1.54	+6.36	+5.58	-4.74	-0.88	-0.09	+0.98	-0.09	-0.37	+0.17	+0.10	+1.78	-1.54	+5.26	+7.27	-7.5	0.0	7
8	+1.48			-4.52			+0.98			+0.16			-1.44				-9.7	8
9	+1.32			-3.89			+0.98			+0.16			-1.12				-1.2	9
10			+6.30			-0.61			-0.45			+2.05			+7.87	-3.0		10
11			+5.61			-0.39			-0.39			+1.95			+7.29	+9.5		11
12			+5.96			-0.57			-0.41			+2.05			+7.60	+2.0		12
13	+0.42		+5.73	-0.68		-0.69	+0.99		-0.41	+0.16		+2.13	+0.40		+7.40	-1.8	-2.3	13
14			+6.17			-0.57			-0.43			+2.13			+7.76	-5.2		14
15	+0.44			-0.64			+0.98			+0.17			+0.45				-1.4	15
16		+6.37			-0.90			-0.08			+0.01			+5.21		-5.5		16
17	+0.44		+6.25	-0.63		-0.78	+0.98		-0.43	+0.17		+2.14	+0.40		+7.83	-5.6	-0.6	17
18		+3.64			-0.49			-0.06			+0.17			+3.20		+8.7		18
19			+3.21				+0.00		-0.20			+1.70			+5.14	+8.1		19
20			+4.75				+0.02		-0.31			+1.71			+6.51	-0.4		20
21	+1.17			-3.33			+0.99			+0.15			-0.89				-1.9	21
22	+1.12			-3.19			+0.99			+0.14			-0.85				-0.5	22
23			+4.22			-0.22			-0.29			+1.89			+6.09	+15.7		23
24			+4.25			-0.22			-0.30			+1.91			+6.14	+5.9		24
25			+4.56			-0.26			-0.32			+1.91			+6.38	+8.3		25
26			+4.26			-0.24			-0.28			+1.92			+6.15	+4.1		26

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $v_0 = 0.31$ ,  $w_0 = 1.29$ . *Vide p. 2.*

## Deflections of the Plumb-line

		EVEREST'S SPHEROID.									
Serial No	Sheet No	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No	
27	39 I	Jharkil T.S.	532	" " "	" " "	" " "	" " "	" " "	" " "	27	
				G 31 21 13'65	G 70 59 44'80	A 208 7 9'2 G 208 7 3'6	Kasain D 0 3	+ 9'2			
28	J	Founsa T.S.	593	G 30 41 51'59	G 70 39 0'13	A 201 7 57'5 G 201 7 41'4	Langawala D 0 11	+ 27'1		28	
29	J	Dera Din Panāh S.	490	A 30 33 59'63 G 30 34 1 87	G 70 56 7'29	A 209 21 14'7 G 209 21 7'3	Sakwala D 0 4	+ 12'5	- 2'2	29	
30	K	Dājil S.	412	G 29 33 20'87	G 70 22 52'98	A 239 26 6'7 G 239 25 53'0	Dalura D 0 6	+ 24'2		30	
31	K	Papra T.S.	316	G 29 5 49'37	G 70 49 45'82	A 273 23 2'0 G 273 22 56'8	Chanikhan D 0 4	+ 9'3		31	
32	N	Mooltan Long. S.	420	A 30 10 56'15 G 30 10 58'70	A 71 26 22'19 G 71 26 27'39			- 4'5	- 2'6	32	
33	O	Lādinsir T.S.	468	A 29 21 39'83 G 29 21 41'58	G 71 59 19'71	A 195 0 23'1 G 195 0 22'1	Guddun D 0 6	+ 1'8	- 1'8	33	
34	40 A	Yūsuf S.	215	G 27 51 8'74	G 68 26 14'75	A 195 51 20'0 G 195 51 16'5	Salur D 0 5	+ 6'6		34	
35	I	Alumkhān T.S.	67	A 24 49 30'50 G 24 49 31'23	G 68 43 47'38	A 174 28 43'3 G 174 28 39'0	Hakimani D 0 4	+ 9'3	- 0'7	35	
36	I	Chūtlī T.S.	72	G 24 46 19'67	G 68 23 40'86	A 141 22 40'1 G 141 22 35'0	R.M.	+ 11'1		36	
37	G	Khori T.S.	63	A 25 0 30'60 G 25 0 31'53	G 69 3 5'32	A 247 8 33'5 G 247 8 32'8	Jan Mahamad D 0 5	+ 1'5	- 0'9	37	
38	H	Chānga H.S.	349	A 24 58 47'25 G 24 58 47'00	G 69 51 23'30	A 238 0 7'4 G 238 0 9'1	Sandohar E 0 0	- 3'6	+ 0'3	38	
39	H	Muirāb-ka-Shahar T.S.	44	G 24 50 10'79	G 69 20 25'56	A 181 11 36'7 G 181 11 34'6	Amisha D 0 5	+ 4'5		39	
40	I	Asu H.S.	479	G 27 10 32'14	G 70 10 59'67	A 201 37 32'6 G 201 37 31'7	Kalu D 0 9	+ 1'8		40	
41	J	Malar H.S.	328	G 26 2 25'80	G 70 3 36'19	A 161 26 22'4 G 161 26 23'4	Ramsar D 0 6	- 2'0		41	
42	L	Lūnki H.S.	588	A 24 58 18'73 G 24 58 23'15	G 70 39 42'32	A 255 9 1'4 G 255 8 58'2	Karebhit D 0 3	+ 6'9	- 4'4	42	
43	L	Rojbra H.S.	518	A 24 57 26'09 G 24 57 26'28	G 70 14 17'90	A 254 1 46'8 G 254 1 44'8	Dharindera D 0 3	+ 4'3	- 0'2	43	
44	I	Virāria H.S.	460	A 24 56 32'64 G 24 56 36'13	G 71 2 58'81	A 106 12 49'8 G 106 12 46'3	Karebhit D 0 1	+ 7'5	- 3'5	44	
45	P	Didāwa H.S.	212	A 24 51 17'32 G 24 51 19'36	G 71 18 57'69	A 72 32 16'7 G 72 32 14'0	Sohagi D 0 2	+ 5'8	- 2'0	45	
46	I	Saria S.	132	G 24 46 44'68	G 71 34 7'48	A 244 27 47'6 G 244 27 43'2	Dawal D 0 3	+ 9'5		46	
47	I	Khankhari S.	362	A 24 36 58'17 G 24 36 56'19	G 71 53 8'91	A 182 0 14'8 G 182 0 15'2	Kosia D 0 6	- 0'9	+ 2'0	47	
48	41 E	Hātbria H.S.	696	G 23 27 14'85	G 69 2 45'83	A 154 56 32'9 G 154 56 34'6	Suru Gandara E 0 1	- 3'9		48	
49	J	Dungarpur H.S.	404	A 22 48 8'85 G 22 48 13'54	G 70 50 39'44	A 199 56 38'7 G 199 56 32'8	Chinarwa D 0 15	+ 14'0	- 4'7	49	
50	K	Kunkāvā T.S.	621	A 21 39 10'31 G 21 39 11'06	G 70 56 8'89	A 161 59 40'0 G 161 59 35'5	Mumaiya D 0 10	+ 11'3	- 1'7	50	
51	L	Dangarvadi H.S.	96	A 20 42 52'01 G 20 43 0'53	G 70 56 5'21				- 8'5	51	
52	N	Ingrodi T.S.	152	A 22 57 2'50 G 22 57 7'58	G 71 48 34'12	A 198 26 44'5 G 198 26 39'4	Por D 0 8	+ 12'0	- 5'1	52	
53	O	Chamardi H.S.	361	A 21 49 23'88 G 21 49 26'65	G 71 55 4'34				- 2'8	53	
54	43 F	Ganga Choti H.S.	9989	G 34 4 31'37	G 73 44 52'15	A 174 38 11'1 G 174 38 24'0	Kafir Khan E 1 7	- 19'1		54	

A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.



XOV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1 \text{ km}$			Case II: $\delta b = 1 \text{ km}$			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200 \text{ metres, } 1/\epsilon = 298.3$					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
27	"	"	+3.92	"	"	0.00	"	"	-0.23	"	"	+1.74	"	"	+5.79	+3.1	"	27
28			+3.65			-0.03			-0.24			+1.77			+5.56	+21.2		28
29	+1.43		+3.52	-4.99		-0.03	+0.99		-0.24	+0.11		+1.78	-1.94		+5.46	+6.7	-0.3	29
30			+3.86			-0.05			-0.26			+1.83			+5.81	+18.1		30
31			+3.67			-0.14			-0.25			+1.87			+5.62	+3.5		31
32	+1.37	+2.17		-4.72	-0.30		+0.99	-0.04		+0.10	+0.06		-1.80	+1.86		-0.4	-0.8	32
33	+1.38		+3.02	-4.14		-0.11	+1.00		-0.20	+0.09		+1.85	-1.37		+5.04	-3.4	-0.4	33
34			+5.05			-0.30			-0.34			+1.93			+6.83	-0.6		34
35	+0.37		+5.25	-0.61		-0.05	+0.99		-0.37	+0.14		+2.15	+0.37		+7.03	+2.1	-1.1	35
36			+5.45			-0.68			-0.38			+2.15			+7.18	+3.7		36
37	+0.40		+5.03	-0.77		-0.61	+0.99		-0.35	+0.14		+2.14	+0.29		+6.86	-5.6	-1.2	37
38	+0.37		+4.57	-0.74		-0.55	+0.99		-0.32	+0.12		+2.14	+0.26		+6.47	-10.3	0.0	38
39			+4.89			-0.61			-0.35			+2.15			+6.73	-2.4		39
40			+4.16			-0.31			-0.28			+1.98			+6.07	-4.5		40
41			+4.38			-0.42			-0.30			+2.06			+6.30	-8.5		41
42	+0.34		+4.10	-0.73		-0.50	+0.99		-0.29	+0.11		+2.14	+0.22		+6.09	+0.6	-4.6	42
43	+0.35		+4.35	-0.72		-0.53	+0.99		-0.31	+0.12		+2.14	+0.25		+6.28	-2.2	-0.5	43
44	+0.32		+3.88	-0.70		-0.48	+0.99		-0.27	+0.11		+2.15	+0.22		+5.92	+1.4	-3.7	44
45	+0.29		+3.73	-0.63		-0.47	+0.99		-0.26	+0.10		+2.16	+0.24		+5.81	-0.2	-2.2	45
46			+3.59			-0.46			-0.25			+2.16			+5.69	+3.6		46
47	+0.21		+3.42	-0.43		-0.45	+1.00		-0.24	+0.09		+2.18	+0.31		+5.57	-0.5	+1.7	47
48			+5.26			-0.82			-0.73			+2.27			+7.06	-11.0		48
49	-0.29		+4.16	+1.12		-0.75	+0.99		-0.30	+0.11		+2.34	+1.01		+6.21	+8.0	-5.7	49
50	-0.66		+4.35	+1.14		-0.89	+0.99		-0.31	+0.11		+2.46	+0.59		+6.43	+5.2	-2.3	50
51	-0.98			+2.79			+1.00			+0.11		+1.61					-10.1	51
52	-0.27		+4.08	+0.99		-0.63	+1.00		-0.22	+0.08		+2.33	+0.94		+6.23	+6.0	-6.0	52
53	-0.63			+1.99			+1.00			+0.09		+1.35					-4.2	53
54			+1.95			+0.07			-0.12			+1.62			+3.90	-23.4		54

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *Fide p. 2.*

TABLE  
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	Height in feet	EVEREST'S SPHEROID.					Name and angular Elevation or Depression of observed station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflection†	Serial No.
				Latitude*	Longitude*	Azimuth*						
				° ' "	° ' "	° ' "	° ' "		"	"		
43	G	No. 451 <i>ut infra</i>										
55	G	Murree h.s.	7250	A 33 54 37.35 G 33 54 57.35		G 73 22 50.15				-20.0	55	
56	G	Jāoli H.S.	1918	G 33 16 48.85		G 73 10 26.50	A 214 27 23.4 G 214 27 22.2	Nerli E O 58	+ 1.8		56	
	J	No. 452 <i>ut infra</i>										
	J	No. 453 <i>ut infra</i>										
	K	No. 454 <i>ut infra</i>										
57	L	Ranjitgarh T.S.	900	A 32 35 6.52 G 32 35 12.11		G 74 37 12.48				-5.6	57	
58	P	Shāhpur T.S.	830	A 32 1 34.23 G 32 1 33.77		G 75 5 34.90				+0.5	58	
59	44 C	Mandresa T.S.	512	G 29 55 9.17		G 72 59 28.42	A 298 34 7.1 G 298 34 5.8	Gajlam D O 1	+ 2.3		59	
60	D	Telu H.S.	470	A 28 56 12.41 G 28 56 11.34		G 72 14 8.80				+ 1.1	60	
61	D	Mugrala H.S.	517	G 28 30 57.06		G 72 22 17.41	A 171 53 31.2 G 171 53 32.0	Habib D O 8	- 1.5		61	
62	D	Khirsar H.S.	603	A 28 29 43.75 G 28 29 40.91		G 72 39 32.34				+ 2.8	62	
63	F	Akbar S.	641	A 30 53 38.53 G 30 53 43.27		G 73 17 13.28	A 216 51 25.8 G 216 51 25.4	Firoz D O 5	+ 0.7	-4.7	63	
64	F	Jhambhera T.S.	630	G 30 5 50.27		G 73 49 16.30	A 185 27 27.5 G 185 27 29.9	Fatehgarh D O 8	- 4.1		64	
65	J	Amritsar Long. S.	770	A 31 38 2.51 G 31 37 58.72		A 74 52 26.46 G 74 52 23.45			+ 2.6	+ 3.8	65	
66	M	Sangatpur T.S.	779	A 31 17 35.42 G 31 17 34.43		G 75 2 19.27	A 61 34 52.8 G 61 34 49.1	Rabza D O 5	+ 6.1	+ 1.0	66	
67	N	Khimūāna T.S.	731	A 30 22 11.74 G 30 22 14.82		G 75 0 42.52				- 3.1	67	
68	O	Sawaipur T.S.	697	A 29 39 13.13 G 29 39 13.96		G 75 3 6.12				- 0.8	68	
69	O	Sirsa S.	738	G 29 31 35.39		G 75 1 14.76	A 17 11 0.2 G 17 10 58.1	Banka D O 6	+ 3.7		69	
70	P	Rām Thal S.	951	A 28 29 38.81 G 28 29 39.27		G 75 0 10.60				- 0.5	70	
71	45 A	Bithnok H.S.	774	A 27 53 24.97 G 27 53 22.03		G 72 39 54.55				+ 2.9	71	
72	A	Jambo H.S.	772	A 27 16 31.94 G 27 16 28.88		G 72 31 5.53	A 153 23 42.9 G 153 23 42.2	Sirad D O 7	+ 1.4	+ 3.1	72	
73	B	Chamu H.S.	1065	A 26 39 53.44 G 26 39 52.74		G 72 35 26.28				+ 0.7	73	
74	B	Thob H.S.	856	A 26 3 2.90 G 26 3 5.85		G 72 22 22.17	A 322 26 25.5 G 322 26 20.4	Samdari D O 8	+ 10.4	- 3.0	74	
75	C	Samdari H.S.	846	A 25 48 59.58 G 25 48 59.55		G 72 34 20.84				0.0	75	
76	D	Gūru Sikkar H.S.	5650	G 24 38 58.39		G 72 46 39.73	A 248 53 38.4 G 248 53 36.1	Belkn D 1 2	+ 5.0		76	
77	D	Oris h.s.	4200	A 24 37 47.63 G 24 37 50.96		G 72 45 39.23				- 3.3	77	
78	D	Birons S.	673	G 24 26 38.64		G 72 13 4.51	A 121 43 10.7 G 121 43 10.9	Sirsa D O 8	- 0.4		78	

\*A - Astronomical Value.  
G - Triangulated or Geodetic Value.

†Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$ .					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
55	+1.65			-7.15			+1.00			+0.07			-3.39				-16.6	55
56			+2.25			+0.06			-0.14			+1.66		+4.24		-2.9		56
57	+1.55			-6.33			+1.00			+0.05			-2.90				-2.7	57
58	+1.52			-6.12			+1.00			+0.04			-2.78				+3.3	58
59			+2.40			-0.06			-0.16			+1.82		+4.53		-2.5		59
60	+1.16			-3.83			+1.00			+0.09			-1.35				+2.5	60
61			+2.87			-0.13			-0.16			+1.90		+4.95		-6.7		61
62	+1.07			-3.50			+1.00			+0.08			-1.20				+4.0	62
63	+1.43		+2.27	-5.21		-0.02	+1.00		-0.15	+0.07	+1.77	-2.06		+4.32		-3.0	-2.6	63
64			+2.02			-0.04			-0.13		+1.81			+4.13		-8.5		64
65	+1.40	+1.67		-5.70	-0.23		+1.00	-0.03		+0.04	+0.13		-2.51	+1.53		+1.1	+6.3	65
66	+1.43		+1.35	-5.48		0.00	+1.00		-0.15	+0.04	+1.76	-2.39		+3.47		+2.3	+3.4	66
67	+1.31			-4.85			+1.00			+0.04			-2.03				-1.1	67
68	+1.22			-3.40			+1.00			+0.04			-1.04				+0.2	68
69			+1.40			-0.02			-0.10		+1.85			+3.64		-0.1		69
70	+1.01			-3.49			+1.00			+0.04			-1.30				+0.8	70
71	+0.96			-3.04			+1.00			+0.08			-0.97				+3.9	71
72	+0.83		+2.86	-2.57		-0.21	+1.00		-0.20	+0.08	+1.90	-0.72		+4.87		-3.7	+3.8	72
73	+0.70			-2.09			+1.00			+0.08			-0.50				+1.2	73
74	+0.57		+3.02	-1.59		-0.37	+1.00		-0.21	+0.08	+2.07	-0.18		+5.12		+5.3	-2.8	74
75	+0.50			-1.41			+1.00			+0.08			-0.18				+0.2	75
76			+2.80			-0.38			-0.20		+2.17			+5.13		-0.1		76
77	+0.19			-0.44			+1.00			+0.08			+0.27				-3.6	77
78			+3.24			-0.44			-0.23		+2.20			+5.43		-5.8		78

\*  $\delta a = 0.743$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . Vide p. 2.

TABLE  
Deflections of the Plumb-line

		EVEREST'S SPHEROID.									
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
				° ' "	° ' "	° ' "	° ' "	° ' "	° ' "	° ' "	
79	45 D	Deesa Tel. Office	s. 443	A 24 15 21.15 G 24 15 29.35	A 72 11 1.26 G 72 11 4.85	A 241 16 15.3 G 241 16 15.5	Jairāj E 1 21	- 0.4	- 8.2	79	
80	D	Chuniāna	H.S. 953	A 24 6 25.39 G 24 6 36.64	G 72 32 19.66				- 11.3	80	
81	H	Kānnagar	H.S. 3607	G 24 58 28.78	G 73 18 59.95	A 266 45 16.1 G 266 45 19.0	Māl Niver E O 2	- 6.2		81	
82	H	Tiki	H.S. 2369	A 24 55 34.52 G 24 55 38.24	G 73 50 44.41	A 106 4 27.1 G 106 4 23.4	Māl Niver E O 57	+ 8.0	- 3.7	82	
83	H	Lakarwas	H.S. 2574	A 24 31 41.05 G 24 31 47.99	G 73 49 43.23				- 6.9	83	
84	J	Rewat	H.S. 1542	A 26 53 54.74 G 26 53 53.08	G 74 16 53.79				+ 0.8	84	
85	J	Jetgarh	H.S. 1967	A 26 18 8.02 G 26 18 6.39	G 74 18 36.01				+ 1.6	85	
86	J	Rājgarh	H.S. 2618	G 26 17 49.31	G 74 35 44.37	A 156 43 41.0 G 156 43 39.9	Kisanpura D O 8	+ 2.2		86	
87	K	Khāmor	H.S. 1393	A 25 45 11.00 G 25 45 15.01	G 74 47 29.19				- 4.0	87	
88	L	Sānd	H.S. 1910	G 24 43 6.11	G 74 32 58.48	A 284 36 7.8 G 284 36 3.9	Mendki D O 7	+ 8.5		88	
89	L	Aramlia	S. 1532	A 24 25 2.66 G 24 25 7.27	G 74 59 5.69	A 244 39 1.5 G 244 38 58.9	Nauka Hūaro E O 5	+ 5.7	- 4.6	89	
90	M	Garinda	S. 1204	A 27 55 30.05 G 27 55 30.55	G 75 1 18.47	A 115 55 45.3 G 115 55 42.2	Biramsir D O 3	+ 5.8	- 0.5	90	
91	F	Rānpūra	H.S. 1920	G 24 28 44.16	G 75 26 52.24	A 260 5 35.8 G 260 5 35.0	Nimthār D O 16	+ 1.8		91	
92	46 A	Kardo	H.S. 807	A 23 57 2.27 G 23 57 10.02	G 72 43 52.88				- 7.8	92	
93	A	Kainath	H.S. 1385	A 23 51 14.99 G 23 51 23.79	G 72 58 51.75				- 8.8	93	
94	A	Dhamanya	T.S. 397	A 23 32 2.65 G 23 32 8.40	G 72 30 56.82				- 5.8	94	
95	A	Morali	H.S. 466	A 23 25 17.47 G 23 25 23.18	G 72 57 44.96				- 5.7	95	
96	A	Sonāda	T.S. 250	A 23 7 15.61 G 23 7 19.80	G 72 46 0.14	A 334 35 18.1 G 334 35 10.2	Mirzāpur D O 5	+ 18.5	- 4.3	96	
97	B	Pāldi	H.S. 208	A 22 53 51.60 G 22 53 57.07	G 72 31 30.86				- 5.5	97	
98	D	Pārnera	H.S. 614	A 20 32 49.83 G 20 32 56.85	G 72 56 56.42	A 349 0 27.3 G 349 0 13.6	Gambhargad E O 17	+ 36.5	- 7.0	98	
99	F	Patāngdi	H.S. 922	G 22 52 15.70	G 73 53 22.34	A 16 47 30.9 G 16 47 26.6	Bhor D O 1	+ 10.2		99	
100	F	Ghorārāo	H.S. 323	A 22 52 8.05 G 22 52 11.17	G 73 21 25.45				- 3.1	100	
101	F	Pāvāgad	H.S. 2721	A 22 27 39.95 G 22 27 44.33	G 73 31 1.07				- 4.4	101	
102	F	Sidhpur	S. 169	A 22 4 11.77 G 22 4 15.21	G 73 28 59.81				- 3.4	102	
103	G	Alamvāni	H.S. 848	A 21 34 30.45 G 21 34 34.13	G 73 30 9.20				- 3.7	103	
104	G	Tarbhār	S. 140	A 21 0 28.36 G 21 0 34.13	G 73 3 49.79				- 5.8	104	
105	H	Sāler	H.S. 5140	G 20 43 18.44	G 73 56 21.93	A 151 26 55.7 G 151 26 50.4	Dopāri D 1 29	+ 14.0		105	
106	M	Deo Dongri	H.S. 1727	A 23 26 43.17 G 23 26 47.79	G 75 32 16.99				- 4.6	106	

• A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $w_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$ .					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical	Deflection in Meridian	
79	+0.11	+3.28	+3.27	-0.13	-0.47	-0.46	+1.00	-0.04	-0.23	+0.09	0.00	+2.21	+0.43	+2.67	+5.46	- 5.9	- 8.6	79
80	+0.07			-0.04			+1.00			+0.09			+0.46				- 11.8	80
81			+2.55			-0.27			-0.18			+2.16			+4.89	- 11.1		81
82	+0.25		+2.24	-0.69		-0.29	+1.00		-0.16	+0.06		+2.15	+0.11		+4.57	+ 3.4	- 3.8	82
83	+0.15			-0.35			+1.00			+0.06			+0.26				- 7.2	83
84	+0.72			-2.27			+1.00			+0.06			-0.64				+ 1.4	84
85	+0.59			-1.80			+1.00			+0.05			-0.42				+ 2.0	85
86			+1.73			-0.17			-0.12			+2.06			+4.09	- 1.9		86
87	+0.45			-1.36			+1.00			+0.05			-0.24				- 3.8	87
88			+1.85			-0.24			-0.13			+2.18			+4.30	+ 4.2		88
89	+0.10		+1.50	-0.26		-0.22	+1.00		-0.11	+0.05		+2.20	+0.26		+4.12	+ 1.6	- 4.9	89
90	+0.92		+1.46	-3.07		-0.07	+1.00		-0.10	+0.06		+1.95	-1.04		+3.78	+ 1.8	+ 0.5	90
91			+1.32			-0.18			-0.09			+2.20			+3.90	- 2.1		91
92	+0.01			+0.13			+1.00			+0.08			+0.52				- 8.3	92
93	-0.02			+0.22			+1.00			+0.07			+0.54				- 9.3	93
94	-0.11			+0.49			+1.00			+0.08			+0.68				- 6.5	94
95	-0.15			+0.59			+1.00			+0.08			+0.70				- 6.4	95
96	-0.24		+3.03	+0.85		-0.52	+1.00		-0.26	+0.08		+2.32	+0.82		+5.32	+ 13.4	- 5.1	96
97	-0.30			+1.04			+1.00			+0.08			+0.91				- 6.4	97
98	-1.09		+3.17	+3.25		-0.75	+1.00		-0.24	+0.08		+2.59	+1.81		+5.63	+ 31.7	- 8.8	98
99			+2.35			-0.42			-0.17			+2.34			+4.74	+ 6.0		99
100	-0.32			+1.07			+1.00			+0.07			+0.90				- 4.0	100
101	-0.46			+1.43			+1.00			+0.07			+1.04				- 5.4	101
102	-0.58			+1.78			+1.00			+0.07			+1.18				- 4.6	102
103	-0.48			+2.23			+1.00			+0.07			+1.61				- 5.3	103
104	-0.59			+2.73			+1.00			+0.07			+1.88				- 7.7	104
105			+2.49			-0.74			-0.18			+2.57			+5.03	+ 9.5		105
106	-0.18			+0.68			+1.00			+0.04			+0.69				- 5.3	106

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $w_0 = 0.31$ ,  $w_0 = 1.20$ . Vide p. 2.

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.											
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
107	46N	Indrāwan T.S.	1834	"	"	"	"	"	"	107	
				G 22 48 48.54	G 75 10 56.60	A 273 34 2.7 G 273 34 1.9	Harnāsa D 0 9	+ 1.9			
108	N	Harnāsa T.S.	1818	A 22 47 26.71 G 22 47 29.91	G 75 33 10.15				- 3.2	108	
109	N	Thikri H.S.	854	A 22 1 3.92 G 22 1 2.77	G 75 24 49.98				+ 1.2	109	
110	P	Vulvādi H.S.	1128	A 20 44 21.27 G 20 44 27.73	G 75 11 7.12	A 166 52 6.2 G 166 51 59.8	Ajnāu D 0 7	+ 16.9	- 6.5	110	
111	47B	Bombay, Colaba Long S.	75	A 18 53 39.15 G 18 53 49.48	A 72 48 55.82 G 72 48 49.10			+ 6.4	- 10.3	111	
112	B	Bombay, Colaba Obsy. S.	63	G 18 53 46.51	G 72 48 47.31	A 288 5 27.7 G 288 5 24.5	Karānja E 1 7	+ 9.3		112	
113	B	Karānja H.S.	997	A 18 51 13.79 G 18 51 24.99	G 72 56 21.88	A 173 10 2.5 G 173 9 56.1	Trombay D 0 4	+ 18.7	- 11.2	113	
114	B	Kankesvar H.S.	1260	A 18 44 17.89 G 18 44 28.16	G 72 55 34.09				- 10.3	114	
115	B	Alibagh Obsy. S.	10	A 18 38 26.35 G 18 38 36.60	G 72 52 12.42				- 10.3	115	
116	E	Kalsubai H.S.	5400	A 19 35 57.89 G 19 36 1.76	G 73 42 35.26	A 73 2 14.5 G 73 2 11.8	Kāmandrug D 1 1	+ 7.6	- 3.9	116	
117	F	Mira Donger H.S.	1863	A 18 40 55.97 G 18 41 1.68	G 73 9 48.88				- 5.7	117	
118	F	Māndvi H.S.	4121	A 18 37 47.94 G 18 37 51.11	G 73 32 21.71	A 271 15 3.9 G 271 15 7.8	Dighi D 0 56	- 11.6	- 3.2	118	
119	G	Mahubaleswar H.S.	4719	A 17 55 9.91 G 17 55 15.55	G 73 40 17.41				- 5.6	119	
120	G	Miryā H.S.	473	A 17 1 29.65 G 17 1 35.92	G 73 15 39.43	A 167 2 11.4 G 167 2 7.4	Achūr D 0 12	+ 13.1	- 6.3	120	
121	J	Khānpisura H.S.	2751	A 18 45 22.60 G 18 45 30.65	G 74 47 49.81	A 191 14 39.3 G 191 14 43.9	Azargaon D 0 2	- 13.5	- 8.1	121	
122	J	Dhaulewar H.S.	2939	A 18 25 42.84 G 18 25 41.64	G 74 9 48.48	A 198 21 22.6 G 198 21 24.4	Bābulsār D 0 31	- 5.4	+ 1.2	122	
123	K	Pāchvad H.S.	3138	G 17 31 1.97	G 74 39 43.71	A 331 12 27.4 G 331 12 29.6	Palsi D 0 13	- 7.0		123	
124	L	Mājala H.S.	2613	A 16 46 55.45 G 16 46 56.82	G 74 26 55.57				- 1.4	124	
125	L	Māvinhūnda H.S.	2582	A 16 25 4.47 G 16 25 4.19	G 74 47 40.38				+ 0.3	125	
126	L	Karabgati H.S.	2544	G 16 7 34.87	G 74 47 56.35	A 179 9 24.9 G 179 9 26.9	Māvinhūnda D 0 7	- 6.9		126	
127	M	Dhārgaon S.	1553	A 19 30 30.82 G 19 30 35.04	G 75 12 43.81				- 4.2	127	
128	N	Kenheri H.S.	2610	A 18 29 21.84 G 18 29 30.75	G 75 43 16.60	A 311 59 50.6 G 311 59 53.3	Garh Dāud D 0 16	- 8.1	- 8.9	128	
129	N	Aisunda H.S.	2165	G 18 26 52.37	G 75 0 35.11	A 227 31 58.4 G 227 32 1.8	Sutāra D 0 4	- 10.2		129	
130	N	Kem H.S.	1951	A 18 10 45.68 G 18 10 48.90	G 75 18 23.92				- 3.2	130	
131	48E	Omukola H.S.	2794	A 15 55 24.94 G 15 55 31.44	G 73 59 21.13	A 166 14 13.4 G 166 14 13.2	Valvan D 0 5	+ 0.7	- 6.5	131	
132	1	Kumbhāri H.S.	2898	A 15 9 4.31 G 15 9 1.80	G 74 17 47.20	A 154 15 30.5 G 154 15 33.4	Saiti D 0 30	+ 11.4	+ 2.5	132	
133	J	Koramūr H.S.	2525	A 14 8 1.71 G 14 8 6.59	G 74 58 24.07	A 235 28 6.8 G 235 28 0.7	Hōnnavalli E 0 3	- 11.5	- 4.9	133	
134	L	Mangalore Long S.	186	A 12 52 17.76 G 12 52 14.76	A 74 50 44.70 G 74 50 42.71	A 205 52 50.8 G 205 52 49.6	Mijār E 0 17	+ 5.2	+ 3.0	134	

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1 \text{ km}$			Case II: $\delta b = 1 \text{ km}$			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200 \text{ metres, } 1/e = 298.8$					
	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
107			+1.55			-0.31			-0.12			+2.35			+4.20	-2.1		107
108	-0.38			+1.15			+1.00			+0.04			+0.85				-4.1	108
109	-0.63			+1.83			+1.00			+0.04			+1.14				+0.1	109
110	-1.06		+1.66	+2.99		-0.39	+1.00		-0.12	+0.04		+2.57	+1.61		+4.53	+12.9	-8.1	110
111	-1.70	+2.89		+4.70	-0.41		+1.00	-0.03		+0.08	-0.09		+2.33	+2.24		+4.2	-12.6	111
112			+3.48			-0.99			-0.32			+2.81			+6.00	+4.2		112
113	-1.72		+3.40	+4.74		-0.97	+1.00		-0.31	+0.08		+2.81	+2.34		+5.95	+13.6	-13.5	113
114	-1.66			+4.85			+1.00			+0.08			+2.47				-12.8	114
115	-1.80			+4.95			+1.00			+0.08			+2.42				-12.7	115
116	-1.45		+2.76	+4.04		-0.73	+1.00		-0.21	+0.06		+2.72	+2.05		+5.44	+3.0	-6.0	116
117	-1.79			+4.91			+1.00			+0.07			+2.39				-8.1	117
118	-1.82		+2.83	+4.64		-0.82	+1.00		-0.21	+0.07		+2.69	+2.16		+5.41	-16.1	-5.4	118
119	-2.10			+5.65			+1.00			+0.06			+2.64				-8.2	119
120	-2.46		+3.44	+6.52		-1.17	+1.00		-0.26	+0.07		+3.11	+2.97		+6.24	+7.8	-9.3	120
121	-1.79		+2.07	+4.84		-0.60	+1.00		-0.16	+0.05		+2.84	+2.32		+5.08	-17.7	-10.4	121
122	-1.91		+2.55	+5.16		-0.76	+1.00		-0.19	+0.06		+2.88	+2.45		+5.45	-10.0	-1.3	122
123			+2.29			-0.75			-0.17			+3.03			+5.40	-11.1		123
124	-2.59			+6.77			+1.00			+0.05			+3.01				-4.4	124
125	-2.75			+7.13			+1.00			+0.05			+3.13				-2.8	125
126			+2.36			-0.92			-0.18			+3.27			+5.65	-11.2		126
127	-1.57			+4.29			+1.00			+0.04			+2.10				-0.3	127
128	-1.92		+1.42	+5.11		-0.20	+1.00		-0.11	+0.03		+2.88	+2.38		+4.84	-12.0	-11.3	128
129			+1.94			-0.36			-0.15			+2.88			+5.27	-14.6		129
130	-2.03			+5.40			+1.00			+0.04			+2.69				-5.9	130
131	-2.95		+3.04	+7.62		-1.15	+1.00		-0.24	+0.06		+3.32	+3.33		+7.08	-5.3	-9.8	131
132	-3.30		+2.91	+8.41		-1.17	+1.00		-0.23	+0.05		+3.49	+3.58		+6.2	-6.0	-1.1	132
133	-3.77		+2.47	+9.45		-1.09	+1.00		-0.19	+0.04		+3.74	+3.91		+6.24	-15.7	-8.8	133
134	-4.36	+1.67	+2.80	+10.77	-0.21	-1.36	+1.00	-0.01	-0.22	+0.05	-0.19	+4.09	+4.34	+1.13	+6.79	+1.0	-1.3	134

$\delta a = 0.924, \delta b = 0.743, u_0 = 0.31, w_0 = 1.29$  Page 2.

TABLE  
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	Height in feet	EVEREST'S SPHEROID.						Meridian Deflection†	Serial No.
				Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot $\lambda$ for azimuth or (A - G) cos $\lambda$ for longitude observations†			
135	48 M	Navalur H.S.	2445	A 15 25 28.48 G 15 25 31.17	G 75 3 15.42					- 2.7	135
136	M	Kundgol H.S.	2145	A 15 15 14.46 G 15 15 15.28	G 75 14 46.64					- 0.8	136
137	N	Hönnavalli H.S.	2777	A 14 16 30.76 G 14 16 32.46	G 75 10 58.48					- 1.7	137
138	53 A	Medwani H.S.	1935	G 31 17 40.45	G 76 11 58.72	A 64 43 36.5 G 64 43 41.6	Hiu D 0 54	- 8.4			138
139	B	Isanpur H.S.	874	A 30 38 16.03 G 30 38 20.01	G 76 6 40.41					- 4.0	139
140	B	Bowra T.S.	855	G 30 20 50.29	G 76 6 39.00	A 208 37 15.2 G 208 37 12.4	Sudhiwal D 0 4	+ 4.8			140
141	B	Kheri T.S.	822	G 30 5 9.30	G 76 5 54.58	A 212 55 16.6 G 212 55 17.2	Khampur D 0 3	- 1.0			141
142	C	Rakhi T.S.	785	A 29 17 20.76 G 29 17 21.28	G 76 6 47.49	A 208 30 58.2 G 208 30 55.5	Barowdha D 0 5	+ 4.8		- 0.5	142
143	E	Lambatach H.S.	10474	A 31 0 34.38 G 31 1 8.46	G 76 54 2.93					-34.1	143
144	F	Bajamara H.S.	9681	A 30 45 27.79 G 30 45 56.20	G 77 54 0.73					-28.4	144
145	F	Amsot H.S.	3140	A 30 22 16.02 G 30 22 44.86	G 77 41 14.77					-28.8	145
146	F	Debra Dun Base-line E. End S.	1967	A 30 16 37.26 G 30 17 7.35	G 77 58 30.74					-30.1	146
147	F	Khujaur s.	2576	A 30 15 56.70 G 30 16 23.63	G 77 52 58.67					-26.9	147
148	F	Shorpur H.S.	2916	A 30 13 15.30 G 30 13 44.43	G 77 57 30.61					-29.1	148
149	F	Hatni h.s.	3069	A 30 12 31.93 G 30 13 1.52	G 77 52 19.58					-29.6	149
150	F	Bulawala h.s.	2432	A 30 6 22.32 G 30 6 51.29	G 77 59 11.27					-29.0	150
151	G	Nojli T.S.	929	A 29 53 14.12 G 29 53 27.76	G 77 40 24.59					-13.6	151
152	G	Godhna T.S.	901	A 29 37 8.73 G 29 37 18.46	G 77 54 2.98					- 9.7	152
153	G	Kuliana s.	828	A 29 30 47.98 G 29 30 54.70	G 77 39 6.03	A 164 18 46.4 G 164 18 46.9	Dahera E 0 1	- 0.9		- 6.7	153
154	H	Dalniri T.S.	767	A 28 43 58.67 G 28 44 4.40	G 77 38 56.31	A 28 44 34.3 G 28 44 34.2	Bostan D 0 6	+ 0.2		- 5.8	154
155	H	Bostan T.S.	758	A 28 30 54.25 G 28 30 59.64	G 77 30 49.08					- 5.4	155
156	H	Chandnos T.S.	699	A 28 5 0.71 G 28 5 1.59	G 77 51 30.60					- 0.9	156
157	I	Kidarkanta H.S.	12509	A 31 0 51.58 G 31 1 21.71	G 78 10 23.37					-30.1	157
158	J	Bahak H.S.	9715	A 30 44 37.60 G 30 45 5.22	G 78 13 36.98					-27.6	158
159	J	Nag Tibba H.S.	9915	A 30 34 41.05 G 30 35 11.57	G 78 9 2.57	A 32 58 41.6 G 32 58 53.9	Eagle's Nest D 8 12	- 20.8		-30.5	159
160	J	Banog H.S.	7433	A 30 28 4.18 G 30 28 36.91	G 78 0 55.96	A 71 5 55.0 G 71 6 8.7	Amsot D 2 28	- 23.3		-32.7	160
161	J	Mussooree Dome Obay. H.S.	6937	A 30 27 4.02 G 30 27 40.55	G 78 4 17.41	A 280 22 46.8 G 280 23 0.5	Top Tibba E 1 7				161
						A 6 17 20.1 G 6 17 35.1	Cole's Satellite Station D 5 26	- 25.5		-36.5	

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.



XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$ .					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
135	-3'19	"	"	+8'13	"	"	+1'00	"	"	+0'04	"	"	+3'46	"	"	"	- 6.2	135
136	-3'27			+8'31			+1'00			+0'04			+3'51				- 4.3	136
137	-3'71			+9'31			+1'00			+0'04			+3'84				- 5.5	137
138		+0'75				0'00			-0'05			+1'76		+2'93		- 11.8		138
139	+1'34			-5'03			+1'00			+0'03			-2'16				- 1.8	139
140		+0'81				-0'02			-0'05			+1'81		+3'05		+ 1.4		140
141		+0'82				-0'02			-0'05			+1'82		+3'08		- 4.4		141
142	+1'15	+0'82		-4'08		-0'03	+1'00		-0'06	+0'03		+1'87	-1'63	+3'13		+ 1.3	+ 1.1	142
143	+1'38			-5'29			+1'00			0'00			-2'35				- 31.7	143
144	+1'35			-5'12			+1'00			0'00			-2'25				- 26.1	144
145	+1'30			-4'85			+1'00			0'00			-2'09				- 26.7	145
146	+1'29			-4'79			+1'00			-0'01			-2'07				- 28.0	146
147	+1'29			-4'78			+1'00			0'00			-2'06				- 24.8	147
148	+1'28			-4'75			+1'00			-0'01			-2'05				- 27.0	148
149	+1'28			-4'74			+1'00			0'00			-2'04				- 27.6	149
150	+1'27			-4'67			+1'00			-0'01			-2'00				- 27.0	150
151	+1'24			-4'51			+1'00			0'00			-1'90				- 11.7	151
152	+1'20			-4'32			+1'00			0'00			-1'81				- 7.9	152
153	+1'18	0'00		-4'24		0'00	+1'00		0'00	0'00		+1'85	-1'5	+2'39		- 3.8	- 4.9	153
154	+1'05	0'00		-3'67		0'00	+1'00		0'00	0'00		+1'90	-1'45	+2'45		- 2.7	- 4.3	154
155	+1'01			-3'51			+1'00			0'00			-1'36				- 4.0	155
156	+0'94			-3'18			+1'00			0'00			-1'21				+ 0.3	156
157	+1'38			-5'30			+1'00			-0'01			-2'35				- 27.7	157
158	+1'35			-5'11			+1'00			-0'01			-2'23				- 25.4	158
159	+1'33	-0'26		-4'99		+0'01	+1'00		+0'02	-0'01		+1'79	-2'18	+2'08		- 23.4	- 28.3	159
160	+1'31	-0'19		-4'92		0'00	+1'00		+0'01	-0'01		+1'80	-2'14	+2'15		- 26.0	- 30.6	160
161	+1'31	-0'22		-4'91		0'00	+1'00		+0'01	-0'01		+1'80	-2'14	+2'13		- 28.1	- 34.4	161

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.39$ . Vide p. 2.

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.										
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.
162	53 J	Jharipani (IX) h.s.	5150	A 30 24 17.55 G 30 25 10.05	0 1 "	A 14 24 59.9 G 14 25 18.0	Dalanwāla D 4 38	- 30.8	- 52.5	162
163	J	Spur point (VIII) h.s.	3850	A 30 23 44.55 G 30 24 37.72	0 1 "	A 17 59 32.4 G 17 59 49.1	Dalanwāla D 2 48	- 28.5	- 53.2	163
164	J	Rajpur h.s.	3500	A 30 23 9.15 G 30 23 56.83	0 1 "	A 24 15 4.4 G 24 15 21.8	Dalanwāla D 2 26	- 29.7	- 47.7	164
165	J	VI	3050	A 30 22 44.90 G 30 23 30.79	0 1 "				- 45.9	165
166	J	V	2980	A 30 22 7.46 G 30 22 51.83	0 1 "				- 44.4	166
167	J	IV	2780	A 30 21 26.78 G 30 22 8.93	0 1 "				- 42.2	167
168	J	III	2660	A 30 21 5.57 G 30 21 46.61	0 1 "				- 41.0	168
169	J	Dehra Dun Obsy. (Old) S.	2289	A 30 19 19.56 G 30 19 57.07	0 1 "	A 165 10 58.8 G 165 11 10.2	Banog E 5 20	- 19.5	- 37.5	169
170	J	Dehra Dun Haig Obsy. S.	2240	A 30 18 51.80 G 30 19 28.73	0 1 "			- 22.1	- 36.9	170
171	J	Lachkuwa h.s.	2674	A 30 4 5.34 G 30 4 34.24	0 1 "				- 28.9	171
172	J	Kānigarh H.S.	7055	A 30 3 34.80 G 30 4 4.47	0 1 "				- 29.7	172
173	K	Harpālsid T.S.	1000	A 29 39 22.24 G 29 39 50.84	0 1 "				- 28.6	173
174	K	Mahesari T.S.	821	A 29 30 8.18 G 29 30 18.21	0 1 "				- 10.0	174
175	K	Sarkāra T.S.	761	A 29 15 35.09 G 29 15 46.91	0 1 "				- 11.8	175
176	L	Sirsa T.S.	739	A 28 54 30.27 G 28 54 39.64	0 1 "	A 149 55 17.1 G 149 55 20.5	Milik D 0 4	- 6.2	- 9.4	176
177	L	Bānsgopāl T.S.	677	A 28 33 23.28 G 28 33 28.08	0 1 "				- 4.8	177
178	L	Sankrāo T.S.	670	A 28 2 28.92 G 28 2 29.00	0 1 "	A 185 44 20.9 G 185 44 18.8	Sakrora D 0 8	+ 3.9	- 0.1	178
179	O	Birond H.S.	6967	A 29 14 29.72 G 29 15 14.15	0 1 "				- 44.4	179
180	P	Kaliānpur T.S.	629	G 28 35 11.10	0 1 "	A 185 30 18.4 G 185 30 17.7	Donno D 0 4	+ 1.3		180
181	54 A	Tāsing H.S.	2050	A 27 52 59.49 G 27 52 59.47	0 1 "	A 77 55 36.5 G 77 55 31.6	Jilo E 0 13	+ 9.5	0.0	181
182	B	Bānskho H.S.	1870	A 26 50 2.37 G 26 50 7.89	0 1 "	A 148 40 55.6 G 148 40 51.9	Rāmgarh E 0 2	+ 7.3	- 5.5	182
183	C	Kānkra H.S.	1652	A 25 37 58.75 G 25 37 59.53	0 1 "	A 145 33 8.7 G 145 33 6.9	Bhojpur D 0 18	+ 3.8	- 0.8	183
184	D	Gurāria H.S.	1360	A 24 25 31.98 G 24 25 32.46	0 1 "	A 300 41 56.8 G 300 41 56.2	Kūsalpura D 0 4	+ 1.3	- 0.5	184
185	D	Māta-ka-hūra H.S.	1645	G 24 14 10.67	0 1 "	A 181 31 35.0 G 181 31 34.3	Sartal D 0 15	+ 1.6		185
186	E	Noh T.S.	710	A 27 50 53.13 G 27 50 53.08	0 1 "	A 50 22 36.5 G 50 22 33.4	Mānpur D 0 6	+ 5.9	+ 0.1	186
187	J	Agra-group W. Point	550	A 27 9 41.43 G 27 9 45.86	0 1 "				- 4.4	187
188	F	Usira H.S.	810	A 26 57 0.50 G 26 57 6.22	0 1 "	A 146 55 27.2 G 146 55 25.9	Madhoni D 0 12	+ 2.6	- 5.7	188
189	G	Kesri H.S.	1487	A 25 46 41.57 G 25 46 35.81	0 1 "	A 206 41 38.8 G 206 41 40.1	Dīn D 0 10	- 2.7	+ 5.8	189

\*A - Astronomical Value.  
G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a=1$ km			Case II: $\delta b=1$ km			Case III: Latitude $u_0=1''$			Case IV: Azimuth $w_0=1''$			$a=6378200$ metres, $1/\epsilon=298.3$					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
162	+1'31		-0'23	-4'88		0'00	+1'00		+0'02	-0'01		+1'80	-2'11		+2'11	-33'4	-50'4	162
163	+1'31		-0'23	-4'87		0'00	+1'00		+0'02	-0'01		+1'80	-2'11		+2'12	-33'1	-51'1	163
164	+1'30		-0'23	-4'86		+0'01	+1'00		+0'02	-0'01		+1'80	-2'11		+2'11	-32'3	-45'6	164
165	+1'30			-4'86			+1'00			-0'01			-2'11				-43'8	165
166	+1'30			-4'85			+1'00			-0'01			-2'10				-42'3	166
167	+1'30			-4'84			+1'00			-0'01			-2'10				-40'1	167
168	+1'30			-4'84			+1'00			-0'01			-2'10				-38'9	168
169	+1'30		-0'21	-4'82		0'00	+1'00		+0'01	-0'01		+1'81	-2'08		+2'13	-22'1	-35'4	169
170	+1'30	-0'24		-4'81	+0'03		+1'00	0'00		-0'01	+0'11		-2'08	-0'05		-22'1	-34'8	170
171	+1'26			-4'64			+1'00			-0'01			-1'98				-26'9	171
172	+1'26			-4'64			+1'00			-0'02			-1'98				-27'7	172
173	+1'20			-4'35			+1'00			-0'02			-1'83				-26'8	173
174	+1'18			-4'23			+1'00			-0'01			-1'75				-8'2	174
175	+1'14			-4'06			+1'00			-0'02			-1'66				-10'1	175
176	+1'09		-0'48	-3'80		+0'02	+1'00		+0'03	-0'02		+1'89	-1'53		+2'02	-8'6	-7'9	176
177	+1'02			-3'54			+1'00			-0'01			-1'40				-3'4	177
178	+0'93		-0'48	-3'15		+0'03	+1'00		+0'03	-0'01		+1'94	-1'19		+2'09	+1'4	+1'1	178
179	+1'15			-4'05			+1'00			-0'03			-1'68				-42'7	179
180			-1'19			+0'06			+0'08			+1'91			+1'43	-0'5		180
181	+0'90		+0'80	-3'03		-0'05	+1'00		-0'06	+0'02		+1'95	-1'08		+3'20	+5'9	+1'1	181
182	+0'69		+0'85	-2'22		-0'07	+1'00		-0'06	+0'03		+2'02	-1'68		+3'33	+3'8	-3'8	182
183	+0'40		+0'89	-1'26		-0'10	+1'00		-0'06	+0'03		+2'11	-0'22		+3'45	+0'3	-0'6	183
184	+0'09		+0'94	-0'26		-0'13	+1'00		-0'07	+0'03		+2'21	-0'23		+3'60	-2'3	-0'3	184
185			+0'63			-0'09				-0'04			+2'22		+3'37	-1'8		185
186	+0'89		+0'01	-3'00		0'00	+1'00		0'00	0'00		+1'95	-1'10		+2'53	+3'0	+1'2	186
187	+0'75			-2'48			+1'00			-0'01			-0'85				-3'5	187
188	+0'71		+0'01	-2'31		0'00	+1'00		0'00	0'00		+2'01	-0'76		+2'61	-0'2	-4'9	188
189	+0'43		-0'01	-1'37		0'00	+1'00		0'00	0'00		+2'10	-0'31		+2'70	-5'6	+6'1	189

\*  $\delta a=0'924$ ,  $\delta b=0'743$ ,  $u_0=0'31$ ,  $w_0=1'29$ . *Vide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.										
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.
190	54 H	Pabārgarh H.S.	1641	A 24 56 6.47 G 24 56 6.92	G 77 41 46.32	A 236 19 20.1 G 236 19 17.8	Nimdānt D O 4	+ 4.9	- 0.5	190
191	H	Paiādhari H.S.	1867	A 24 38 18.79 G 24 38 17.59	G 77 39 49.31				+ 1.2	191
192	H	Salot H.S.	1834	G 24 14 52.08	G 77 15 2.41	A 175 58 10.5 G 175 58 10.7	Hutni D O 8	- 0.4		192
193	H	Sūrāntal H.S.	1802	A 24 14 21.36 G 24 14 20.42	G 77 40 43.91				+ 0.9	193
194	H	Sironj Base-line N.E. End S.	1481	A 24 8 55.45 G 24 8 53.57	G 77 50 41.14				+ 1.9	194
195	H	Kaliānpur H.S.	1765	A 24 7 10.97 G 24 7 11.26	A 77 39 17.57	A 190 27 6.29 G 190 27 5.10	Sūrāntāl D O 0	+ 2.7	- 0.3	195
		(Origin)‡	1765	A 27 7 11.57 G 27 7 11.26	A 77 39 17.57	A 190 27 6.39 G 190 27 5.10	Surantal	+ 2.9	+ 0.3	
196	H	Tinsia H.S.	1776	A 24 6 29.05 G 24 6 27.97	G 77 18 30.70				+ 1.1	196
197	H	Losalli S.	1749	A 24 6 18.19 G 24 6 19.17	G 77 33 14.11	A 149 5 52.1 G 149 5 50.4	Rāmpur D O 2	+ 3.8	- 1.0	197
198	I	Salimpur T.S.	645	A 27 46 36.23 G 27 46 36.46	G 78 30 48.70				- 0.2	198
199	I	Agra-group N Point	550	A 27 14 10.31 G 27 14 14.10	G 78 1 4.71				- 3.8	199
200	I	Agra Long. S.	550	A 27 9 34.62 G 27 9 39.93	A 78 1 7.49 G 78 1 1.89			+ 5.0	- 5.3	200
201	I	Agra-group E. Point	550	A 27 9 16.21 G 27 9 21.00	G 78 6 3.64				- 4.8	201
202	I	Agra Parade Point	550	A 27 8 52.18 G 27 8 57.47	G 78 1 9.70				- 5.3	202
203	I	Agra-group S. Point	550	A 27 5 32.95 G 27 5 38.51	G 78 1 2.38				- 5.6	203
204	J	Gūrmī T.S.	575	A 26 36 5.97 G 26 36 3.63	G 78 30 49.82	A 155 50 8.0 G 155 50 8.8	Panāhat D O 3	- 1.8	+ 2.3	204
205	J	Mujhār H.S.	1028	A 26 6 20.39 G 26 6 17.00	G 78 28 17.73				+ 3.3	205
206	K	Algi H.S.	854	A 25 29 48.16 G 25 29 46.19	G 78 21 30.98				+ 2.0	206
207	L	Audhārī H.S.	1330	A 24 41 11.31 G 24 41 6.78	G 78 13 48.99				+ 4.5	207
208	L	Bhaorāsa H.S.	1387	A 24 8 5.13 G 24 8 3.74	G 78 0 40.73				+ 1.4	208
209	L	Budhon H.S.	1867	A 24 5 8.99 G 24 5 8.41	G 78 31 11.89	A 205 22 28.1 G 205 22 27.7	Tinsmal D O 6	+ 0.9	+ 0.6	209
210	M	Mohammadābād T.S.	565	G 27 18 24.05	G 79 25 39.80	A 291 59 0.9 G 291 58 51.5	Chandanpur D O 8	+ 18.2		210
211	P	Dargawa H.S.	1152	A 24 37 17.32 G 24 37 13.21	G 79 1 24.63				+ 4.1	211
212	P	Kangīr (old) S.	1184	A 24 0 19.28 G 24 0 20.37	G 79 25 59.25	A 106 1 11.0 G 106 1 24.2	Tinsmal E O 11	- 29.6	- 1.1	212
213	55 E	Kāmkhēra H.S.	1780	A 23 59 42.89 G 23 59 44.93	G 77 43 6.85				- 2.0	213
214	E	Ahmadpur H.S.	1713	A 23 36 18.42 G 23 36 20.88	G 77 40 48.26	A 185 10 55.0 G 185 10 53.8	Kāmkhēra D O 9	+ 2.7	- 2.5	214
215	E	Lādi H.S.	1875	A 23 8 39.10 G 23 8 44.13	G 77 42 30.87				- 5.0	215
216	F	Bhimbhat H.S.	2120	G 22 50 2.06	G 77 37 15.53	A 194 34 0.7 G 194 33 58.6	Lādi D O 16	+ 5.0		216

\* A - Astronomical Value.  
G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.  
‡ Derived from group of stations surrounding Kaliānpur.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $w_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$ .			Deflection in Prime Vertical	Deflection in Meridian	
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$			
190	+0.22		-0.02	-0.09		0.00	+1.00		0.00	0.00		+2.16	0.00		+2.77	+2.1	-0.5	190
191	+0.14			-0.43			+1.00			0.00			+0.12				+1.1	191
192			+0.24			-0.04			-0.02			+2.22			+3.05	-3.5		192
193	+0.03			-0.10			+1.00			0.00			+0.27				+0.6	193
194	+0.01			-0.02			+1.00			0.00			+0.30				+1.6	194
195	0.00	0.00	0.00	0.00	0.00	0.00	+1.00	0.00	0.00	0.00	0.00	+2.23	+0.30	0.00	+2.88	+0.1	-0.6	195
(Origin)	0.00	0.00	0.00	0.00	0.00	0.00	+1.00	0.00	0.00	0.00	0.00	+2.23	+0.30	0.00	+2.88	0.0	0.0	(Origin)
196	0.00			+0.01			+1.00			+0.01			+0.33				+0.8	196
197	-0.01		+0.06	+0.02		-0.01	+1.00		0.00	0.00		+2.24	+0.32		+2.93	+0.9	-1.3	197
198	+0.88			-2.95			+1.00			-0.01			-1.09				+0.9	198
199	+0.77			-2.53			+1.00			-0.01			-0.67				-3.1	199
200	+0.75	-0.22		-2.48	+0.03		+1.00	0.00		-0.01	+0.05		-0.85	-0.11		+5.1	-4.4	200
201	+0.75			-2.47			+1.00			-0.01			-0.85				-3.9	201
202	+0.75			-2.47			+1.00			-0.01			-0.84				-4.5	202
203	+0.74			-2.42			+1.00			-0.01			-0.82				-4.8	203
204	+0.63		-0.49	-2.04		+0.04	+1.00		+0.03	-0.01		+2.03	-0.64		+2.20	-4.2	+2.9	204
205	+0.52			-1.64			+1.00			-0.01			-0.46				+3.8	205
206	+0.37			-1.14			+1.00			-0.01			-0.22				+2.2	206
207	+0.16			-1.47			+1.00			-0.01			+0.09				+4.4	207
208	0.00			-0.01			+1.00			-0.01			+0.29				+1.1	208
209	-0.01		-0.52	+0.03		+0.08	+1.00		+0.04	-0.01		+2.24	+0.34		+2.48	-1.6	+0.3	209
210			-0.99			+0.07			+0.07			+1.98			+1.72	+16.3		210
211	+0.14			-0.42			+1.00			-0.02			+0.10				+4.0	211
212	-0.02		-1.08	+0.10		+0.16	+1.00		+0.08	-0.03		+2.24	+0.32		+2.04	-31.6	-1.4	212
213	-0.04			+0.11			+1.00			0.00			+0.36				-2.4	213
214	-0.15		-0.01	+0.44		0.00	+1.00		0.00	0.00		+2.28	+0.50		+2.93	-0.2	-3.0	214
215	-0.29			+0.84			+1.00			0.00			+0.67				-5.7	215
216			+0.02			-0.01			0.00			+2.35			+3.06	+2.1		216

\*  $\delta a = 0.024$ ,  $\delta b = 0.743$ ,  $v_0 = 0.31$ ,  $w_0 = 1.29$ . *Vide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.										
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflection†	Serial No.
217	55 G	Nilgarh H.S.	2533	G 21 45 50.12	G 77 39 18.82	A 321 4 43.7 G 321 4 44.2	Sālbaldi E O 1	- 1.3	"	217
218	G	Takulkhera s.	1094	A 21 5 50.17 G 21 5 56.76	G 77 38 24.94				- 6.6	218
219	H	Rāngrui s.	1046	A 20 48 7.16 G 20 48 14.68	G 77 35 53.83				- 7.5	219
220	H	Badgaon H.S.	1128	A 20 44 15.54 G 20 44 23.06	G 77 36 31.79	A 183 9 0.9 G 183 8 59.5	Ashti D O 8	+ 3.7	- 7.5	220
221	H	Dhānura s.	1135	A 20 44 3.35 G 20 44 10.84	G 77 41 43.27				- 7.5	221
222	H	Dotra s.	1140	A 20 41 22.25 G 20 41 28.91	G 77 32 45.66				- 6.7	222
223	H	Sakri H.S.	1810	G 20 0 14.11	G 77 42 7.32	A 175 24 37.7 G 175 24 35.4	Kopdi D O 20	+ 6.3		223
224	I	Saugor H.S.	2033	A 23 49 48.71 G 23 49 48.07	G 78 46 18.16				+ 0.6	224
225	I	Nāharmau H.S.	1940	A 23 30 13.14 G 23 30 18.15	G 78 49 49.13				- 5.0	225
226	M	Karaundi H.S.	1025	A 23 10 45.07 G 23 10 40.02	G 79 50 43.34	A 206 22 35.6 G 206 22 38.4	Lora D O 1	- 6.5	+ 5.1	226
227	M	Jabalpur Long. s.	...	G 23 10 10.10	A 79 56 52.42 G 79 57 2.61				- 9.4	227
228	P	Bhimsain H.S.	1490	A 20 57 28.54 G 20 57 35.96	G 79 46 7.40	A 297 55 2.8 G 297 55 2.3	Partābgurh E O 0	+ 1.3	- 7.4	228
229	P	Rājūli H.S.	1070	A 20 12 51.25 G 20 12 55.45	G 79 44 49.27				- 4.2	229
230	56B	Nitali H.S.	2289	A 18 17 2.74 G 18 17 7.16	G 76 16 23.32	A 239 23 1.3 G 239 23 5.7	Harangal D O 10	- 13.3	- 4.4	230
231	B	Achola H.S.	2274	A 18 14 44.87 G 18 14 48.12	G 76 59 20.51	A 272 47 57.4 G 272 47 58.5	Mangunāl D O 11	- 3.3	- 3.3	231
232	E	Halda s.	1335	A 19 9 24.41 G 19 9 29.38	G 77 41 1.39				- 5.0	232
233	E	Voi s.	1439	A 19 7 14.69 G 19 7 19.89	G 77 34 46.88				- 5.2	233
234	E	Somtana H.S.	1714	G 19 5 0.52	G 77 39 16.29	A 186 51 46.9 G 186 51 48.3	Terbāu D O 5	- 4.0		234
235	E	Mandāla s.	1294	A 19 2 42.84 G 19 2 48.24	G 77 43 35.14				- 5.4	235
236	E	Talegaon s.	1233	A 19 1 21.65 G 19 1 26.64	G 77 37 16.75				- 5.0	236
237	F	Dāmargida Obsy. S.	1941	A 18 3 14.92 G 18 3 17.35	G 77 40 4.41	A 188 11 59.1 G 188 11 50.8	Burgāpāli D O 15	- 2.1	- 2.4	237
238	G	Devanūr s.	1593	A 17 10 56.88 G 17 11 0.43	G 77 41 7.41				- 3.6	238
239	G	Akampalle h.s.	1557	A 17 10 50.39 G 17 10 53.96	G 77 34 29.30				- 3.6	239
240	G	Kodangal S.	1906	A 17 7 53.74 G 17 7 57.35	G 77 38 25.73	A 62 29 16.3 G 62 29 17.8	Nēlugat E O 1	- 4.9	- 3.6	240
241	G	Linganupalle h.s.	1815	A 17 7 13.40 G 17 7 16.66	G 77 42 28.61				- 3.3	241
242	G	Pulmudi s.	1869	A 17 4 1.06 G 17 4 6.05	G 77 36 22.06				- 5.0	242
243	H	Tōnalgutta s.	1133	A 16 18 2.36 G 16 18 6.91	G 77 34 49.44				- 4.6	243
244	H	Pēddapād s.	1090	A 16 17 14.13 G 16 17 20.38	G 77 44 30.54				- 6.3	244

\* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$ .						
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
217	"	"	0'00	"	"	0'00	"	"	0'00	"	"	2'46	"	"	3'18	- 4'2	"	217	
218	-0'95			+2'66			+1'00			0'00			+1'41				- 8'0	218	
219	-1'05			+2'93			+1'00			0'00			+1'52				- 9'0	219	
220	-1'08		+0'03	+2'99		-0'01	+1'00		0'00	0'00		+2'58	+1'53		+3'36	+ 0'6	- 9'0	220	
221	-1'08			+3'00			+1'00			0'00			+1'56				- 9'1	221	
222	-1'09			+3'04			+1'00			0'00			+1'56				- 8'3	222	
223			-0'01			0'00			0'00			+2'67		+3'43		+ 3'4		223	
224	-0'08			+0'25			+1'00			-0'03			+0'40				+ 0'2	224	
225	-0'18			+0'53			+1'00			-0'02			+0'52				- 5'5	225	
226	-0'26		-1'45	+0'81		+0'25	+1'00		+0'07	-0'04		+2'32	+0'62		+1'85	- 8'3	+ 4'5	226	
227		-1'38			+0'20			+0'02		-0'02				-1'15			- 8'3	227	
228	-0'99		-1'41	+2'79		+0'32	+1'00		+0'10	-0'03		+2'55	+1'43		+2'26	- 0'7	- 8'8	228	
229	-1'25			+3'48			+1'00			-0'03			+1'69				- 5'9	229	
230	-1'99		+1'03	+5'30		-0'31	+1'00		-0'08	+0'02		+2'91	+2'44		+4'45	- 16'9	- 6'8	230	
231	-2'00		+0'52	+5'34		-0'16	+1'00		-0'04	+0'01		+2'92	+2'44		+4'11	- 6'5	- 5'7	231	
232	-1'65			+4'47			+1'00			0'00			+2'11				- 7'1	232	
233	-1'67			+4'51			+1'00			0'00			+2'12				- 7'3	233	
234			+0'01			0'00			0'00			+2'79		+3'61		- 7'0		234	
235	-1'70			+4'57			+1'00			0'00			+2'14				- 7'5	235	
236	-1'71			+4'60			+1'00			0'00			+2'15				- 7'2	236	
237	-2'09		0'00	+5'53		+0'01	+1'00		0'00	0'00		+2'95	+2'49		+3'50	- 5'0	- 4'9	237	
238	-2'45			+6'38			+1'00			0'00			+2'79				- 6'4	238	
239	-2'45			+6'30			+1'00			0'00			+2'73				- 6'3	239	
240	-2'47		+0'01	+6'43		0'00	+1'00		0'00	0'00		+3'10	+2'81		+4'01	- 7'6	- 6'4	240	
241	-2'47			+6'44			+1'00			0'00			+2'81				- 6'1	241	
242	-2'50			+6'50			+1'00			0'00			+2'83				- 7'8	242	
243	-2'89			+7'42			+1'00			0'00			+3'15				- 7'8	243	
244	-2'83			+7'27			+1'00			0'00			+3'10				- 9'4	244	

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *Vide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.											
Serial No	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
245	56 B	Darur H.S.	1796	G 16 13 35.40	G 77 39 36.51	A 132 35 57.2 G 132 35 59.2	Köttappalle D 0 15	- 6.9	"	245	
246	H	Tuugat h.s.	1450	A 16 9 46.73 G 16 9 51.66	G 77 34 11.59				- 4.9	246	
247	II	Gattinārayantippa h.s.	1225	A 16 7 48.95 G 16 7 54.81	G 77 45 51.00				- 5.9	247	
248	II	Devaragat h.s.	1332	A 16 6 31.98 G 16 6 37.27	G 77 41 26.21				- 5.3	248	
249	K	Pirmulo H.S.	2093	A 17 52 58.32 G 17 53 2.81	G 78 35 50.98	A 105 0 48.0 G 105 0 49.0	Narsula D 0 12	- 3.1	- 4.5	249	
250	K	Bolarum P.W.D. Office Long. S.	1971	A 17 30 7.36 G 17 30 13.41	A 78 31 7.84 G 78 31 11.12	A 25 57 35.8 G 25 57 35.8	Hyderābād Naubat-pahar D 0 13	0.0	- 6.1	250	
251	M	Diwai H.S.	967	A 19 49 26.87 G 19 49 32.57	G 79 32 28.62	A 154 17 54.2 G 154 17 55.1	Ambāgarh D 0 8	- 2.5	- 5.7	251	
252	M	Ankora H.S.	1463	A 19 24 26.63 G 19 24 34.75	G 79 36 27.70				- 8.1	252	
253	N	Burgpalli H.S.	983	A 18 54 3.48 G 18 54 7.20	G 79 41 36.96	A 142 8 7.5 G 142 8 8.8	Rechni D 0 8	- 3.8	- 3.7	253	
254	N	Kāmgir H.S.	1772	A 18 35 26.90 G 18 35 26.12	G 79 31 42.36				+ 0.8	254	
255	O	Bolikonda H.S.	1363	A 17 42 29.08 G 17 42 35.82	G 79 47 57.98				- 6.7	255	
256	O	Vānākonda H.S.	1694	A 17 36 0.22 G 17 36 6.87	G 79 22 20.70	A 180 4 14.5 G 180 4 15.3	Yarābali D 0 1	- 2.5	- 6.7	256	
257	O	Niālamari H.S.	1144	A 17 1 25.93 G 17 1 33.63	G 79 43 29.78				- 7.7	257	
258	57 A	Bellary Long. s.	...	G 15 8 33.06	A 76 55 38.89 G 76 55 39.58			- 0.7		258	
259	B	Yērragunta h.s.	1698	A 14 48 27.31 G 14 48 23.26	G 76 58 17.56				+ 4.1	259	
260	C	Nughallibēta H.S.	3140	G 13 1 32.95	G 76 28 32.46	A 54 31 39.1 G 54 31 41.7	Sātanhalli E 0 3	- 11.2		260	
261	E	Namtnabad s.	1169	A 15 5 51.75 G 15 5 52.40	G 77 36 26.91				- 0.7	261	
262	F	Chikalgurki s.	1516	A 14 59 5.16 G 14 59 4.53	G 77 11 6.39				+ 0.6	262	
263	F	Bandūr s.	1447	A 14 57 44.41 G 14 57 42.32	G 77 0 36.91				+ 2.1	263	
264	F	Hōnnūr H.S.	1579	A 14 55 22.20 G 14 55 18.96	G 77 6 2.60				+ 3.2	264	
265	F	Nimbāgal s.	1565	A 14 51 56.14 G 14 51 52.43	G 77 11 51.78				+ 3.7	265	
266	F	Pavagada H.S.	3022	A 14 6 18.80 G 14 6 15.39	G 77 16 42.43				+ 3.4	266	
267	G	Bōmmasandra s.	2005	A 13 59 42.63 G 13 59 36.34	G 77 28 43.96				+ 6.3	267	
268	G	Bangalore Base-line N.E. End S.	3016	A 13 4 53.17 G 13 4 56.05	G 77 39 16.23	A 44 32 19.7 G 44 32 19.3	Bangalore Base-line S.W. End E 0 8	+ 1.7	- 2.9	268	
269	G	Bangalore Base-line S.W. End S.	3126	A 13 0 36.12 G 13 0 40.91	A 77 34 57.29 G 77 35 0.19	A 224 31 21.7 G 224 31 21.6	Bangalore Base-line N.E. End D 0 13	+ 0.4	- 4.8	269	
270	II	Dōddagunta s.	3003	A 12 59 51.52 G 12 59 55.76	G 77 37 33.25				- 4.1	270	
271	M	Dānapa H.S.	130	A 15 55 59.69 G 15 56 0.14	G 79 56 7.39	A 265 47 36.0 G 265 47 39.4	Babbāpalle D 0 22	- 11.9	- 0.5	271	
272	M	Deventippa s.	195	A 15 0 33.52 G 15 0 36.47	G 79 55 12.92				- 3.0	272	

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.



XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $\nu_n = 1''$			Case IV: Azimuth $\omega_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical	Deflection in Meridian	
245	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	245
246	-2.88			+7.39			+1.00			0.00			+3.14				-8.0	246
247	-2.89			+7.43			+1.00			0.00			+3.16				-9.1	247
248	-2.90			+7.45			+1.00			0.00			+3.18				-8.5	248
249	-2.16		-0.71	+5.70		+0.22	+1.00		+0.05	-0.03		+2.97	+2.51		+3.37	-5.6	-7.0	249
250	-2.31	-0.52	-0.66	+6.06	+0.07	+0.22	+1.00	+0.01	+0.05	-0.01	-0.11	+3.04	+2.67	-0.58	+3.49	-2.5	-8.8	250
251	-1.40		-1.30	+3.83		+0.34	+1.00		+0.10	-0.03		+2.69	+1.83		+2.55	-4.5	-7.5	251
252	-1.55			+4.23			+1.00			-0.03			+1.98				-10.1	252
253	-1.74		-1.45	+4.71		+0.42	+1.00		+0.11	-0.03		+2.82	+2.16		+2.63	-5.8	-5.9	253
254	-1.87			+5.05			+1.00			-0.03			+2.29				-1.5	254
255	-2.22			+5.86			+1.00			-0.04			+2.56				-9.3	255
256	-2.27		-1.31	+5.97		+0.43	+1.00		+0.10	-0.03		+3.02	+2.61		+3.04	-4.6	-9.3	256
257	-2.50			+6.53			+1.00			-0.03			+2.81				-10.5	257
258		+0.43		-0.06				0.00			-0.16		+0.16				-0.9	258
259	-3.48			+8.77			+1.00			+0.01			+3.62				+0.5	259
260			+1.17			-0.56				-0.09		+4.05			+5.85	-14.5		260
261	-3.35			+8.47			+1.00			0.00			+3.51				-4.2	261
262	-3.40			+8.59			+1.00			+0.01			+3.56				-3.0	262
263	-3.41			+8.61			+1.00			+0.01			+3.57				-1.5	263
264	-3.43			+8.65			+1.00			+0.01			+3.58				-0.4	264
265	-3.45			+8.71			+1.00			+0.01			+3.60				+0.1	265
266	-3.80			+9.49			+1.00			+0.01			+3.86				-0.5	266
267	-3.85			+9.60			+1.00			0.00			+3.88				+2.4	267
268	-4.28		0.00	+10.55		0.00	+1.00		0.00	0.00		+4.03	+4.19		+5.20	-0.9	-7.1	268
269	-4.32	+0.04	+0.07	+10.63	-0.01	-0.04	+1.00	0.00	-0.01	0.00	-0.19	+4.06	+4.22	-0.21	+5.26	-2.3	-9.0	269
270	-4.32			+10.64			+1.00			0.00			+4.22				-8.4	270
271	-2.97		-1.90	+7.62		+0.72	+1.00		+0.15	-0.04		+3.32	+3.18		+3.12	-13.6	-3.7	271
272	-3.38			+8.55			+1.00			-0.04			+3.50				-6.5	272

$\delta a = 0.124$   $\delta b = 0.743$   $\omega_0 = 0.31$   $\omega_0 = 1.29$ . *Fide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.											
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
273	57 N	Kistama H.S.	458	A 14 27 12.28 G 14 27 14.56	° ' "	G 79 45 18.51	A 80 1 54.1 G 80 1 55.6	Pallakōndu E 0 16	- 5.8	- 2.3	273
274	P	Anandalamalai H.S.	923	G 12 55 50.73	° ' "	G 79 23 46.76	A 171 57 36.3 G 171 57 37.6	Pullur E 0 23	- 5.7		274
275	58 B	Yēttimalai S.	617	A 11 3 52.10 G 11 3 50.00	° ' "	G 77 50 47.10				+ 2.1	275
276	F	Pachapālaiyam s.	970	A 10 59 40.81 G 10 59 39.88	° ' "	G 77 37 25.80	A 167 34 2.4 G 167 34 1.2	Chennimalai E 0 40	+ 6.2	+ 0.9	276
277	F	Kātpālaiyam s.	878	A 10 56 36.66 G 10 56 35.97	° ' "	G 77 40 50.63				+ 0.7	277
278	G	Shūlukarai s.	333	A 9 32 15.53 G 9 32 13.28	° ' "	G 77 56 51.22				+ 2.3	278
279	H	Rādihāpuram S.	167	A 8 17 1.75 G 8 16 59.44	° ' "	G 77 42 7.71	A 5 55 25.4 G 5 55 24.1	Kudankulam Obsy. D 0 4	+ 8.9	+ 2.3	279
280	H	Tanakurakulam S.	176	A 8 13 57.50 G 8 13 55.39	° ' "	G 77 38 53.81				+ 2.1	280
281	H	Arasākulam S.	55	A 8 13 41.96 G 8 13 39.52	° ' "	G 77 44 30.98				+ 2.4	281
282	H	Vijayāpati S.	90	A 8 12 10.67 G 8 12 8.34	° ' "	G 77 46 35.58				+ 2.3	282
283	H	Nagarkoil Long. S.	110	G 8 11 25.30	° ' "	A 77 26 1.82 G 77 26 3.56			- 1.7		283
284	H	Kudankulam Obsy. S.	175	A 8 10 23.41 G 8 10 21.55	° ' "	G 77 41 26.26	A 185 55 18.8 G 185 55 18.6	Rādihāpuram D 0 5	+ 1.4	+ 1.9	284
285	H	Punnae Obsy. S.	48	A 8 9 29.92 G 8 9 27.79	° ' "	G 77 37 35.33				+ 2.1	285
286	I	Kanjamalai H.S.	3236	G 11 36 55.92	° ' "	G 78 3 36.52	A 38 11 59.1 G 38 12 0.1	Morur D 1 3	- 4.9		286
287	K	Manēgandi S.	56	G 9 46 15.13	° ' "	G 78 55 20.84	A 178 0 47.2 G 178 0 50.3	Manikamkota D 0 1	- 18.0		287
288	K	Black s.	346	A 9 31 4.22 G 9 31 1.30	° ' "	G 78 2 58.77				+ 2.9	288
289	K	Kutipārai S.	347	A 9 28 47.09 G 9 28 44.87	° ' "	G 78 0 37.76	A 25 17 6.2 G 25 17 6.8	Koilpati D 0 4	+ 3.6	+ 2.2	289
290	K	Pandulngudi s.	217	A 9 23 30.55 G 9 23 27.69	° ' "	G 78 5 54.11				+ 2.9	290
291	K	Rāmnad S.	48	G 9 21 51.96	° ' "	G 78 49 17.66	A 57 57 54.9 G 57 57 56.2	Uttarakoshamangai E 0 0	- 7.9		291
292	M	Kāllapat Trestle S.	199	G 11 57 12.30	° ' "	G 79 33 52.96	A 214 44 19.0 G 214 44 20.0	Pērunukkal D 0 1	- 4.7		292
293	M	Tiruvēndipuram s.	...	A 11 44 43.40 G 11 44 37.64	° ' "	G 79 42 45.80				+ 5.8	293
294	M	Naynīpuriyān Trestle S.	158	G 11 7 49.06	° ' "	G 79 20 51.19	A 152 57 0.1 G 152 56 57.4	Kachiperumāl E 0 12	+ 13.7		294
295	N	Pātharakōta S.	120	G 10 28 2.31	° ' "	G 79 12 43.59	A 179 40 40.6 G 179 40 42.9	Kakkarakōta D 0 4	- 12.4		295
296	62 D	Rāmnapur (old) T.S.	541	A 28 22 0.10 G 28 22 11.04	° ' "	G 80 28 38.33	A 302 56 33.7 G 302 56 30.9	Rāmnagar D 0 5	+ 5.2	- 10.9	296
297	63 A	Jarāra T.S.	536	A 27 59 50.22 G 27 59 55.94	° ' "	G 80 28 10.95				- 5.7	297
298	A	Nimkār T.S.	486	A 27 21 8.16 G 27 21 8.09	° ' "	G 80 29 3.67	A 178 58 28.0 G 178 58 20.7	Darawal D 0 6	+ 14.1	+ 0.1	298
299	B	Etorā T.S.	429	A 26 54 22.63 G 26 54 17.85	° ' "	G 80 39 38.26				+ 4.8	299
300	B	Dewarān T.S.	439	A 26 15 58.32 G 26 15 52.89	° ' "	G 80 18 14.46				+ 5.4	300

\* A - Astronomical Value.

G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical	Deflection in Meridian	
273	-3'63		-1'90	+9'13		+0'82	+1'00		+0'15	-0'03		+3'65	+3'69		+3'61	-7'5	-6'0	273
274			-1'74			+0'84			+0'14			+4'08			+4'31	-7'8		274
275	-5'27			+12'69			+1'00			0'00			+4'87				-2'8	275
276	-5'31		+0'04	+12'76		-0'02	+1'00		-0'01	0'00		+4'79	+4'89		+6'19	+3'6	-4'0	276
277	-5'33			+12'82			+1'00			0'00			+4'90				-4'2	277
278	-6'05			+14'34			+1'00			-0'01			+5'36				-3'1	278
279	-6'71		-0'08	+15'70		+0'06	+1'00		+0'01	0'00		+6'34	+5'78		+8'15	+6'2	-3'5	279
280	-6'74			+15'76			+1'00			0'00			+5'80				-3'7	280
281	-6'74			+15'76			+1'00			0'00			+5'79				-3'4	281
282	-6'75			+15'79			+1'00			0'00			+5'80				-3'5	282
283		+0'13			-0'01			0'00			-0'27			-0'24		-1'5		283
284	-6'77		-0'06	+15'82		+0'04	+1'00			0'00		+6'43	+5'81		+8'27	-1'3	-3'9	284
285	-6'78			+15'84			+1'00			0'00			+5'82				-3'7	285
286			-0'45			+0'23			+0'03			+4'53			+5'62	-7'1		286
287			-1'64			+0'98			+0'13			+5'38			+6'20	-20'1		287
288	-6'06			+14'36			+1'00			-0'01			+5'37				-2'5	288
289	-6'08		-0'47	+14'40		+0'29	+1'00		+0'04	-0'01		+5'54	+5'38		+6'94	-0'9	-3'2	289
290	-6'13			+14'49			+1'00			-0'01			+5'41				-2'5	290
291			-1'57			+0'96			+0'13			+5'61			+6'55	-10'3		291
292			-2'05			+1'06			+0'16			+4'41			+4'64	-6'0		292
293	-4'93			+11'96			+1'00			-0'03			+4'61				+1'2	293
294			-1'94			+1'06			+0'15			+4'73			+5'15	+11'5		294
295			-1'88			+1'08			+0'15			+5'02			+5'59	-14'8		295
296	+1'01		-1'53	-3'40		+0'08	+1'00		+0'10	-0'05		+1'92	-1'34		+1'15	+3'6	-0'6	296
297	+0'94			-3'12			+1'00			-0'05			-1'20				-4'5	297
298	+0'81		-1'57	-2'63		+0'12	+1'00		+0'11	-0'05		+1'90	-0'95		+1'23	+12'5	+1'1	298
299	+0'72			-2'28			+1'00			-0'05			-0'78				+5'0	299
300	+0'57			-1'77			+1'00			-0'04			-0'54				+5'9	300

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *Fide* p. 2.

TABLE  
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	Height in feet	EVEREST'S SPHEROID.						Meridian Deflection†	Serial No.
				Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†			
301	63 C	Kānākhera T.S.	416	A 25 51 25.97 G 25 51 20.95	G 80 25 31.61					+ 5.0	301
302	C	Pavia H.S.	481	A 25 27 21.18 G 25 27 17.39	G 80 44 12.26					+ 3.8	302
303	D	Potenda S.	993	A 24 37 24.71 G 24 37 23.04	G 80 57 7.18					+ 1.7	303
304	E	Dadaura T.S.	420	A 27 43 3.51 G 27 43 18.33	G 81 42 44.29					- 14.8	304
305	E	Māsi T.S.	406	A 27 38 14.79 G 27 38 25.17	G 81 23 8.97	A 153 5 50.5 G 153 5 52.2	Bela D O 5	- 3.2		- 10.4	305
306	E	Imlia T.S.	428	A 27 19 17.83 G 27 19 18.00	G 81 7 37.37					- 1.1	306
307	F	Utiāmau T.S.	386	A 27 0 1.02 G 26 59 57.08	G 81 12 17.24					+ 4.5	307
308	F	Parewa T.S.	380	A 26 38 11.44 G 26 38 4.00	G 81 12 11.14					+ 7.4	308
309	F	Sora T.S.	400	A 26 17 26.39 G 26 17 18.83	G 81 12 23.12	A 239 43 3.6 G 239 42 55.9	Jansi D O 4	+ 15.6		+ 7.6	309
310	G	Pariāou T.S.	346	A 25 50 11.59 G 25 50 5.26	G 81 22 16.31					+ 6.3	310
311	G	Pabhosa H.S.	565	G 25 21 17.32	G 81 19 8.40	A 187 38 4.1 G 187 38 4.7	Karra D O 14	- 1.3			311
312	H	Karāra H.S.	1966	A 24 4 42.20 G 24 4 42.01	G 81 15 47.29	A 269 18 28.7 G 269 18 34.9	Marwās D O 16	- 13.9		+ 0.2	312
313	I	Manchank T.S.	360	A 27 36 28.91 G 27 36 48.14	G 82 5 3.16					- 19.2	313
314	I	Pathārdi T.S.	320	A 27 25 56.11 G 27 26 14.77	G 82 45 2.97					- 18.7	314
315	I	Bāsadela T.S.	377	A 27 23 50.71 G 27 24 3.24	G 82 16 50.44	A 106 15 8.7 G 106 15 7.9	Saibara D O 8	+ 1.5		- 12.5	315
316	J	Orejhār S.	392	G 26 46 55.54	G 82 12 7.60	A 308 36 18.9 G 308 36 17.7	Bisaul D O 7	+ 2.4			316
317	J	Fyzabad Long. S.	...	G 26 46 40.66	A 82 8 7.60 G 82 8 8.15					- 0.5	317
318	J	Bisaul T.S.	342	G 26 40 37.38	G 82 20 54.43	A 128 40 15.9 G 128 40 14.8	Orejhār D O 1	+ 2.2			318
319	K	Mārār T.S.	370	G 25 41 17.20	G 82 14 19.00	A 42 20 13.2 G 42 20 13.1	Buriu D O 7	+ 0.2			319
320	L	Gurwāni H.S.	2083	A 24 1 28.93 G 24 1 25.71	G 82 17 28.34	A 210 29 53.8 G 210 29 49.4	Pokra D O 5	+ 9.9		+ 3.3	320
321	M	Ghāna T.S.	296	A 27 20 48.34 G 27 21 5.08	G 83 5 38.81					- 16.7	321
322	N	Rajabāri T.S.	296	G 26 54 3.04	G 83 15 35.49	A 104 47 9.8 G 104 47 9.9	Nandaaur D O 5	- 0.2			322
323	N	Samenda T.S.	285	G 26 0 23.97	G 83 13 30.67	A 304 8 50.2 G 304 8 48.7	Chit Bisram D O 6	+ 3.1			323
324	O	Hirdepur T.S.	289	G 25 24 23.05	G 83 14 15.46	A 304 4 33.1 G 304 4 33.8	Barhāni D O 6	- 1.5			324
325	O	Gora H.S.	1828	G 24 4 55.71	G 83 14 13.47	A 282 48 33.9 G 282 48 27.5	Sewādhi D O 10	- 8.1			325
326	34 A	Amha H.S.	2113	A 23 59 57.02 G 23 59 56.24	G 80 29 17.26	A 200 4 21.4 G 200 4 19.0	Lakanpura D O 19	+ 5.4		+ 0.8	326
327	A	Lora H.S.	1923	A 23 29 46.30 G 23 29 41.53	G 80 9 56.85					+ 4.8	327
328	B	Sarand Pat H.S.	1627	G 23 13 18.98	G 80 3 5.98	A 159 45 20.8 G 159 45 20.5	Tālla E O 6	+ 0.7			328

\* A = Astronomical Value  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

· · XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*				Serial No.	
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$ .					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical		Deflection in Meridian
301	+0.47	"	"	-1.44	"	"	+1.00	"	"	-0.05	"	"	-0.38	"	"	"	+5.4	301
302	+0.38			-1.11			+1.00			-0.05			-0.24				+4.0	302
303	+0.16			-0.43			+1.00			-0.05			+0.08				+1.6	303
304	+0.90			-2.91			+1.00			-0.07			-1.11				-13.7	304
305	+0.88	-2.06		-2.85	+0.14		+1.00	+0.14		-0.06	+1.96		-1.07	+1.78		-5.4	-9.3	305
306	+0.81			-2.60			+1.00			-0.06			-0.95				-0.1	306
307	+0.75			-2.35			+1.00			-0.06			-0.82				+5.3	307
308	+0.66			-2.06			+1.00			-0.06			-0.68				+8.1	308
309	+0.59	-2.02		-1.79	+0.19		+1.00	+0.14		-0.06	+2.06		-0.55	+0.98		+14.4	+8.2	309
310	+0.48			-1.43			+1.00			-0.06			-0.38				+6.7	310
311		-2.13			+0.25			+0.15			+2.13			+1.01		-2.5		311
312	+0.02	-2.18	+0.03	+0.32	+1.00	+0.15	-0.06	+2.23	+0.27	+1.16	-15.1	-0.1	312					
313	+0.89			-2.83			+1.00			-0.07			-1.06				-18.1	313
314	+0.87			-2.69			+1.00			-0.08			-1.00				-17.7	314
315	+0.89	-2.57		-2.67	+0.19	+1.00	+0.18	-0.07	+1.98	-0.95	+0.37	+0.9	-11.5	315				
316		-2.58			+0.22	+0.18	+2.02	+0.44	+1.7	316								
317		-2.69		+0.38	+0.04	+0.05	-2.13	+1.6	317									
318		-2.65		+0.23	+0.18	+2.03	+0.40	+1.6	318									
319		-2.65		+0.29	+0.10	+2.10	+0.53	-0.5	319									
320	+0.02	-2.80	+0.07	+0.41	+1.00	+0.20	-0.08	+2.24	+0.29	+0.66	+9.2	+2.9	320					
321	+0.85			-2.63	+1.00	-0.09	-0.97	-15.7	321									
322		-3.14		+0.26	+0.22	+2.01	-0.06	-0.3	322									
323		-3.19		+0.32	+0.22	+2.07	+0.03	+3.1	323									
324		-3.24		+0.37	+0.23	+2.12	+0.08	-1.6	324									
325		-3.36		+0.49	+0.24	+2.23	+0.21	-8.3	325									
326	-0.01	-1.71	+0.10	+0.25	+1.00	+0.12	-0.05	+2.24	+0.31	+1.54	+3.9	+0.5	326					
327	-0.17		+0.53	+1.00	-0.04	+0.50	+4.3	327										
328		-1.53	+0.30	+0.11	+2.41	+1.96	-1.1	328										

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *I* vide p. 2.

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.											
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	Serial No.	
329	64 B	Sarey Khan Lat. S.	1409	A 22 12 50.66 G 22 12 55.61	0 0 0	G 80 2 49.79			- 5.0	329	
330	C	Lingmāra H.S.	1400	A 21 42 55.36 G 21 43 3.07	0 0 0	G 80 7 36.30			- 7.7	330	
331	C	Sitāpār H.S.	1237	A 21 24 43.83 G 21 24 50.54	0 0 0	G 80 19 26.36			- 6.7	331	
332	J	Dalea H.S.	1622	A 22 19 30.25 G 22 19 33.62	0 0 0	G 82 1 31.25			- 3.4	332	
333	K	Pathāidi T.S.	879	A 21 48 43.06 G 21 48 45.96	0 0 0	G 82 16 46.96	A 198 23 42.8 G 198 23 43.3	Konārgurh D 0 5	- 1.2	- 2.9	333
334	L	Kamai H.S.	1313	A 20 56 50.31 G 20 56 51.47	0 0 0	G 82 8 18.55	A 223 15 23.2 G 223 15 23.8	Khalāri E 0 9	- 1.6	- 1.2	334
335	P	Sindur H.S.	2918		0 0 0	G 83 39 42.83	A 201 20 4.4 G 201 20 10.3	Lakh Parbat D 0 35	- 16.0		335
336	65 C	Singāwāram H.S.	714	A 17 45 8.71 G 17 45 10.38	0 0 0	G 80 56 9.04	A 249 3 4.9 G 249 3 6.5	Nārākonda E 0 29	- 5.0	- 1.7	336
337	D	Dhūlipalla S.	245	A 16 25 53.47 G 16 25 56.75	0 0 0	G 80 5 29.59	A 125 53 37.6 G 125 53 39.9	Kachulboru E 0 47	- 7.8	- 3.3	337
338	G	Parampudi H.S.	684	A 17 12 32.63 G 17 12 38.28	0 0 0	G 81 12 10.06	A 114 12 9.2 G 114 12 13.2	Nāgaldurgam E 0 15	- 12.9	- 5.7	338
339	I	Hātvena H.S.	2600	A 19 51 42.60 G 19 51 43.34	0 0 0	G 82 1 25.96			+ 0.3		339
340	I	Kurāi H.S.	2014	A 19 12 2.67 G 19 12 5.98	0 0 0	G 82 7 7.97	A 201 43 17.4 G 201 43 17.9	Motūson E 0 4	- 1.4	- 3.3	340
341	K	Kālingkonda H.S.	4634		0 0 0	G 82 18 40.67	A 189 41 25.0 G 189 41 24.8	Kaurālbiding D 0 25	+ 0.6		341
342	K	Sānjib H.S.	2142	A 17 31 12.32 G 17 31 18.68	0 0 0	G 82 41 24.30	A 135 38 16.0 G 135 38 15.9	Dhār E 0 55	+ 0.3	- 6.4	342
343	N	Rāwal H.S.	874	A 18 32 4.73 G 18 32 9.22	0 0 0	G 83 33 11.63	A 317 29 5.0 G 317 29 5.0	Pindi D 0 11	0.0	- 4.5	343
344	N	Vizagapatām Base-line N. End S.	181	A 18 0 56.66 G 18 1 2.93	0 0 0	G 83 13 43.36	A 203 44 24.5 G 203 44 24.5	Bor E 0 12	0.0	- 6.3	344
345	O	Waltair Long. S.	200	A 17 43 20.44 G 17 43 29.31	0 0 0	A 83 19 0.17 G 83 19 3.52			- 3.2	- 8.9	345
346	66 A	Ongole H.S.	250	A 15 29 52.87 G 15 29 56.82	0 0 0	G 80 2 27.77			- 4.0		346
347	B	Gudan H.S.	292	A 14 1 10.65 G 14 1 9.45	0 0 0	G 80 1 13.36			+ 1.2		347
348	C	Madras Observatory Long. S.	54	A 14 4 8.97 G 13 4 4.17	0 0 0	A 80 14 47.06 G 80 14 54.33			- 7.1	+ 4.8	348
349	C	St. Thomas's Mount Trestle S.	250	A 13 0 20.64 G 13 0 14.79	0 0 0	G 80 11 41.38	A 12 30 5.3 G 12 30 6.2	Nannangalam D 0 7	- 3.9	+ 5.9	349
350	D	Injamākam H.S.	29		0 0 0	G 80 15 11.23	A 99 4 39.1 G 99 4 40.6	Nannangalam E 0 23	- 6.5		350
351	72 B	Nannangurthi T.S.	344		0 0 0	G 84 23 46.86	A 107 52 43.1 G 107 52 47.8	rukwa D 0 7	- 9.2		351
352	B	Jatāpur T.S.	232	A 26 3 45.56 G 26 3 39.42	0 0 0	G 84 23 9.46	A 111 52 41.5 G 111 52 40.0	Kurwārpur D 0 3	+ 3.1	+ 6.1	352
353	C	Nuāon T.S.	251	A 25 34 45.64 G 25 34 37.94	0 0 0	G 84 14 15.86			+ 7.7		353
354	C	Beumpar T.S.	315	A 25 5 22.35 G 25 5 14.02	0 0 0	G 84 22 6.95	A 215 46 30.0 G 215 46 33.5	Beumpar D 0 6	- 7.5	+ 8.3	354
355	D	Teona H.S.	740	A 24 34 49.76 G 24 34 38.94	0 0 0	G 84 10 26.42			+ 10.8		355
356	D	Hurālong H.S.	1378	A 14 2 16.74 G 14 2 5.99	0 0 0	G 84 21 50.58	A 128 18 18.3 G 128 18 24.0	Khairā Pāndu D 0 4	- 12.8	+ 10.8	356

\* A - Astronomical Value.

G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $w_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6376200$ metres, $1/\epsilon = 298.3$ .					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
329	-0.56	"	"	+1.65	"	"	+1.00	"	"	-0.04	"	"	+0.97	"	"	"	- 6.0	329
330	-0.72			+2.10			+1.00			-0.04			+1.15				- 8.9	330
331	-0.83			+2.37			+1.00			-0.04			+1.25				- 8.0	331
332	-0.50			+1.55			+1.00			-0.07			+0.91				- 4.3	332
333	-0.60	-2.98		+2.01		+0.60	+1.00		+0.22	-0.07		+2.45	+1.10		+0.92	- 1.9	- 4.0	333
334	-0.95	-2.98		+2.79		+0.67	+1.00		+0.22	-0.07		+2.55	+1.41		+1.10	- 2.4	- 2.6	334
335		-4.09				+1.00			+0.30			+2.62			+0.44	- 15.9		335
336	-2.19	-2.48		+5.81		+0.79	+1.00		+0.19	-0.05		+2.99	+2.54		+2.21	- 6.3	- 4.2	336
337	-2.75	-1.97		+7.12		+0.71	+1.00		+0.15	-0.04		+3.22	+3.01		+2.92	- 9.3	- 6.3	337
338	-2.40	-2.75		+6.34		+0.92	+1.00		+0.21	-0.06		+3.08	+2.73		+2.19	- 13.8	- 8.4	338
339	-1.35			+3.79			+1.00			-0.07			+1.80				- 1.5	339
340	-1.59	-3.17		+4.42		+0.87	+1.00		+0.24	-0.07		+2.77	+2.03		+1.36	- 2.4	- 5.3	340
341		-3.51				+1.11			+0.27			+2.97			+1.50	0.0		341
342	-2.25	-3.85		+6.03		+1.25	+1.00		+0.29	-0.08		+3.02	+2.61		+1.36	- 0.2	- 9.0	342
343	-1.82	-4.31		+5.04		+1.27	+1.00		+0.33	-0.09		+2.86	+2.26		+0.75	+ 0.1	- 6.8	343
344	-2.03	-4.16		+5.54		+1.20	+1.00		+0.31	-0.09		+2.94	+2.44		+1.00	- 0.1	- 8.7	344
345	-2.15	-3.39		+5.83		+0.40	+1.00		+0.03	-0.09	-0.11		+2.53	-2.83		- 0.4	- 11.4	345
346	-3.16			+8.06			+1.00			-0.04			+3.33				- 7.3	346
347	-3.83			+9.57			+1.00			-0.04			+3.83				- 2.6	347
348	-4.28	-1.54		+10.56		+0.20	+1.00		+0.01	-0.04	-0.19		+4.16	-1.52		- 5.6	+ 0.6	348
349	-4.31	-2.52		+10.63		+1.20	+1.00		+0.20	-0.04		+4.05	+4.18		+3.85	- 5.6	+ 1.7	349
350		-2.60		+1.25			+1.00		+0.21			+4.08			+3.86	- 8.2		350
351		-3.77		+0.30			+1.00		+0.26			+2.00			-0.60	- 8.8		351
352	+0.61	-3.84		-1.62		+0.38	+0.99		+0.27	-0.11		+2.06	-0.41		-0.53	+ 3.6	+ 6.5	352
353	+0.48			-1.22			+0.99			-0.11			-0.30				+ 8.0	353
354	+0.36			-0.82		+0.49	+0.99		+0.28	-0.11		+2.14	-0.11		-0.42		+ 8.4	354
355	+0.22			-0.40			+0.99			-0.10			+0.09				+ 10.7	355
356	+0.08			-0.06		+0.58	+0.99		+0.29	-0.11		+2.22	+0.28		-0.34	- 12.3	+ 10.5	356

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $w_0 = 0.31$ ,  $w_0 = 1.29$ . *Fide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.										
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed Station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observation†	Meridian Deflection†	Serial No.
				° ' "	° ' "	° ' "		"	"	
357	72 E	Kaulia H.S.	7051	A 27 48 25.5 G 27 48 58.6	G 85 14 20.7				-33.1	357
358	E	Mahudeo Pokra H.S.	7095	A 27 40 53.6 G 27 41 31.5	G 85 31 10.0				-37.9	358
359	F	Pota T.S.	222		G 85 26 20.33	A 180 4 5.0 G 180 4 8.3	Madanpur D O 4	- 6.7		359
360	F	Pahladpur T.S.	175	A 26 4 27.24 G 26 4 21.01	G 85 27 13.16				+ 6.2	360
361	G	Dūbauli T.S.	189	A 25 40 22.99 G 25 40 16.23	G 85 20 16.53				+ 6.8	361
362	G	Bihar H.S.	391	A 25 12 39.27 G 25 12 26.05	G 85 30 31.33				+13.2	362
363	H	Mahar H.S.	1606	A 24 44 31.12 G 24 44 20.88	G 85 9 55.13				+10.2	363
364	K	Bichwi H.S.	321		G 86 8 3.99	A 357 49 29.7 G 357 49 32.4	Ekgora E O 40	- 5.7		364
365	N	Chūni T.S.	197		G 87 2 52.80	A 185 49 39.4 G 185 49 49.0	Minai D O 4	- 19.5		365
366	O	Sirkanda T.S.	132		G 87 8 23.57	A 145 34 17.2 G 145 34 21.8	Pureni D O 4	- 9.7		366
367	73 A	Bulbul H.S.	3352	A 23 37 53.44 G 23 37 44.63	G 84 26 13.94				+ 8.8	367
368	A	Mahwari H.S.	3153	A 23 26 9.28 G 23 26 4.96	G 84 54 1.75				+ 4.3	368
369	A	Uhursu H.S.	2271		G 84 44 19.21	A 149 58 57.7 G 149 59 0.8	Bagru E O 38	- 7.2		369
370	C	Andhari H.S.	1442		G 84 14 57.17	A 17 40 39.6 G 17 40 43.1	Garpati D O 1	- 8.7		370
371	E	Chendwār (old) H.S.	2817	A 23 57 16.82 G 23 57 13.75	G 85 26 9.29	A 92 35 20.3 G 92 35 20.5	Kasiātu D O 16	- 0.5	+ 3.1	371
372	H	Outtack H.S.	133	A 20 28 52.05 G 20 29 0.68	G 85 52 1.43	A 155 35 54.6 G 155 35 54.3	Kuplās E 1 21	+ 0.8	- 8.6	372
373	I	Pārasnāth H.S.	4481		G 86 8 10.86	A 145 7 21.0 G 145 7 22.9	Bāmāni D 1 2	- 4.3		373
374	I	Tilubani H.S.	1329	G 23 57 34.89	G 86 33 14.64	A 272 58 23.5 G 272 58 23.1	Sūsīniā D O 8	+ 0.9		374
375	M	Malūncha H.S.	970	A 23 54 29.64 G 23 54 29.02	G 87 5 41.86	A 74 46 32.3 G 74 46 35.7	Durgapur D O 1	- 7.7	+ 0.6	375
376	M	Madhpur T.S.	180		G 87 44 37.29	A 206 49 9.1 G 206 49 5.5	Parhat D O 8	+ 8.4		376
377	N	Kalsibhānga T.S.	303		G 87 8 19.19	A 115 7 20.2 G 115 7 17.7	Kalābani D O 1	+ 6.1		377
378	O	Dariāpur T.S.	63	A 21 47 28.82 G 21 47 27.95	G 87 52 3.32				+ 0.9	378
379	O	Patna T.S.	80	A 21 47 17.28 G 21 47 20.83	G 87 11 45.53	A 207 38 56.0 G 207 38 58.5	Dāntūn D O 3	- 6.3	- 3.6	379
380	O	Chandipur T.S.	53	A 21 26 34.03 G 21 26 36.99	G 87 2 3.66	A 96 49 54.7 G 96 49 55.1	Nilgiri E O 59	- 1.0	- 3.0	380
381	74 A	Khundābolo H.S.	3115	A 19 51 7.03 G 19 51 12.90	G 84 58 17.43	A 196 41 21.2 G 196 41 22.9	Chiklikhāi D O 27	- 4.7	- 5.9	381
382	B	Deodonger H.S.	4534		G 84 3 35.84	A 146 26 29.1 G 146 26 33.0	Thaladi D 1 21	- 11.4		382
383	B	Mai H.S.	483	A 18 47 6.75 G 18 47 16.97	G 84 30 44.31				-10.2	383
384	78 A	Phallut h.s.	11815	A 27 12 4.30 G 27 12 40.86	G 88 1 0.96				-36.6	384

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.



XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/s = 298.3$ .					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
357	+1.01			-2.99			+0.99			-0.12			-1.13				-32.0	357
358	+1.00			-2.90			+0.99			-0.13			-1.08				-36.8	358
359			-4.41			+0.40			+0.30			+2.04		-1.05		-5.6		359
360	+0.64			-1.63			+0.99			-0.12			-0.47				+6.7	360
361	+0.54			-1.30			+0.99			-0.12			-0.32				+7.1	361
362	+0.43			-0.92			+0.99			-0.13			-0.14				+13.3	362
363	+0.29			-0.54			+0.99			-0.12			+0.03				+10.2	363
364			-4.94			+0.58			+0.35			+2.12		-1.28		-4.2		364
365			-5.33			+0.50			+0.37			+2.04		-1.82		-17.9		365
366			-5.48			+0.60			+0.38			+2.10		-1.79		-7.9		366
367	-0.04			+0.41			+0.99			-0.11			+0.43				+8.4	367
368	-0.08			+0.57			+0.99			-0.12			+0.50				+3.8	368
369			-4.36			+0.71			+0.31			+2.29		-0.43		-6.6		369
370			-4.23			+0.83			+0.31			+2.43		-0.06		-8.4		370
371	+0.09			-4.69	+0.12	+0.68	+0.99		+0.33	-0.12		+2.23	+0.32	-0.85		+0.6	+2.8	371
372	-1.02			-5.54	+3.20	+1.30	+0.99		+0.41	-0.13		+2.58	+1.58	-0.69		+2.0	-10.2	372
373				-5.10		+0.74			+0.37			+2.23		-1.19		-2.9		373
374				-5.44		+0.86			+0.39			+2.27		-1.35		+2.5		374
375	+0.13			-5.69	+0.15	+0.83	+0.99		+0.41	-0.15		+2.22	+0.35	-1.63		-5.9	+0.2	375
376				-6.21		+1.02			+0.44			+2.28		-1.88		+10.5		376
377				-5.99		+1.11			+0.43			+2.37		-1.53		+8.1		377
378	-0.49				+2.00		+0.98			-0.16			+1.13			-0.2		378
379	-0.52			-6.13	+2.00	+1.22	+0.99		+0.45	-0.15		+2.43	+1.12	-1.49		-4.3	-4.7	379
380	-0.64			-6.10	+2.32	+1.27	+0.99		+0.45	-0.15		+2.47	+1.24	-1.37		+0.9	-4.2	380
381	-1.28			-5.05	+3.79	+1.29	+0.99		+0.37	-0.12		+2.66	+1.79	-0.16		-3.9	-7.7	381
382				-4.60		+1.30			+0.34			+2.80		+0.44		-10.9		382
383	-1.70				+4.80		+0.99			-0.11			+2.16				-12.4	383
384	+1.00				-2.53		+0.98			-0.17			-0.86				-35.7	384

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *Fide p. 2.*

TABLE  
Deflections of the Plumb-line

EVEREST'S SPHEROID.										
Serial No.	Sheet No.	Observed at	Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection	Serial No.
385	78 A	Tonglu h.s.	10073	A 27 1 11'30 G 27 1 53'54	" " "	G 88 5 2'93	" " "	" " "	-42'2	385
386	B	Nenchal h.s.	8600	A 26 58 33'01 G 26 50 8'25	" " "	G 88 17 44'78	" " "	" " "	-35'2	386
387	B	Kurseong h.s.	4428	A 26 51 15'05 G 26 52 5'56	" " "	G 88 15 54'68	" " "	" " "	-50'5	387
388	B	Siliguri s.	401	A 26 41 18'10 G 26 41 40'37	" " "	G 88 24 49'54	" " "	" " "	-22'3	388
389	B	Jalpaiguri Long. s.	280	A 26 31 11'44 G 26 31 17'39	" " "	A 88 43 52'42 G 88 44 12'77	Dharampur D 0 2	- 15'4	- 6'0	389
390	B	Rānganj T.S.	249	G 26 18 55'51	" " "	G 88 17 30'43	Kanchābari D 0 2	- 24'9		390
391	B	Lohārgara T.S.	205	A 26 2 14'17 G 26 2 12'02	" " "	G 88 21 56'60	" " "	" " "	+ 2'2	391
392	C	Chanduria T.S.	160	A 25 44 31'93 G 25 44 27'47	" " "	G 88 22 17'15	" " "	" " "	+ 4'5	392
393	D	Charaldānga T.S.	149	A 24 52 45'36 G 24 52 43'95	" " "	G 88 23 4'21	" " "	" " "	+ 1'4	393
394	F	Ataro Bānki T.S.	133	G 26 4 50'62	" " "	G 89 28 3'10	Chandrapur D 0 3	- 24'7		394
395	G	Alangjāni T.S.	143	G 25 59 6'81	" " "	G 89 45 41'19	Sānding E 0 1	- 22'8		395
396	G	Halkāchar T.S.	103	G 25 9 55'94	" " "	G 89 42 48'42	Kānchupāra D 0 5	- 23'2		396
397	H	Aloākāndi T.S.	88	G 24 45 29'80	" " "	G 89 38 42'19	Gaborgram D 0 4	- 13'2		397
398	J	Kaikusni H.S.	80.3	G 26 8 11'37	" " "	G 90 39 47'24	Bhairaber Chura E 0 34	- 23'0		398
399	O	Rāngsanobo H.S.	4455	G 25 15 19'60	" " "	G 91 43 20'86	Mosingi E 1 21	- 14'0		399
400	79 A	Madhupur T.S.	92	A 23 56 42'82 G 23 56 38'97	" " "	G 88 29 7'66	Imānagar D 0 3	- 14'0	+ 3'9	400
401	A	Anandbās T.S.	67	G 23 21 19'24	" " "	G 88 22 40'30	Jeodhāra D 0 3	- 7'9		401
402	B	Aknāpur T.S.	98	G 22 54 22'85	" " "	G 88 3 6'66	Hākistāpur D 0 4	- 2'8		402
403	B	Calcutta Base-line S. End T.S.	13	G 22 36 55'68	" " "	G 88 22 54'43	Calcutta Base-line N. End D 0 2	- 7'2		403
404	B	Calcutta Long. s.	18	A 22 32 55'58 G 22 32 54'67	" " "	A 88 21 17'84 G 88 21 29'10	" " "	- 10'4	+ 0'9	404
405	K	Tepri T.S.	67	G 23 57 24'45	" " "	G 89 52 11'99	Bungaon D 0 4	- 6'5		405
406	E	Daulatpur T.S.	60	G 23 8 43'76	" " "	G 89 42 57'76	Maheshpur D 0 5	+ 4'7		406
407	I	Lakhinagar T.S.	51	G 23 0 39'73	" " "	G 90 45 43'08	Kānshpur D 0 4	+ 12'0		407
408	J	Gangapur T.S.	54	G 22 59 34'77	" " "	G 90 27 28'63	Malgaon D 0 5	- 0'2		408
409	M	Dawa H.S.	205	G 23 45 17'63	" " "	G 91 20 16'63	Lambusara D 0 5	- 3'9		409
410	N	Neimu Tān H.S.	226	G 22 48 38'48	" " "	G 91 47 38'49	Bhalti Moin E 1 22	- 4'0		410
411	N	Nagārkhūna H.S.	290	A 22 22 57'08 G 22 22 56'40	" " "	G 91 48 30'42	Chandranath E 0 22	- 8'0	+ 0'7	411
412	N	Chittagong Long. S.	...	G 22 20 18'43	" " "	A 91 50 4'91 G 91 50 16'68	" " "	- 10'9		412

\* A - Astronomical Value.  
G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical	Deflection in Meridian	
385	+0.96	"	"	-2.40	"	"	+0.98	"	"	-0.17	"	"	-0.79	"	"	"	-41.4	385
386	+0.96			-2.36			+0.98			-0.17			-0.78				-34.4	386
387	+0.94			-2.27			+0.98			-0.17			-0.74				-49.8	387
388	+0.90			-2.13			+0.98			-0.17			-0.67				-21.6	388
389	+0.88	-6.63	-0.23	-2.00	+0.93	+0.54	+0.98	+0.09	+0.43	-0.18	+0.03	+2.01	-0.60	-5.37	-2.64	-13.0	-5.4	389
390			-6.02			+0.55			+0.42			+2.02			-2.41	-22.7		390
391	+0.75			-1.61			+0.98			-0.17			-0.43				+2.6	391
392	+0.68			-1.37			+0.98			-0.17			-0.32				+4.8	392
393	+0.46			-0.67			+0.98			-0.17			0.00				+1.4	393
394			-6.71			+0.64			+0.47			+2.03			-2.97	-21.9		394
395			-6.89			+0.67			+0.48			+2.04			-3.09	-19.7		395
396			-7.00			+0.81			+0.49			+2.10			-3.01	-20.2		396
397			-7.04			+0.88			+0.50			+2.13			-2.95	-10.2		397
398			-7.37			+0.69			+0.51			+2.02			-3.54	-19.5		398
399			-8.13			+0.91			+0.57			+2.07			-4.00	-10.0		399
400	+0.20		-6.51	+0.12	+0.94	+0.98	+0.40	-0.17	+2.21	+0.36		-2.32	-11.7	+3.5				400
401			-6.56		+1.04		+0.47		+2.26			-2.22	-5.5					401
402			-6.44		+1.10		+0.46		+2.31			-2.02	-0.6					402
403			-6.71		+1.19		+0.49		+2.33			-2.16	-4.8					403
404	-0.23	-6.41		+1.32	+0.90	+0.98	+0.07	-0.17	-0.03	+0.85	-5.27		-5.1	0.0				404
405			-7.40		+1.05		+0.52		+2.20			-3.06	-3.3					405
406			-7.51		+1.23		+0.54		+2.27			-2.93	+7.8					406
407			-8.08		+1.34		+0.58		+2.28			-3.30	+15.4					407
408			-7.90		+1.32		+0.57		+2.28			-3.20	+3.0					408
409			-8.26		+1.19		+0.59		+2.20			-3.73	-0.2					409
410			-8.86		+1.28		+0.63		+2.29			-4.08	+0.1					410
411	-0.10		-9.02	+1.45	+1.32	+0.97	+0.64	-0.22	+2.33	+1.00		-4.15	-3.8	-0.3				411
412			-8.49		+1.20		+0.09		-0.04			-6.97	-3.0					412

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *vide p. 2.*

TABLE  
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	Height in feet	EVEREST'S SPHEROID.						Meridian Deflection†	Serial No.
				Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†			
413	83 H	Loijing H.S.	6610	° ' "	° ' "	° ' "	Pangkibot D 0 48	- 8.7		413	
				G 24 44 28.17	G 93 43 44.79	A 120 58 12.0 G 120 58 16.0					
414	K	Thyoliching H.S.	6566	G 25 0 5.76	G 94 43 46.98	A 113 3 3.9 G 113 3 7.3	Sirohifurur E 1 15	- 7.3		414	
415	L	Tamunja H.S.	3387	G 24 39 9.33	G 94 36 48.01	A 116 36 26.7 G 116 36 26.9	Khambiching E 1 34	- 0.4		415	
416	P	Seikpa H.S.	3857	G 24 35 37.84	G 95 45 33.72	A 258 32 25.1 G 258 32 37.6	Mudhun E 0 17	- 27.3		416	
417	P	Thonbinzin H.S.	1932	G 24 14 3.03	G 95 58 8.16	A 277 46 13.1 G 277 46 21.6	Katha E 0 37	- 18.9		417	
418	84 C	Fi Tān H.S.	563	G 21 49 20.78	G 92 8 16.02	A 256 23 22.7 G 256 23 20.2	Laraintong E 1 34	- 16.2		418	
419	D	Akyab Long. S.	20	A 20 8 14.87 G 20 8 13.10	A 92 53 38.63 G 92 53 49.63			- 10.3	+ 1.8	419	
420	H	Yeponetaung H.S.	2819	A 20 14 51.83 G 20 14 55.91	G 93 41 49.34	A 74 17 19.6 G 74 17 23.3	Rongdong D 1 1	- 10.0	- 4.1	420	
421	H	Dattaung H.S.	455	G 20 13 14.44	G 93 1 9.09	A 171 27 28.3 G 171 27 31.0	Bengara E 0 37	- 7.3		421	
422	M	Ubyetaung H.S.	2766	G 23 40 52.06	G 95 57 42.75	A 303 38 45.7 G 303 38 50.1	Tagaungtaung D 0 17	- 10.0		422	
423	M	Male H.S.	848	G 23 2 53.30	G 95 57 18.09	A 316 31 54.4 G 316 32 0.8	Wapyadaung E 1 12	- 15.0		423	
424	N	Sheinmaga H.S.	456	G 22 16 33.89	G 95 58 15.79	A 354 23 23.8 G 354 23 31.5	Mingun E 0 31	- 18.8		424	
425	N	Mingun H.S.	1343	G 22 3 0.71	G 95 59 41.41	A 174 23 57.7 G 174 24 5.6	Sheinmaga D 0 43	- 19.5		425	
426	P	Taungpila H.S.	1012	G 20 41 52.71	G 95 53 4.50	A 240 23 15.5 G 240 23 17.0	Yuba E 0 50	- 4.0		426	
427	85 E	Retkamauk H.S.	1582	A 19 47 37.32 G 19 47 38.55	G 93 28 13.32	A 229 44 59.4 G 229 44 50.2	Ingrautauung D 0 14	+ 0.6	- 1.2	427	
428	N	Kyunggyi S.	...	G 18 49 20.95	G 95 12 55.40	A 109 26 42.1 G 109 26 46.0	Prome E 3 45	- 11.4		428	
429	N	Prome Long. S.	100	A 18 49 18.62 G 18 49 14.28	A 95 12 42.20 G 95 12 57.44			- 14.4	+ 4.3	429	
430	92 G	Kumon Bum H.S.	7970	G 25 38 13.48	G 97 3 34.06	A 308 54 40.0 G 308 54 46.8	Maran Bum D 2 7	- 14.2		430	
431	H	Kumtum Bum H.S.	1833	G 24 46 44.32	G 97 9 17.40	A 210 24 26.5 G 210 24 32.9	Maran Bum D 0 17	- 13.9		431	
432	93 A	Sinpitaung H.S.	2649	G 23 29 48.36	G 96 45 45.79	A 162 2 20.3 G 162 2 25.8	Taungkalat D 0 32	- 12.7		432	
433	E	Loi Hpa Laing H.S.	3591	G 23 14 13.07	G 97 37 24.38	A 188 35 42.8 G 188 35 49.4	Loi Song E 1 11	- 15.4		433	
434	J	Loi Hpatan H.S.	6419	G 22 55 35.51	G 98 0 52.58	A 210 28 36.0 G 210 28 43.6	Loi Taow E 0 13	- 18.0		434	
435	O	Loi Kiipma H.S.	3792	G 22 1 56.87	G 98 4 30.30	A 154 30 44.7 G 154 30 53.0	Loi Hsimu E 0 2	- 20.5		435	
436	94 A	Loi Heam Heum H.S.	6472	G 21 41 45.04	G 99 53 57.51	A 115 53 15.6 G 115 53 22.4	R. M.	- 17.1		436	
437	B	Letpataung H.S.	3975	G 19 34 7.27	G 96 28 38.76	A 174 46 24.5 G 174 46 31.5	Byingye E 0 36	- 19.7		437	
438	B	Toungoo S.	186	G 18 56 1.54	G 96 25 55.21	A 30 46 36.9 G 30 46 42.5	Bhondan E 0 49	- 16.3		438	
439	B	Myayabeingkyo H.S.	1411	G 18 21 33.93	G 96 22 53.46	A 169 33 42.4 G 169 33 43.8	Khengdun E 0 32	- 4.2		439	
440	H	Martaban h.s.	273	A 16 31 39.6 G 16 31 33.08	G 97 36 59.57				+ 6.5	440	

\* A - Astronomical Value.  
G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$ .					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
413	"	"	-9.40	"	"	+1.14	"	"	+0.66	"	"	+2.09	"	"	-4.93	-4.0	413	
414			-9.94			+1.14			+0.70			+2.06			-5.43	-2.1	414	
415			-9.93			+1.23			+0.70			+2.09			-5.34	+4.7	415	
416			-10.60			+1.32			+0.75			+2.08			-5.90	-21.6	416	
417			-10.82			+1.45			+0.76			+2.11			-5.96	-13.1	417	
418			-9.48			+1.33			+0.67			+2.37			-4.49	-11.9	418	
419	-0.79	-9.11		+3.48	+1.29		+0.97	+0.09		-0.24	-0.08		+1.85	-7.52		-2.8	-0.1	419
420	-0.69		-11.20	+3.37		+1.70	+0.96		+0.80	-0.25		+2.54	+1.83		-5.57	-5.2	-5.9	420
421			-10.77			+1.59			+0.77			+2.55			-5.24	-2.9		421
422			-11.01			+1.56			+0.78			+2.16			-5.99	-4.2		422
423			-11.28			+1.05			+0.80			+2.21			-6.09	-9.1		423
424			-11.63			+1.75			+0.83			+2.29			-6.24	-13.1		424
425			-11.75			+1.78			+0.84			+2.31			-6.31	-13.7		425
426			-12.38			+1.98			+0.88			+2.45			-6.54	+2.0		426
427	-0.88		-11.29	+3.80		+1.71	+0.96		+0.81	-0.25		+2.60	+1.99		-5.56	+5.4	-3.2	427
428			-13.04			+2.18			+0.94			+2.70			-6.66	-5.6		428
429	-1.12	-10.48		+4.70	+1.45		+0.95	+9.10		-0.28	-0.11		+2.40	-8.71		-5.7	+1.9	429
430			-11.05			+1.09			+0.77			+1.99			-6.59	-7.8		430
431			-11.34			+1.35			+0.80			+2.05			-6.58	-7.5		431
432			-11.55			+1.66			+0.82			+2.16			-6.43	-6.5		432
433			-12.17			+1.78			+0.87			+2.18			-6.84	-8.8		433
434			-12.55			+1.87			+0.90			+2.20			-7.00	-11.1		434
435			-13.01			+2.03			+0.93			+2.28			-7.28	-13.4		435
436			-14.28			+2.31			+1.02			+2.03			-8.54	-8.9		436
437			-13.41			+2.27			+0.97			+2.58			-7.05	-13.2		437
438			-13.79			+2.41			+0.99			+2.60			-7.20	-9.7		438
439			-14.15			+2.51			+1.02			+2.75			-7.35	+2.3		439
440	-1.87			+6.90			+0.94			-0.31			+3.29				+3.2	440

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . Vide p. 2.

TABLE  
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.	
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection		
441	94 H	Moulmein Long. S.	90	A 16 30 2.97 G 16 29 54.90	A 97 37 23.41 G 97 37 40.04				- 15.9	+ 8.1	441
442	H	Taungzun H.S.	854	A 16 25 56.2 G 16 25 48.55	G 97 40 18.59	A 31 16 18.9 G 31 16 22.7	Konlah D 0 34		- 12.9	+ 7.7	442
443	95 G	Southern Moscos H.S.	1186	G 13 49 59.67	G 97 55 4.49	A 162 20 54.5 G 162 20 55.6	Middle Moscos D 0 9		- 4.5		443
444	L	Sandawat H.S.	719	A 12 28 0.1 G 12 27 51.87	G 98 40 31.46					+ 8.2	444
445	L	Nutkalintaung H.S.	888	G 12 25 33.42	G 98 43 33.26	A 127 46 35.9 G 127 46 37.6	Sandawat D 0 22		- 7.7		445
446	L	Mergui Base-line E. End T.S.	20	A 12 22 29.5 G 12 22 21.17	G 98 46 36.49	A 72 29 47.9 G 72 29 49.3	Mergui Base-line W. End D 0 2		- 6.4	+ 8.3	446
447	L	Mergui Base-line W. End T.S.	18	A 12 21 41.3 G 12 21 32.57	G 98 43 59.73	A 252 29 14.0 G 252 29 15.7	Mergui Base-line E. End D 0 2		- 7.8	+ 8.7	447
448	95 L	Minthangtaung H.S.	3850	A 12 19 44.2 G 12 19 35.05	G 98 47 47.88	A 157 5 40.9 G 157 5 43.2	Mergui Base-line E. End D 0 25		- 10.5	+ 9.1	448
A d d e n d a .											
449	30 C	Robat S.	3095	A 29 49 9.16 G 29 48 58.75	G 60 55 10.90					+ 10.4	449
450	34 L	Zawa H.S.	7922	G 28 57 44.43	G 66 35 18.70	A 178 40 55.7 G 178 40 50.0	Zebra E 0 6		+ 10.3		450
451	43 G	Murree Obey. S.	7458	G 33 54 57.36	G 73 24 25.67	A 227 39 50.4 G 227 39 54.1	Nerh D 1 6		- 5.5		451
452	43 J	Rustamgarhi h.s.	5351	G 34 4 38.70	G 74 49 53.68	A 30 52 6.8 G 30 52 27.1	Gogipatri E 1 24		- 30.0		452
453	43 J	Poshkar H.S.	8323	A 34 2 2.78 G 34 1 48.98 ⊙	G 74 29 51.22 ⊙	A 318 14 0.7 G 318 13 54.0 ⊙	Gogipatri D 0 30		+ 9.9	+ 13.8	453
454	43 K	Gogipatri H.S.	7752	A 33 51 46.90 G 33 51 43.87 ⊙	G 74 40 38.57 ⊙	A 222 17 18.0 G 222 17 11.7 ⊙	Zebanwan E 0 29		+ 9.4	+ 3.0	454
459	43 J	Zebanwan H.S.	8799	A 34 3 33.50 G 34 3 59.14	G 74 54 2.94					- 25.6	459
460	43 K	Reban H.S.	5447	A 33 45 17.74 G 33 45 25.76	G 74 59 52.05					- 8.0	460
461	78 B	Dumdanzi T.S.	507	A 26 28 25** G 26 28 31	G 88 17 35					- 6**	461
462	78 B	Thakurganj T.S.	264	A 26 24 53** G 26 25 2	G 88 7 45					- 9**	462

\* A = Astronomical Value.  
G = Triangulated or Geodetic Value

⊙ Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

⊙ Adjusted value.

\*\*Provisional value.

.. XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I: $\delta a = 1 \text{ km}$			Case II: $\delta b = 1 \text{ km}$			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200 \text{ metres, } 1/e = 298.3.$					
	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	u	v cos $\lambda$	w cot $\lambda$	Deflection in Prime Vertical	Deflection in Meridian	
441	-1.88	-11.86		+6.93	+1.66		+0.94	+0.10		-0.31	-0.15		+3.31	-9.89		-6.0	+4.8	441
442	-1.90		-16.69	+7.00		+3.38	+0.94		+1.21	-0.31		+3.04	+3.33		-8.61	-6.0	+4.4	442
443			-19.85			+4.50			+1.45			+3.58			-9.88	+3.4		443
444	-3.64			+11.05			+0.93			+0.33			+4.72				+3.5	444
445			-22.77			+5.76			+1.67			+3.96			-11.13	+1.1		445
446	-3.68		-22.91	+11.14		+5.82	+0.93		+1.68	-0.33		+3.98	+4.74		-11.19	+2.5	+3.6	446
447	-3.69		-22.90	+11.16		+5.81	+0.93		+1.68	-0.33		+3.98	+4.75		-11.18	+1.1	+3.9	447
448	-3.70		-23.01	+11.19		+5.87	+0.93		+1.69	-0.33		+3.99	+4.76		-11.23	-1.6	+4.3	448
A d d e n d a .																		
449	+1.85			-4.49			+0.96			+0.26			-0.98				+11.4	449
450			+5.92			-0.22			-0.40			+1.85			+7.58	+2.5		450
451			+2.12			+0.08			-0.13			+1.63			+4.08	-10.0		451
452			+1.91			+0.05			-0.09			+1.63			+3.87	-34.3		452
453	+1.63		+1.58	-7.22		+0.06	+1.00		-0.10	+0.05		+1.6	-3.48		+3.56	+5.9	+17.3	453
454	+1.62		+1.49	-7.12		+0.05	+1.00		-0.09	+0.05		+1.63	-3.42		+3.49	+5.5	+6.4	454
459	+1.63			-7.24			+1.00			+0.04			-3.51				-22.1	459
460	+1.61			-7.05			+1.00			+0.04			-3.39				-4.6	460
461	+0.85			-1.96			+0.98			-0.17			-0.39				-5.8*	461
462	+0.83			-1.91			+0.98			-0.17			-0.57				-8.8*	462

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.29$ . *Fide* p. 2.      \*\* Provisional value.

To complete the statement of data concerning gravity values of the residuals, observed *minus* theoretical values are now given. These have been taken from Professional Paper No. 15 and for a full explanation reference should be made thereto. It is only necessary to explain that  $\gamma_a, \gamma_b, \gamma_c$  are the theoretical values of gravity based on Helmert's formula:—  $\gamma_0 = 978 \cdot 030 (1 + 0 \cdot 005302 \sin^2 \phi - 0 \cdot 000007 \sin^2 2\phi)$  and assuming the corrections according to the Free Air, Bouguer and Hayford hypotheses respectively.

TABLE XCVI.

No.	Name	Latitude	Longitude	Height	$g - \gamma_a - \cdot 011$	$g - \gamma_b + \cdot 030$	$g - \gamma_c - \cdot 011$
		° ' "	° ' "	feet	dynes	dynes	dynes
1	Agra ...	27 10	78 1	535	-012	+011	+006
2	Aligarh ...	27 54	78 1	612	-039	-019	-018
3	Allahābād ...	25 26	81 55	288	-023	+008	-002
4	Amgaon ...	21 22	80 28	1032	-015	-009	-014
5	Amraoti ...	20 56	77 46	1123	+014	+017	+015
6	Arrah ...	25 34	84 39	188	-067	-032	-039
7	Asarori ...	30 14	77 58	2467	-061	-101	...
8	Asirgarh ...	21 28	76 18	2077	+046	+023	+019
9	Badnūr ...	21 54	77 54	2103	+045	+015	+027
10	Bangalore ...	13 1	77 35	3118	+014	-050	...
11	Bassein ...	16 47	94 44	23	+006	+046	...
12	Bhopāl ...	23 16	77 25	1630	+018	+004	+011
13	Bilāspur ...	22 4	82 12	878	-006	+005	+002
14	Bina ...	24 11	78 12	1355	+015	+010	+015
15	Buxar ...	25 35	83 59	207	-051	-017	-025
16	Chātra ...	24 13	88 23	64	-025	+014	-006
17	Colāba ...	18 54	72 49	34	+052	+092	+052
18	Cuttack ...	20 29	85 52	92	-005	+033	-005
19	Daltonganj ...	24 2	84 4	707	-004	+013	+014
20	Damoh ...	23 50	79 26	1213	-012	-012	-007
21	Darjeeling ...	27 3	88 16	6966	+044	-124	...
22	Dehra Dūn ...	30 19	78 3	2239	-085	-115	-005
23	Dera Ghāzi Khān ...	30 4	70 46	397	-108	-080	...
24	Dholpur ...	26 42	77 55	577	-030	-008	-016
25	Edgar Shaft (Surface) ...	12 56	78 16	2945	+053	-005	...
26	Ellichpur ...	21 18	77 31	1314	+019	+016	+020
27	Fatehpur ...	30 26	77 44	1434	-085	-089	...
28	Ferozepur ...	30 56	74 37	647	-004	+015	...
29	Gaya ...	24 48	85 0	361	-031	-002	-008
30	Gesupur ...	28 33	77 42	691	-031	-013	-006
31	Goonā ...	24 39	77 19	1569	+015	+003	+008
32	Gorakhpur ...	26 45	83 23	257	-127	-095	-081
33	Gwalior ...	26 14	78 13	658	-030	-011	-018



TABLE XCVI.—(Continued).

No.	Name	Latitude	Longitude	Height	$g-\gamma_s - \cdot 011$	$g-\gamma_s + \cdot 030$	$g-\gamma_s - \cdot 011$
		°	'	feet	dynes	dynes	dynes
34	Hardwār ...	29 56	78 9	949	-·117	-·106	...
35	Hāthras ...	27 37	78 3	587	-·020	+·001	000
36	Henzada ...	17 39	95 27	46	-·031	+·008	...
37	Hoshangābād ...	22 45	77 44	1002	·000	+·007	+·010
38	Jacobābād ...	28 17	68 27	183	+·003	+·038	+·027
39	Jalgaon ...	21 0	75 34	760	·000	+·015	+·009
40	Jalpaiguri ...	26 31	88 44	268	-·124	-·091	-·031
41	Japla ...	24 32	84 0	474	-·031	-·006	-·009
42	Jhānsi ...	25 27	78 34	858	-·004	+·008	+·003
43	Jubbulpore ...	23 9	79 59	1467	+·017	+·009	+·019
44	Kaliāna ...	29 31	77 39	810	-·065	-·051	-·018
45	Kaliānpur ...	24 7	77 39	1763	+·039	+·021	+·028
46	Kālka ...	30 50	76 56	2202	-·045	-·074	...
47	Kālsi ...	30 31	77 50	1684	-·084	-·090	...
48	Katni ...	23 50	80 26	1254	-·010	-·011	-·004
49	Kesarbāri ...	26 8	88 31	204	-·071	-·037	...
50	Khandwa ...	21 50	76 22	1014	+·033	+·040	+·036
51	Khurjā ...	28 14	77 52	649	-·054	-·035	-·030
52	Kisnapur ...	25 2	88 28	113	+·001	+·038	+·028
53	Kodaikānal ...	10 14	77 28	7665	+·156	-·049	...
54	Kurseong ...	26 53	88 17	4913	-·011	-·117	...
55	Lalitpur ...	24 41	78 24	1199	-·016	-·015	-·013
56	Ludhiāna ...	30 55	75 51	835	-·053	-·040	...
57	Mach ...	29 52	67 18	3522	-·032	-·103	...
58	Madras ...	13 4	80 15	20	-·024	+·016	-·064
59	Maihar ...	24 16	80 48	1161	-·020	-·018	-·014
60	Majhauri Rāj ...	26 18	83 58	219	-·105	-·071	-·068
61	Mandalay ...	22 0	96 6	244	-·028	+·006	...
62	Maymyo ...	22 1	96 28	3495	+·050	-·026	...
63	Meiktila ...	20 51	95 52	799	-·003	+·011	...
64	Meerut ...	29 0	77 42	734	-·035	-·019	...
65	Mhow ...	22 33	75 46	1903	-·002	-·025	-·026
66	Mīān Mir ...	31 32	74 23	708	-·004	+·013	+·029
67	Moghal Sarai ...	25 17	83 6	257	-·040	-·008	-·016
68	Mogok ...	22 55	96 30	3685	+·063	-·020	...
69	Mohan ...	30 11	77 55	1660	-·031	-·093	...
70	Monghyr ...	25 23	86 28	154	-·067	-·031	-·036
71	Montgomery ...	30 40	73 6	557	-·011	+·011	+·008
72	Mortakka ...	22 13	76 3	576	-·022	·000	-·006

TABLE XCVI—(Continued).

No.	Name	Latitude	Longitude	Height	$g-\gamma_0-.011$	$g-\gamma_0+.030$	$g-\gamma_0-.011$
		° ' "	° ' "	feet	dynes	dynes	dynes
73	Mukhtiāra ...	22 24	75 59	926	-.039	-.029	-.030
74	Multān ...	30 11	71 26	404	-.066	-.039	...
75	Mussooree (Camel's Back)	30 28	78 5	6924	+.074	-.093	+.042
76	Mussooree (Dunseverick)	30 27	78 4	7129	+.076	-.098	...
77	Muttra ...	27 28	77 42	562	-.015	+.007	+.004
78	Muzaffarpur ...	26 7	85 25	179	-.091	-.056	-.053
79	Myingyan ...	21 29	95 24	248	-.020	+.013	...
80	Mysore ...	12 19	76 40	2501	+.003	-.010	...
81	Nojli ...	29 53	77 40	879	-.099	-.087	...
82	Ootacamund ...	11 25	76 42	7395	+.184	-.016	+.001
83	Pathānkot ...	32 17	75 39	1088	-.175	-.169	-.087
84	Pendra ...	22 47	82 0	1996	+.010	-.016	-.003
85	Prome ...	18 50	95 14	101	-.027	+.011	...
86	Pyinmana ...	19 44	96 12	409	-.014	+.014	...
87	Quetta ...	30 12	67 1	5520	+.020	-.123	-.004
88	Raipur ...	21 14	81 41	996	-.013	-.005	-.014
89	Rājpur ...	30 24	78 6	3321	-.051	-.113	+.015
90	Rāmchāndpur ...	25 41	88 33	132	-.031	+.006	...
91	Rānchi ...	23 23	85 19	2167	+.040	+.008	+.019
92	Rangoon ...	16 48	96 9	164	+.010	+.045	...
93	Roorkee ...	29 52	77 54	867	-.112	-.099	-.055
94	Salem ...	11 40	78 9	948	-.047	-.038	-.059
95	Sandakphu ...	27 6	88 0	11766	+.178	-.125	+.037
96	Sasarām ...	24 57	83 59	340	-.025	+.005	-.002
97	Saugor ...	23 52	78 48	1757	+.010	-.008	.000
98	Seoni ...	22 5	77 29	2032	+.041	+.014	+.025
99	Shāhpur ...	22 12	77 54	1286	+.006	+.004	+.012
100	Sibi ...	29 33	67 53	434	-.137	-.109	-.070
101	Siliguri ...	26 42	88 25	387	-.160	-.130	-.050
102	Simla ...	31 6	77 10	7043	+.080	-.100	...
103	Sīpri ...	25 26	77 39	1533	+.027	+.016	+.018
104	Sultānpur ...	26 16	82 5	314	-.064	-.034	...
105	Toungoo ...	18 56	96 27	159	-.011	+.025	...
106	Ujjain ...	23 11	75 47	1612	-.013	-.026	-.022
107	Umaria ...	23 32	80 54	1499	+.016	+.007	+.018
108	Yercaud ...	11 47	78 12	4493	+.072	-.027	-.044

## CHAPTER X.

### Deflections of the Plumb-line and values of "g" derived in Turkistan (Ferghana) by the Russian observations.

1. The problem of the origin of the Himalayas has already been attacked from the geodetic point of view making use of the deflection results and gravity anomalies obtained in India. All these results relate to points south of the main chain. It is now possible to put certain results obtained by the Russian surveyors into the same terms: and it has accordingly been considered suitable to do this, so that all geodetic data closely related to the Himalayas will be conveniently available in one volume.

2. Values of deflection can be found for Osh base, N.W. end, the starting point of the Russian triangulation which links up with the Indian Pamir triangulation. The Russian triangulation emanates from a base near Osh latitude  $40^{\circ} 31'$ , longitude  $42^{\circ} 30'$  E. of Pulkowa. Latitude and azimuth were observed astronomically while the longitude of Osh was found by electric telegraph from Tashkent, and transferred from Osh to the north-west end of the base by chronometer. Presumably the longitude of Tashkent is in terms of astronomical longitude measured from Pulkowa (Poulcovo) which is  $2^h 1^m 18^s \cdot 57$  of time E. of Greenwich (*vide* Nautical Almanac). This converted becomes  $30^{\circ} 19' 38'' \cdot 55$  which must be added to the values of longitudes expressed in Russian terms. Calculations of the Russian triangulation were performed on Bessel's spheroid. Suppose the deflections at Osh are  $\xi$ ,  $\eta$  in prime vertical and meridian (positive if S. or W.). Then the geodetic elements are found by deducting  $\eta = -u$ ,  $\xi \sec \lambda = -v$ ,  $\xi \tan \lambda = -w$  from the astronomic values of the latitude, longitude and azimuth respectively. The Astronomic elements of the N.W. end Osh base are:—

Latitude	$40^{\circ} 37' 16'' \cdot 67$
Longitude	$72^{\circ} 56' 11'' \cdot 17^*$
Azimuth	not stated.

From this astronomic origin and a measured base triangulation was carried to a station Kukhtek of the Indian triangulation, the deduced latitude being  $37^{\circ} 17' 43'' \cdot 94$  and the longitude  $74^{\circ} 59' 55'' \cdot 53^*$ .

It is necessary to express these in terms of the geodetic value of Osh and the Helmert spheroid. This can be done approximately by means of the tables already prepared with reference to Kalianpur as origin. Kukhtek is  $2^{\circ} 3' 44'' \cdot 36$  east of the Osh origin, and as the tables prepared for Kalianpur are shown in absolute longitude the corresponding longitude required is that of Kalianpur increased by this *i.e.*  $77^{\circ} 39' 17'' \cdot 57 + 2^{\circ} 3' 44'' \cdot 36 \doteq 79^{\circ} 43' 2'' \doteq 79^{\circ} \cdot 7$ .

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\* This includes  $30^{\circ} 19' 38'' \cdot 55$ , the difference of longitude of Greenwich and Pulkowa.

3. Suppose corresponding to changes  $u, v, w$  at Osh, changes  $u', v', w'$  occur at Kukhtek and that both are due to imaginary changes  $u_0, v_0, w_0$  at an origin O on the longitude of Osh and at the latitude of Kalianpur together with a change of axes from those of Bessel to those of Helmert for which  $\delta a = +.803$  and  $\delta b = +.739$  (*vide* Appendix).

From tables XVII-XX a change at O of  $u_0, v_0$  and  $w_0$  causes changes of  $u_0, v_0 + .370 w_0, 1.20 w_0$  at Osh and from the axes change the further changes (from tables XXIX-XXXIV) at Osh are  $+.803 \times 1.24 - .739 \times 10.56, .803 \times 0 + .739 \times 0, .803 \times 0 + .739 \times 0$  i.e.  $-6.808, 0, 0$ . The total changes at Osh accordingly are  $u_0 - 6.808, v_0 + .370 w_0, 1.20 w_0$ . The changes at Kukhtek,  $2^\circ 3' 44''$  E. of meridian origin and latitude  $37^\circ.3$  may be found in the same tables under longitude  $79^\circ.7$ .

They are 
$$\begin{matrix} u_0 - .032 w_0 & \left| \begin{matrix} r_0 + .027 u_0 + .284 w_0 \\ -.803 \times 1.525 + .739 \times .191 \end{matrix} \right| & \begin{matrix} w_0 + .045 u_0 + 1.145 w_0 \\ -.803 \times .752 - .739 \times .056 \\ - .604 - .041 \end{matrix} \\ + .803 \times 1.626 - .739 \times 9.033 & & \\ 1.306 - 6.675 & \left| \begin{matrix} -1.225 + .141 \end{matrix} \right| & \end{matrix}$$
 which reduce to  $u' = u_0 - .032 w_0 - 5.369$   $v' = v_0 + .027 u_0 + .284 w_0 - 1.084$   $w' = w_0 + .045 u_0 + 1.145 w_0 - .645$

Now at Osh  $u = u_0 - 6.808, v = v_0 + .370 w_0, w = 1.2 w_0$   
 In terms of vertical deflection  $u = -\eta, v = -\xi \sec \lambda, w = -\xi \tan \lambda$  which determine  $u_0, v_0, w_0$  in terms of  $\xi$  and  $\eta$  as follows:—

$$u_0 = -\eta + 6.808, w_0 = -\frac{1}{1.2} \xi \tan \lambda, v_0 = -\xi \sec \lambda + .308 \xi \tan \lambda \quad \text{when } \lambda = 40^\circ 37'$$

$$u_0 = -\eta + 6.808, w_0 = -.715 \xi, \quad v_0 = -1.317 \xi + .264 \xi = -1.053 \xi$$

Hence  $u' = -\eta + 6.808 + .023 \xi - 5.369, v' = -1.053 \xi + .027 (-\eta + 6.808) - .203 \xi - 1.084$   
 and  $w' = -.715 \xi + .045 (-\eta + 6.808) - .818 \xi - .645$   
 $= u' = -\eta + -.023 \xi + 1.439, v' = -1.256 \xi - .027 \eta - .900, w' = -1.533 \xi - .045 \eta - .339.$

These are the corrections which have to be applied to the Russian values of coordinates of Kukhtek, to bring them into terms of the Indian triangulation.

The Indian values have to be corrected to bring into terms of the Helmert spheroid and the observed values of the elements at Kalianpur, corresponding to  $u_0 = .31, v_0 = 0^*, w_0 = 1.29, \delta a = .924, \delta b = .743$ . Denoting the corrections by  $u'', v'', w''$  the corrections at Kukhtek latitude  $37^\circ.3$  and longitude  $75^\circ.0$  are:—

$$\begin{matrix} u'' = .31 \times .999 + 1.29 \times .043 & \left| \begin{matrix} v'' = -.31 \times .036 + 1.29 \times .282 \\ + .924 \times 1.590 - .743 \times 9.032 \\ = .310 + .055 + 1.470 - 6.711 \\ = -4.877 \end{matrix} \right| & \begin{matrix} w'' = -.31 \times .058 + 1.29 \times 1.144 \\ + .924 \times .979 + .743 \times .072 \\ = -.018 + 1.476 + .905 + .053 \\ = +2.416 \end{matrix} \\ & & \end{matrix}$$

The values obtained by the Indian triangulation for latitude and longitude of Kukhtek are:—

	$37^\circ 17' 32''.97$		$75^\circ 0' 12''.19$
$u''$	$-4.88$	$v''$	$+2.00$
Values in terms of Helmert spheroid	$37 17 28.09$		$75 0 14.19$
Values obtained by Russian triangulation are	$37^\circ 17' 43''.94$		$74^\circ 59' 55''.53$

and to bring these into accord with the Indian values it is necessary to apply  $-15''.85$  to the latitude and  $+18''.66$  to the longitude.

\* This is zero because the deflection at Kalianpur in prime vertical had never been taken account of, although it was implied by the values of azimuths adopted.

The following equations are formed giving the quantities  $\xi$  and  $\eta$  at Osh :—

$$-\eta + \cdot 023 \xi + 1\cdot 439 = -15\cdot 85$$

$$-\cdot 027 \eta - 1\cdot 256 \xi - \cdot 900 = +18\cdot 66$$

whence

$$\eta = + \cdot 023 \xi + 17\cdot 29$$

$$1\cdot 256 \xi = - \cdot 027 \eta - 19\cdot 56$$

$$= - \cdot 001 \xi - \cdot 467 - 19\cdot 56$$

$$\xi = - \frac{20\cdot 03}{1\cdot 257} = -15\cdot 94$$

$$\eta = 17\cdot 29 - \cdot 37 = +16\cdot 92$$

} at Osh.

The values found by the Russian observers in terms of Tashkent vertical are  $\xi = -6\cdot 04$  and  $\eta = 23\cdot 43$ , *vide* table *XCVIII*.

4. It is possible to bring these results into agreement with the values deduced from the Indian side by supposing that the vertical at Tashkent, latitude  $41^\circ 21'$  and longitude  $69^\circ 18'$ , is deflected with reference to Kalianpur.

Consider the effect of changes  $u_0, v_0, w_0$  at Tashkent longitude and Kalianpur latitude, and from Bessel to Helmert spheroid for which  $\delta a = +\cdot 803, \delta b = +\cdot 739$ .

From tables *XVII-XX* a change at origin of  $u_0, v_0, w_0$  causes changes of  $u_0, v_0 + \cdot 372 w_0, 1\cdot 204 w_0$  at Tashkent and from axes changes the further changes from tables *XXIX-XXXIV* at Tashkent are  $-7\cdot 962, 0, 0$ . The total changes at Tashkent accordingly are  $u_0 - 7\cdot 962, v_0 + \cdot 372 w_0, 1\cdot 204 w_0$ .

The total changes at Osh  $\lambda 40^\circ 37', L 81^\circ 17', \left\{ = 77^\circ 39' + (72^\circ 56' - 69^\circ 18') \right\}$  calculated from the same tables are :—

$$+ \cdot 998 u_0 - \cdot 059 w_0 - 6\cdot 791 \left| v_0 + \cdot 052 u_0 + \cdot 357 w_0 - 2\cdot 015 \right| + \cdot 081 u_0 + 1\cdot 197 w_0 - 1\cdot 283$$

The following equations are formed :—

$$+ \cdot 998 u_0 - \cdot 059 w_0 - 6\cdot 791 = u = 6\cdot 5 \quad (1)$$

$$v_0 + \cdot 052 u_0 + \cdot 357 w_0 - 2\cdot 015 = v = 9\cdot 9 \sec \lambda = 13\cdot 042 \quad (2)$$

$$+ \cdot 081 u_0 + 1\cdot 197 w_0 - 1\cdot 283 = w = 9\cdot 9 \tan \lambda = 8\cdot 490 \quad (3)$$

in which the numerical quantities 6·5 and 9·9 are the corrections necessary to  $\xi$  and  $\eta$  as determined by the Russian observers, to bring them into agreement with the Kalianpur terms.

Hence

$$u_0 = \cdot 059 w_0 + 13\cdot 318 \quad (1)$$

$$1\cdot 197 w_0 = -\cdot 081 u_0 + 9\cdot 773 \quad (3)$$

$$= -\cdot 005 w_0 - 1\cdot 079 + 9\cdot 773$$

$$\text{or } 1\cdot 202 w_0 = 8\cdot 694$$

$$\text{or } w_0 = 7\cdot 23$$

$$\therefore u_0 = \cdot 427 + 13\cdot 318 = 13\cdot 75 \quad (1)$$

$$\text{and } v_0 = -\cdot 715 - 2\cdot 581 + 2\cdot 015 + 13\cdot 042 = 11\cdot 76 \quad (2)$$

5. With these origin changes the corrections at certain degree squares have been computed from the tables and the results are exhibited in table *XCVII*.

TABLE *XCVII*.

$\lambda \backslash L$		$L$				
		$69^\circ$	$70^\circ$	$71^\circ$	$72^\circ$	$73^\circ$
$u$	$40^\circ$	7·20	7·09	6·98	6·87	6·76
		14·41	14·05	13·70	13·34	12·99
		8·63	8·59	8·55	8·52	8·48
$u$	$41^\circ$	6·78	6·67	6·56	6·45	6·34
		14·59	14·24	13·89	13·54	13·20
		8·75	8·71	8·66	8·62	8·57

6. The results of the Russian observations\* are now given first in terms of Tashkent and Bessel's spheroid and then corrected to the Kalianpur vertical and Helmert's spheroid. The stations are shown in chart No. VI.

TABLE XCVIII.

Serial No.	Station	Latitude	Longitude in Greenwich Terms	Bessel's Spheroid and Tashkent Vertical		Corrections		Helmert's Spheroid and Kalianpur Vertical	
				Deflection in Prime Vertical	Deflection in Meridian	-u	-u cot λ	Deflection in Prime Vertical	Deflection in Meridian
	Tashkent ...	41° 21'	69° 18'	0"00	0"00	"	"	- 9"9	- 6"5
1	Chodschent ...	40 17	37	+ 13.06	+ 3.83	- 6.94	- 10.12	+ 2.9	- 3.1
2	Karatschekum ...	16	70 5	+ 3.71	+ 22.52	- 6.89	- 10.09	- 6.4	+ 15.6
3	Kanybadam ...	19	26	- 4.63	+ 22.66	- 6.83	- 10.05	- 14.7	+ 15.8
4	Tschil-Machram ...	33	33	...	+ 2.36	- 6.72	...	...	- 4.4
5	Begowat (south) ...	19	43	...	+ 42.82	- 6.80	...	...	+ 36.0
6	Puntan ...	44	43	...	- 8.13	- 6.62	...	...	- 14.8
7	Sary-Kurgan ...	20	71 2	+ 9.47	+ 30.34	- 6.77	- 10.03	- 0.6	+ 23.6
8	Pap ...	54	4	...	- 20.89	- 6.53	...	...	- 27.4
9	Begowat (north) ...	38	14	...	+ 8.92	- 6.60	...	...	+ 2.3
10	Karaul-tjube ...	31	15	+ 18.14	+ 16.64	- 6.65	- 9.97	+ 8.2	+ 10.0
11	Warsyk ...	41 7	16	...	- 26.87	- 6.40	...	...	- 33.3
12	Katput ...	40 16	20	...	+ 41.13	- 6.76	...	...	+ 34.4
13	Gurt-tjube ...	50	28	...	- 4.13	- 6.51	...	...	- 10.6
14	Kassan ...	41 15	36	...	- 20.30	- 6.31	...	...	- 26.6
15	Chalmion ...	40 11	39	...	+ 49.41	- 6.76	...	...	+ 42.7
16	Namangan ...	41 0	41	- 2.94	- 9.73	- 6.41	- 9.83	- 12.8	- 16.1
17	Martelan ...	40 23	47	+ 3.53	+ 32.56	- 6.65	- 9.99	- 6.5	+ 25.9
18	Bjalaja ...	41 4	50	...	- 13.08	- 6.37	...	...	- 19.5
19	Kara-tjube ...	40 37	50	...	+ 17.17	- 6.57	...	...	+ 10.6
20	Belyktschi ...	53	52	...	- 1.87	- 6.46	...	...	- 8.3
21	Utsch-Kurgan I ...	41 6	72 3	...	- 10.63	- 6.33	...	...	- 17.0
22	Utach-Kurgan II ...	40 14	4	...	+ 41.63	- 6.70	...	...	+ 34.9
23	Kuwa ...	31	4	+ 0.42	+ 28.40	- 6.58	- 9.94	- 9.5	+ 21.8
24	Tsch-tjube ...	40	10	...	+ 18.97	- 6.51	...	...	+ 12.5
25	Tschumbagsysch ...	54	13	...	- 0.38	- 6.41	...	...	- 6.8
26	Isbaken ...	41 2	21	...	- 12.36	- 6.34	...	...	- 18.7
27	Mim-tjube ...	40 29	22	- 6.84	+ 33.74	- 6.55	- 9.94	- 16.8	+ 27.2
28	Andischan ...	47	24	...	+ 12.63	- 6.43	...	...	+ 6.2
29	Kisyl-Kurgan ...	20	24	...	+ 32.39	- 6.63	...	...	+ 25.8
30	Salb-tjube ...	38	34	...	+ 25.85	- 6.50	...	...	+ 19.4
31	Massy ...	41 5	39	...	- 15.71	- 6.27	...	...	- 22.0
32	Chodscha-Syrjan ...	40 48	43	...	+ 11.49	- 6.38	...	...	+ 5.1
33	Tjulka-tjube ...	18	46	...	+ 32.71	- 6.59	...	...	+ 26.1
34	Osh ...	31	49	- 6.04	+ 23.43	- 6.50	- 9.90	- 15.9	+ 16.9
35	Chasret Ujunys ...	46	56	...	+ 13.70	- 6.38	...	...	+ 7.3
36	Mady ...	34	56	...	+ 22.80	- 6.47	...	...	+ 16.3
37	Deschalabad ...	55	73 1	...	+ 4.38	- 6.32	...	...	- 1.9

\* c. f. Comptes Rendus de L' Association Geodesique Internationale for 1898 (Annexe A IIc, p. 268).  
 † Converted from Pulkowa Longitude by applying + 30° 20' (more accurately 30° 19' 38".55).

The following description is extracted from *Comptes-Rendus de L' Association Geodesique Internationale* for 1896 (Annexe B XI p. 309):—

“The researches recently completed on the deviation of the plumb-line in Ferghana (Turkistan) are of special interest. This valley lying between  $40^{\circ} 15'$  and  $41^{\circ} 15'$  in latitude and  $39^{\circ} 30'$  and  $42^{\circ} 45'$  in longitude, east of Pulkowa is a deep depression the walls of which are pierced in their western part by the narrow bed of the Syr Daria. The bottom of the valley has an approximately elliptic figure with its major axis 250 km. long following the direction of the parallel, and its minor axis 110 km. that of the meridian. On the north the valley is enclosed by chains of mountains of an average height of 2500 to 3500 m. and on the south are the Alai, the Trans Alai, the Pamirs and the Hindu Kush. Such a position leads one to expect considerable deviation particularly in latitude and explains the investigations which have been made in order to verify this supposition. To this end 37 determinations of latitude have been made of points equally distributed over the district and 10 of longitude at points nearly on the same parallel. Taking Bessel's Ellipsoid, and the point Balyktschi as zero, we obtain for the deviation in latitude values for A—G from  $-25''$  on the north up to  $+51''$  on the south of the valley, in longitude the deviation of opposite sign amounts to  $25''$ ”.

7. A table of gravity residuals in the same district\* is also given for the sake of completeness.

TABLE XCIX.

	No.	Station	Latitude	Longitude	Height in Metres	$\gamma_0 - \gamma_0$ $C_m$ 10-3 x
TURKISTAN	1	Pamir Post ... ..	$38^{\circ} 10' 0''$	$73^{\circ} 58' 2''$	3700	- 30
	2	Kala-i-Wanj ... ..	$38 22 \cdot 2$	$71 27 \cdot 0$	1795	- 177
	3	Sar-i-pul ... ..	$38 24 \cdot 5$	$70 5 \cdot 5$	1500	- 100
	4	Kala-i-Chamb ... ..	$38 27 \cdot 3$	$70 46 \cdot 5$	1345	- 152
	5	Rabat Ak Baital ... ..	$38 29 \cdot 7$	$73 51 \cdot 5$	4100	+ 74
	6	Rabat Maskol ... ..	$38 42 \cdot 0$	$73 31 \cdot 7$	4200	+ 169
	7	Kara Kul Lake ... ..	$39 6 \cdot 4$	$73 31 \cdot 2$	3920	+ 35
	8	Irkeshtan Fort ... ..	$39 41 \cdot 9$	$73 55 \cdot 5$	2850	- 56
	9	Ak-bossaga ... ..	$39 48 \cdot 6$	$73 13 \cdot 7$	2875	- 128
	10	Sufi Kurgan ... ..	$40 1 \cdot 5$	$73 30 \cdot 0$	2115	- 91
	11	Karaul Kishlak ... ..	$40 2 \cdot 2$	$72 6 \cdot 0$	1300	- 157
	12	Gultsha ... ..	$40 19 \cdot 0$	$73 25 \cdot 7$	1583	- 126
	13	New Marghilan ... ..	$40 23 \cdot 7$	$71 46 \cdot 7$	581	- 159
	14	Langar ... ..	$40 24 \cdot 6$	$73 5 \cdot 7$	1685	- 67
	15	Osh ... ..	$40 31 \cdot 4$	$72 46 \cdot 6$	1021	- 106
	16	Andijan ... ..	$40 45 \cdot 8$	$72 20 \cdot 6$	530	- 185
	17	Tashkent ... ..	$41 19 \cdot 5$	$69 17 \cdot 7$	478	- 50
	18	Wysokoji Khojand ... ..	$42 30 \cdot 9$	$70 33 \cdot 9$	1060	- 1
	19	Chodient ... ..	$40 17 \cdot 1$	$69 34 \cdot 7$	320	- 140
	20	Namangan ... ..	$40 59 \cdot 7$	$71 38 \cdot 7$	440	- 178

\* c. f. C. R. 1911—Volume III pp. 156-168.





## APPENDIX.

### Various Determinations of the Axes of the Earth.

For convenience of reference the principal values of the elements of the figure of the earth obtained from time to time are given below, expressed in units of 1000 feet and kilometers. It is to be observed that the datum of height in different continents is only in the same terms on the assumption that the geoid is identical with the spheroid. The quantities determined really refer to the several concentric spheroids through these sea level datum points.

#### 1. RADIUS OF THE EARTH CONSIDERED AS A SPHERE.

Reference No.	Authority	Date	RADIUS		Data used
			in 1,000 feet	in kilometres	
1	Eratosthenes ...	250 B. C.	24370	7428	Arc from Syene (Upper Egypt) to Alexandria.
2	Posidonius ...	80 ..	23190	7069	"
3	Richard Norwood...	1637 A. D.	21038	6413	Arc from London to York. Mean Lat. 52° 42' 30".
4	Jean Picard ...	1669 ..	20806	6372	Arc from Paris to Amiens. Mean Lat. 49° 30'.

#### 2. AXES OF THE EARTH CONSIDERED AS A SPHEROID.

Jean Richer (d. 1696) pointed out that the Earth was not a sphere.

Reference No.	Authority	Date	RADIUS OF CURVATURE IN MEAN LATITUDE OF ARC		DEDUCED		Data used	
			in 1,000 feet	in kilometres	in 1,000 feet	$\frac{1}{\epsilon}$		
5	J. and D. Cassini ...	1684-1718	20889.2	6300.8	20988.5 (=6397.3km)	216.82	Arc from Paris to Dunkirk. Mean Lat. 49° 56' 9".	The immediate inference was that the degree diminishing with the increasing latitude, the Earth must be a <i>prolate spheroid</i> .  Bouguer, De la Coudamine, Maupertuis, Clairault, De Thury and De Lacaille proved that the Earth was an <i>oblate and not a prolate spheroid</i> .
6	J. and D. Cassini ...	...	20910.4	6376.1			Arc from Paris to Collioure. Mean Lat. 45° 40' 42".	
7	Bouguer and De la Coudamine	1735-1761	20795.4	6338.3			Arc in Pern. Mean Lat. 1° 31' 0" S.	
8	Maupertuis and Clairault	1736	21038.4	6412.4			Arc in Finland. Mean Lat. 66° 10' 35".	
9	De Lacaille ...	1762	20.97.4	6369.4		Arc at Cape of Good Hope. Mean Lat. 33° 18' 30" S.		

## 3. AXES OF THE EARTH CONSIDERED AS A SPHEROID DETERMINED FROM A GROUP OF ARCS, GRAVITY ETC.

Quantities given by authority named are shown in roman figures; deduced quantities are in italics.

Reference No.	Authority	Date	SEMI MAJOR AXIS (=a)		SEMI MINOR AXIS (=b)		$\frac{1}{\epsilon}$	Data used etc.
			in 1,000 feet	in kilometres	in 1,000 feet	in kilometres		
10	Laplace ...	1790	20919.768	6376.340	20862.822	6355.935	312.20	Arcs in Peru, India, France, England, and Sweden.
11	Everest ...	1830	22.84095	77.276	53.28403	56.075	300.6017	Arc from Damargida to Kalianpur Mean Lat. 21° 5' 13".
			(22.93180)	...	(53.37458)	...	...	As expressed by Everest in terms of Indian 10-foot bar A (=0.99995658 feet).
12	Airy ...	1830	23.713	76.542	53.610	56.236	299.33	14 meridian arcs and 4 arcs of parallel.
13	Bessel ...	1841	23.237 ± .702	77.397 ± .214	53.296	56.079	299.15	From 10 meridian arcs.
14	Clarke ...	1857	26.348	78.345	55.233	56.669	294.26	Arcs: Anglo-Gallic, Russian, Indian, Prussian, Peruvian, Hanoverian, Danish.
15	Pratt ...	1863	26.189	78.297	55.316	56.695	295.26	Semi axes and ellipticity of the Mean Figure of the Earth. From a comparison of the Anglo-Gallic, Russian and Indian arcs &c.
16	Clarke ...	1866	26.062	78.258	55.121	56.635	294.98	Arcs: Anglo-Gallic (rejecting 21 lat. stations), 2nd Indian, Russian, Peruvian, Cape.
17	Clarke ...	1880	26.202	78.301	54.895	56.871	293.47	
18	Clarke-Bessel ...	...	26.203	78.301	56.252	[56.980]	299.15	
19	Clarke-Bessel ...	...	26.839	78.190	56.888	[56.869]	299.15	From Clarke's value of 1866 with an old conversion factor.
20	Darwin ...	1899	...	...	...	...	296.4	From consideration of precession.
21	.....	...	25.329	78.035	55.381	56.715	299.15	(1.0001) × Bessel's value; adopted by the Central Bureau, International Geodetic Association.
22	Hayford (C.andG.S.) ...	1906	26.144 ± .112	78.283 ± .034	56.885	56.668	297.8 ± 0.9	
23	Helmert ...	1907	26.871	78.200	56.721	56.818	298.3	From gravity determinations.
24	Hayford (C.andG.S.) ...	1909	26.488 ± .059	78.388 ± .018	56.019	56.909	297.0 ± 0.5	Adopted for International $\frac{1}{M}$ map.
25	Helmert-Hayford ...	...	26.436	78.372	56.978	56.896	297	Obtained by S. Wellich taking Helmert's value with weight=unity and the modified Hayford values with weight=4.
26	International Map Committee ...	1909	26.002	78.24	54.874	56.56	[294.2]	
27	E.W. Brown ...	1914	...	...	...	...	293.7 ± 0.3	From Lunar theory. The value will make the observed motions of perigee and node agree with the theoretical values.
28	Nautical Almanac ...	1911	...	...	...	...	297	Adopted in the conference of Nautical Almanac directors.
29	Crommelin ...	...	...	...	...	...	294.4 ± 1.5	From Moon's parallax at Greenwich and Cape. Obtained by a hundred pairs of simultaneous observations at the Cape and Greenwich Observatories by a comparison between theoretical and observed values of the Moon's parallax.

4. AXES OF THE EARTH CONSIDERED AS AN ELLIPSOID.

Reference No.	Authority	Date	Equatorial Axes				Polar Axis		Longitude of major axis E. of Greenwich	Data used
			a		b		c			
			in 1,000 feet	in kilometres	in 1,000 feet	in kilometres	in 1,000 feet	in kilometres		
80	Schubert ...	About 1860	20027.397 (3272871 toises)	6378.665	20025.044 (3272303 toises)	6377.048	20856.750 (3201468 toises)	6356.630	41° 4'	Arcs: French, English, Russian, Indian, Cape, Prussian and Peruvian.
81	Clarke ...	1860	„ 26.629	„ 78.431	„ 25.105	„ 77.986	„ 53.477	„ 56.439	15° 34'	Arcs: French, English, Indian, Russian, Prussian, Peruvian, Cape.

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For some other values see *La Figure de La Terre*, Paris, 1901, by Capt. G. Pezrier.

Conversion factors used— \*1 metre = 3.28084275 feet. log = 0.5159854152  
 or 1 foot = 0.30479973 metre. „ = 1.4840145780  
 1 toise = 6.39459252 feet „ = 0.8058128753

\* *Vide* "Determination du Rapport du yard au metre" by M. Benoit, Paris, 1896; also *Text Book of Topographical and Geographical Surveying*, p. 359 by Colonel C. F. Close, C. M. G., R. E.



## Notes

## Poshkar H. S. and Gogipatri H. S.

Poshkar H. S. and Gogipatri H. S. are two geodetic stations of Capt. T. G. Montgomerie's Kashmir Principal Series which was adjusted in 1921 in terms of the Gilgit Series and its extension, the Indo-Russian Connection, involving small corrections to the values of their geodetic coordinates.

*Latitude observations.*

In 1860 Capt. Montgomerie observed astronomical latitudes both at Poshkar and Gogipatri with a 14-inch vernier theodolite by circum-meridian altitudes; but as the results by N. stars differed by  $20''\cdot76$  and  $20''\cdot90$  respectively from those by S. stars, they were not accepted as reliable.

In 1922 Major K. Mason reobserved at both these stations by Talcott method with the Zenith telescope. His deflection results (S.  $13''\cdot8$ ) at Poshkar differed by  $1''$  from the mean of Montgomerie's (S.  $14''\cdot8$ ). At Gogipatri Major Mason's result (S.  $3''\cdot03$ ) was almost identical with that of Montgomerie (S.  $3''\cdot04$ ).

## Pamir Boundary Pillar No. 1

*Latitude observations.*

During the Pamir Boundary Commission in 1895, Col. R. A. Wahab (Wauhope) observed astronomical latitudes of this point.

*Longitude observations.*

Col. Zalesky of the Russian Commission during the same work obtained the longitude ( $73^{\circ} 46' 32''$ ) of the same point by comparison of local time with that shown by six chronometers brought from Osh the longitude of which had been determined telegraphically from Pulkowa. He considered his probable error not greater than 5 seconds of arc.

*The trigonometrical observations.*

This point was fixed trigonometrically by the British Commission. To bring all the Boundary Commission work in conformity with the adjusted triangulation of India, mean corrections of  $-4''\cdot7$  in  $\lambda$ ,  $-3''\cdot3$  in  $L$ , and  $+176$  feet in height have been applied.

Though the results thus obtained are not highly accurate yet they show that no great plumb-line deflections exist here.

## Pyr. de la base S. W. (Kizil-Rabate).

In 1912-13 during the Indo-Russian Connection, the Russians observed astronomical latitude and astronomical azimuth at their station Pyr. de la base S. W. The triangulated values have been expressed in Indian terms through Kukhtek and Sarblock. (*vide pp. 211-12*).

## Depsang Transit S. and Leh Transit S.

*Latitude and Longitude observations.*

In 1914 during the De Filippi Expedition astronomical latitudes and longitudes of Depsang, Leh and other six stations were observed. Longitudes were determined by means of wireless signals transmitted from Lahore which were simultaneously timed and recorded at Dehra Dun, in the Transit room adjoining the Dome Observatory (new).

*The trigonometrical observations.*

The point at Leh was intersected from two stations of Kashmir triangulation and that at Depsang was resected from known peaks, by Major H. Wood (Survey of India) attached to the expedition.

## Dumdangi H. S. and Thakurganj H. S.

*Latitude observations.*

In 1922 Major H. T. Morshead observed astronomical latitudes at these two stations with a 6-inch Transit theodolite. As the work was done with a small instrument, the values of deflection are given to the nearest second.

## Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A-G) cot $\lambda$ for azimuth or (A-G) cos $\lambda$ for longitude observations†	Meridian Deflection†	
449	30 C	Robat 1908 S.	3099	A 29 49 9.16 G 29 48 58.75	" " "	G 60 55 10.90	" " "	" " "	+ 10.4	449
450	34 L	Zawa 1904 H.S.	7922	G 28 57 44.43	G 66 35 18.70	A 178 40 55.7 G 178 40 50.0	Zebra E 0 6	+ 10.3		450
451	42 G	Pamir Boundary Pillar No. 1 1895	13574	A 37 26 33 G 37 26 27	A 73 46 32 G 73 46 30			+ 2	+ 6	451
452	K	Pyr. de la base S.W. (Kizil-Rabate) 1912		A 37 26 40.28 G 37 26 38.80	G 74 44 24.05	A 255 45 22.7 G 255 45 24.7		- 2.6	+ 1.5	452
453	43 G	Murree Obey. 1860 s.	7458	G 33 54 57.36	G 73 24 25.67	A 227 39 50.4 G 227 39 54.1	Nerh D 1 6	- 5.5		453
454	J	Rustamgarhi 1962 h.s.	5351	G 34 4 38.70	G 74 49 53.68	A 30 52 6.8 G 30 52 27.1	Gogipatri E 1 24	- 30.0		454
455	J	Poshkar 1922 H.S.	8323	A 34 2 2.78 G 34 1 48.98	G 74 29 51.23	A 318 14 0.7 G 318 13 54.0	Gogipatri D 0 30	+ 9.9	+ 13.8	455
456	J	Zebanwan 1922 H.S.	8799	A 34 3 33.50 G 34 3 59.14	G 74 54 2.94				- 25.6	456
457	K	Gogipatri 1922 H.S.	7752	A 33 51 46.90 G 33 51 43.87	G 74 40 38.57	A 222 17 18.0 G 222 17 11.7	Zebanwan E 0 29	+ 9.4	+ 3.0	457
458	K	Reban 1922 H.S.	5447	A 33 45 17.74 G 33 45 25.76	G 74 59 52.05				- 8.0	458
459	52 E	Depang Transit 1914 s.	17591	A 35 17 20.77 G 35 17 23.59	A 77 58 17.85 G 77 58 23.82			- 4.9	- 2.8	459
460	F	Leh Transit 1914 s.	11554	A 34 9 54.10 G 34 10 8.6	A 77 34 53.89 G 77 35 4.4			- 8.7	- 14.5	460
461	63 I	Kopa 1924 T.S.	365	A 27 7 1.24 G 27 7 3.74	G 82 12 48.22				- 2.50	461
462	J	Sirwara 1924 T.S.	348	A 26 16 33.74 G 26 16 23.86	G 82 7 30.01				+ 9.88	462
463	J	Bisaul 1924 T.S.	342	A 26 40 42.76 G 26 40 37.38	G 82 20 54.43				+ 5.38	463
464	K	Ramapura 1924 T.S.	356	A 25 45 3.23 G 25 44 55.09	G 82 5 40.58				+ 8.14	464
465	M	Saunbarsa 1924 T.S.	315	A 27 11 16.36 G 27 11 26.33	G 83 21 18.45				- 9.97	465
466	M	Mathia 1923 T.S.	334	A 27 7 54.10 G 27 8 4.37	G 83 51 32.13				- 10.27	466
467	N	Rajabari 1923 T.S.	267	A 26 54 2.13 G 26 54 3.04	G 83 15 35.49				- 0.92	467
468	N	Baniapar 1924 T.S.	267	A 26 15 15.50 G 26 15 7.73	G 83 23 2.28				+ 7.78	468
469	O	Kanann 1924 T.S.	270	A 25 43 13.38 G 25 43 3.62	G 83 23 51.38				+ 9.76	469
470	72 A	Kannagar 1923 T.S.	312	A 27 8 50.43 G 27 9 4.09	G 84 19 35.56				- 13.66	470
471	A	Sakta 1923 T.S.	267	A 27 1 31.97 G 27 1 44.04	G 84 40 54.56				- 12.07	471
472	F	Minaria 1923 T.S.	239	A 26 45 6.27 G 26 45 12.55	G 85 15 54.25				- 6.28	472
473	F	Shahpur 1923 T.S.	173	A 26 24 42.73 G 26 24 42.17	G 85 47 25.37				+ 0.36	473
474	78 b	Dumlangi 1922 T.S.	307	A 26 28 15 G 26 28 11	G 88 17 35				- 6	474
475	B	Thakorganj 1922 T.S.	264	A 26 24 53 G 26 24 1	G 88 7 45				- 9	475

DEFLECTIONS OF THE PLUMB-LINE

XCV. (Continued from pp. 206 & 207)  
in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*					Serial No.
	Case I : $\delta a = 1 \text{ km}$			Case II : $\delta b = 1 \text{ km}$			Case III : Latitude $u_0 = 1''$			Case IV : Azimuth $w_0 = 1''$			$a = 6378200 \text{ metres}, 1/\epsilon = 298.3.$					
	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	$u$	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian	
449	+1.85	"	"	-4.49	"	"	+0.96	"	"	+0.26	"	"	-0.98	"	"	"	+ 11.4	449
450			+5.92			-0.22			-0.40			+1.85			+7.58	+ 2.5		450
451	+1.60	+2.30		-9.10	-0.29		+1.00	-0.04		+0.06	+0.23		-4.90	+2.20		0	+ 11	451
452	+1.59		+1.41	-9.11		+0.11	+1.00		-0.08	+0.05		+1.50	-4.93		+3.29	- 6.6	+ 6.4	452
453			+2.12			+0.08			-0.13			+1.63			+4.08	- 10.0		453
454			+1.91			+0.05			-0.09			+1.63			+3.87	- 34.3		454
455	+1.63		+1.58	-7.22		+0.06	+1.00		-0.10	+0.05		+1.63	-3.48		+3.56	+ 5.9	+ 17.3	455
456	+1.63			-7.24			+1.00			+0.04			-3.51				- 22.1	456
457	+1.62		+1.49	-7.12		+0.05	+1.00		-0.09	+0.05		+1.63	-3.42		+3.49	+ 5.5	+ 6.4	457
458	+1.61			-7.05			+1.00			+0.04			-3.39				- 4.6	458
459	+1.63	-0.19		-7.95	+0.02		+1.00	0.00		-0.01	+0.19		-4.10	+0.09		- 5.0	+ 1.3	459
460	+1.61	+0.04		-7.30	-0.01		+1.00	0.00		0.00	+0.17		-3.63	+0.16		- 9.0	- 10.9	460
461	+0.79			-2.45			+1.00			-0.07			-0.87				- 1.63	461
462	+0.60			-1.78			+1.00			-0.07			-0.55				+10.43	462
463	+0.69			-2.10			+1.00			-0.08			-0.72				+ 6.10	463
464	+0.47			-1.36			+1.00			-0.07			-0.36				+ 8.50	464
465	+0.83			-2.51			+1.00			-0.09			-0.90				- 9.07	465
466	+0.83			-2.46			+0.99			-0.10			-0.88				- 9.39	466
467	+0.76			-2.28			+1.00			-0.09			-0.80				- 0.12	467
468	+0.62			-1.76			+1.00			-0.09			-0.54				+ 8.32	468
469	+0.50			-1.33			+1.00			-0.09			-0.33				+ 10.09	469
470	+0.85			-2.48			+0.99			-0.11			-0.89				- 12.77	470
471	+0.83			-2.38			+0.99			-0.11			-0.84				- 11.33	471
472	+0.79			-2.17			+0.99			-0.12			-0.73				- 5.55	472
473	+0.73			-1.90			+0.99			-0.13			-0.60				+ 0.96	473
474	+0.85			-1.96			+0.98			-0.17			-0.59				- 3	474
475	+0.83			-1.91			+0.98			-0.17			-0.57				- 8	475

\*  $\delta a = 0.924$ ,  $\delta b = 0.743$ ,  $u_0 = 0.31$ ,  $w_0 = 1.10$ . Vide p. 2.





DEFLECTIONS OF THE PLUMB-LINE.  
*XCVI*  
 TABLE ~~XVI~~—(Continued).

No.	Name	Latitude	Longitude	Height	$g-\gamma_a-.011$	$g-\gamma_b+.030$	$g-\gamma_c-.011$
<i>Season 1913-14.</i>							
		° ' "	° ' "	<i>feet</i>	<i>dynes</i>	<i>dynes</i>	<i>dynes</i>
109	Abu	... 24 36	72 43	3836	+ .103	+ .022	+ .018
110	Ahmedabad	... 23 1	72 34	156	+ .020	+ .056	+ .025
111	Alibagh	... 18 39	72 52	12	- .016	+ .025	- .014
112	Baroda	... 22 19	73 11	109	- .026	+ .011	- .019
113	Broach	... 21 42	72 59	51	- .002	+ .037	+ .004
114	Daman	... 20 25	72 50	15	+ .032	+ .073	+ .038
115	Deesa	... 24 15	72 12	465	+ .032	+ .057	+ .036
116	Erinpura	... 25 9	73 4	872	+ .005	+ .017	+ .012
117	Pali-Marwar	... 25 48	73 19	719	- .001	+ .016	+ .009
118	Surat	... 21 10	72 48	30	+ .016	+ .056	+ .021
<i>Season 1923-24.</i>							
119	Bagaha Ghat	... 27 8	84 3	298	- .164	- .133	- .099
120	Bahraich	... 27 34	81 36	403	- .132	- .105	- .073
121	Etawah	... 26 47	79 1	492	- .046	- .022	- .024
122	Fatehgarh	... 27 22	79 38	493	- .063	- .039	- .040
123	Gainsari	... 27 32	82 36	364	- .167	- .138	- .102
124	Gonda	... 27 8	81 56	352	- .134	- .105	- .096
125	Motihari	... 26 39	84 55	220	- .164	- .130	- .110
126	Pilibhit	... 28 39	79 50	610	- .125	- .105	- .066
127	Shahjahanpur	... 27 54	79 56	510	- .084	- .060	- .050
128	Sitapur	... 27 33	80 41	449	- .101	- .075	- .065
129	Sonaripur	... 28 28	80 44	514	- .153	- .129	- .078

To be pasted over the column below  $g-\gamma_c-.011$  on pages  
208, 209 and 210 respectively.

<i>dynes</i>	<i>dynes</i>	<i>dynes</i>
+ .006	- .028	- .030
+ .018	.000	- .042
- .002	- .029	+ .042
		...
- .014	+ .011	+ .004
+ .015	+ .027	- .053
- .039	+ .009	
		+ .005
- .007	- .031	- .034
+ .021	- .009	- .040
+ .027	+ .003	
		+ .001
- .036	+ .019	- .087
- .006	- .018	- .002
+ .011	+ .028	
		- .015
+ .002	+ .007	+ .015
+ .015	+ .022	- .004
- .025	- .004	
		- .014
- .006	- .008	+ .012
+ .052	+ .036	+ .010
- .005	- .030	
		+ .019
+ .014	+ .028	+ .005
- .006	- .043	- .052
+ .021	- .001	
		- .059
- .005	- .013	+ .037
- .075	- .009	- .002
- .016	- .014	
		.000
- .002	- .064	+ .025
+ .021	- .014	+ .012
+ .008	- .068	
		- .070
+ .027	+ .017	- .050
- .008	+ .007	+ .024
- .006	+ .012	
		+ .018
+ .008	- .003	- .038
- .081	- .026	+ .015
- .018	+ .029	
		- .022
	- .016	+ .018
	+ .020	- .044
	- .008	
	- .036	
	+ .008	
	- .005	

